

# Constraining Long-Lived Sparticles Using Simplified Models

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*UFABC - Santo André*

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*ICTP-SAIFR - October 29th, 2015*

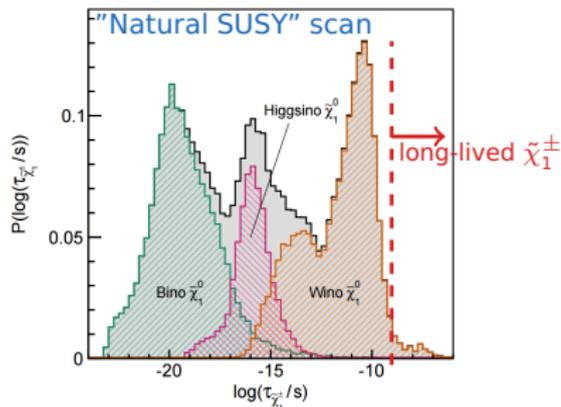
\*Work done in collaboration with J. Heisig and L. Quertenmoint

- Why long-lived charged sparticles?
  - ▶ Cosmological constraints and motivations
- Overview of experimental searches (CMS)
- Simplified models for HSCPs
- Application to the CMSSM
- Conclusions

# Why Long-Lived Charged (S)Particles?

- In SUSY, heavy stable charged particles (HSCPs) typically appear when:

$m_{NLSP} \simeq m_{LSP}$   
(degenerate spectra)

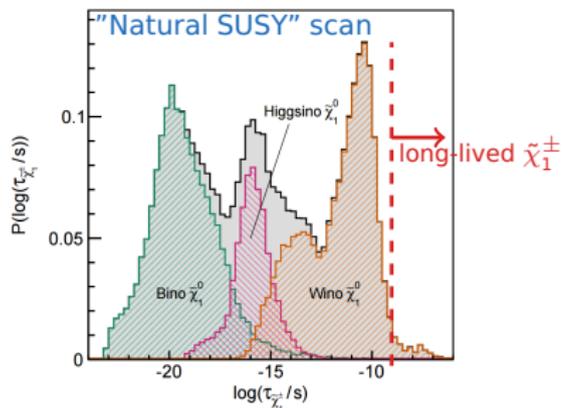


N.-E. Bomark et al., JHEP 05, 007 (2014)

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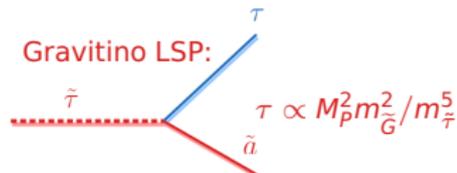
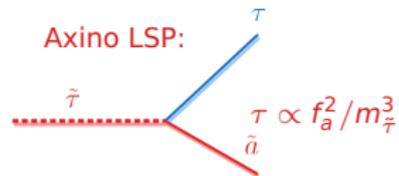
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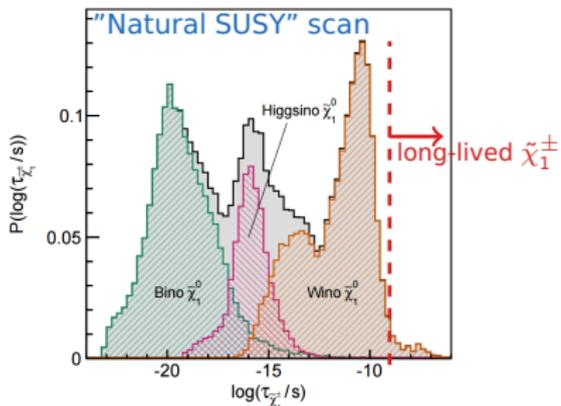
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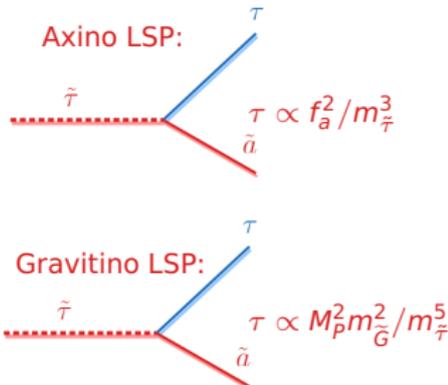
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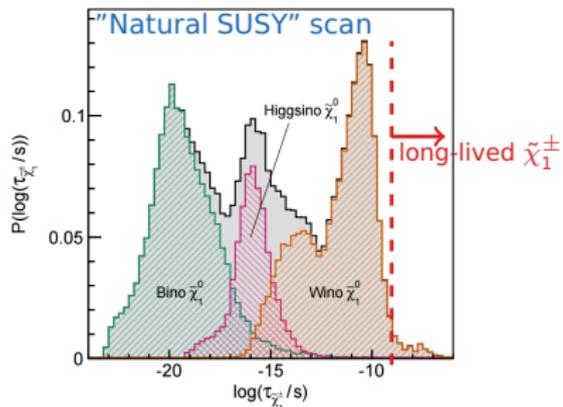


