

# Scattering off of Black Holes, Isomonodromy and Painlevé

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# Importance of Scattering Theory

- Astrophysical phenomena: detection of gravitational waves
- Stability criteria of gravitational solutions
- AdS/CFT applications: quark-gluon plasma and condensed matter systems
- Quantum description of black holes

# Scalar Field Perturbation

- Non-minimally coupled massless scalar field  $\phi(x)$

$$(\nabla^2 + \xi R)\phi(x) = 0, \quad \nabla^2\phi \equiv \frac{1}{\sqrt{-g}}\partial_a(\sqrt{-g}g^{ab}\partial_b\phi)$$

- Radial and Angular equations for Kerr-NUT-(A)dS metric

$$\partial_r(P_r(r)\partial_r\phi_{\omega\ell m}) - Q_r(r)\phi_{\omega\ell m} = 0$$

$$\partial_\theta(P_\theta(\theta)\partial_\theta S_{\omega\ell m}) - Q_\theta(\theta)S_{\omega\ell m} = 0$$

- Angular eigenvalues from (A)dS-spheroidal harmonics

# Complex ODEs and Monodromy

- Radial equation

$$\partial_z(U(z)\partial_z\phi(z)) - V(z)\phi(z) = 0, \quad z \in \mathbb{CP}^1$$

- Ingoing and outgoing solutions

$$\phi_i^\pm(z) = (z - z_i)^{\pm\theta_i/2} (1 + \mathcal{O}(z - z_i))$$

- Singular points = Branch points  $\Rightarrow$  Monodromy

$$\phi_i^\pm(ze^{2\pi i}) = e^{\pm i\pi\theta_i} \phi_i^\pm(z)$$

# Monodromies and Gauge Connection

- Gauge connection formulation

$$(\partial_z - A(z))\Phi(z) = 0 ,$$

$$A(z) = \begin{pmatrix} 0 & U^{-1} \\ V & 0 \end{pmatrix} , \quad \Phi(z) = \begin{pmatrix} \phi_1 & \phi_2 \\ U\partial_z\phi_1 & U\partial_z\phi_2 \end{pmatrix}$$

- Monodromy matrix

$$\Phi_\gamma(z) = \mathcal{P} \exp \left( \oint_\gamma A \right) \Phi(z) =: \Phi(z) M_\gamma$$

# Monodromies and Frobenius solutions

- Loop around only one pole  $z = z_i \Rightarrow \Phi_{\gamma_i} = \Phi M_i$
- Loop enclosing all poles gives monodromy identity

$$M_1 M_2 \dots M_n = \mathbb{1}$$

- Monodromy matrix in arbitrary basis

$$M_i = g_i^{-1} \begin{pmatrix} e^{i\pi\theta_i} & 0 \\ 0 & e^{-i\pi\theta_i} \end{pmatrix} g_i$$

# Scattering Amplitudes and Connection Matrix

- Change of basis matrix = Connection matrix

$$\mathcal{M}_{i \rightarrow j} = \Phi_i^{-1} \Phi_j = g_i g_j^{-1}$$

- For purely imaginary  $\theta_i \notin i\mathbb{Z}$

$$\mathcal{M}_{i \rightarrow j} = \begin{pmatrix} \frac{1}{\mathcal{T}} & \frac{\mathcal{R}}{\mathcal{T}} \\ \frac{\mathcal{R}^*}{\mathcal{T}^*} & \frac{1}{\mathcal{T}^*} \end{pmatrix}, \quad |\mathcal{R}|^2 + |\mathcal{T}|^2 = 1$$

Castro *et al* arxiv:1304.3781

# Transmission between Two Regular Singular Points

Let  $g_i \in \mathrm{SL}(2, \mathbb{C})$  and one of the  $M_i$  be diagonal. If we write

$$m_{ij} = \mathrm{Tr} M_i M_j = 2 \cos \pi \sigma_{ij}$$

then

$$|\mathcal{T}|^2 = \frac{\sin \pi \theta_i \sin \pi \theta_j}{\sin \frac{\pi}{2} (\sigma_{ij} + \theta_i - \theta_j) \sin \frac{\pi}{2} (\sigma_{ij} - \theta_i + \theta_j)}$$

