



Probing neutralino properties in minimal supergravity with bilinear R-parity violation

Diego Restrepo¹

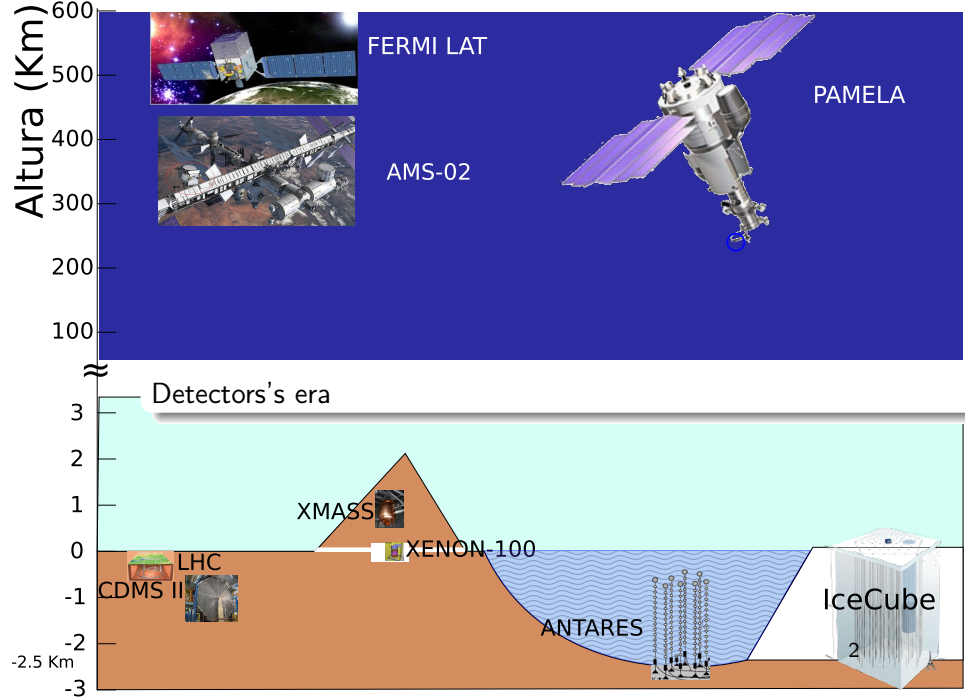
¹Instituto de Física
Universidad de Antioquia

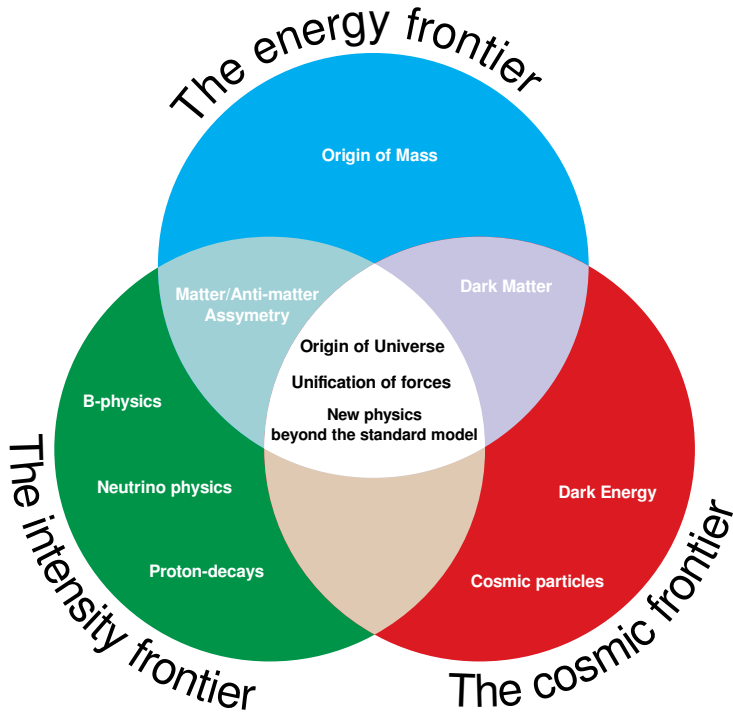


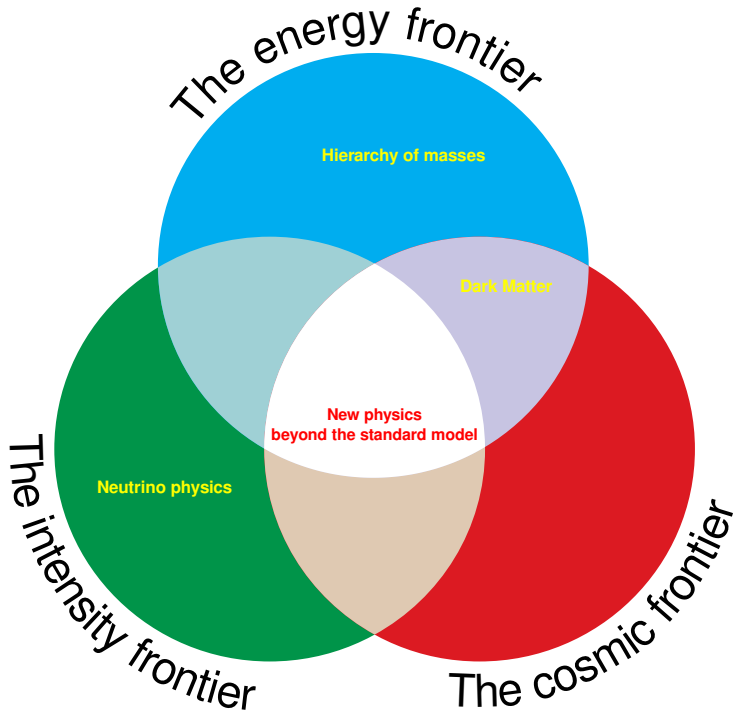
GRUPO DE
FENOMENOLOGÍA DE INTERACCIONES FUNDAMENTALES



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Bilinear R-parity Violation (BR_{pV})