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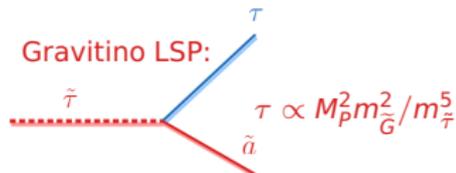
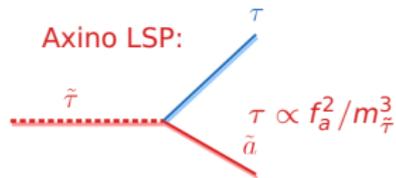
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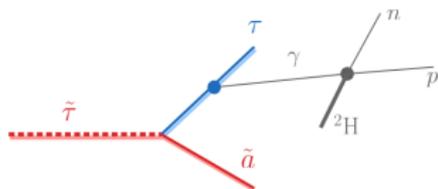
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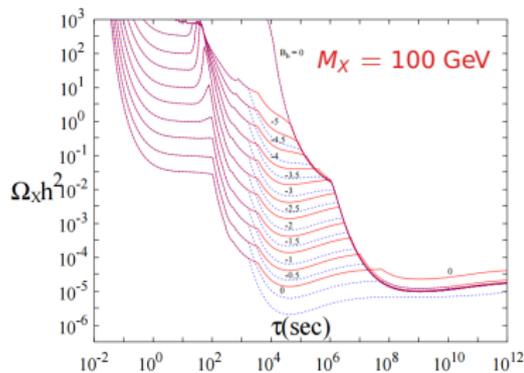
- Most of these scenarios are DM-motivated
- $\tau \gtrsim 1 - 10$  ns  $\rightarrow$  long-lived

# HSCPs and Cosmology: Big-Bang Nucleosynthesis

- BBN constraints:



(Depletion of light elements)

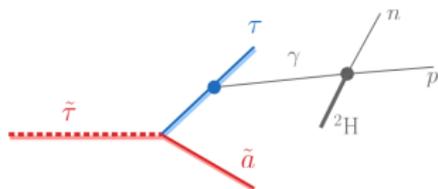


K. Jedamzik, Phys. Rev. D74, 103509, 2006

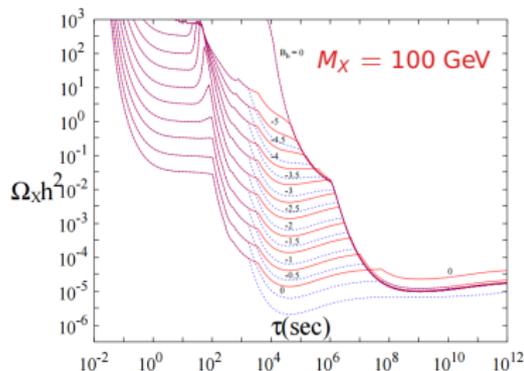
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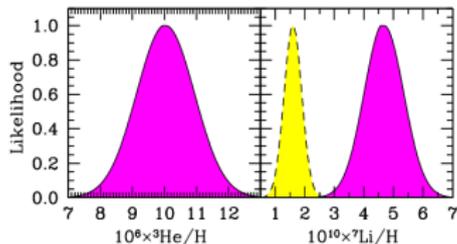
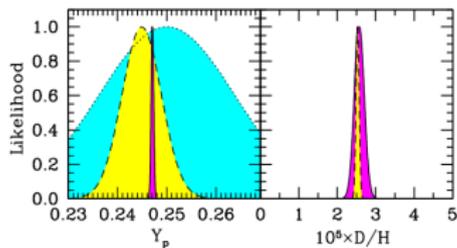
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- SBBN predictions:



B. Cyburt, B. Fields, K. Olive and T.-H. Yeh, arXiv:1505.01076 (2015)

$$\left(\frac{{}^7\text{Li}}{\text{H}}\right)_{\text{theo}} = (4.68 \pm 0.67) \times 10^{-10}$$

$$\left(\frac{\text{Li}}{\text{H}}\right)_{\text{exp}} = (1.6 \pm 0.3) \times 10^{-10}$$