# $D = 4$ Kerr-(A)dS Black Hole

$$ds^2 = -\frac{Q(r)}{r^2 + p^2}(dt + p^2 d\phi)^2 + \frac{P(p)}{r^2 + p^2}(dt - r^2 d\phi)^2 \\ + \frac{r^2 + p^2}{Q(r)} dr^2 + \frac{r^2 + p^2}{P(p)} dp^2$$

$$P(p) = -\frac{\Lambda}{3}p^4 - \epsilon p^2 + k$$

$$Q(r) = -\frac{\Lambda}{3}r^4 + \epsilon r^2 - 2Mr + k$$

# Separation of Variables in KG equation

- Radial equation (5 regular singular points)

$$\partial_r(Q(r)\partial_r R(r)) + \left( -4\Lambda\xi r^2 + \frac{(\Psi_0 r^2 + \Psi_1)^2}{Q(r)} \right) R(r) = C_\ell R$$

- Conformally Coupled Case  $\xi = 1/6$

$$y'' + \left( \frac{1 - \theta_0}{z} + \frac{1 - \theta_1}{z - 1} + \frac{1 - \theta_{t_0}}{z - t_0} \right) y' + \left( \frac{1 + \theta_\infty}{z(z - 1)} - \frac{t_0(t_0 - 1)K_0}{z(z - 1)(z - t_0)} \right) y = 0$$

Heun equation (4 regular singular points)

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Heun equation (4 regular singular points)

How to find  $\sigma_{ij}$  ?

# Deformed Heun and Apparent Singularity

- *Deformed* Heun equation with one apparent singularity

$$\partial_z^2 y + \left( \frac{1 - \theta_0}{z} + \frac{1 - \theta_1}{z - 1} + \frac{1 - \theta_t}{z - t} - \frac{1}{z - \lambda} \right) \partial_z y + \left( \frac{\kappa}{z(z - 1)} - \frac{t(t - 1)K}{z(z - 1)(z - t)} + \frac{\lambda(\lambda - 1)\mu}{z(z - 1)(z - \lambda)} \right) y = 0$$

- Initial condition for our Heun

$$\lambda(t_0) = t_0, \quad \mu_0 = -\frac{K_0}{\theta_t}$$

and  $\theta_t \rightarrow \theta_t - 1$

# Isomonodromic Hamiltonian System

- $z = \lambda$  is an apparent singularity if

$$K(\lambda, \mu, t) = \frac{1}{t(t-1)} [\lambda(\lambda-1)(\lambda-t)\mu^2 - \{\theta_0(\lambda-1)(\lambda-t) + \theta_1\lambda(\lambda-t) + (\theta_t-1)\lambda(\lambda-1)\}\mu + \kappa(\lambda-t)]$$

- Hamiltonian System

$$\frac{d\lambda}{dt} = \frac{\partial K}{\partial \mu}, \quad \frac{d\mu}{dt} = -\frac{\partial K}{\partial \lambda}$$

generates isomonodromic flow  $(\lambda(t), \mu(t), K(\lambda, \mu, t))$

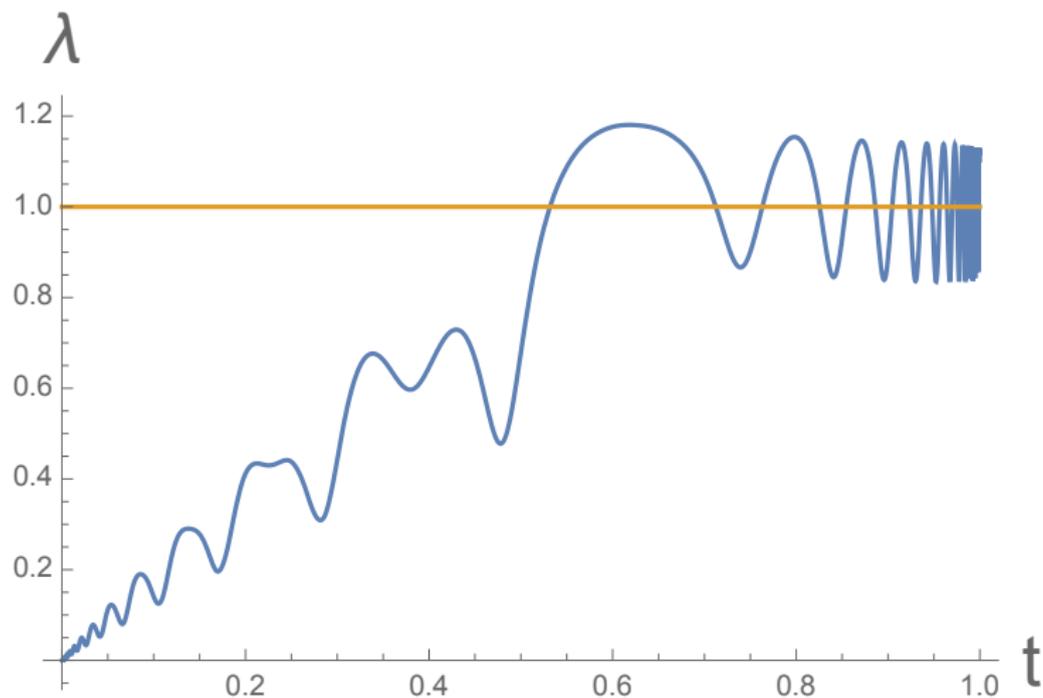
- Second-order equation for  $\lambda(t) = \text{Painlevé VI}$

