

Searching for new paradigms in particle physics

Towards the origin of the electroweak scale

ICTP-SAIFR, Sao Paulo

**Alex Pomarol,
IFAE &UAB (Barcelona)**

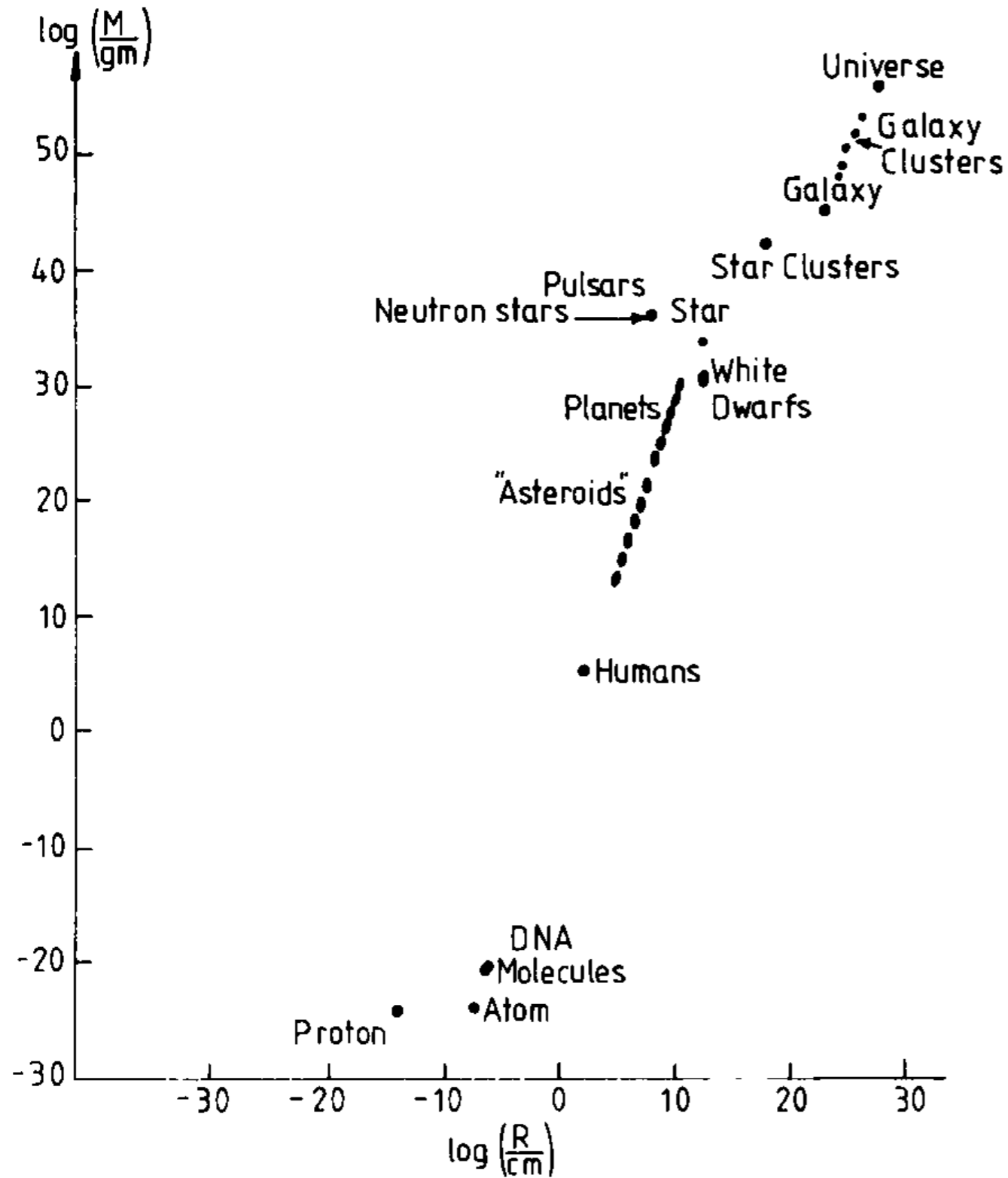
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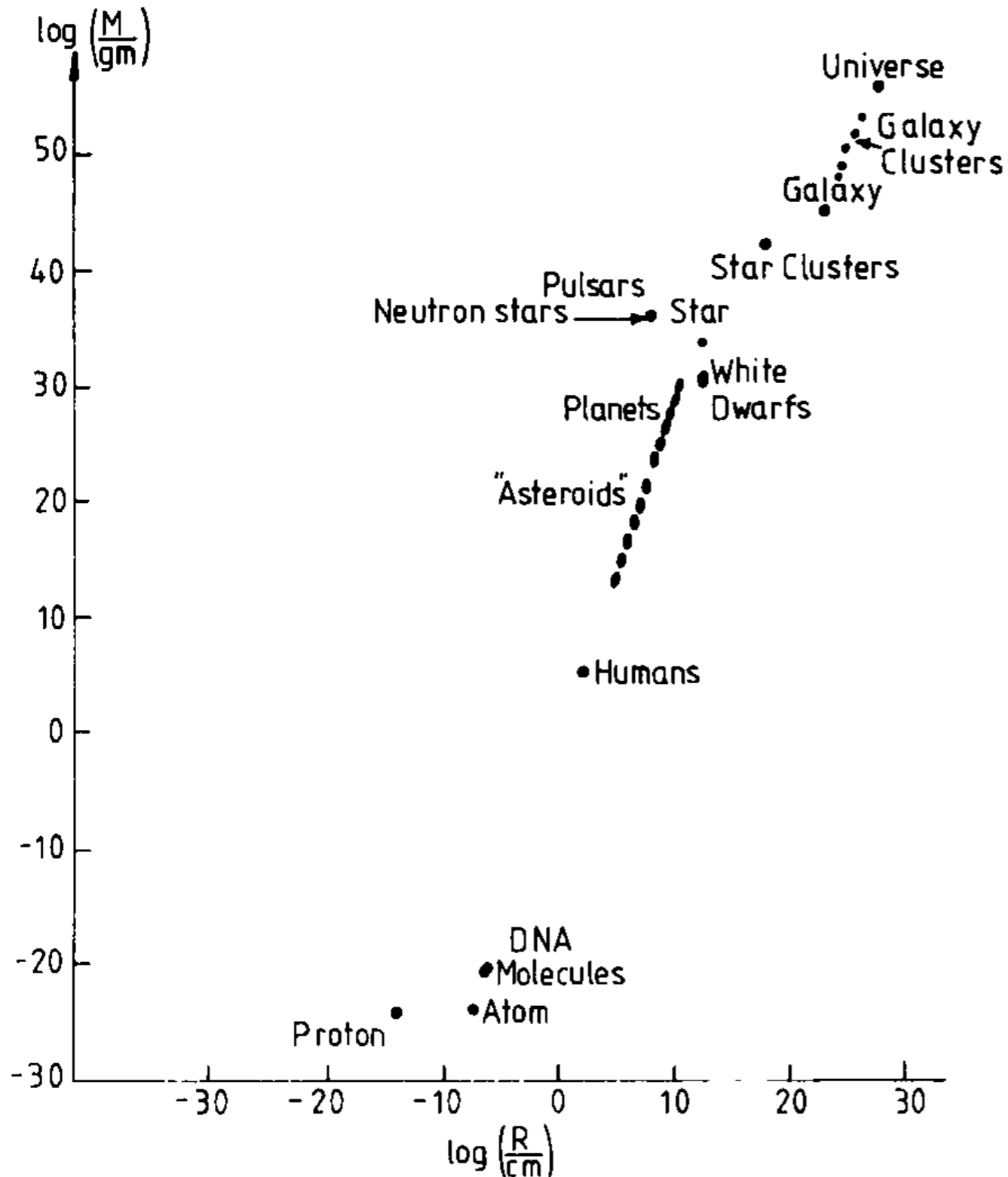
ICTP-SAI FR, Sao Paulo