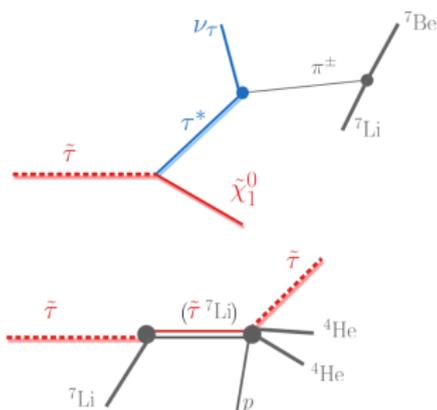
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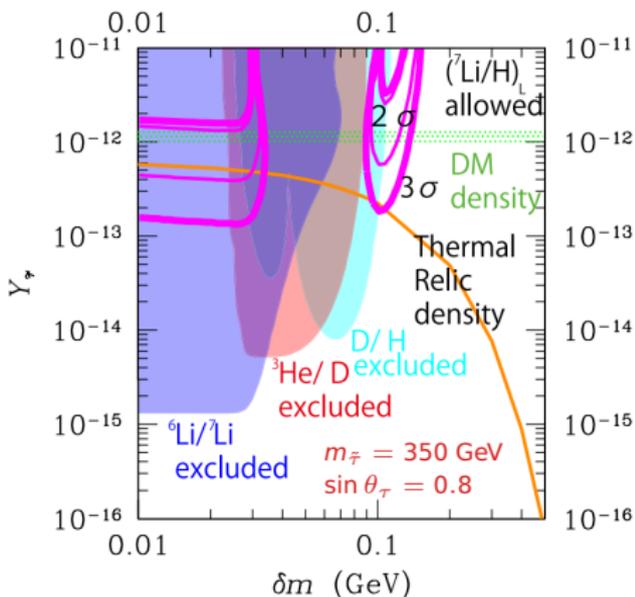
How to deplete the primordial Lithium abundance?

- ${}^7\text{Li}$  depletion with long-lived  $\tilde{\tau}$ :



- A solution is possible with:

- ▶  $Y_{\tilde{\tau}} > 10^{-13}$
- ▶  $\tau_{\tilde{\tau}} > 1 - 100\text{s}$



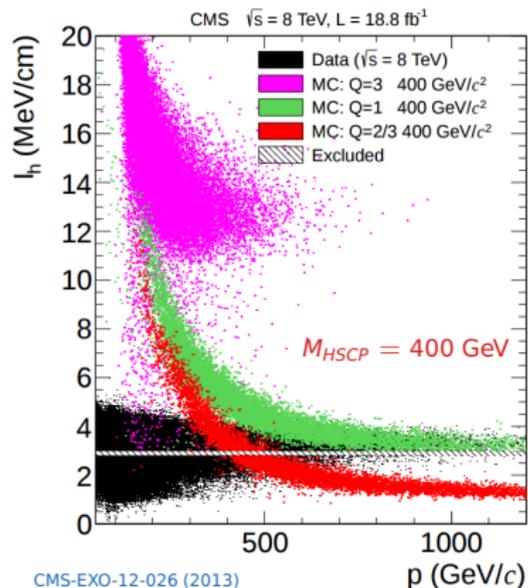
T. Jittoh et al., Phys.Rev. D84 035008 (2011)

How to look for HSCPs @ LHC?

# LHC Searches for HSCPs

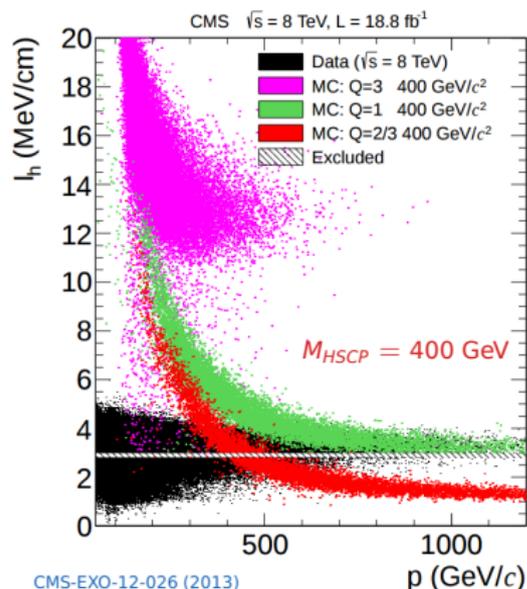
## How to look for HSCPs @ LHC?

