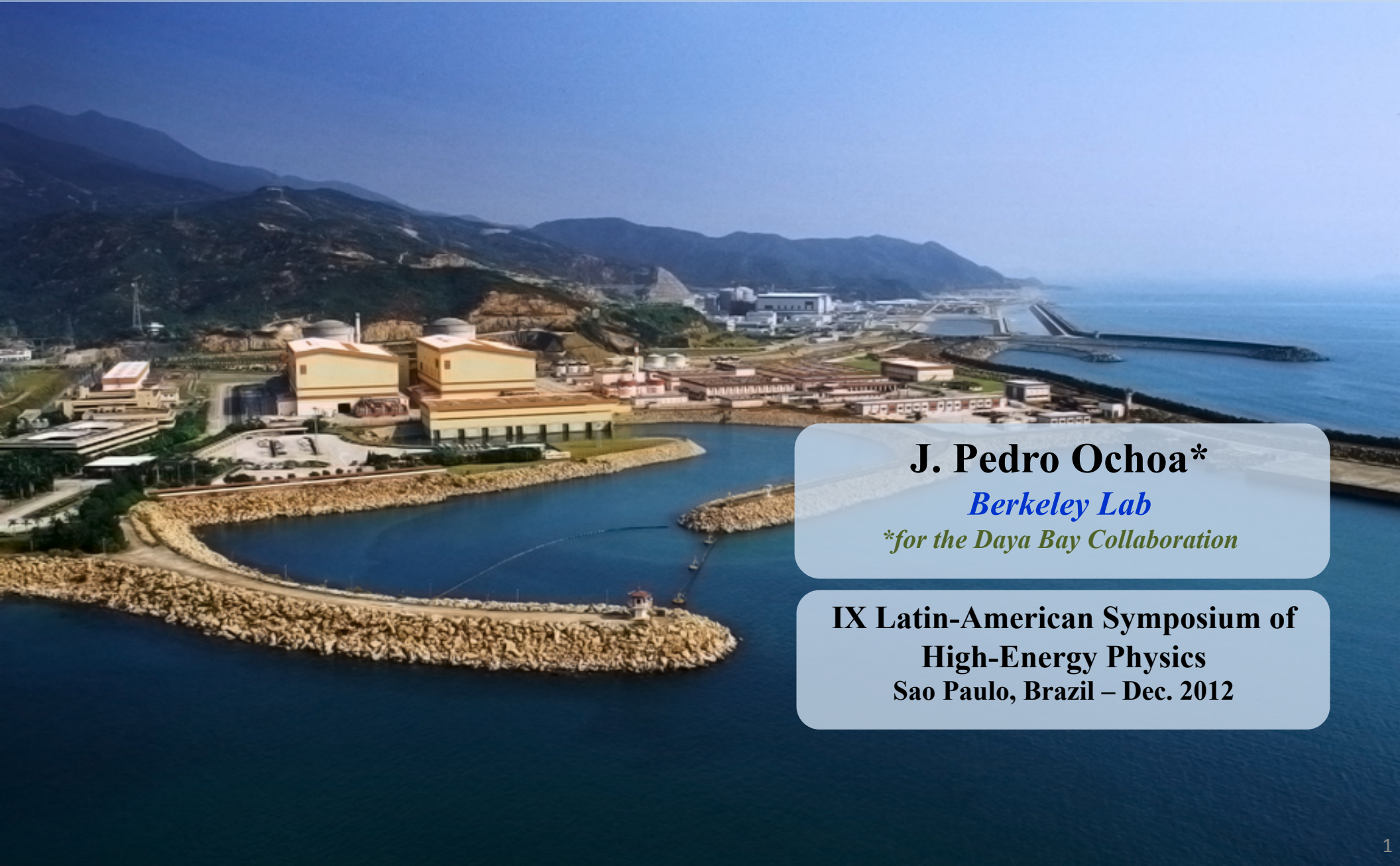
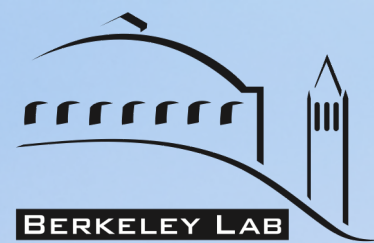




# Results from the Daya Bay Reactor Neutrino Experiment



**J. Pedro Ochoa\***

*Berkeley Lab*

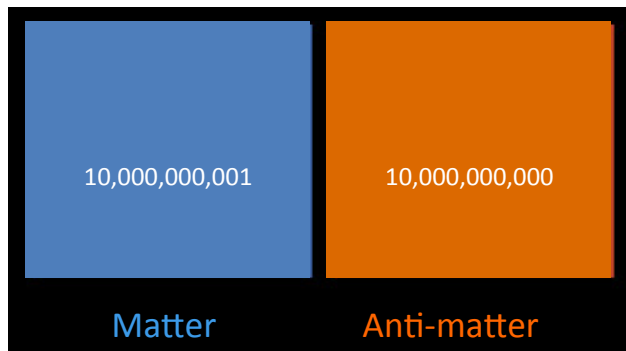
*\*for the Daya Bay Collaboration*

**IX Latin-American Symposium of  
High-Energy Physics  
Sao Paulo, Brazil – Dec. 2012**

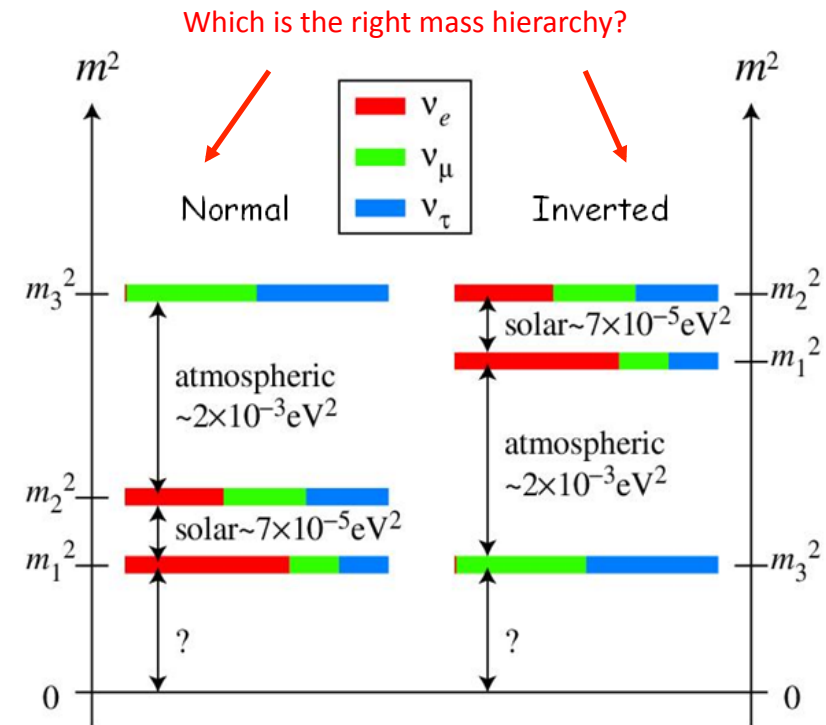
# The search for $\theta_{13}$

- ❖ The study of neutrino oscillations has now entered a precision era:
  - In particular, the neutrino mass differences are now known to a few percent
- ❖ Until recently however, one piece of the puzzle was missing: the  $\theta_{13}$  mixing angle
  - Important parameter in Standard Model, key input to future extensions
  - A non-zero  $\theta_{13}$  could also open the door to answering many other questions:

- Is there CP violation in the lepton sector?
- Why is there more matter than anti-matter in our universe?



- What is the mass hierarchy of the neutrino sector?