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Understanding scales: The driving force of physics



Understanding scales: The driving force of physics



$$M_{\text{star}} \sim \left(\frac{M_{\text{P}}}{m_{\text{N}}} \right)^3 m_{\text{N}}$$

$$R_{\text{planet}} \sim \frac{m_{\text{N}}}{M_{\text{P}}} \frac{1}{\sqrt{\alpha m_{\text{e}}}}$$

$$T_{\text{room}} \sim 10^{-3} \alpha^2 m_{\text{e}}$$

$$r_{\text{atom}} \sim \frac{1}{\alpha m_{\text{e}}}$$

$$r_{\text{proton}} \sim \frac{1}{m_{\text{N}}}$$

Based on few underlying fundamental interactions



**EM interaction
& Gravity**

**Strong
interaction**

**Weak
interaction**

“Coulomb phase”

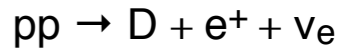
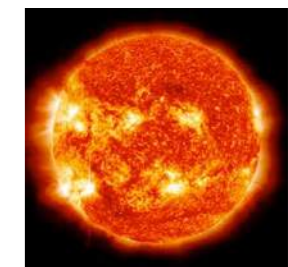
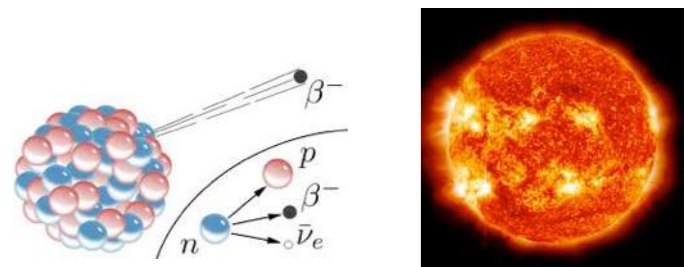
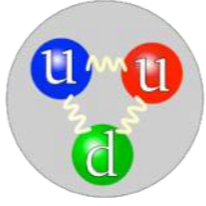
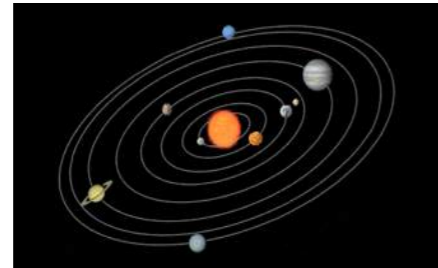
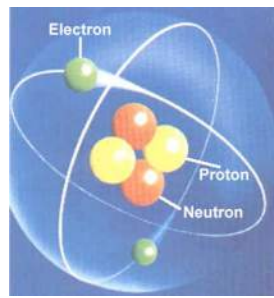
$$V(\mathbf{r}) \sim \frac{1}{r}$$

“Confining phase”

$$V(\mathbf{r}) \sim r$$

“Higgs phase”