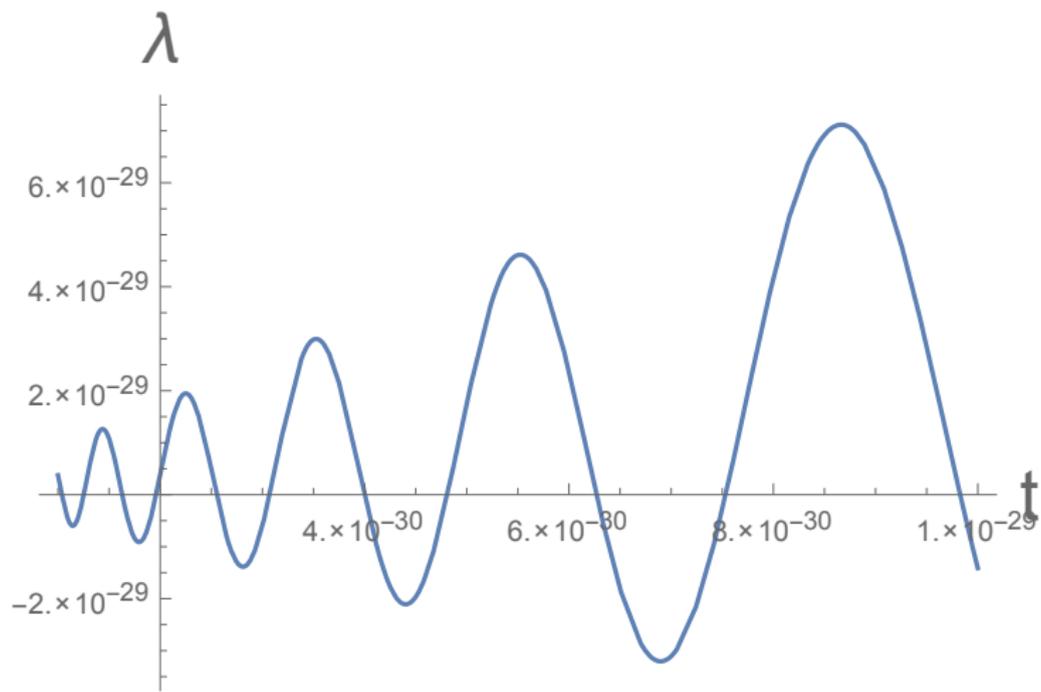
# Painlevé VI Asymptotics

- $P_{VI}$  asymptotics for  $0 < \operatorname{Re} \sigma_{ij} < 1$

$$\lambda(t) = \begin{cases} a_0 t^{1-\sigma_0 t} (1 + O(t^\delta)), & |t| < r, \\ 1 + a_1 (1-t)^{1-\sigma_1 t} (1 + O((1-t)^\delta)), & |t-1| < r, \\ a_\infty t^{\sigma_0 1} (1 + O(t^{-\delta})), & |1/t| < r, \end{cases}$$