- ❖ Until less than a year ago we had some indications that  $\theta_{13}$  might be non-zero:

$$\sin^2(2\theta_{13}) < 0.15 \text{ (90 C.L.)}$$

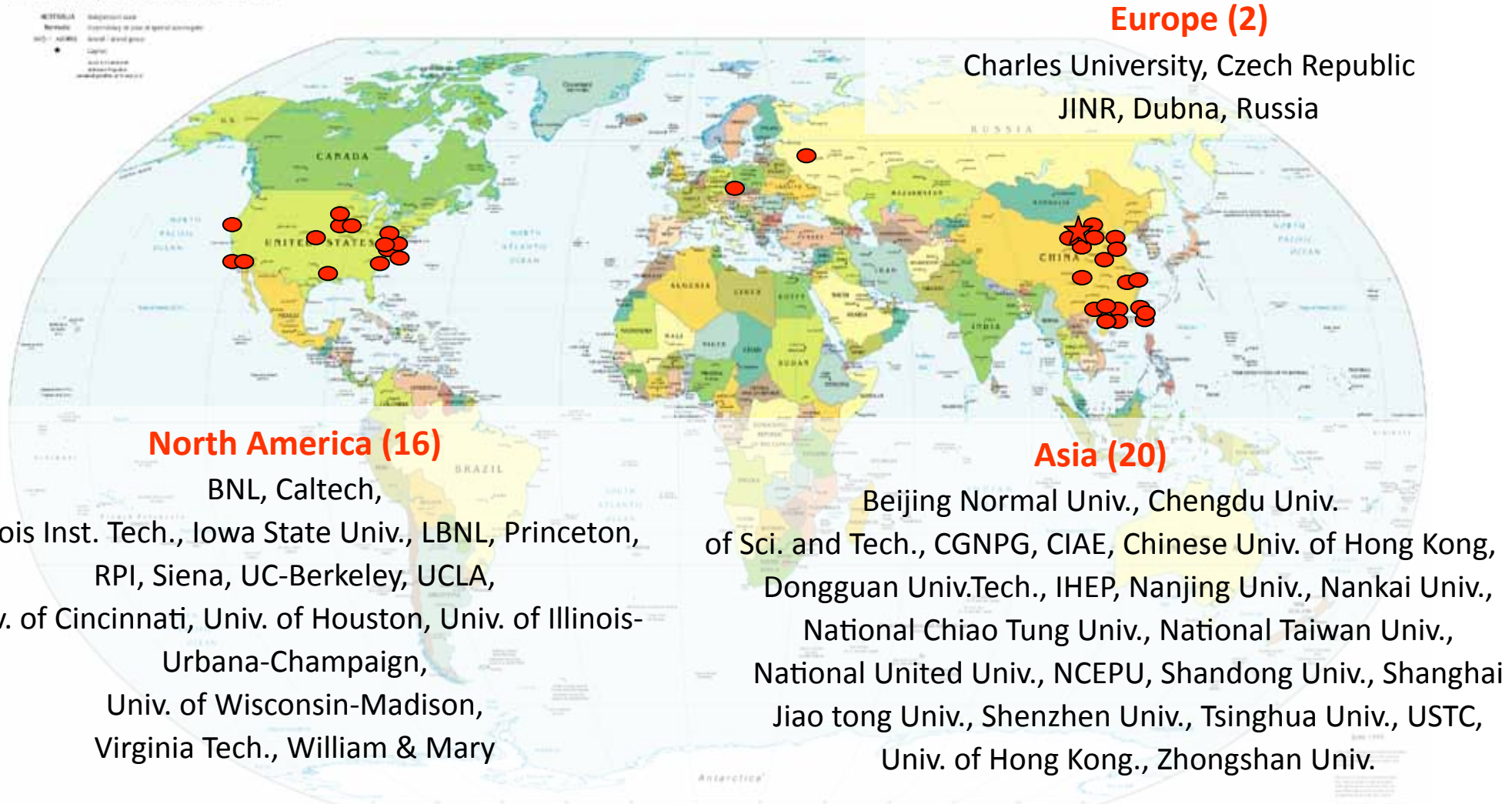
← from PDG 2011

- The Daya Bay Experiment was designed to make a precise measurement of  $\theta_{13}$

# The Daya Bay Reactor Neutrino Experiment

## ❖ The Daya Bay Collaboration:

Political Map of the World, June 1999



~230 Collaborators

# Daya Bay Experimental Layout

- ❖ Electron anti-neutrinos are produced in copious amounts in nuclear reactors.

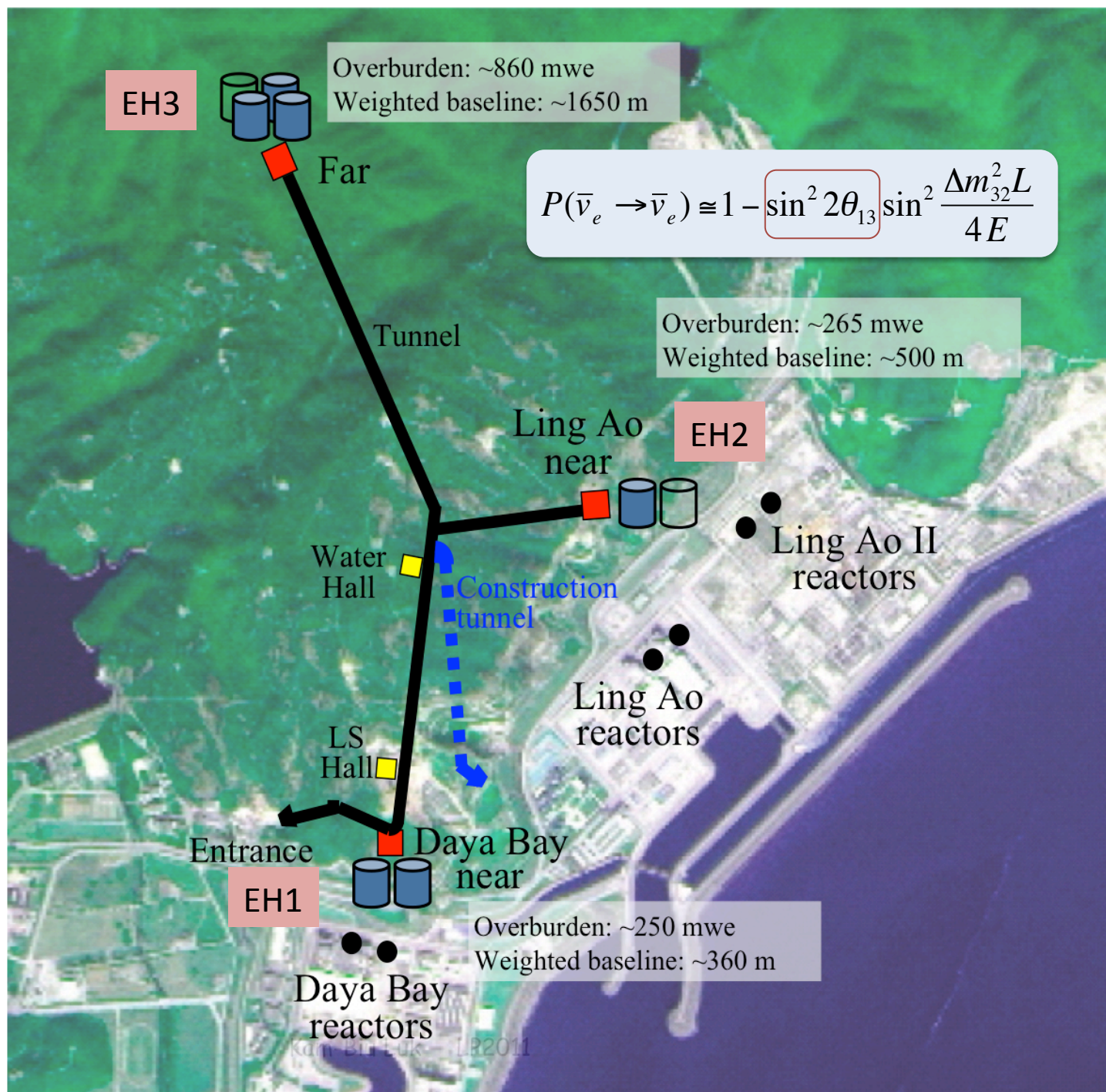
We position our detectors around the Daya Bay Power Plant in China, among the most powerful in the world.