- For  $\tau > 10^{-8}$  s  $\rightarrow$  charged tracks
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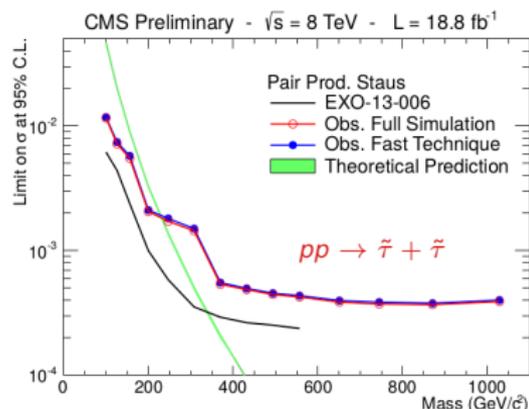
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- CMS Event selection:

- ▶ Charged track
- ▶  $|\eta| < 2.1$ ,  $p_T > 45$  GeV
- ▶ Energy deposit ( $I_h > 3$  MeV)
- ▶ Isolation (in  $\Delta R < 0.3$ ):  
Charged particles:  $(\sum p_T) < 50$  GeV  
Visible particles:  $(\sum \frac{E}{|\vec{p}|}) < 0.3$

- *Partial efficiencies are provided as a function of  $p_{HSCP}$*

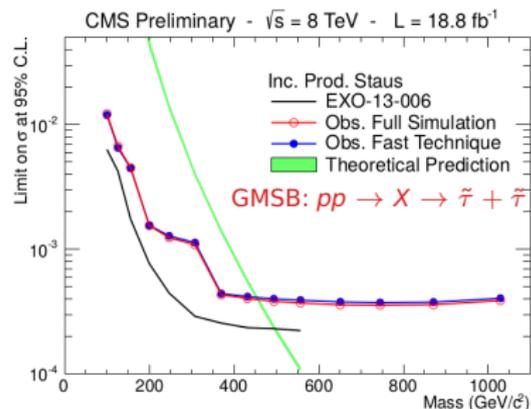
- CMS results (EXO-13-006):



Exclusive production:

$$m_{\tilde{\tau}} > 300 \text{ GeV}$$

~ Model independent



Inclusive production:

$$m_{\tilde{\tau}} > 450 \text{ GeV}$$

Model dependent

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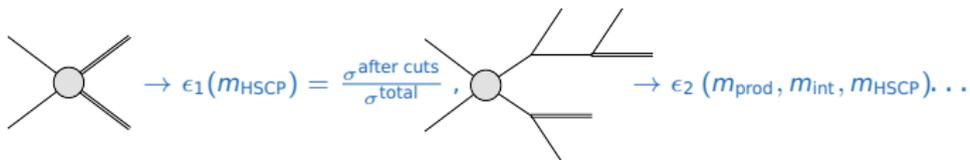
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- Is there a more efficient way of doing it?
  - **Simplified Models**

# Simplified Models for HSCPs

- **Basic idea:**

1. Compute efficiencies for classes of simplified models (SMS)

→ efficiencies database



$\rightarrow \epsilon_1(m_{\text{HSCP}}) = \frac{\sigma_{\text{after cuts}}}{\sigma_{\text{total}}}, \quad \rightarrow \epsilon_2(m_{\text{prod}}, m_{\text{int}}, m_{\text{HSCP}}), \dots$