$$V(\mathbf{r}) \sim e^{-r}$$



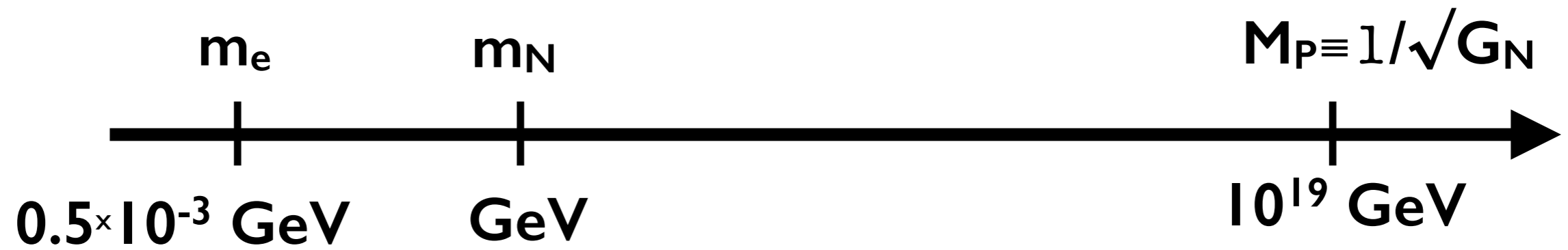
responsible
for atoms

solar system

responsible
for nuclei

responsible
for sun's ignition

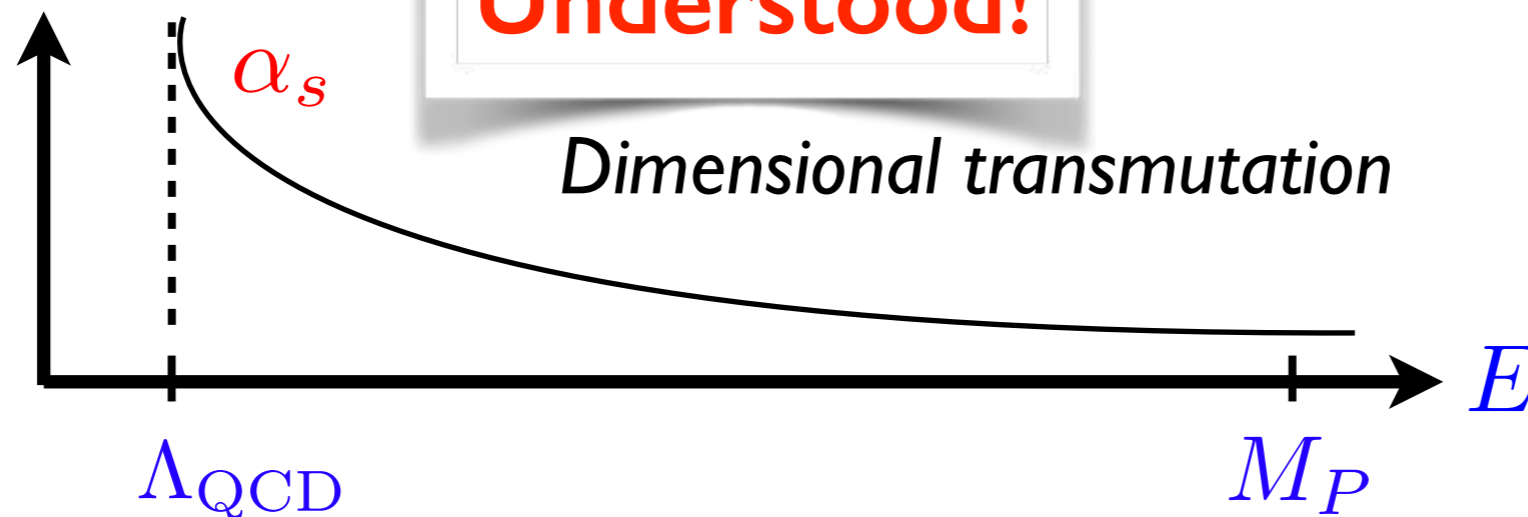
Underlying mass scales in the universe



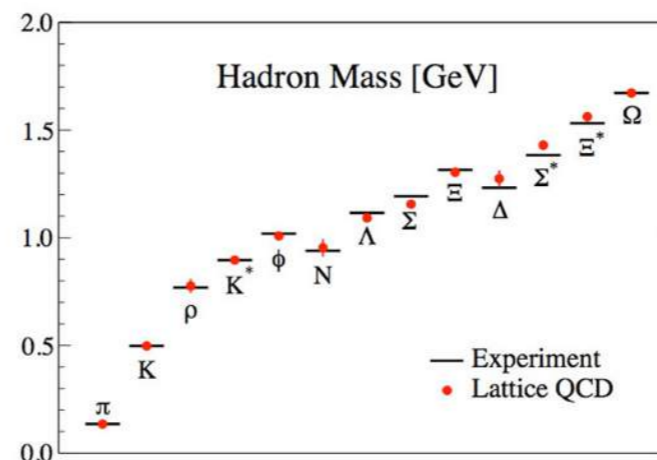
Underlying mass scales in the universe



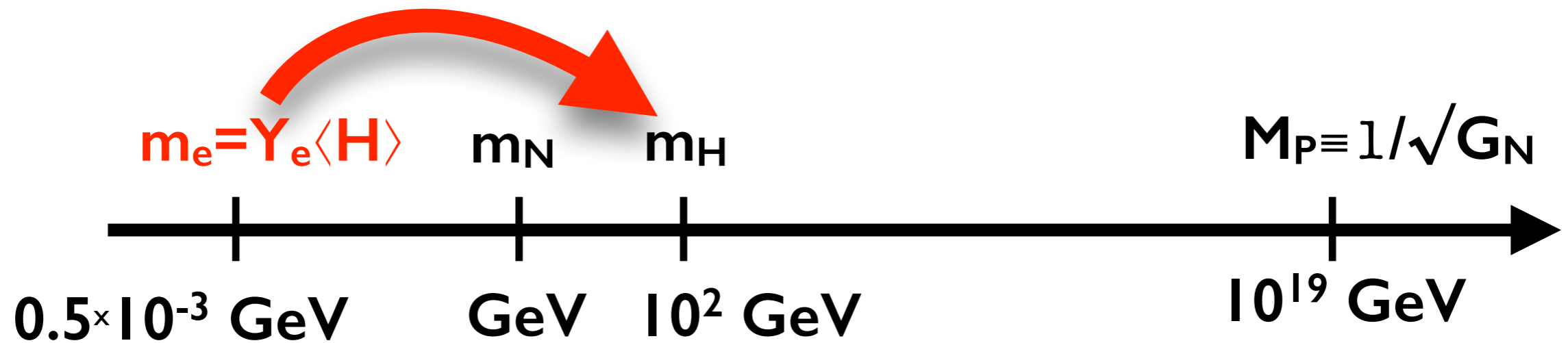
Understood!



Explains why $\Lambda_{\text{QCD}} \ll M_P$
and the origin
of most hadron masses



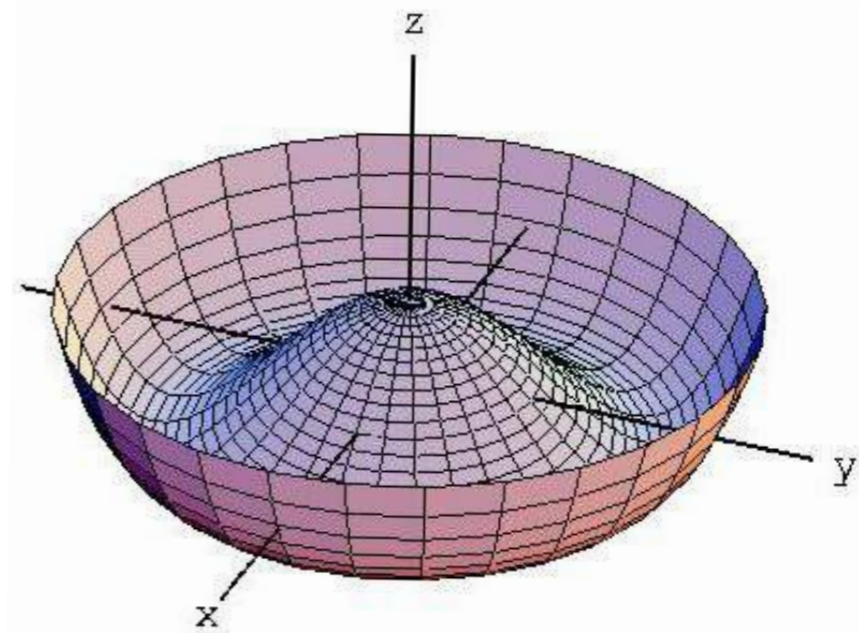
Underlying mass scales in the universe



Standard Model
masses from the Higgs VEV:

$$\langle H \rangle \sim m_H / \sqrt{\lambda}$$

Reinforced by
the 2012 Higgs discovery

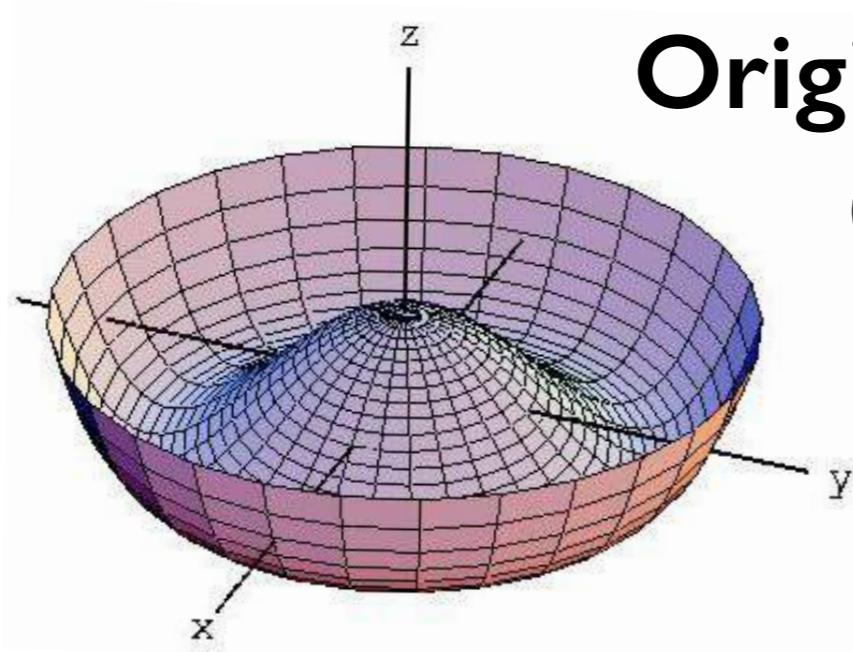


$$V(\mathbf{H}) = -m_H^2 |\mathbf{H}|^2 + \lambda |\mathbf{H}|^4$$

Underlying mass scales in the universe



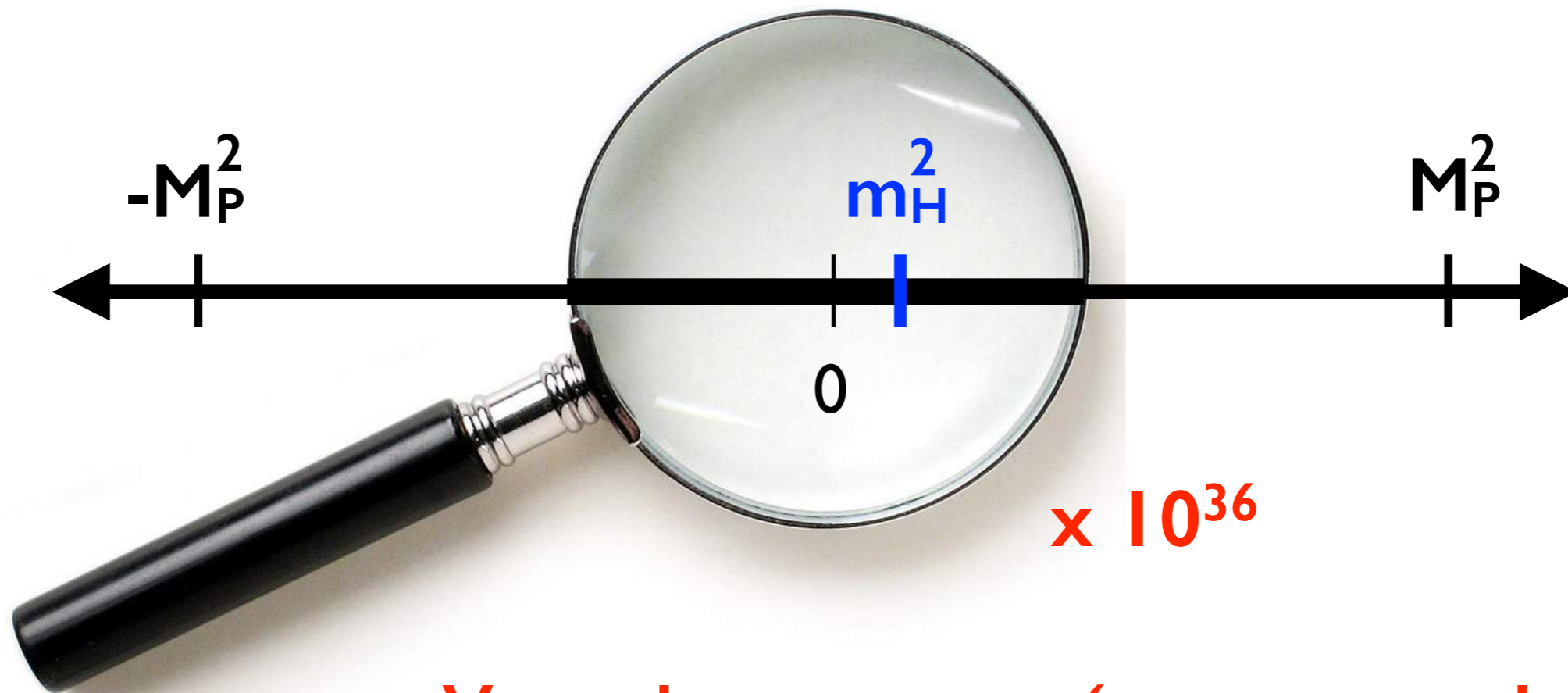
Remaining mystery!



**Origin of the Higgs potential
(electroweak scale)**

$$V(\mathbf{H}) = -m_H^2 |\mathbf{H}|^2 + \lambda |\mathbf{H}|^4$$

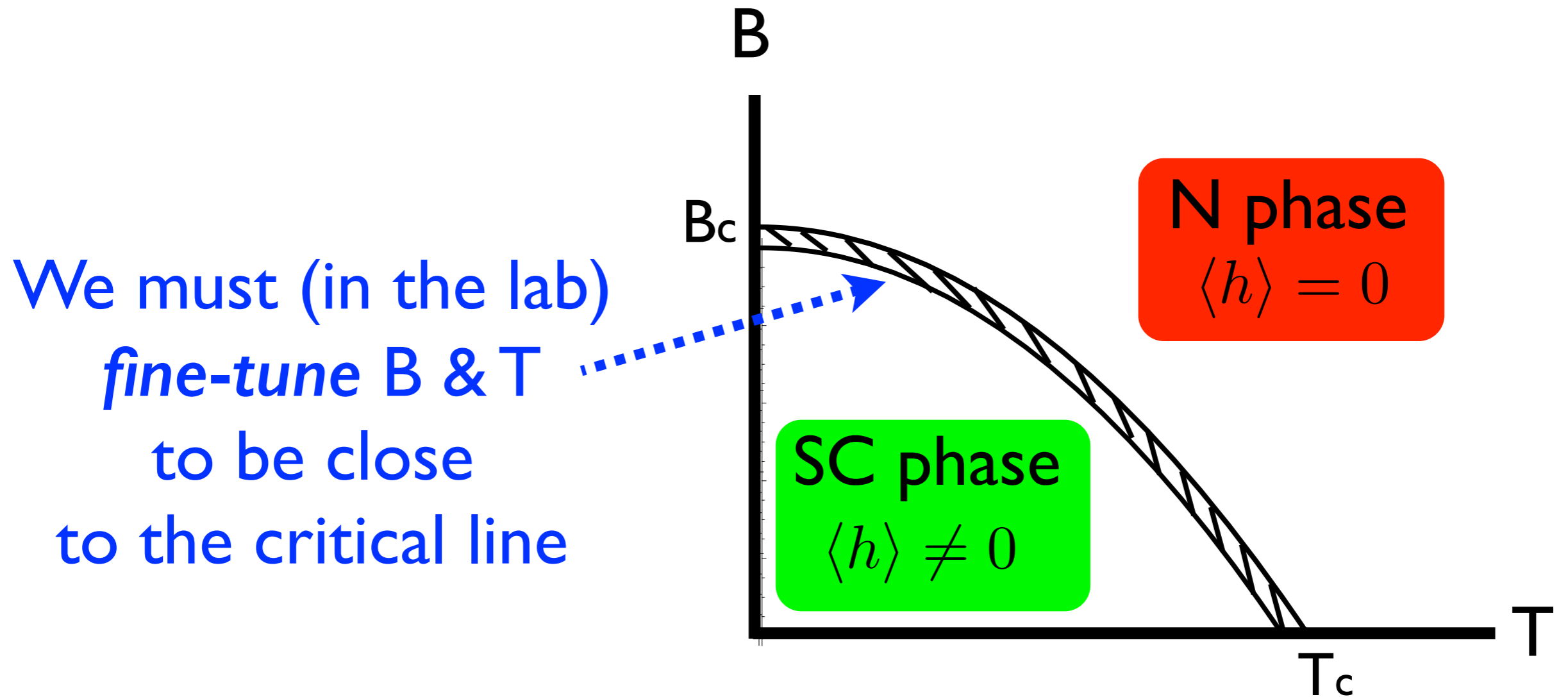
Hierarchy problem



Very close to zero (as compared to M_P)