- ❖ **Main principle:**

(i) sample the reactor anti-neutrino flux in the near and far locations, and

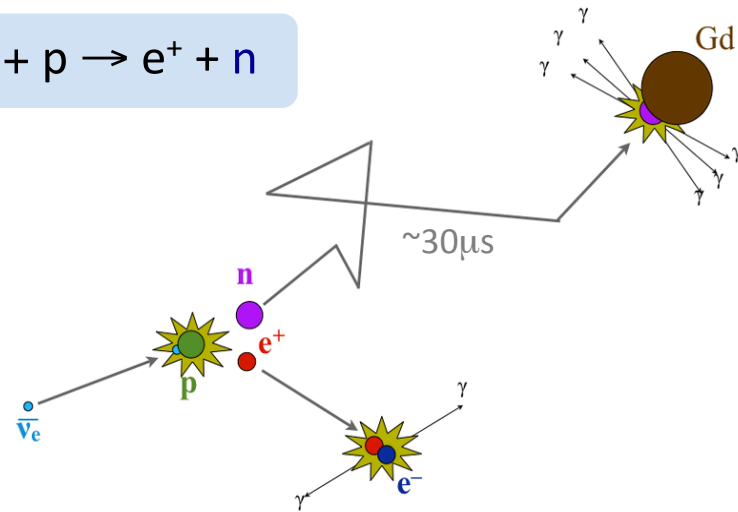
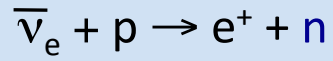
(ii) look for evidence of disappearance

Note: 6 / 8 detectors started taking physics data since Dec. 2011



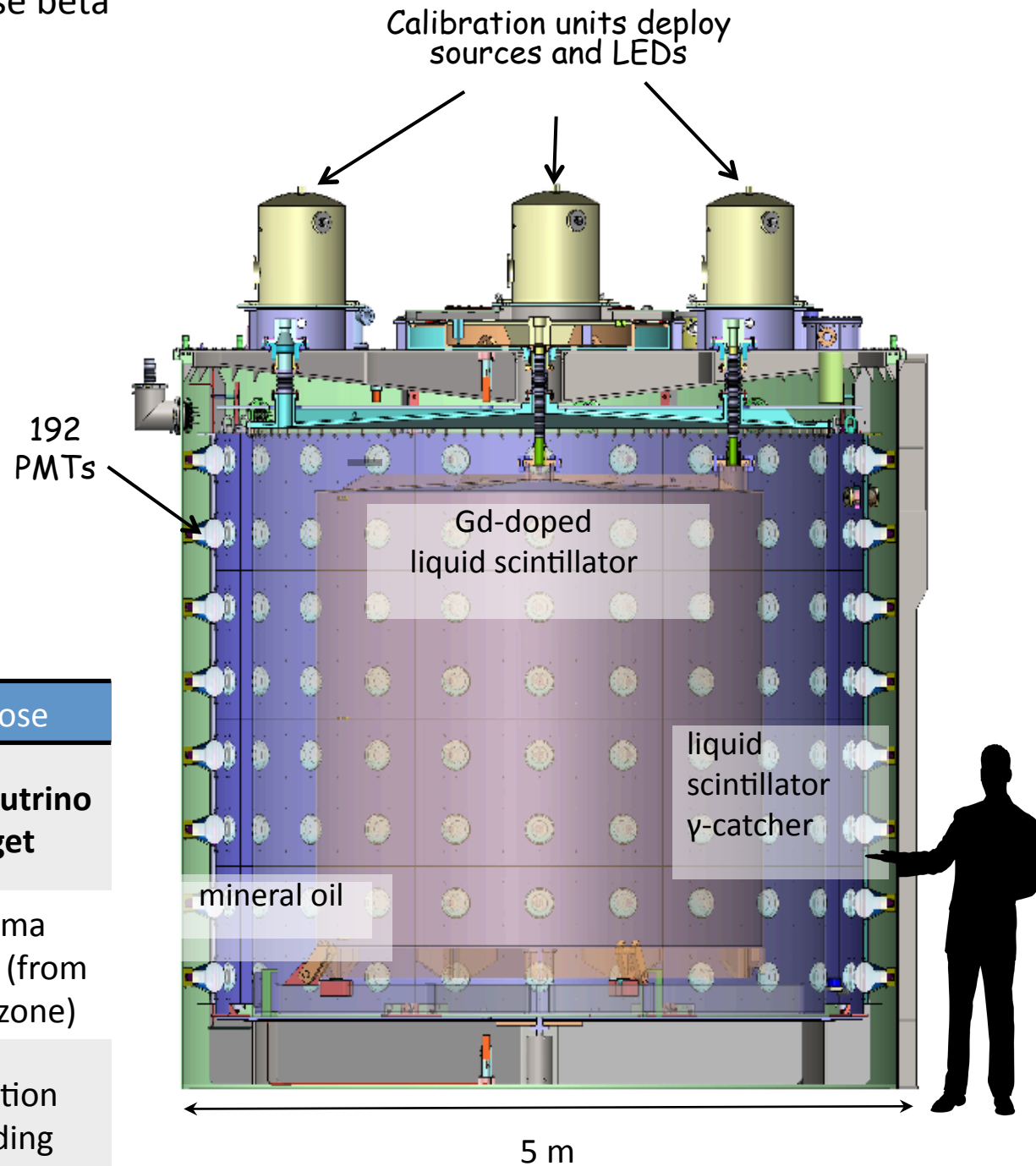
# The Detectors

- ❖ Anti-neutrinos are detected via the inverse beta decay (IBD) reaction:



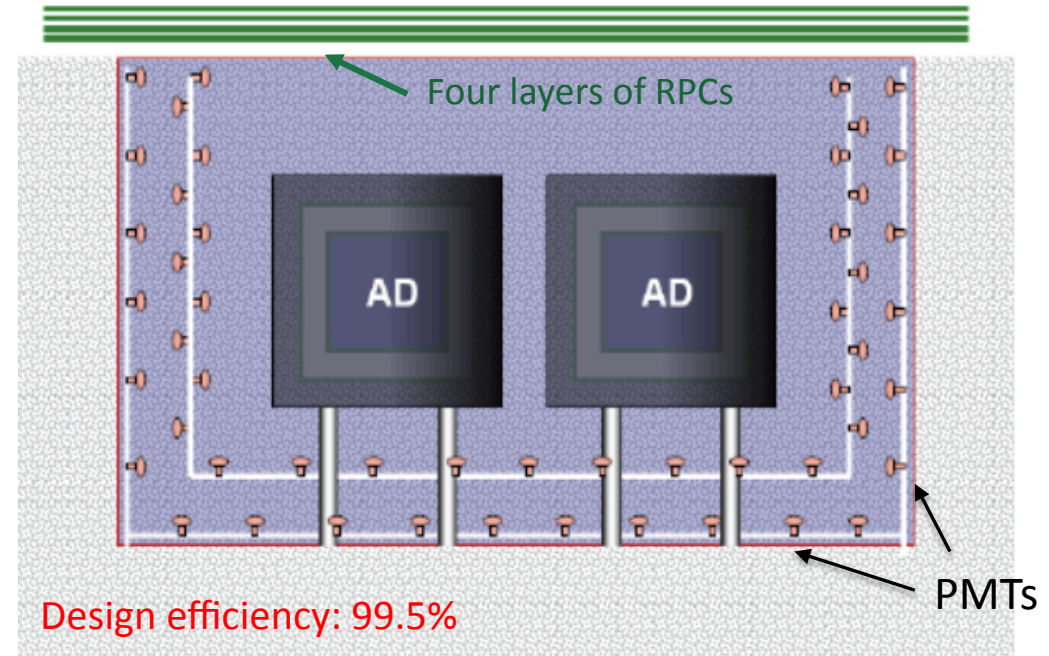
- ❖ The detectors are  $\sim 100$ ton three-zone cylindrical modules:

Zone	Mass	Liquid	Purpose
Inner acrylic vessel	20 t	Gd-doped liquid scintillator	<b>Anti-neutrino target</b>
Outer acrylic vessel	20 t	Liquid scintillator	Gamma catcher (from target zone)
Stainless steel vessel	40 t	Mineral Oil	Radiation shielding

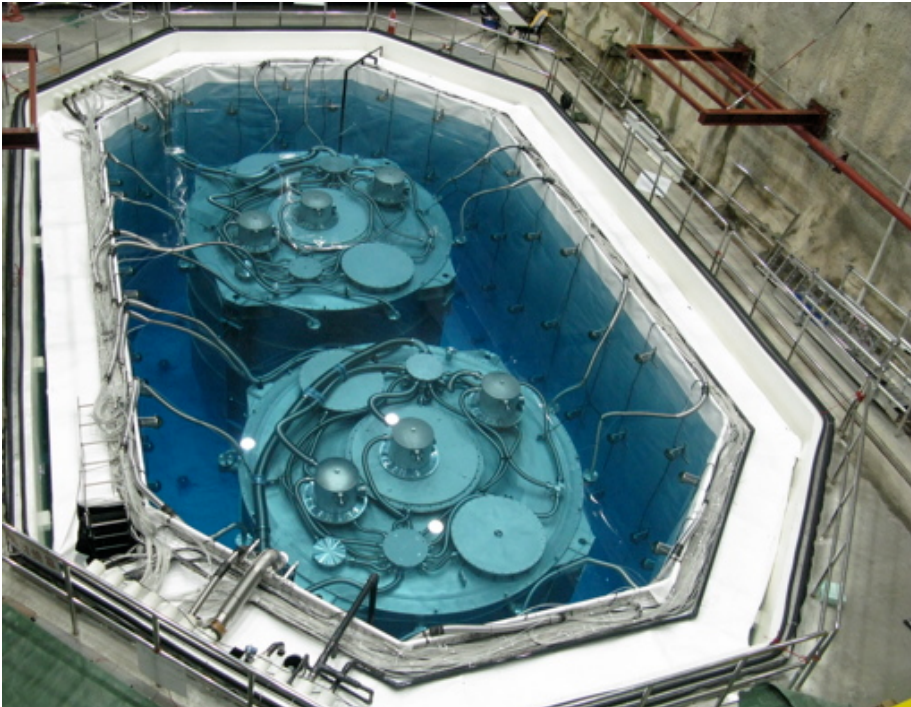


# The Water Pool

- ❖ The detectors are immersed in an instrumented water pool:
  - Double purpose:
    - ✓ Attenuates gammas from ambient radioactivity as well as neutrons produced by cosmic rays
    - ✓ Serves as a Cerenkov detector to tag cosmic ray muons (thus reducing backgrounds)