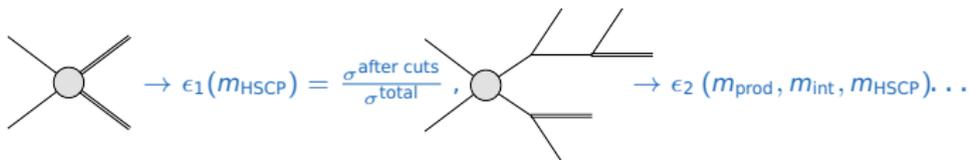
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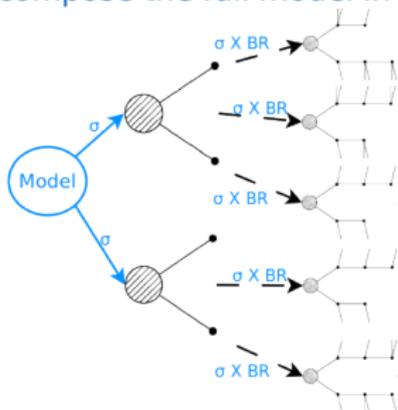
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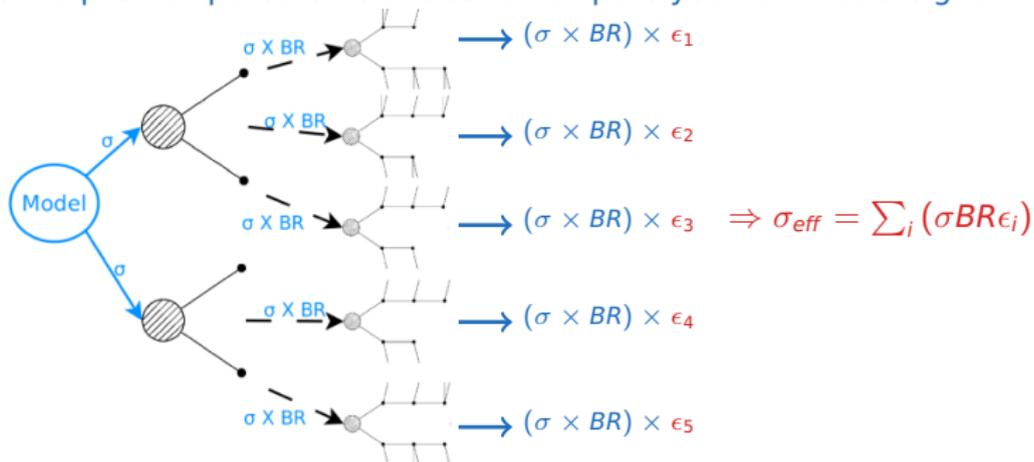
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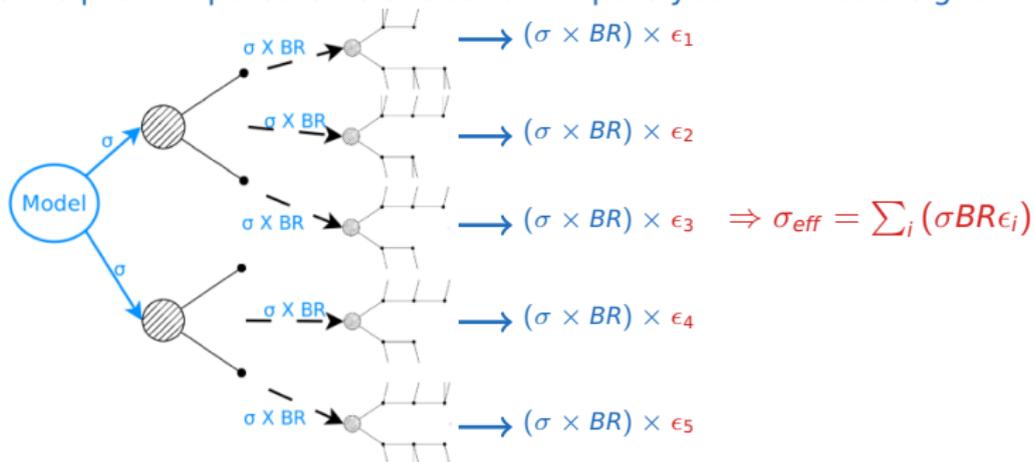
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# Simplified Models for HSCPs

- Basic idea:

3. Use the pre-computed efficiencies to compute your full model signal



4. Compare to the experimental UL:

$\sigma_{eff} > \sigma_{UL} \rightarrow$  the model is excluded

# Simplified Models for HSCPs

- **Basic idea:**

1. Compute efficiencies for classes of simplified models (SMS)
2. Decompose the full model in a coherent sum of SMS
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4. Compare to the experimental UL

- **Pros:**

- ▶ Can be applied to any model
- ▶ No MC simulation needed
- ▶ Very fast
- ▶ Decomposition tools available:

**SM**odels

The logo for 'SModels' features the word 'SModels' in a black, sans-serif font. The letter 'o' is replaced by a red circular icon containing a white particle detector-like structure with wavy lines extending from its base.

- **Cons:**

- ▶ Neglects specific model dynamics:  
 $\epsilon \simeq \epsilon(\text{masses}, \text{topology})$
- ▶ Limited by the simplified models used to compute  $\epsilon$
- ▶ Decomposition can be slow in special cases

# Simplified Models for HSCPs

Computing efficiencies...

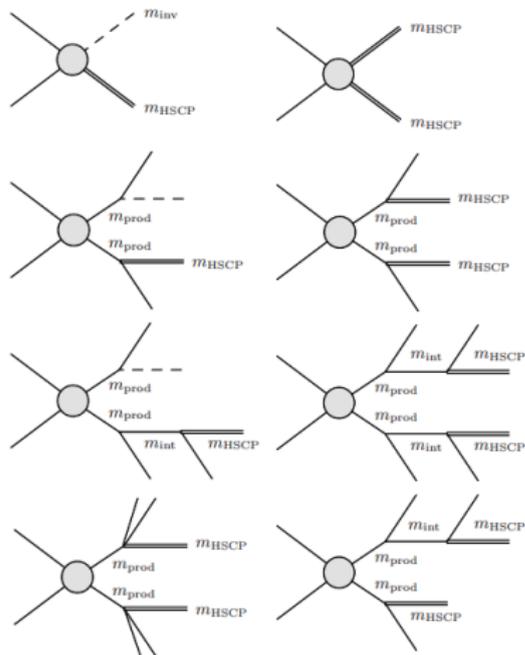
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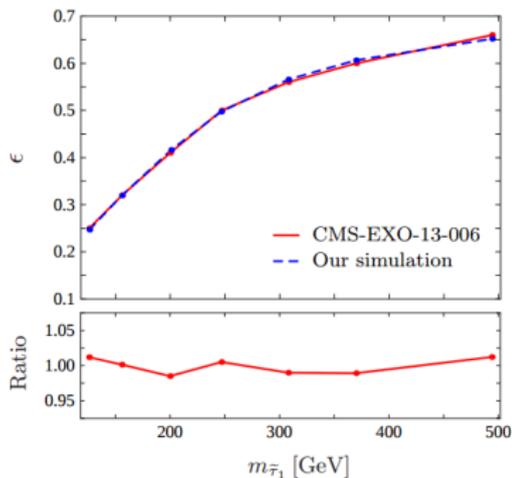