Analogy with Superconductivity

$$\langle h \rangle = \langle e^- e^- \rangle$$



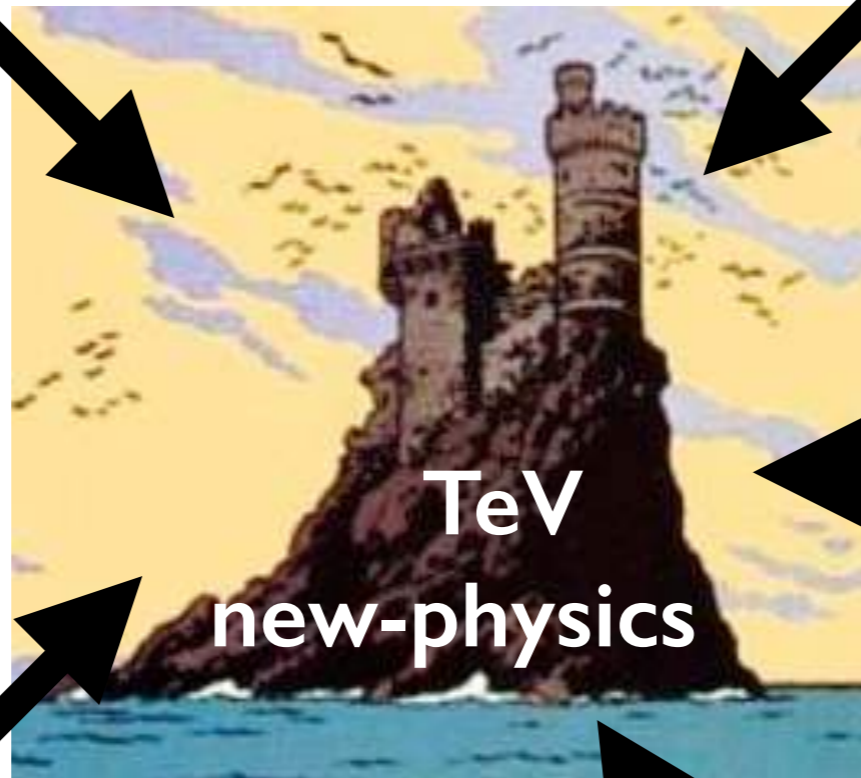
But who is tuning $m_H \ll M_P$?

Most of our effort in experimental particles physics has been to *attack* the TeV territory from different fronts

Looking for deviations in SM couplings



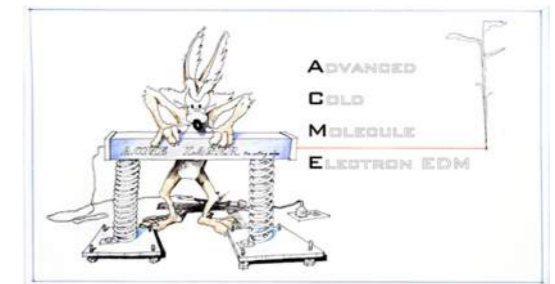
Looking for new particles



Looking for new flavor-transitions



Looking for Electric Dipole Moments



Looking for WIMPs (Dark Matter)

Theoretical proposals for the origin of the Electroweak Scale ($m_H \ll M_P$)

*Idealized models have a useful role to play,
as ways to clarify your thinking*

Paul Krugman

& status of their *health* (experimental) *checkups*



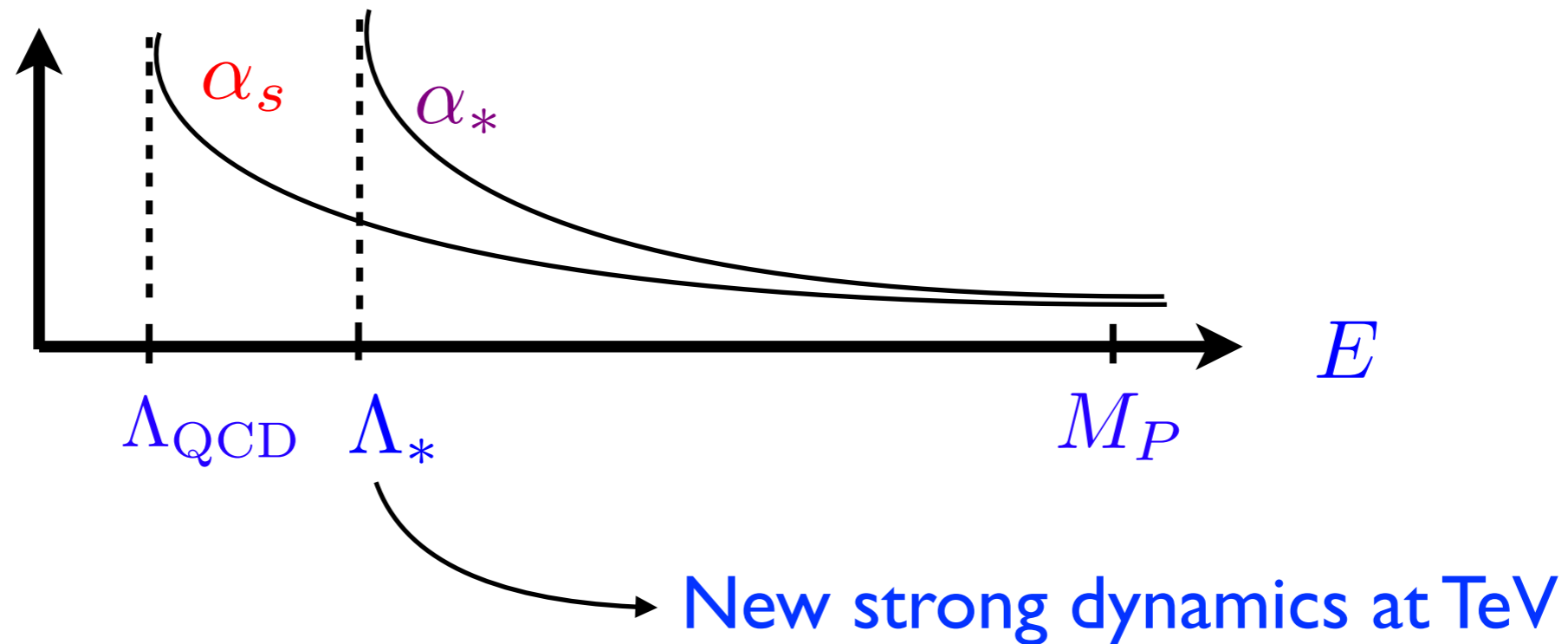
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QCD approach: Compositeness