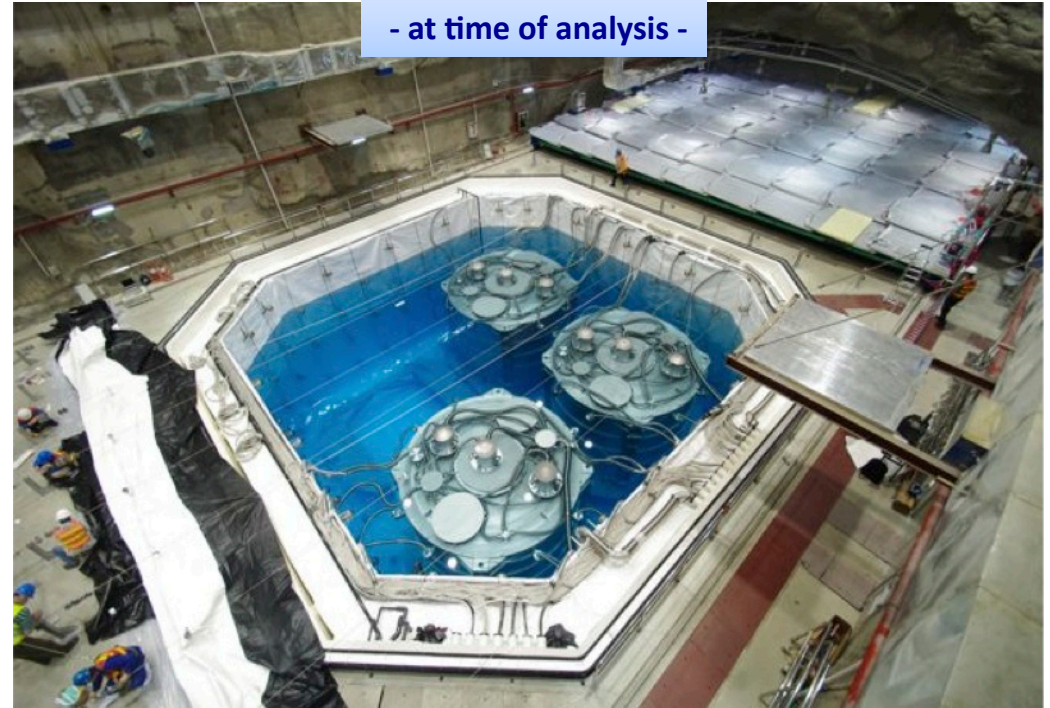


**EH1 (Daya Bay Near Hall)**



**EH3 (Far Hall)**

- at time of analysis -

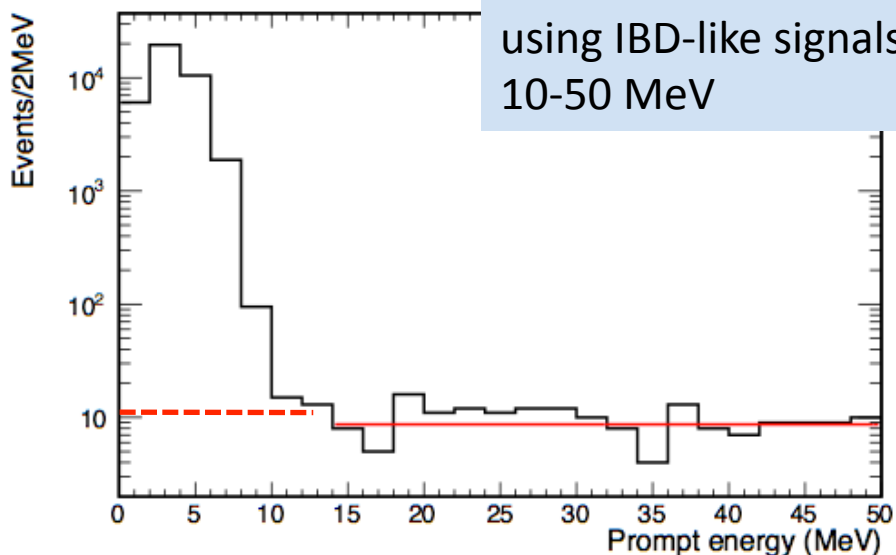


# Backgrounds

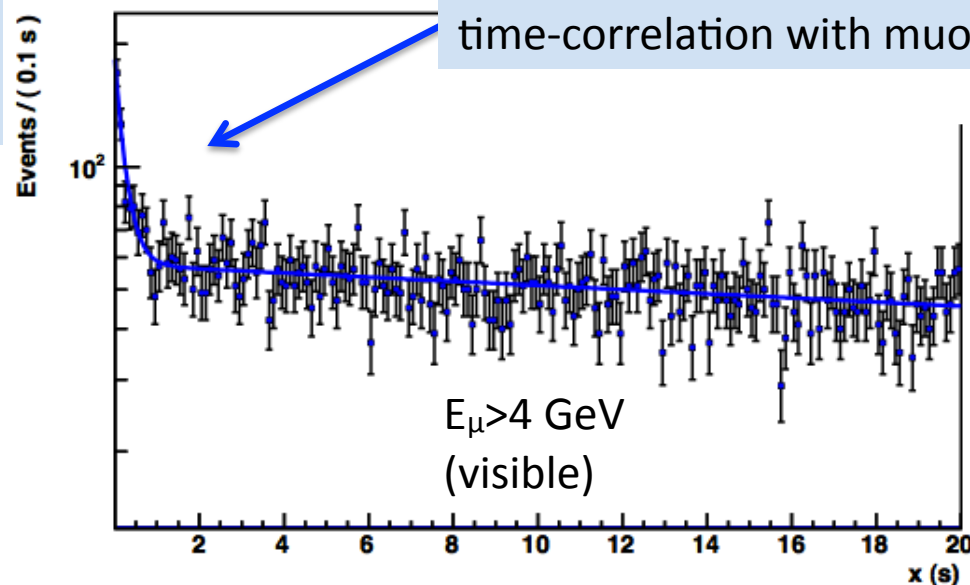
- ❖ Not quite a background free experiment, but close:
  - Total backgrounds are **5% (2%)** in far(near) halls
  - Background uncertainties are **0.3% (0.2%)** in far (near) halls
- ❖ The backgrounds are all estimated using data-driven methods:
  - The largest source of background (accidental coincidences) can be measured to  $\sim 1\%$

	Near Halls		Far Hall	
	B/S %	$\sigma_{B/S}$ %	B/S %	$\sigma_{B/S}$ %
<b>Accidentals</b>	<b>1.5</b>	0.02	<b>4.0</b>	0.05
<b>Fast neutrons</b>	0.12	0.05	0.07	0.03
<b><math>{}^9\text{Li}/{}^8\text{He}</math></b>	0.4	<b>0.2</b>	0.3	<b>0.2</b>
<b><math>{}^{241}\text{Am}-{}^{13}\text{C}</math></b>	0.03	0.03	0.3	<b>0.3</b>
<b><math>{}^{13}\text{C}(\alpha, n){}^{16}\text{O}</math></b>	0.01	0.006	0.05	0.03

EH1 Prompt energy, AD#1



${}^9\text{Li}/{}^8\text{He}$  Fit



# Dataset for oscillation analysis

## ❖ Dataset:

✓ December 24<sup>th</sup> 2011 to May 11<sup>th</sup> 2012 ← less than six months!