where  $r, \delta > 0$  and  $a_i$  are functions of monodromy data

Numerical Integration of  $P_{VI}$ 

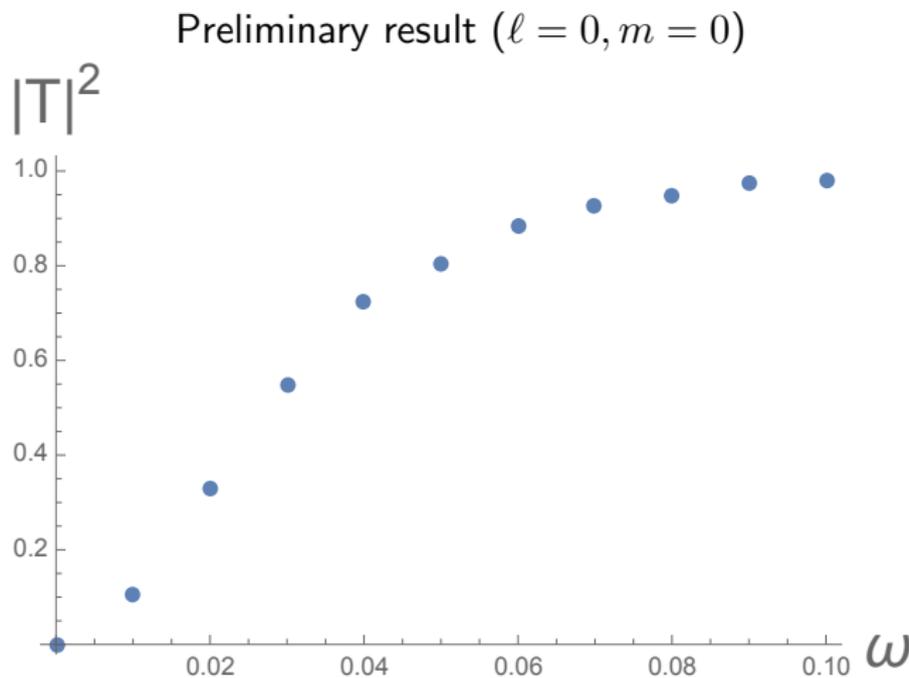
Numerical Integration of  $P_{VI}$  near  $t = 0$ 

# Kerr-dS Greybody Factor

$$\gamma_{\ell}(\omega, m) = \frac{\sinh\left(\frac{\omega - \Omega_H m}{2T_H}\right) \sinh\left(\frac{\omega - \Omega_C m}{2T_C}\right)}{\cosh\left(\frac{\omega - \Omega_H m}{2T_H} + \frac{\omega - \Omega_C m}{2T_C}\right) - \cosh(2\pi\nu_{HC})}$$

- $\nu_{HC}$  encodes all global information
- $\Omega_H m < \omega < \Omega_C m \Rightarrow$  *superradiance*

# Kerr-dS Greybody Factor



# Conclusions

- Monodromy technique is the most powerful way to treat scattering problems
- General formula for scattering amplitudes between two regular singular points
- Conformally coupled case is easier
- Valid for higher-dimensional Kerr-NUT-(A)dS black holes
- Flat case can be recovered by confluence

# Conclusions

- Monodromy technique is the most powerful way to treat scattering problems
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Thank you

# Perspectives

- Higher-spin modes and gravitational stability
- Higher-dimensional Kerr-(A)dS and SUGRA backgrounds
- Recover literature via  $\Lambda \rightarrow 0$  confluence. Irregular singular points ( $P_V$  and  $P_{III}$ )
- Quasinormal modes and plasma thermalization
- CFT dual of extremal Kerr-(A)dS conformal modes?
- Twistorial and geometrical interpretation of isomonodromic hidden symmetry

# Acknowledgments



# Hidden Symmetry of Isomonodromic Flow

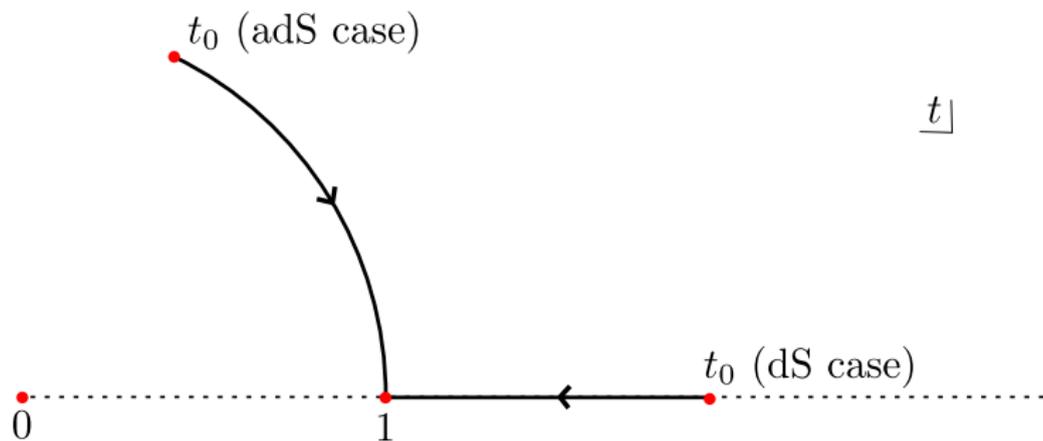
- Limit  $t \rightarrow 1$  of Garnier ODE gives hypergeometric equation

$$\partial_z^2 y + \left( \frac{1 - \theta_0}{z} + \frac{1 - \theta_1 - \tilde{\theta}_t}{z - 1} \right) \partial_z y + \left( \frac{\kappa}{z(z - 1)} + \frac{L_1}{z(z - 1)^2} \right) y = 0$$