Supersymmetric Model

$$W = W_{\text{MSSM}} + \epsilon_i L_i H_u$$

$$i = 1, 2, 3$$



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\downarrow
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Supersymmetric Model

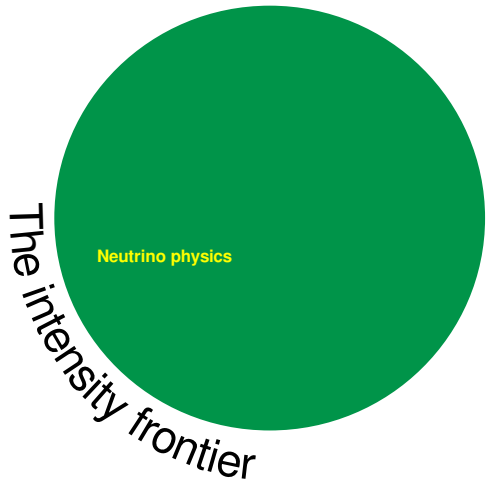
$$W = W_{\text{MSSM}} + \epsilon_i L_i H_u$$

$$i = 1, 2, 3$$

$$\downarrow \\ v_i$$

$$\frac{\Lambda_i}{\mu v_d} = \frac{\epsilon_i}{\mu} + \frac{v_i}{v_d}$$

6 new parameters

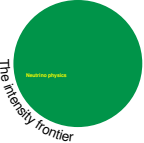


The intensity frontier

Neutrino physics

Parámetro	1σ
$\Delta m_{32}^2 [10^{-3} \text{eV}^2]$	$2.50^{+0.09}_{-0.16}$
$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$	$7.59^{+0.20}_{-0.18}$
$\sin^2 \theta_{23}$	$0.52^{+0.06}_{-0.07}$
$\sin^2 \theta_{12}$	$0.312^{+0.017}_{-0.015}$
$\sin^2 \theta_{13}$	$0.013^{+0.007}_{-0.005}$

Valle et al, arXiv:1108.1376



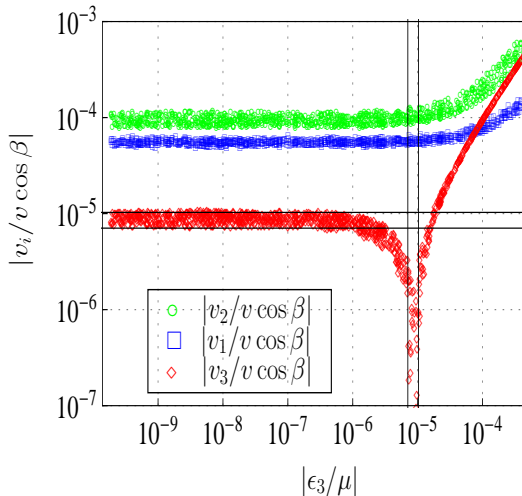
$$\Delta m_{32} = \frac{M_{\tilde{\gamma}}}{4|M_{\chi^0}^0|} |\Lambda|^2$$

$$\tan^2 \theta_{23}^{\text{BRpV}} = \frac{\Lambda_2^2}{\Lambda_3^2}$$

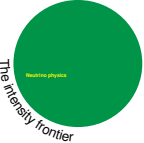
$$\Delta m_{21} = \frac{3}{8\pi^2} \sin 2\theta_{\tilde{b}} \frac{m_b^3}{v^2 c_\beta^2} \Delta B_0^{\tilde{b}_2 \tilde{b}_1} \times \frac{(\bar{\epsilon}_1^2 + \bar{\epsilon}_2^2)}{\mu^2}$$

$$\tan^2 \theta_{12} = \frac{\bar{\epsilon}_1^2}{\bar{\epsilon}_2^2}$$

$$\bar{\epsilon}_i = f_i(\Lambda_i, \epsilon_i)$$



Susy point with $\tilde{\nu}_\tau$ (N)LSP:
Aristizabal, D.R., Spinner, *in progress*



Fully implemented in SPheno [by W. Porod]