## ❖ Have accumulated more than 200k antineutrino interactions:

	AD1	AD2	AD3	AD4	AD5	AD6
Antineutrino candidates	69121	69714	66473	9788	9669	9452
DAQ live time (day)	127.5470		127.3763		126.2646	
Efficiency	0.8015	0.7986	0.8364	0.9555	0.9552	0.9547
Accidentals (/day)	9.73±0.10	9.61±0.10	7.55±0.08	3.05±0.04	3.04±0.04	2.93±0.03
Fast neutron (/day)	0.77±0.24	0.77±0.24	0.58±0.33	0.05±0.02	0.05±0.02	0.05±0.02
<sup>8</sup> He/ <sup>9</sup> Li (/day)	2.9±1.5		2.0±1.1		0.22±0.12	
Am-C corr. (/day)	0.2±0.2					
<sup>13</sup> C(α, n) <sup>16</sup> O (/day)	0.08±0.04	0.07±0.04	0.05±0.03	0.04±0.02	0.04±0.02	0.04±0.02
<b>Antineutrino rate (/day)</b>	<b>662.47 ±3.00</b>	<b>670.87 ±3.01</b>	<b>613.53 ±2.69</b>	<b>77.57 ±0.85</b>	<b>76.62 ±0.85</b>	<b>74.97 ±0.84</b>

**Consistent rates for side-by-side detectors**

Uncertainty currently dominated by statistics



# Summary of Uncertainties

	Detector		Uncorrelated
	Efficiency	Correlated	
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed energy cut	90.9%	0.6%	0.12%
Prompt energy cut	99.88%	0.10%	0.01%
Multiplicity cut		0.02%	<0.01%
Capture time cut	98.6%	0.12%	0.01%
Gd capture ratio	83.8%	0.8%	<0.1%
Spill-in	105.0%	1.5%	0.02%
Livetime	100.0%	0.002%	<0.01%
Combined	78.8%	1.9%	0.2%

**For near/far oscillation, only uncorrelated uncertainties are used.**

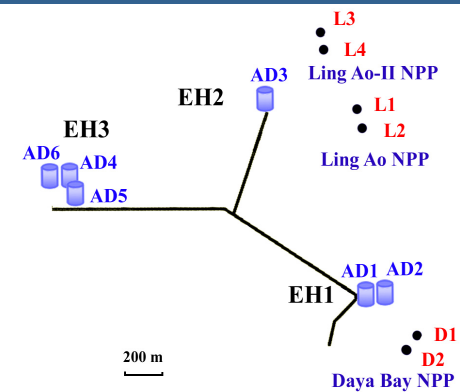
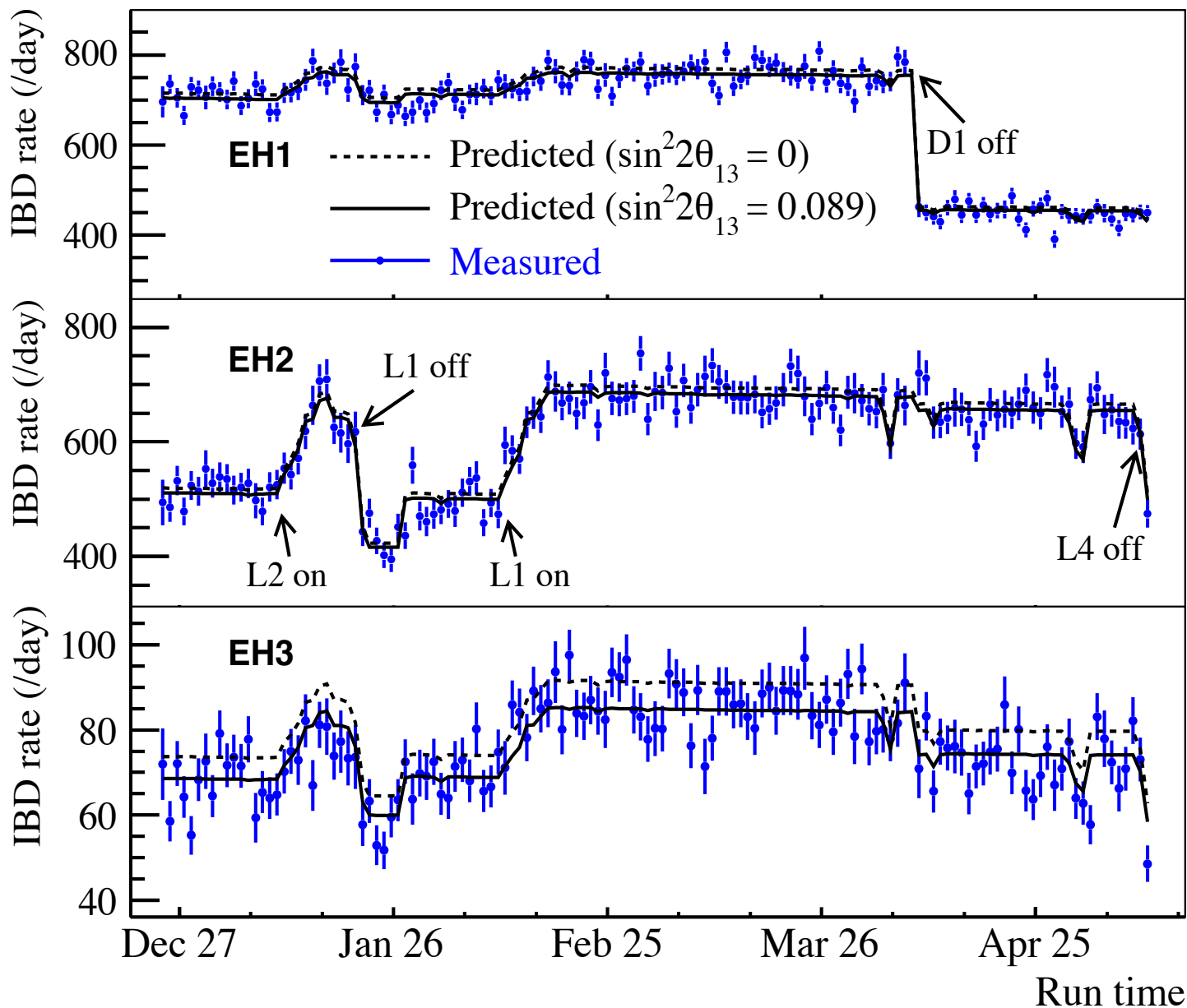
Largest systematics are smaller than far site statistics (~1%)

Reactor			
Correlated		Uncorrelated	
Energy/fission	0.2%	Power	0.5%
$\bar{\nu}_e$ /fission	3%	Fission fraction	0.6%
		Spent fuel	0.3%
Combined	3%	Combined	0.8%

Influence of uncorrelated reactor systematics reduced by far vs. near measurement.