- Fixed point of isomonodromic flow corresponds to some (near-horizon) extremal black hole

# Hidden Symmetry of Isomonodromic Flow

- This suggests that scattering data of non-extremal black holes is equivalent in some sense to extremal black hole scattering



# Relation with Fuchsian Equation

- Fuchsian ODE normal form with  $n$  finite singular points

$$\psi''(z) + T(z)\psi(z) = 0, \quad T(z) = \sum_{i=1}^n \left( \frac{\delta_i}{(z - z_i)^2} + \frac{c_i}{z - z_i} \right),$$

$$\sum_{i=1}^n c_i = 0, \quad \sum_{i=1}^n (c_i z_i + \delta_i) = 0, \quad \sum_{i=1}^n (c_i z_i^2 + 2\delta_i z_i) = 0$$

- Local monodromies:  $\delta_i = (1 - \theta_i^2)/4$
- Accessory parameters  $c_i$  have global properties
- $2(n - 3)$  independent parameters:  $(c_i, z_i)$

# Symplectic Structure of Flat $SL(2, \mathbb{C})$ Connections

- Moduli space of flat connections  $A \sim$  moduli space of monodromy group
- Atiyah-Bott symplectic structure

$$\Omega = \sum_{i=1}^{n-3} dc_i \wedge dz_i = \sum_{i=1}^{n-3} d\nu_i \wedge d\mu_i$$

where  $(\nu_i, \mu_i)$  are trace coordinates (Nekrasov et al 2011)

- Canonical transformation connects both set of coordinates
- Suggests analytical approach to find composite monodromies
- Relation with classical conformal blocks of 2D CFT

# Recurrence Relations

- Taylor solution  $y(z) = \sum_{n=0}^{\infty} g_n z^{n/2}$ ,  $|z| < 1$

$$-(Q_0 + q)g_0 + R_0g_1 = 0,$$

$$P_n g_{n-1} - (Q_n + q)g_n + R_n g_{n+1} = 0, \quad (n > 0)$$

$$P_n = (n - 1 + \alpha_+)(n - 1 + \alpha_-),$$

$$Q_n = n((t + 1)(n - 1 + \gamma) + t\delta + \epsilon),$$

$$R_n = t(n + 1)(n + \gamma)$$

- Solved using Leaver's continued-fraction method (Leaver 1985, Berti, Cardoso and Will (2006))

# Continued-fraction Method

- Augmented convergence for  $|z| \geq 1$  if

$$\lim_{n \rightarrow \infty} \left| \frac{g_{n+1}}{g_n} \right| = |t|^{-1} = \hat{a}^2 \Rightarrow a < L$$

- Recurrence relation in terms of  $v_n = g_{n+1}/g_n$

$$v_{n-1} = \frac{P_n}{(Q_n + q) - R_n v_n}$$

- Equivalent to continued-fraction

$$(Q_0 + q) - \frac{R_0 P_1}{(Q_1 + q) -} \frac{R_1 P_2}{(Q_2 + q) -} \dots = 0$$

- Solve numerically with  $v_N = \hat{a}^2$  for some large integer  $N$

# Schlesinger System Asymptotics

- Near  $t = 0$

$$A_0 \approx t^\Lambda A_0^0 t^{-\Lambda} \quad \text{and} \quad A_t \approx t^\Lambda A_t^0 t^{-\Lambda}, \quad \text{where } \Lambda = A_0^0 + A_t^0$$

- Schlesinger system degenerates to two hypergeometric connections

$$\frac{dY_0}{dz} = \left( \frac{\Lambda}{z} + \frac{A_1^0}{z-1} \right) Y_0, \quad \frac{dY_1}{dz} = \left( \frac{A_0^0}{z} + \frac{A_t^0}{z-1} \right) Y_1$$