$$\Delta m_{32} = \frac{M_{\tilde{\gamma}}}{4|M_{\chi^0}^0|} |\Lambda|^2$$

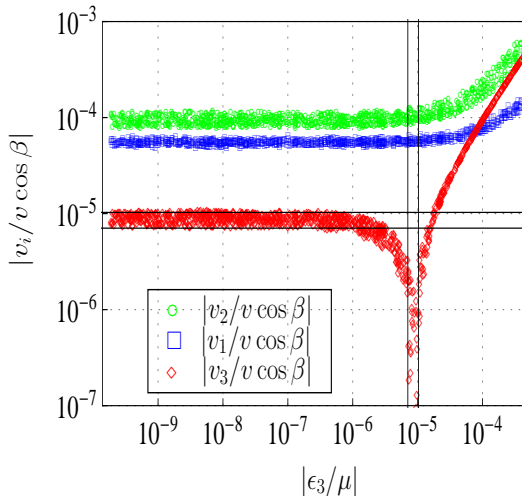
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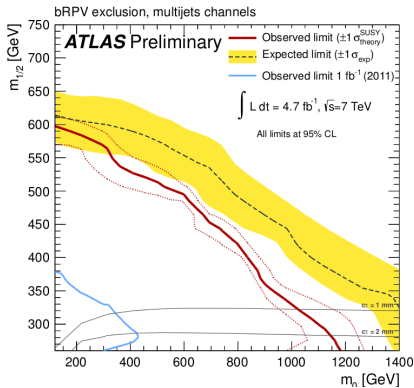
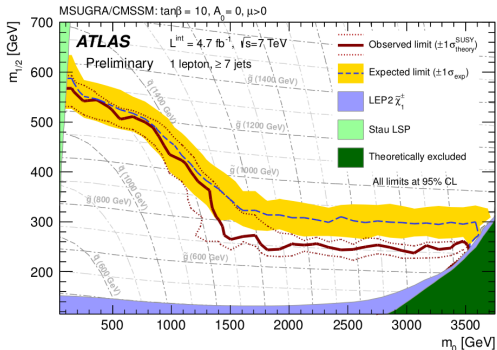
The energy frontier

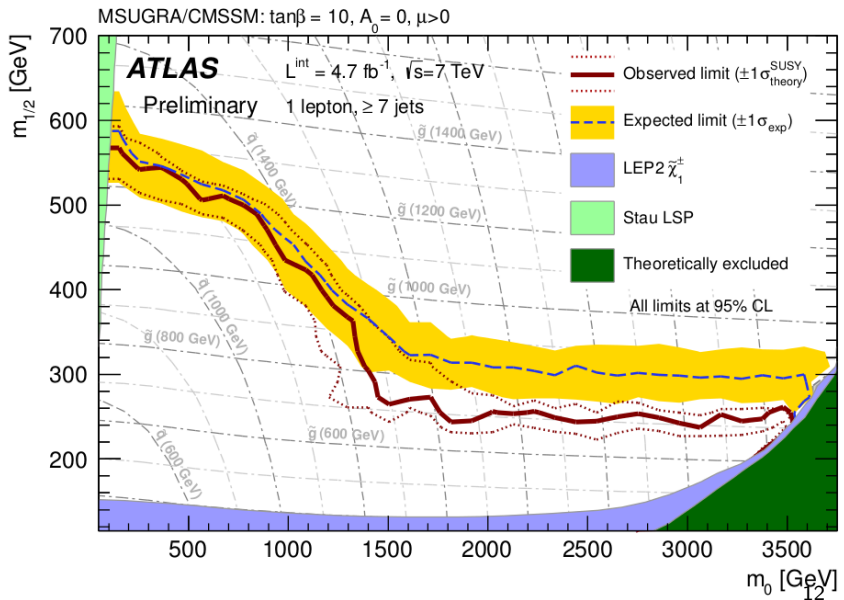
**New physics
beyond the standard model**