# Latest Results

## ❖ Measured anti-neutrino event rates vs. time:



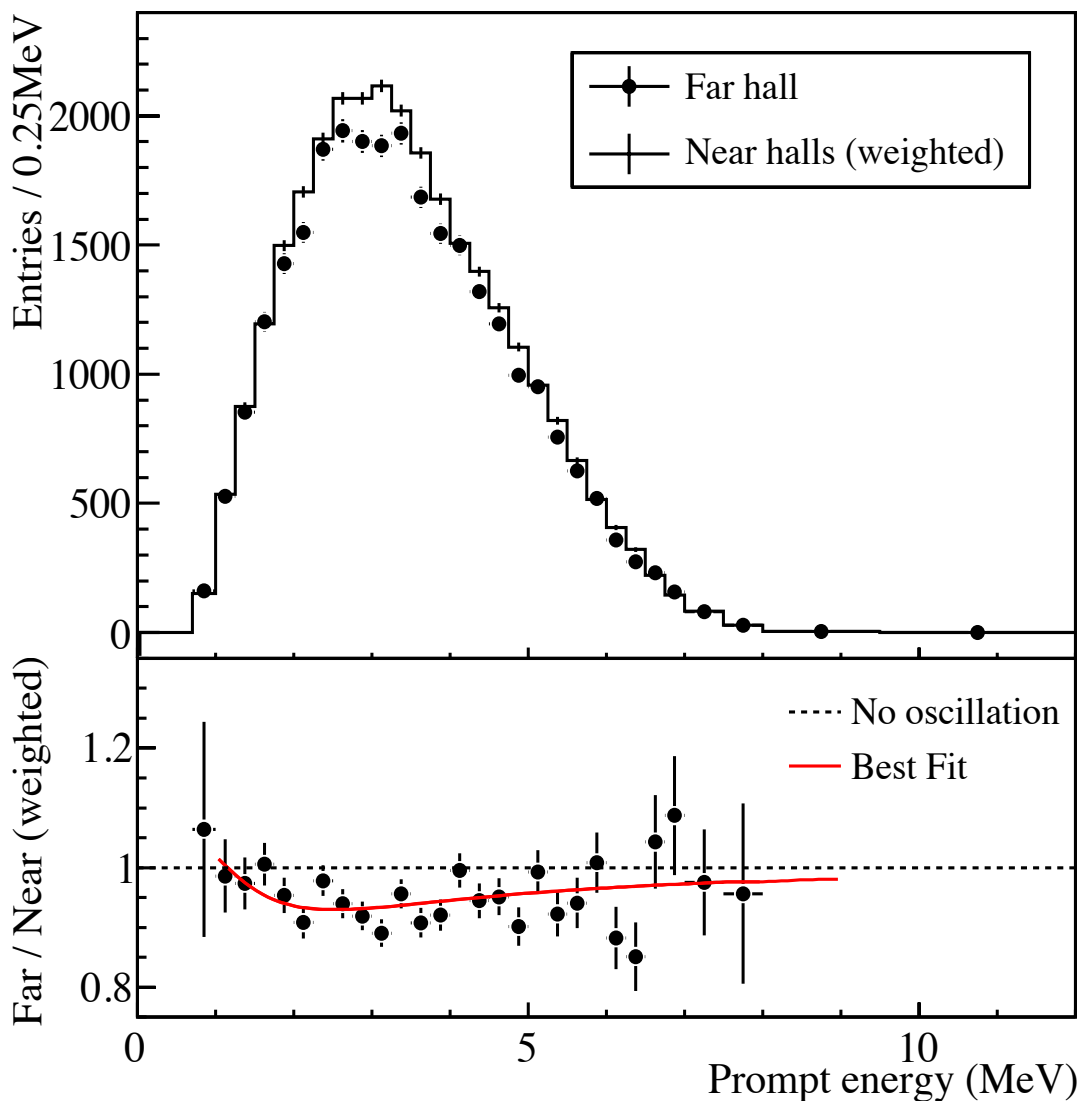
- DX and LX denote the DayaBay and LingAo reactors respectively
- The reactors come on and off due to refueling.
- Detector rates strongly correlated with reactor flux expectations.

### Notes:

- Normalization is determined by fit to near-hall data.
- Absolute normalization is within a few percent of expectations.

# Near/Far Comparison

Compare the far/near measured rates and spectra



$$R = \frac{Far_{measured}}{Far_{expected}} = \frac{M_4 + M_5 + M_6}{\sum_{i=4}^6 (\alpha_i(M_1 + M_2) + \beta_i M_3)}$$

$M_n$  are the measured rates in each detector.  
Weights  $\alpha_i, \beta_i$  are determined from baselines and reactor fluxes.

$$R = 0.944 \pm 0.007 \text{ (stat)} \pm 0.003 \text{ (syst)}$$

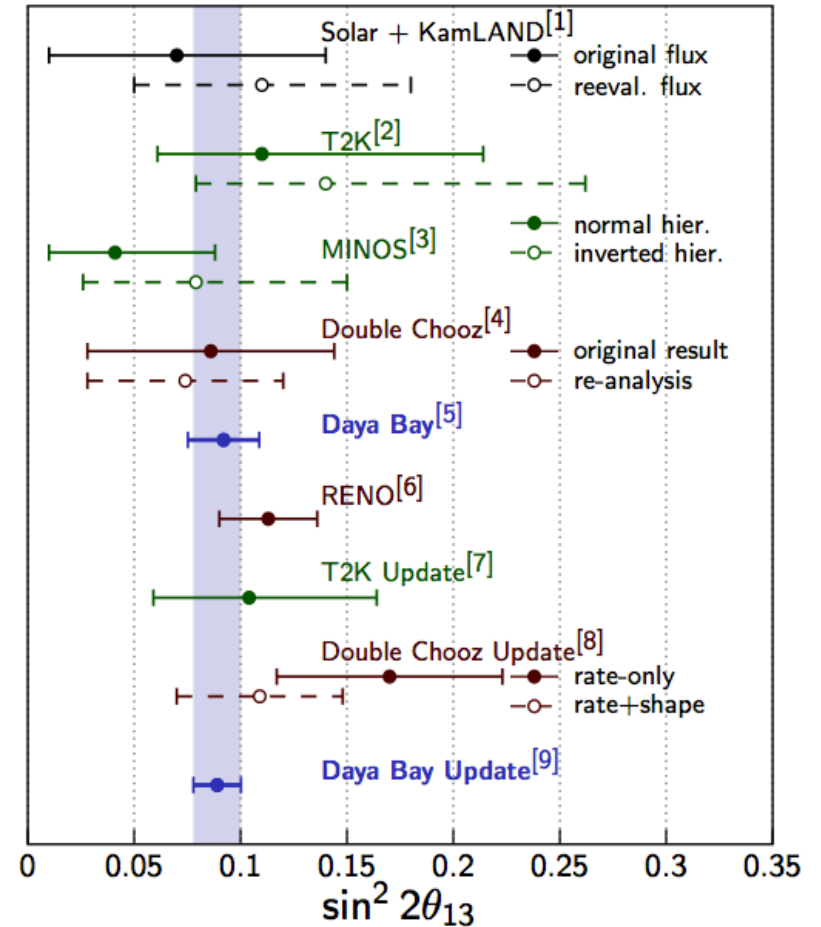
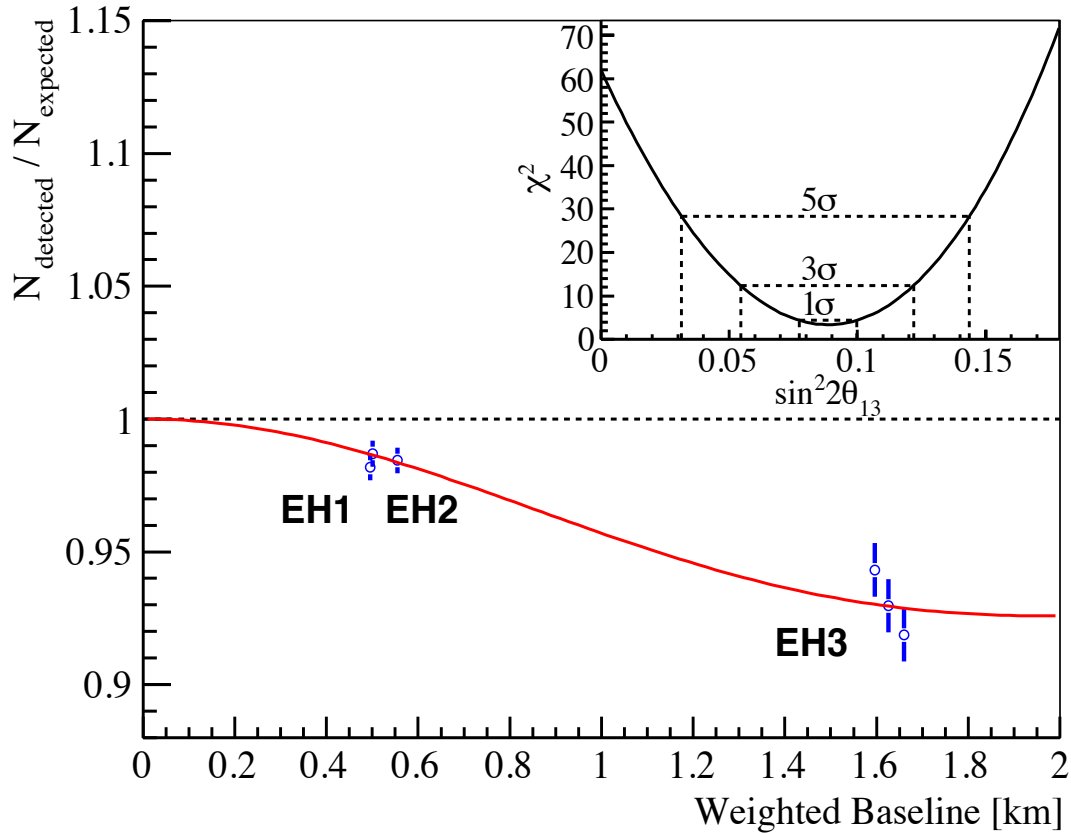
Clear observation of far site deficit.