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## Computing efficiencies...

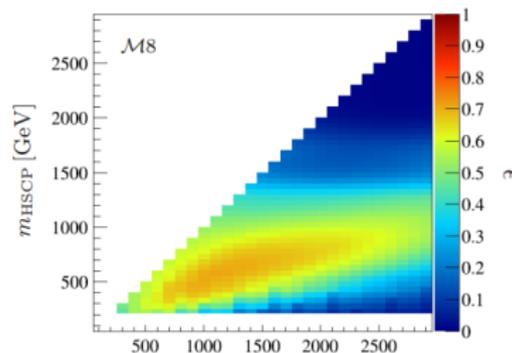
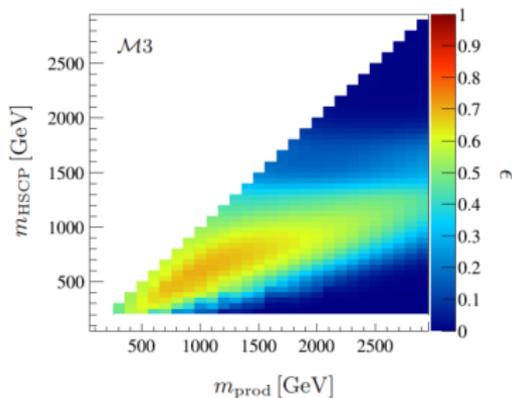
(MadGraph + Pythia + CMS probabilities)

### ● Validation:



agreement within  $\lesssim 5\%$

### ● Results for $\epsilon$ :

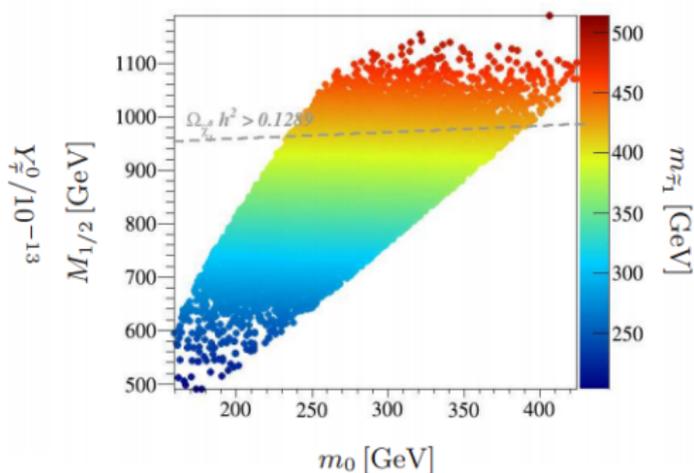
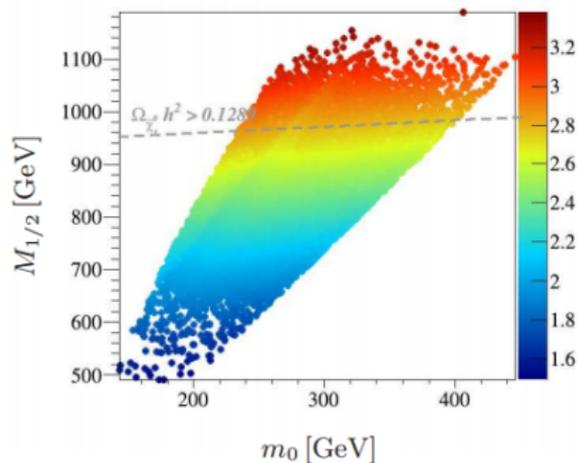


- Application to the CMSSM

- **Application to the CMSSM+** Solution to the Lithium-7 Problem:
  - ▶ Neutralino LSP, stau NLSP
  - ▶  $Y_{\tilde{\tau}} > 10^{-13}$
  - ▶  $\tau_{\tilde{\tau}} > 1s$  ( $m_{\tilde{\tau}} - m_{N1} < m_{\tau}$ )
  - ▶  $\tan \beta = 10$