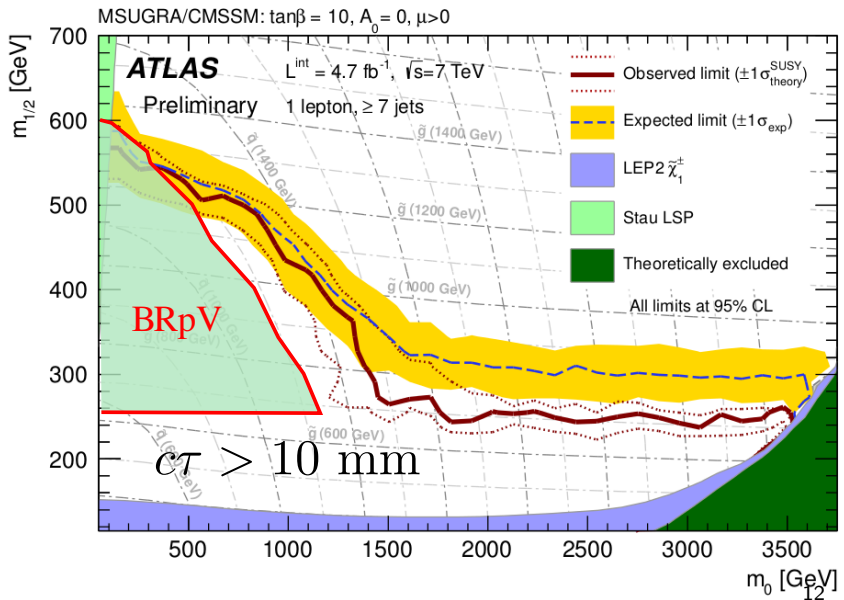


ATLAS-CONF-2012-140 (Oct): $\sqrt{s} = 7$ TeV,

$$\int L dt = 4.7 \text{ fb}^{-1}$$



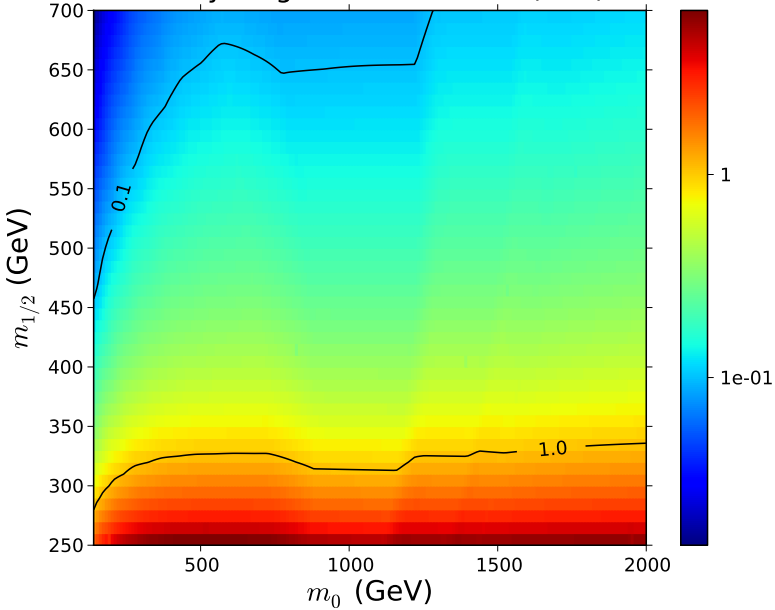






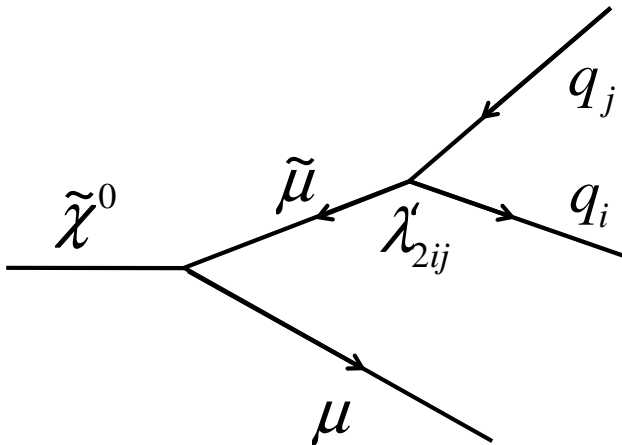
Neutralino decay length BRpV

Decay length without boost (mm)

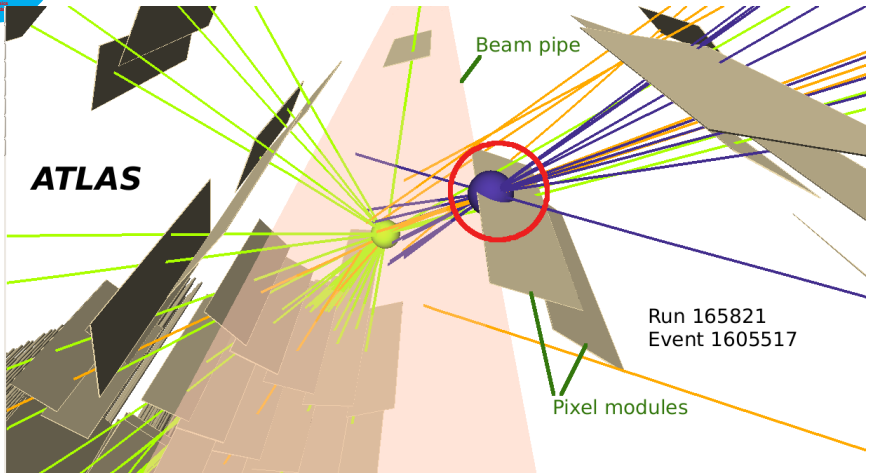




ATLAS displaced vertices [ATLAS-CONF-2012-103] (August)

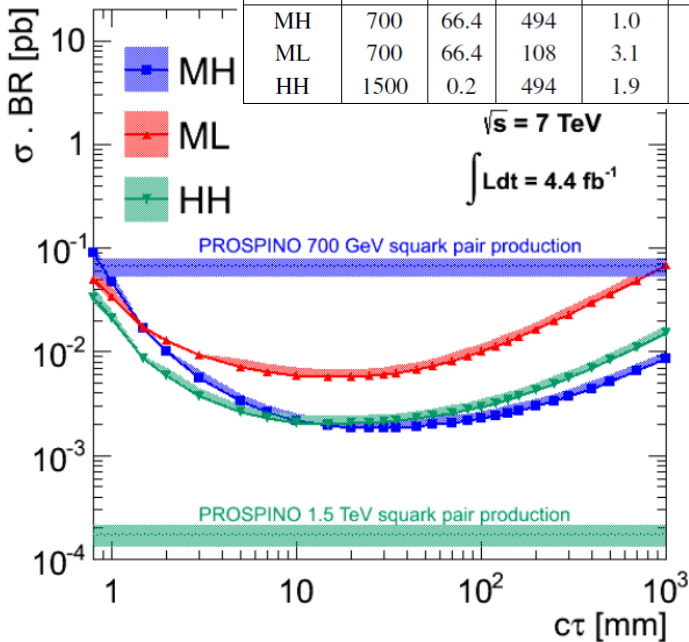


ATLAS displaced vertices [ATLAS-CONF-2012-103] (August)





Sample	$m_{\tilde{q}}$ [GeV]	σ [fb]	$m_{\tilde{\chi}_1^0}$ [GeV]	$\langle \gamma\beta \rangle_{\tilde{\chi}_1^0}$	$c\tau_{MC}$ [mm]	λ'_{211}
MH	700	66.4	494	1.0	78	3.0×10^{-6}
ML	700	66.4	108	3.1	101	1.5×10^{-5}
HH	1500	0.2	494	1.9	82	1.5×10^{-5}