Spectral distortion consistent with oscillation.\*

\* Caveat: Spectral systematics not fully studied;  
 $\theta_{13}$  value from shape analysis is not recommended.

# Oscillation Analysis

Estimate  $\theta_{13}$  using measured rates in each detector.



Uses standard  $\chi^2$  approach.

Far vs. near relative measurement (Absolute rate is not constrained)

Consistent results obtained by independent analyses, different reactor flux models.

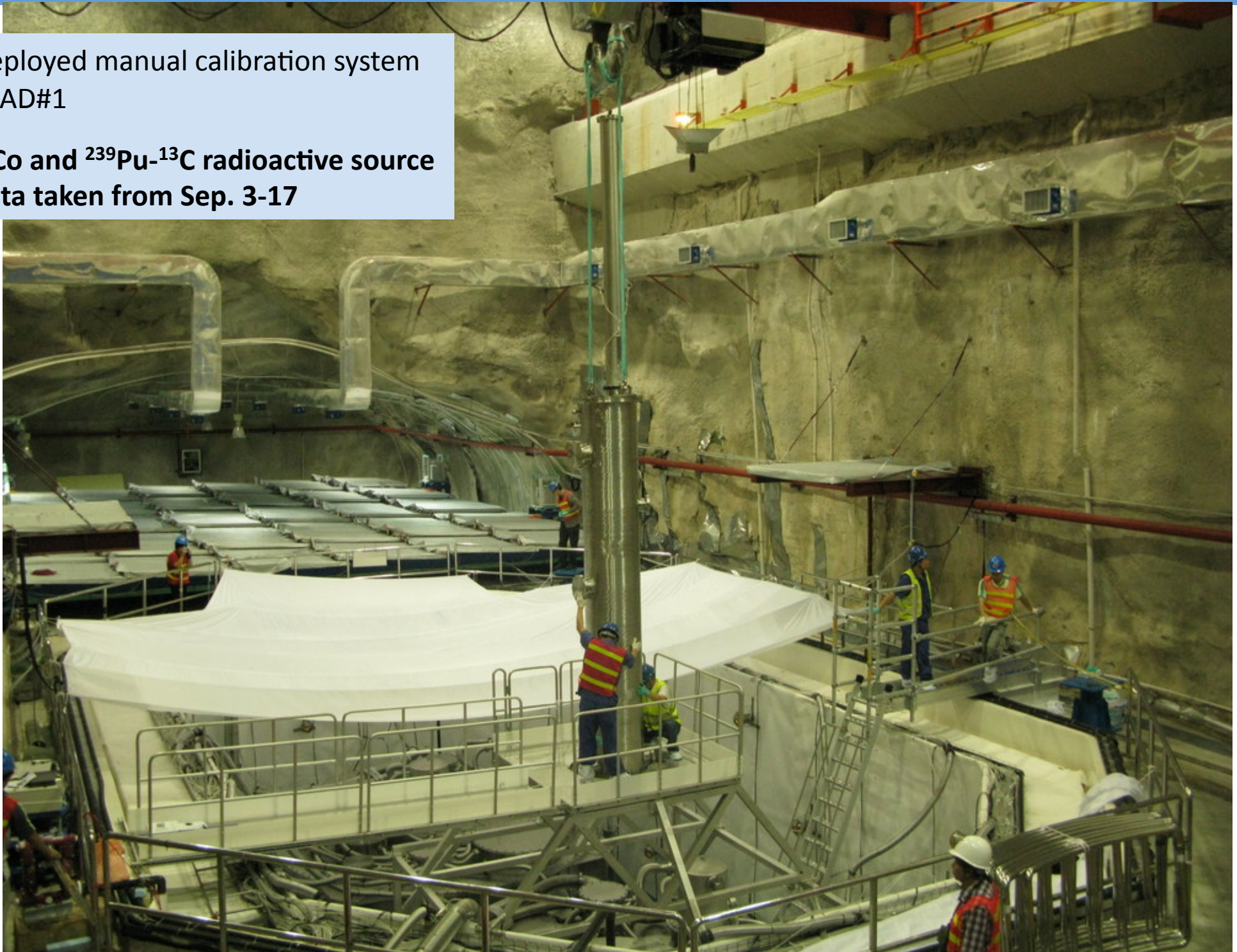
**Most precise measurement of  $\sin^2 2\theta_{13}$  to date.**

$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 \text{ (stat)} \pm 0.005 \text{ (syst)}$$

# Recent Progress

Deployed manual calibration system  
in AD#1

$^{60}\text{Co}$  and  $^{239}\text{Pu}$ - $^{13}\text{C}$  radioactive source  
data taken from Sep. 3-17



# Current Status

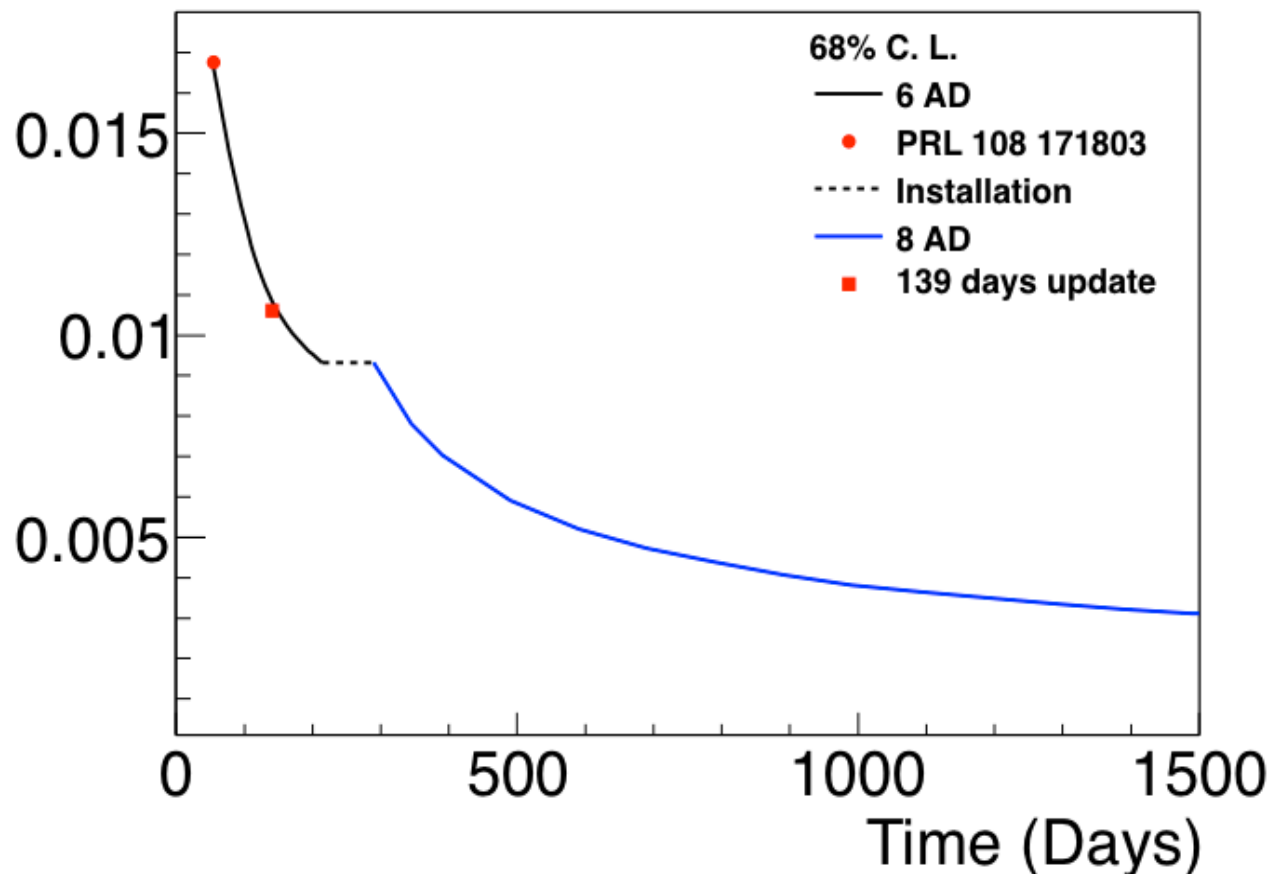
All eight detectors have been taking data since October 2012



# Prospects

- ❖ Prospects for  $\theta_{13}$  are very encouraging:
  - Current  $\theta_{13}$  uncertainty is still largely dominated by statistics
  - **Expect a ~5% measurement of  $\sin^2(2\theta_{13})$  in less than 3 years**
- ❖ Daya Bay will also make significant contributions in other areas:

Projected Daya Bay's Sensitivity of  $\sin^2 2\theta_{13}$



- Location of oscillation “dip” in energy spectrum constraints  $\Delta m^2_{ee}$  (i.e. combination of  $\Delta m^2_{31}$  and  $\Delta^2 m_{32}$ )
- High-statistics IBD samples from Daya Bay can yield most precise measurement of reactor anti-neutrino spectrum

→ in progress!