Follow the path of QCD:

Lets try the same for the EW scale

Assume that there is a **New Strong sector** at around the TeV-scale:

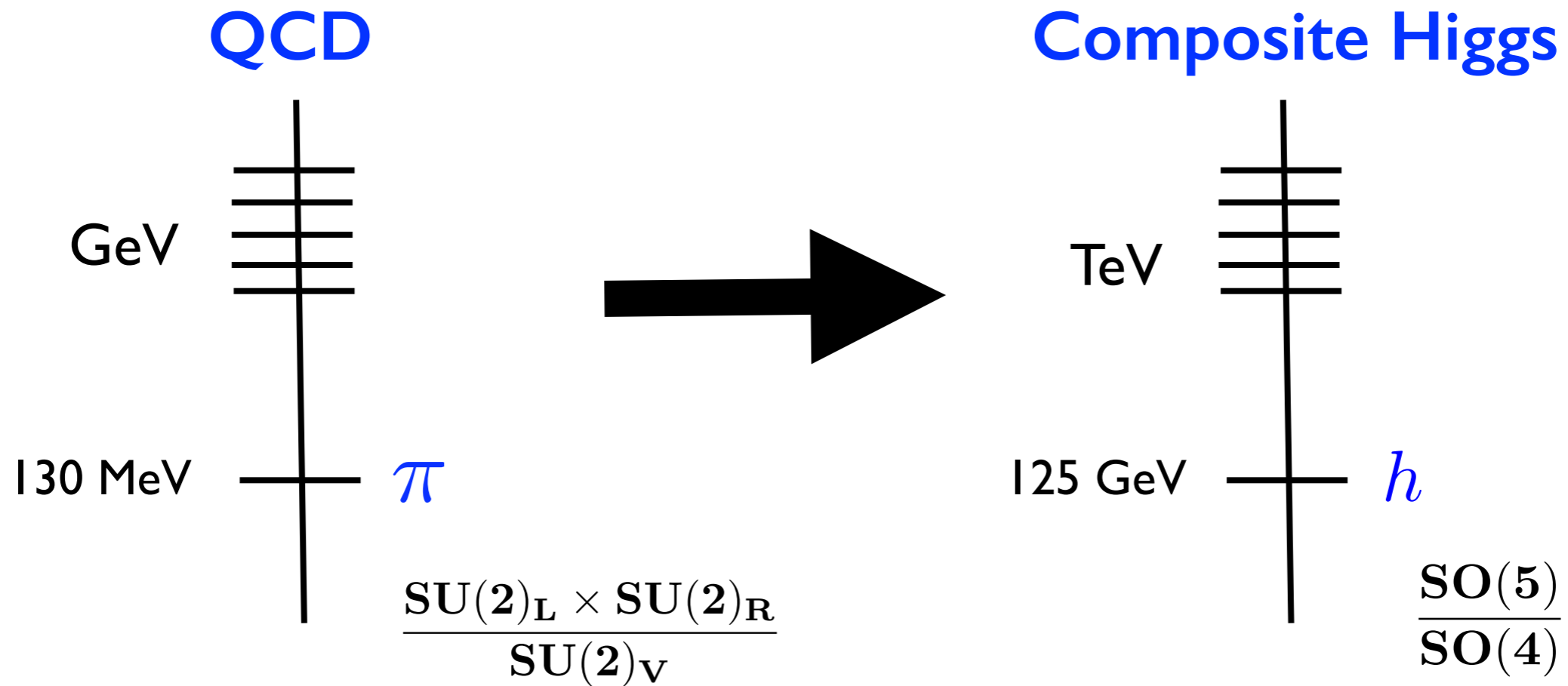


It could explain why $m_H \lesssim \Lambda_* \sim \text{TeV} \ll M_P$

Composite Higgs: $H = \bar{u}u$

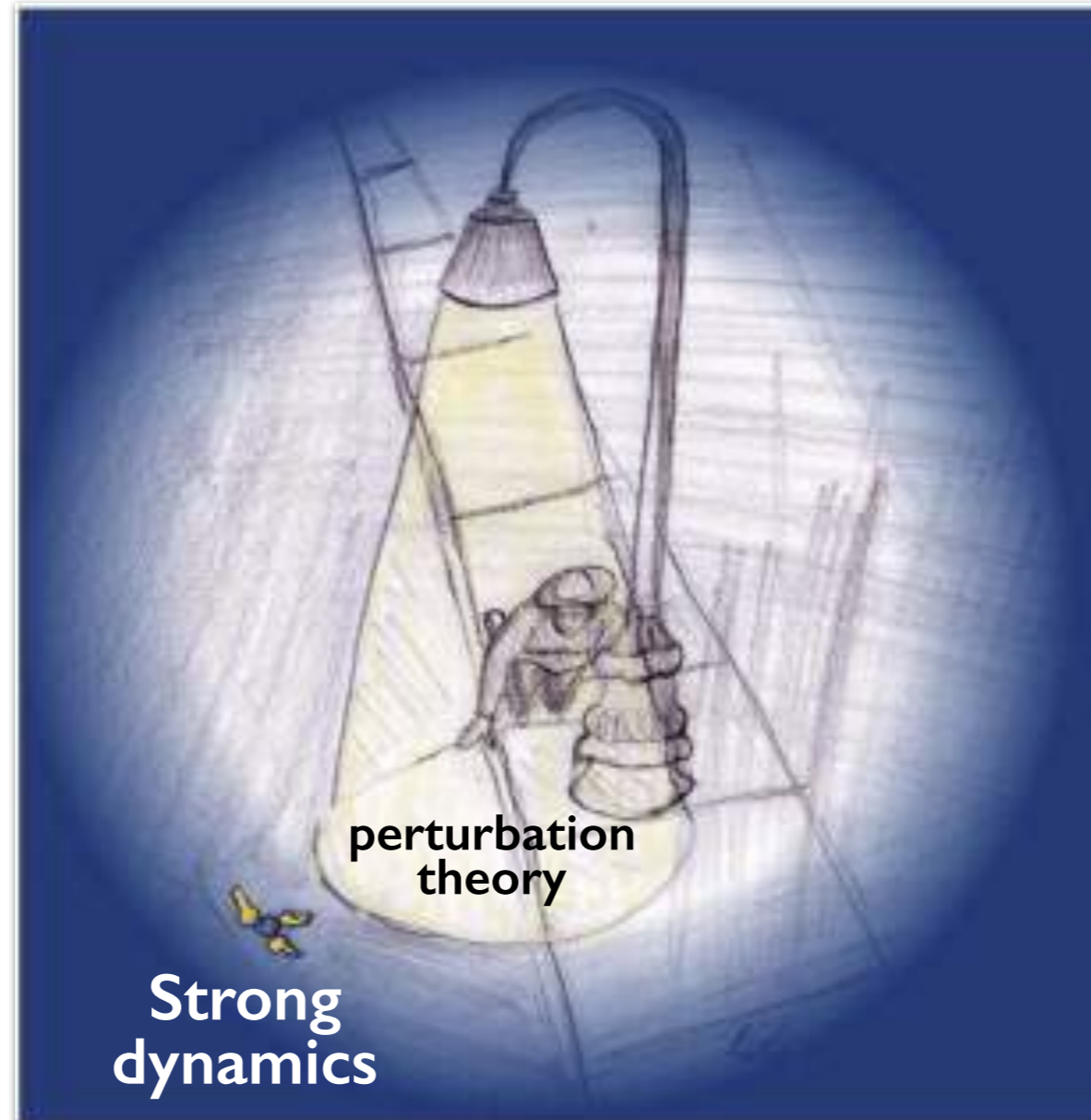


The Higgs, the lightest of the new strong resonances, as pions in QCD: they are Pseudo-Goldstone Bosons (PGB)