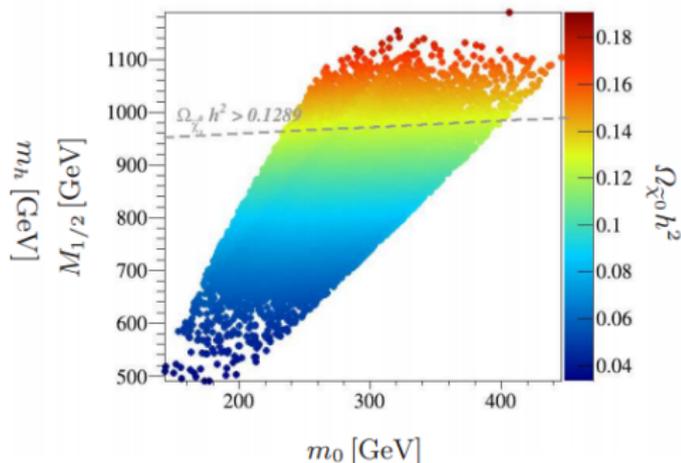
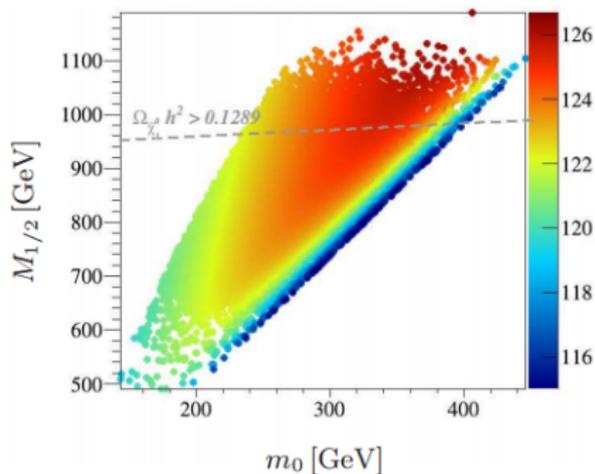
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  - ▶  $\tan \beta = 10$
- Scan over  $m_0, M_{1/2}, A_0$  ( $\mu > 0$ ):



# CMSSM with long-lived $\tilde{\tau}$ s

- Higgs and Dark Matter constraints:



- We require:

- ▶  $120 \text{ GeV} < m_h < 130 \text{ GeV}$

# CMSSM with long-lived $\tilde{\tau}$ s: LHC Constraints

- LHC Constraints:

- ▶ MET signatures:  $\tilde{q} + \tilde{q} \rightarrow qq + \tilde{\chi}_1^0 + \tilde{\chi}_1^0 \sim 70\%$
- ▶ HSCP signatures:  $\tilde{\chi}_1^\pm + \tilde{\chi}_1^\pm \rightarrow \nu_\tau + \tilde{\tau}_1^\pm + \nu_\tau + \tilde{\tau}_1^\pm \sim 10\%$
- ▶ Mixed signatures:  $\tilde{\chi}_1^\pm + \tilde{\chi}_2^0 \rightarrow \nu_\tau + \tilde{\tau}_1^\pm + Z + \tilde{\chi}_1^0 \sim 20\%$