# Summary & Conclusions

- ❖ We observe an unambiguous observation of electron antineutrino disappearance at a distance of  $\sim 2\text{km}$  from the source:

$$R = 0.944 \pm 0.007 \text{ (stat)} \pm 0.003 \text{ (syst)}$$

- The interpretation in terms of neutrino oscillations rules out  $\theta_{13}=0$  at more than  $7\sigma$ :

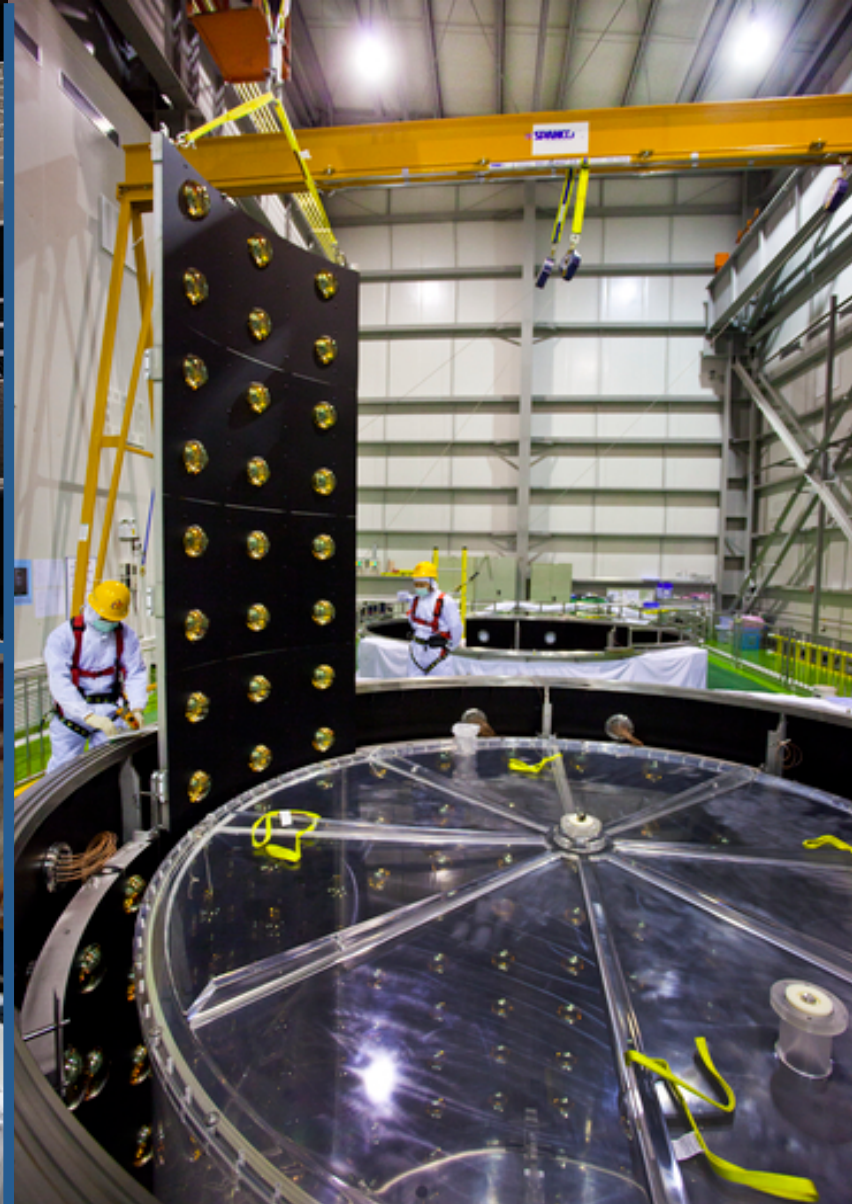
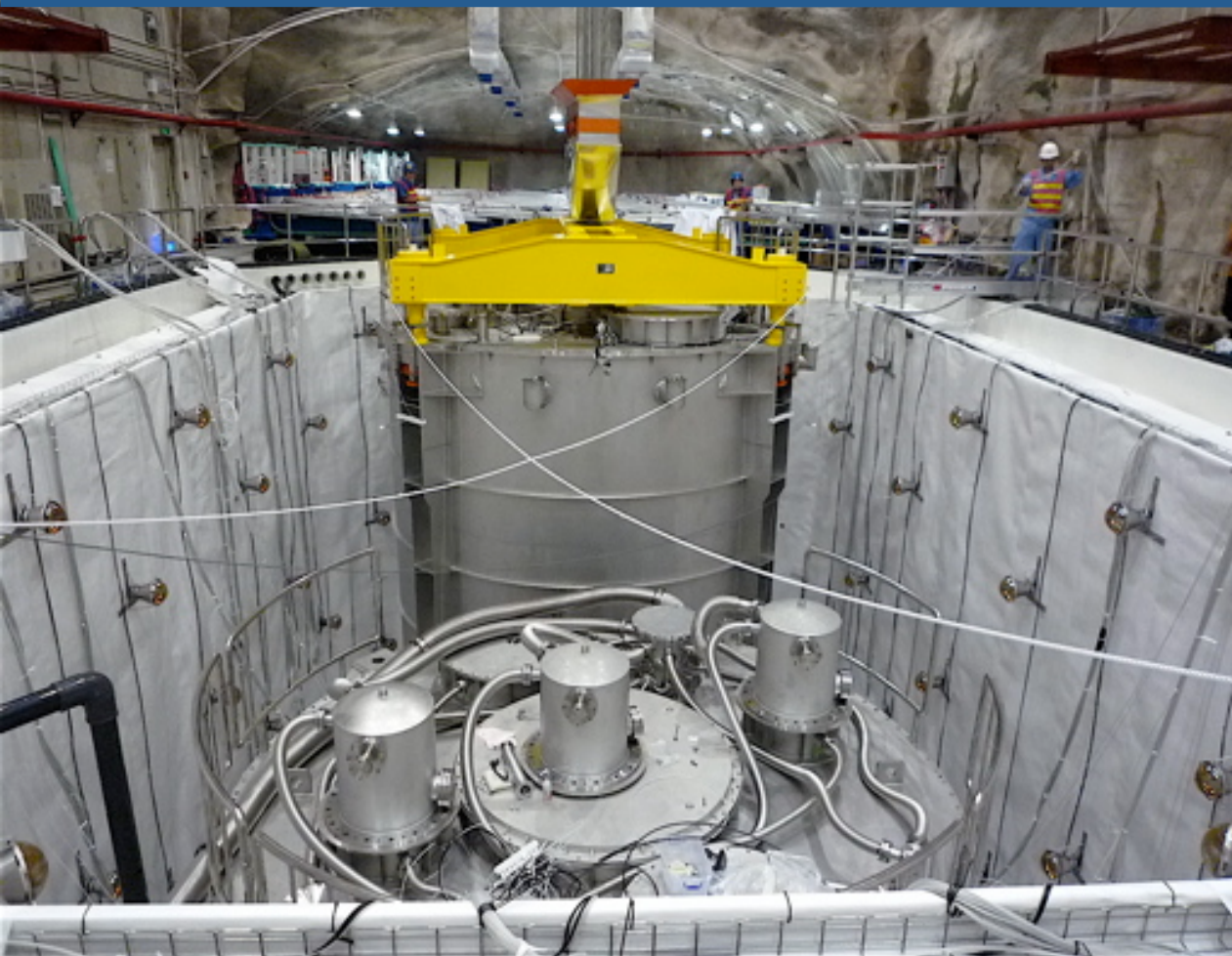
$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 \text{ (stat)} \pm 0.005 \text{ (syst)}$$

- The result has important consequences for particle physics and cosmology

- ❖ Future looks very positive for Daya Bay:

- Last two detectors already installed and commissioned
- Daya Bay will continue to have the best sensitivity to  $\theta_{13}$  among all the other experiments in operation or in construction.
- Other results expected soon, such as precise reactor flux and spectra measurements,  $\Delta m_{ee}^2$  measurement, and others.





Thank you for  
your attention!