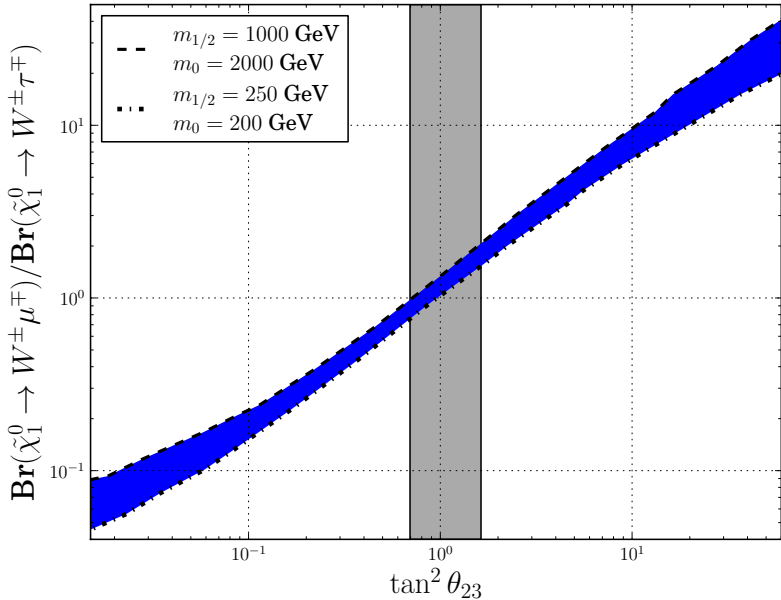


The energy frontier

The intensity frontier

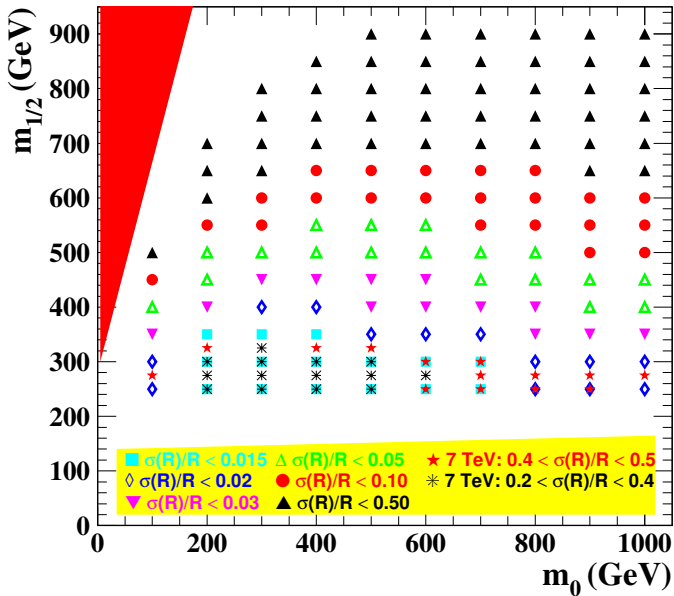
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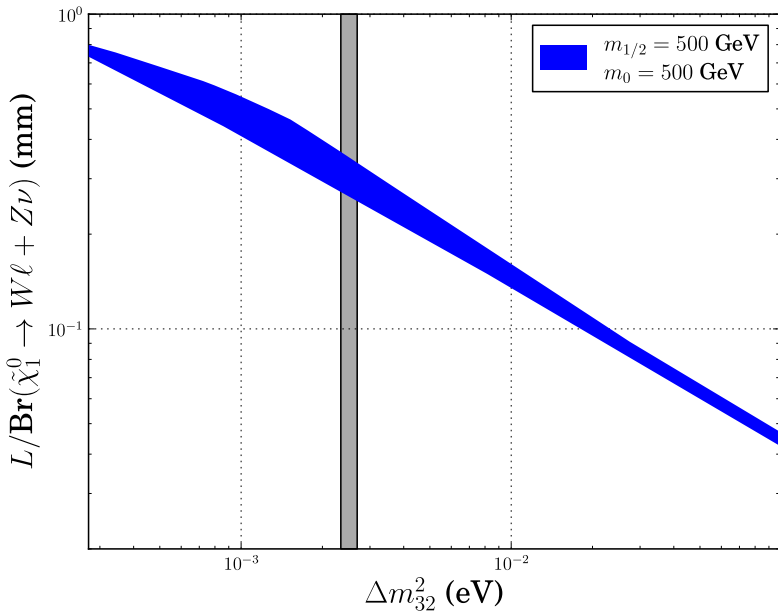


$$\sqrt{s} = 14\text{TeV} \int L dt = 100 \text{ fb}^{-1} \quad (A_0 = -100 \text{ GeV}, \tan \beta = 10, \text{sgn}(\mu)=+1)$$





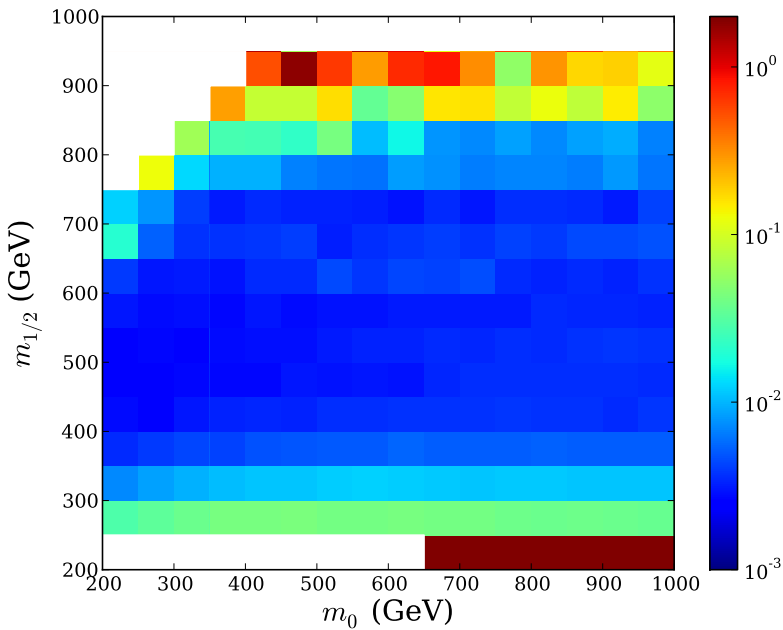
D.R., *et al* arXiv:1206.3605 [PRD]