# Schlesinger System Asymptotics

- Using that  $\det A_i^0 = -\theta_i^2/4$  and  $\det \Lambda = -\sigma_{0t}^2/4$

$$\Lambda + \frac{1}{2}\sigma\mathbb{1} = \frac{1}{4\theta_\infty} \begin{pmatrix} (-\theta_\infty - \theta_1 + \sigma)(\theta_\infty - \theta_1 - \sigma) & (-\theta_\infty - \theta_1 + \sigma)(\theta_\infty + \theta_1 + \sigma) \\ (\theta_\infty - \theta_1 + \sigma)(\theta_\infty - \theta_1 - \sigma) & (\theta_\infty - \theta_1 + \sigma)(\theta_\infty + \theta_1 + \sigma) \end{pmatrix}$$

$$A_1^0 + \frac{1}{2}\theta_1\mathbb{1} = \frac{1}{4\theta_\infty} \begin{pmatrix} -(\theta_\infty - \theta_1)^2 + \sigma^2 & (\theta_\infty + \theta_1)^2 - \sigma^2 \\ -(\theta_\infty - \theta_1)^2 + \sigma^2 & (\theta_\infty + \theta_1)^2 - \sigma^2 \end{pmatrix}$$

$$A_0^0 + \frac{1}{2}\theta_0\mathbb{I} = G_1 \frac{1}{4\sigma} \begin{pmatrix} (\theta_0 - \theta_t + \sigma)(\theta_0 + \theta_t + \sigma) & (\theta_0 - \theta_t + \sigma)(-\theta_0 - \theta_t + \sigma) \\ (\theta_0 - \theta_t - \sigma)(\theta_0 + \theta_t + \sigma) & (\theta_0 - \theta_t - \sigma)(-\theta_0 - \theta_t + \sigma) \end{pmatrix} G_1^{-1}$$

$$A_t^0 + \frac{1}{2}\theta_t\mathbb{I} = G_1 \frac{1}{4\sigma} \begin{pmatrix} (\theta_t + \sigma)^2 - \theta_0 & -(\theta_t - \sigma)^2 + \theta_0^2 \\ (\theta_t + \sigma)^2 - \theta_0 & -(\theta_t - \sigma)^2 + \theta_0^2 \end{pmatrix} G_1^{-1}.$$

# Quasinormal Modes

- Modes purely ingoing at  $r_H$  and purely outgoing at  $r_C$
- Possible only for complex  $\omega$
- In this case,

$$\mathcal{M}_{C \rightarrow H} = \begin{pmatrix} \frac{1}{\mathcal{T}} & \frac{\mathcal{R}}{\mathcal{T}} \\ \frac{\mathcal{R}'}{\mathcal{T}'} & \frac{1}{\mathcal{T}'} \end{pmatrix} = \begin{pmatrix} 0 & -1 \\ 1 & \frac{1}{\mathcal{T}'} \end{pmatrix}$$

- Poles of transcendental equation

$$\nu_{HC}(\omega, \ell, m) = \frac{\omega - \Omega_H m}{2T_H} + \frac{\omega - \Omega_C m}{2T_C} + 2\pi i n, \quad n \in \mathbb{Z}$$

# Properties of Greybody Factor

- *Conjectured* scattering regimes

$$\begin{cases} \omega > \Omega_H m & \text{or} & \Omega_C m > \omega & \text{Normal scattering} \\ \Omega_H m > \omega > \Omega_C m & & & \text{Superradiant scattering} \end{cases}$$

- Poles of scattering matrix (resonances)

$$\omega = \begin{cases} m\Omega_H - 2\pi i n T_H \\ m\Omega_C + 2\pi i n T_C \end{cases} \quad (n \in \mathbb{Z}^+)$$

- We expect that

$$\gamma_l(\omega) \rightarrow 1, \quad \text{as } \omega \rightarrow \infty$$

$$\gamma_l(\omega) \rightarrow 0 \text{ or constant} \quad \text{as } \omega \rightarrow 0$$

# Superradiant Scattering

- Superradiance = wave analog of Penrose process
- In terms of the classical impact parameter  $b = \mathcal{L}/\mathcal{E} \sim \ell/\omega$

$$\frac{\omega}{m} = \frac{\omega \ell}{\ell m} \sim \frac{1}{b} \frac{\mathcal{L}}{\mathcal{L}_z}$$

- Problem: Greybody factor pole even for real  $\omega$  in the superradiant range! Meaning?
- Perturbative analysis of  $\nu_{HC}$  might shed some light