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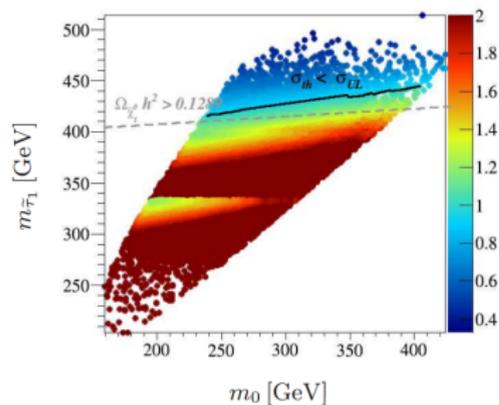
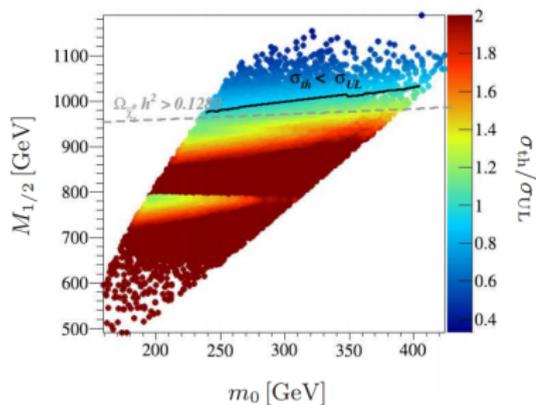
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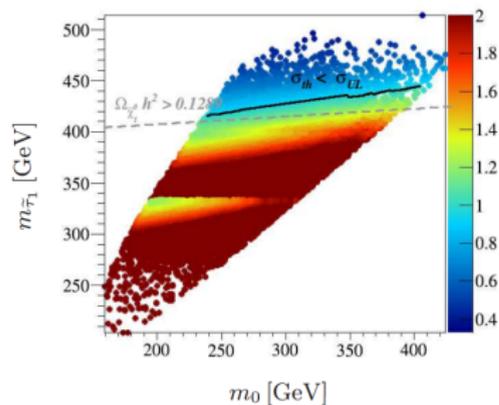
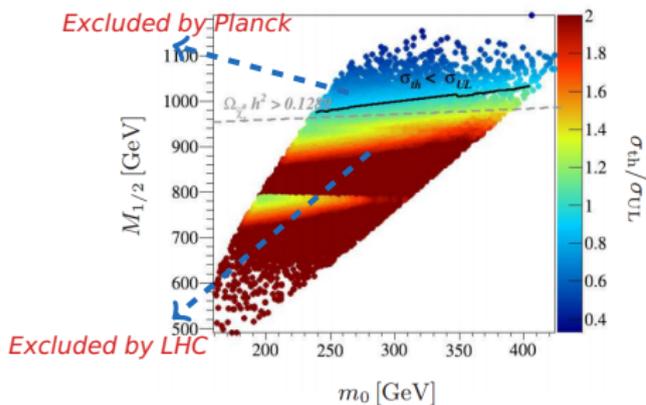
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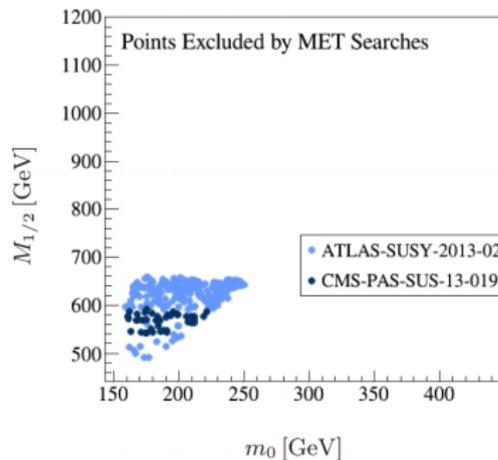
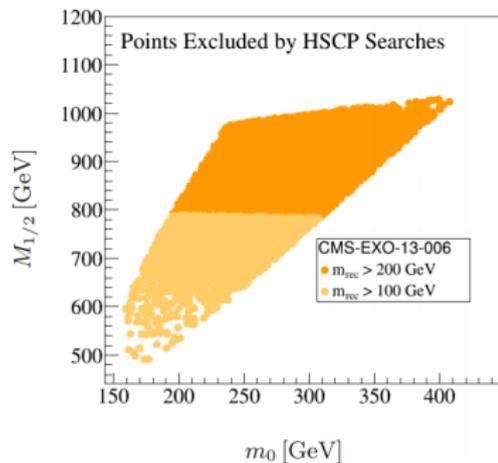
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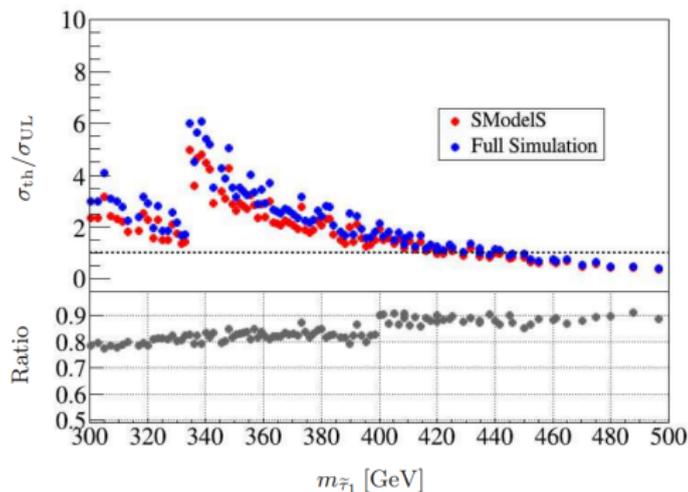
- Results: MET vs HSCP



- HSCP constraints dominate (even though the HSCP signal is only  $\sim 30\%$ )
- MET constraints are smaller than in the usual CMSSM (MET signal  $\sim 70\%$ )

# CMSSM with long-lived $\tilde{\tau}$ s: LHC Constraints

- How does SMS + efficiencies compare with the full sim?



→ Signal coverage  $\sim 90\%$

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- Simplified Models approach  $\sim$  Full simulation

# Thanks!