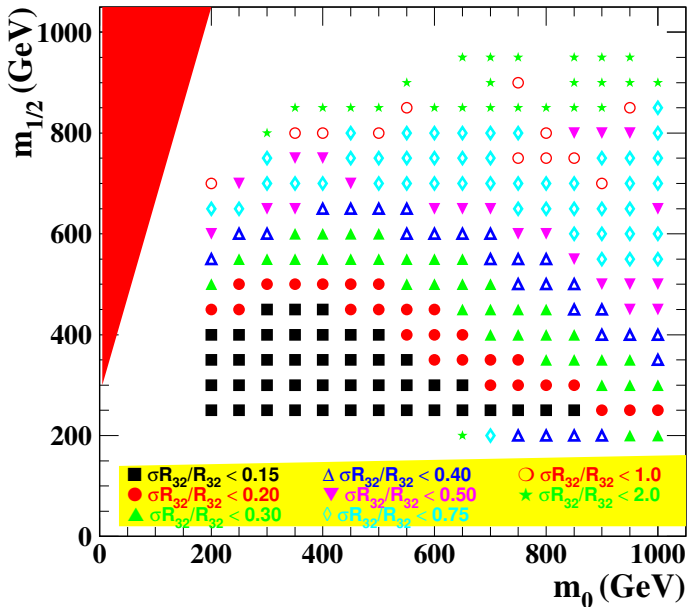
($\sqrt{s} = 14\text{TeV}$ $\int Ldt = 100 \text{ fb}^{-1}$ $A_0 = -100 \text{ GeV}$, $\tan \beta = 10$, $\text{sgn}(\mu)=+1$)

σ_{L_0}/L_0





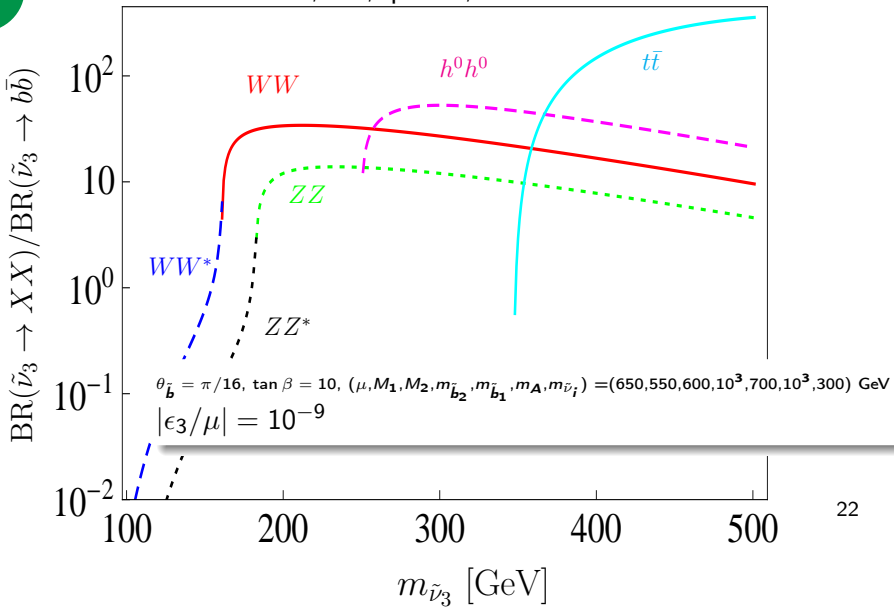
$$R_{32} = \frac{L_0}{\text{Br}(\tilde{\chi}_1^0 \rightarrow IW + Z\nu)} \approx \frac{1}{\Gamma(\tilde{\chi}_1^0 \rightarrow IW + Z\nu)}$$





$\tilde{\nu}$ (N)LSP

Aristizabal, D.R, Spinner, **PRELIMINARY**





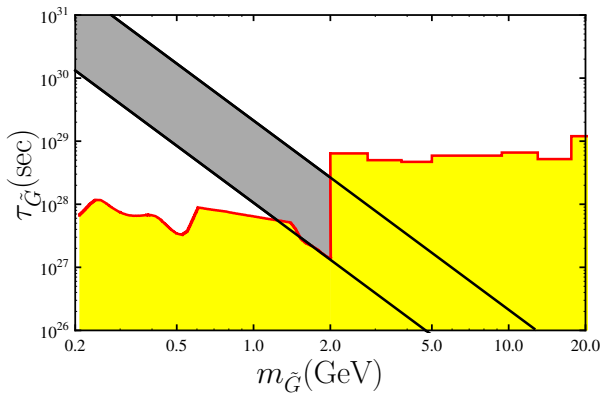
Dark Matter

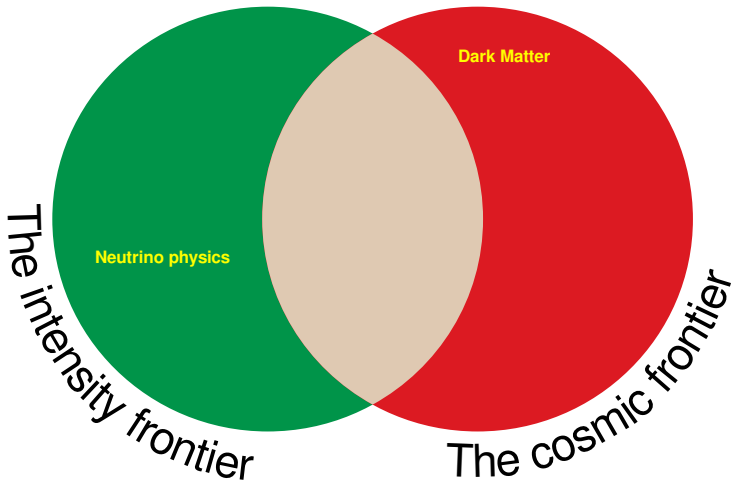
The cosmic frontier

Gravitino decay

$$\Gamma = \Gamma(\tilde{G} \rightarrow \sum_i \nu_i \gamma) \simeq \frac{1}{32\pi} |U_{\tilde{\gamma}\nu}|^2 \frac{m_{\tilde{G}}^3}{M_P^2}$$

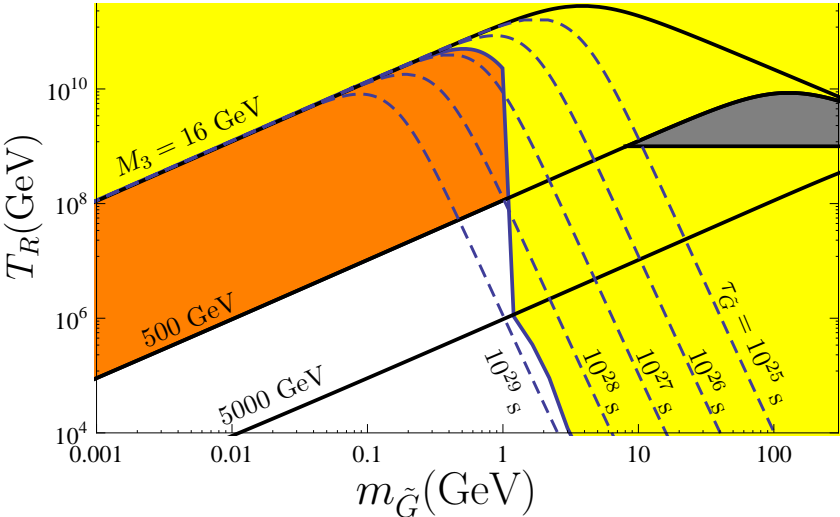
$$|U_{\tilde{\gamma}\nu}|^2 \simeq \frac{\mu^2 g^2 \sin^2 \theta_W}{4 |M_{\chi}^0|} (M_2 - M_1)^2 |\vec{\lambda}|^2,$$





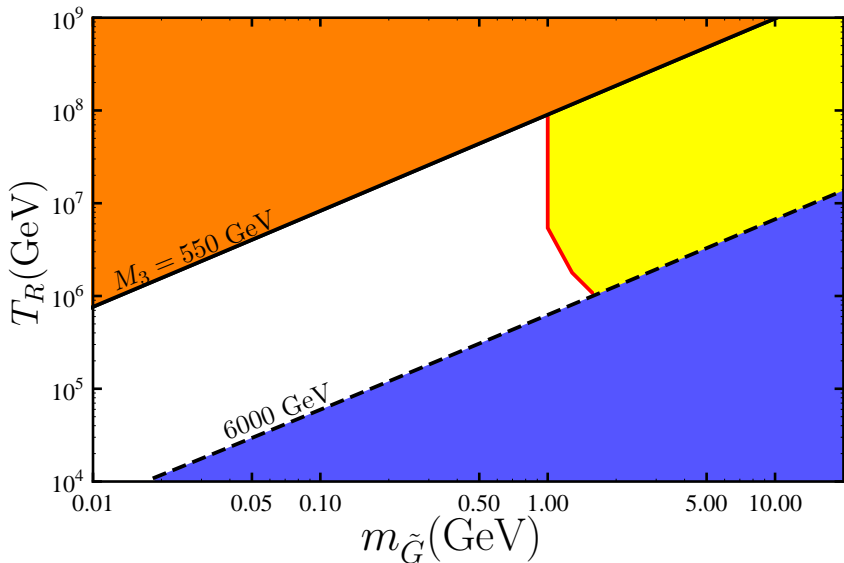
Gravitino DM & neutrino masses

D.R, Taoso, Valle, Zapata, 1109.0512 [PRD] $\implies m_{\tilde{G}} < 2 \text{ GeV}$

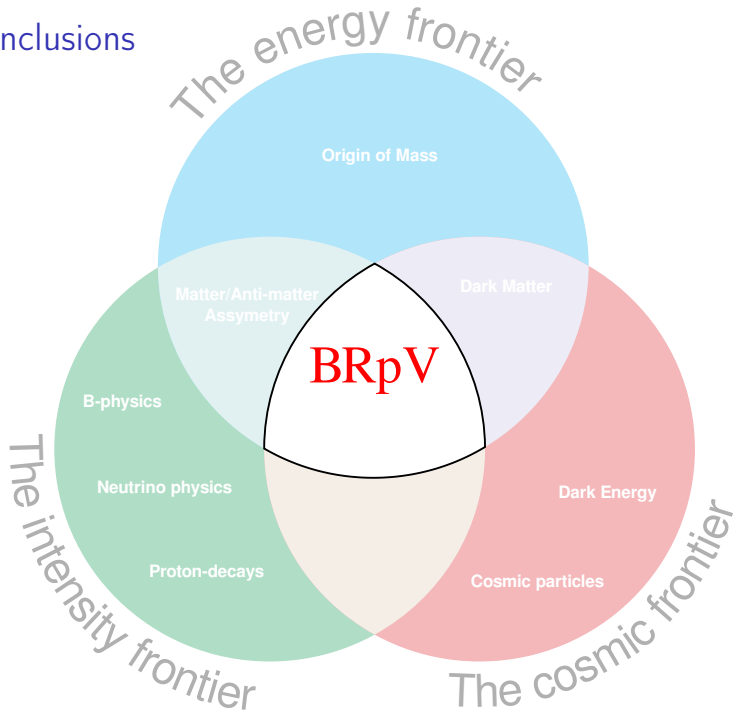


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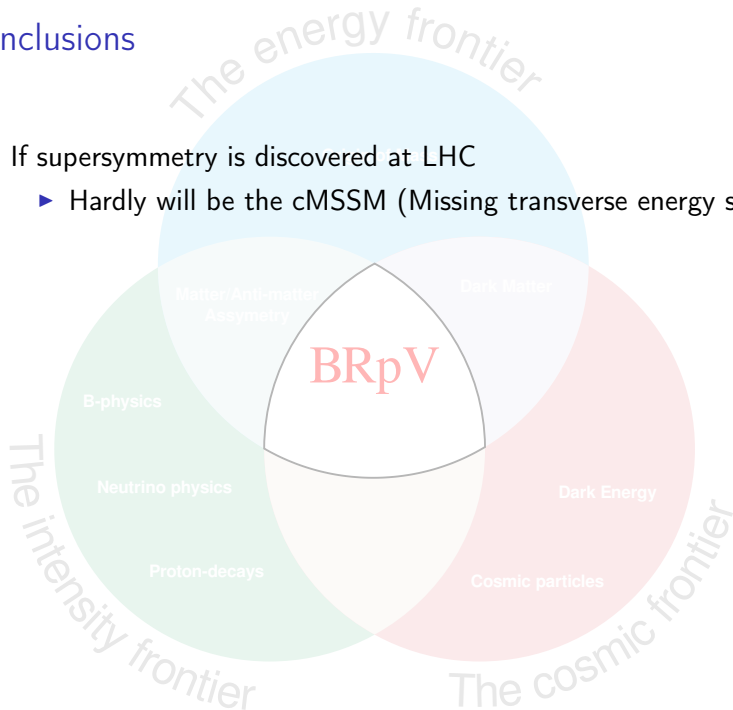
Conclusions



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If supersymmetry is discovered at LHC

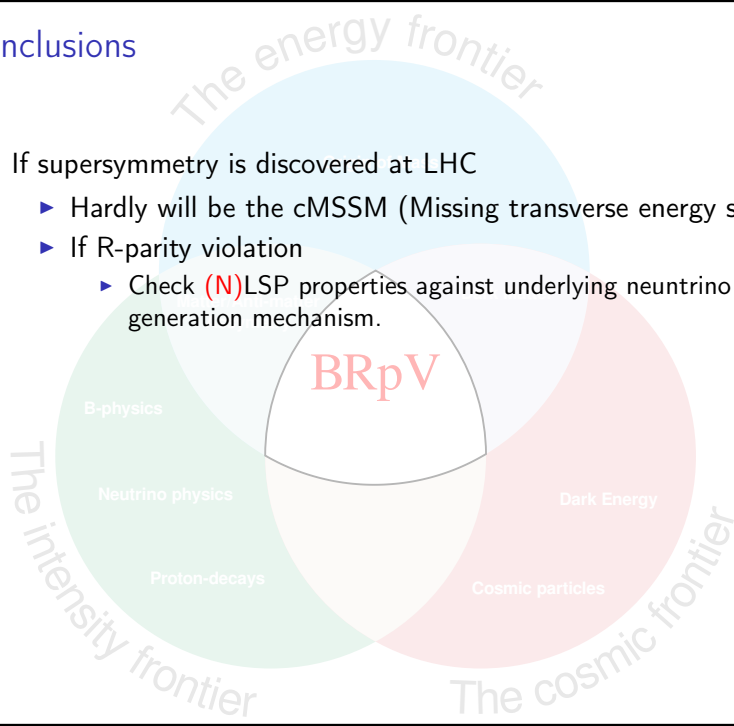
- ▶ Hardly will be the cMSSM (Missing transverse energy signal)



Conclusions

If supersymmetry is discovered at LHC

- ▶ Hardly will be the cMSSM (Missing transverse energy signal)
- ▶ If R-parity violation
 - ▶ Check (N)LSP properties against underlying neutrino mass generation mechanism.



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If supersymmetry is discovered at LHC

- ▶ Hardly will be the cMSSM (Missing transverse energy signal)
- ▶ If R-parity violation
 - ▶ Check (N)LSP properties against underlying neutrino mass generation mechanism.
 - ▶ Check against cosmic ray results:
 - ▶ PAMELA anomaly,
 - ▶ 130 GeV gamma-ray line, interpreted as Dark Matter,
 - ▶ Some Direct Dark Matter (DM) signal would exclude BRpV as the explanation for neutrino masses and mixings.

The energy frontier

The intensity frontier

The cosmic frontier

B-physics

Neutrino physics

Proton-decays

BRpV

Dark Energy

Cosmic particles