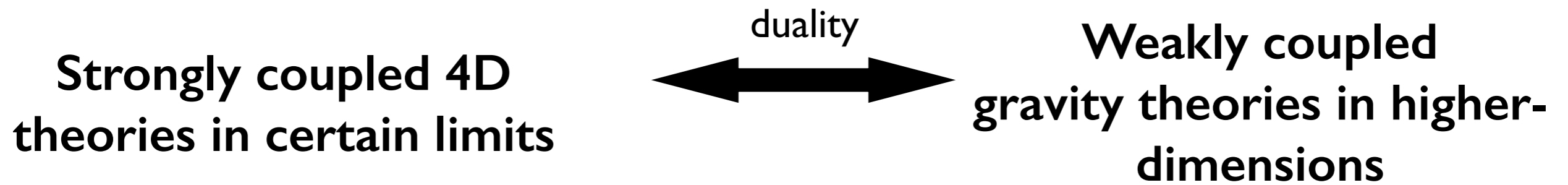
Dealing with strong dynamics....

Beyond the lamp-post:



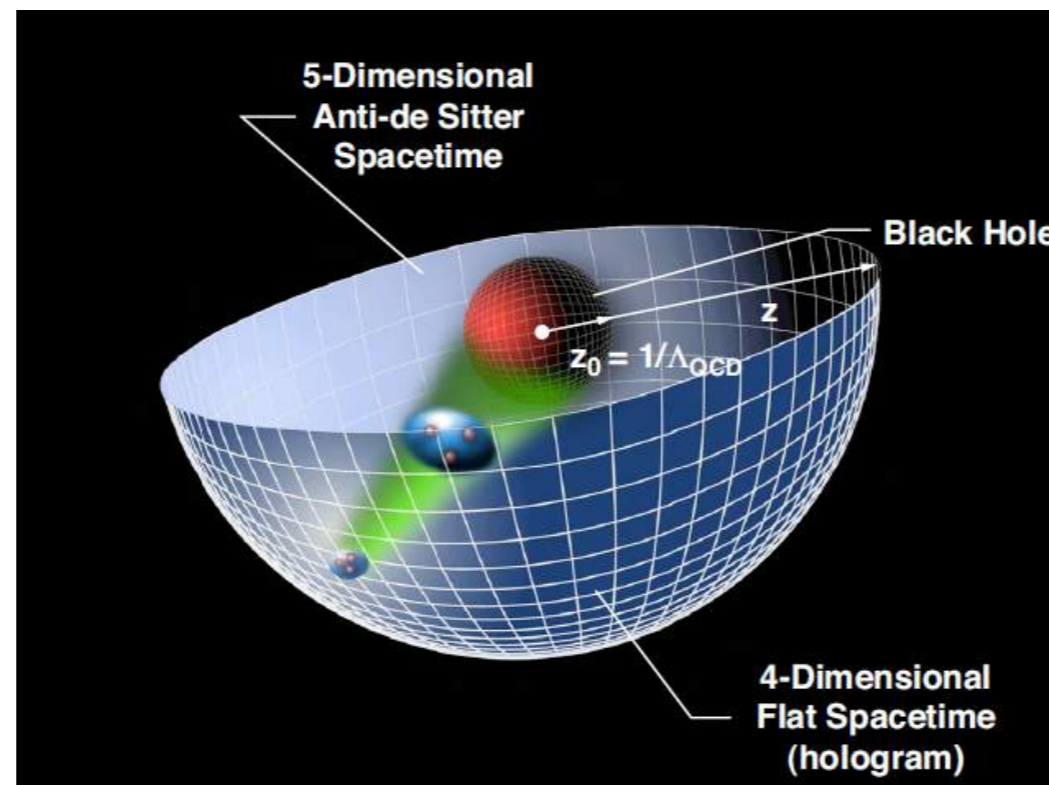
The AdS/CFT correspondence

Maldacena 97



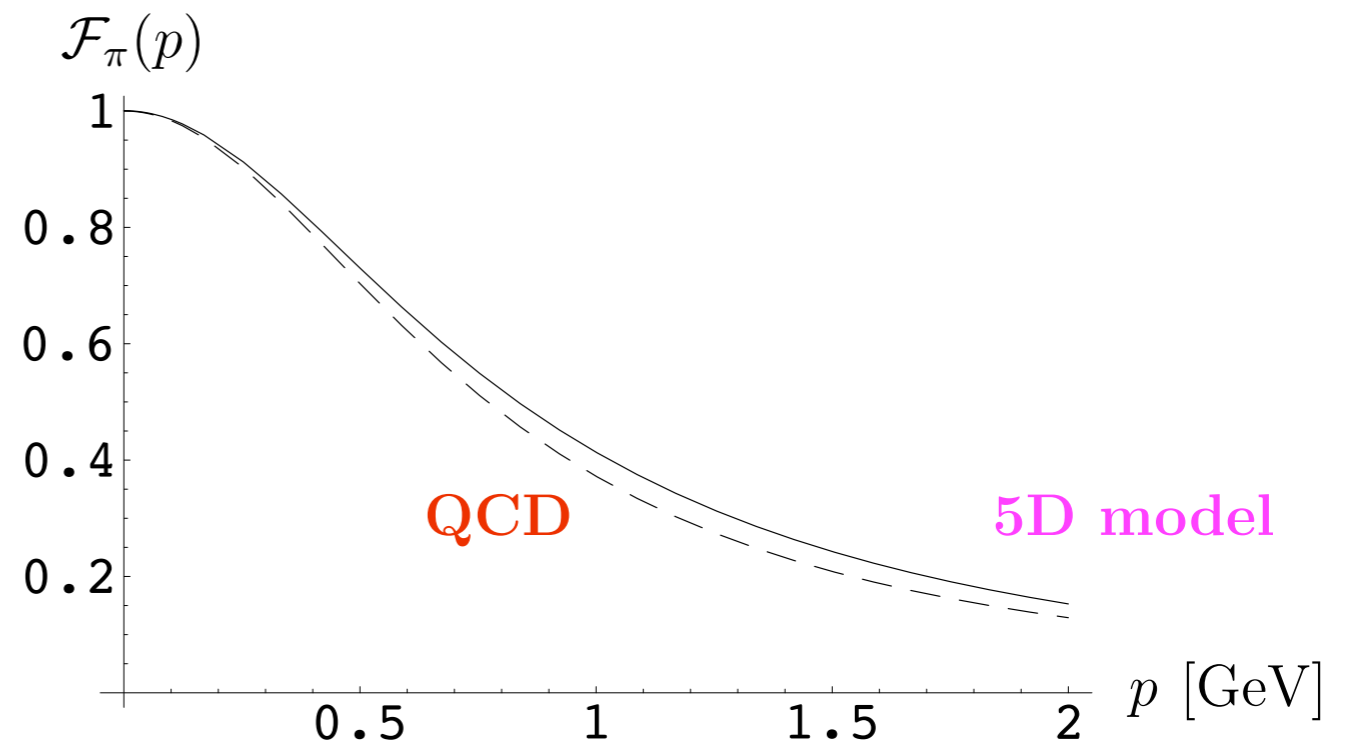
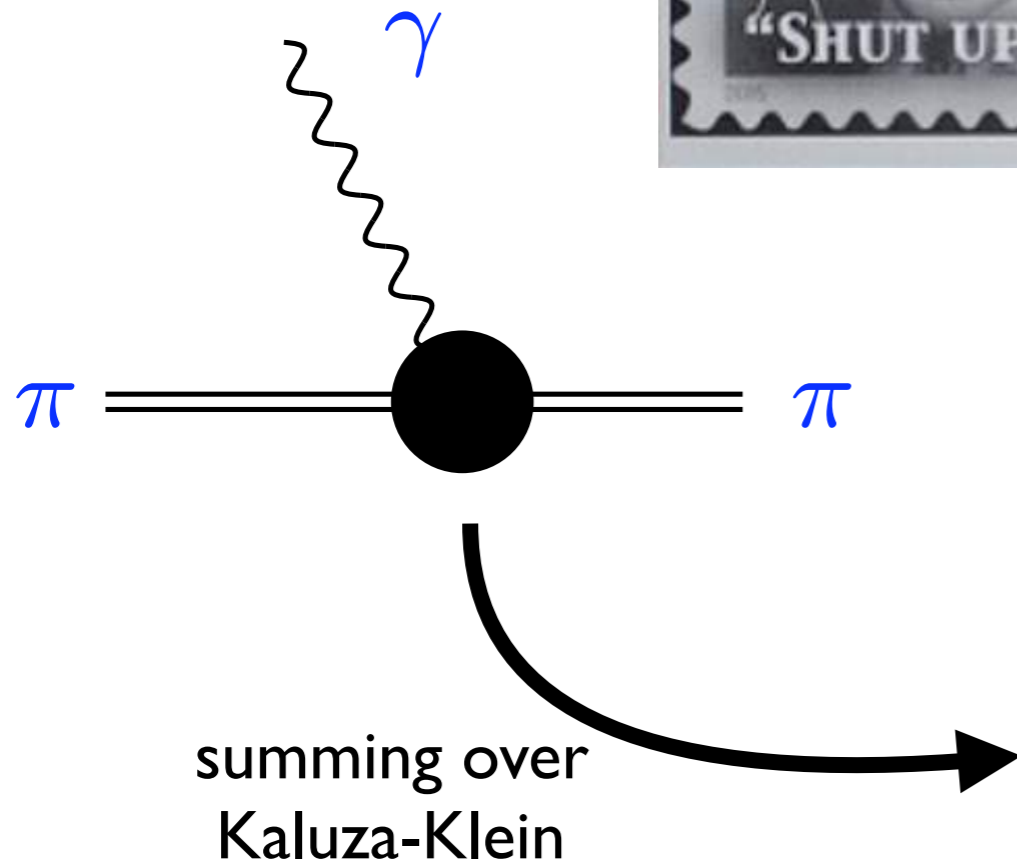
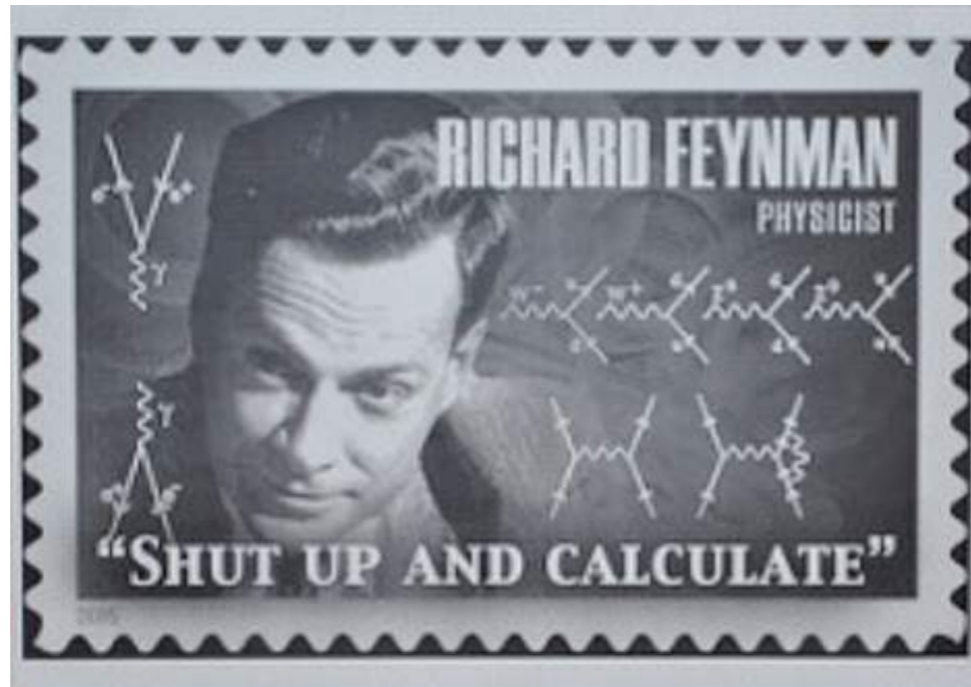
👉 Holographic models:

5D models with the properties of a composite Higgs



Picture from
G.F. de Teramond

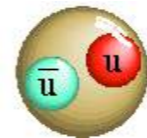
Why extra-dimensional models are models of compositeness?



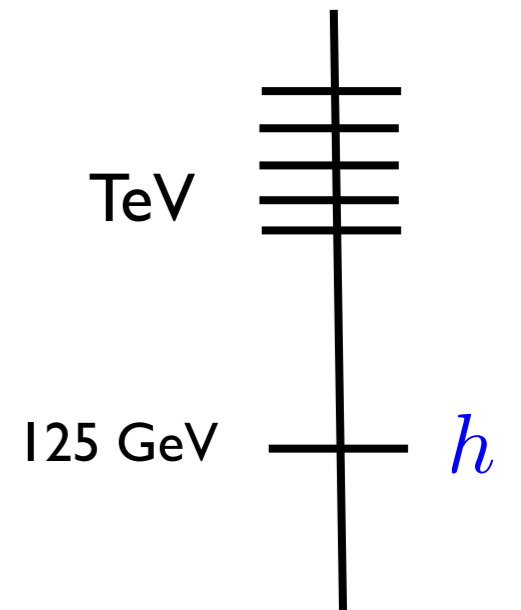
Physical implications of TeV strong-dynamics

New flavor-violating
& CP-violating
transitions

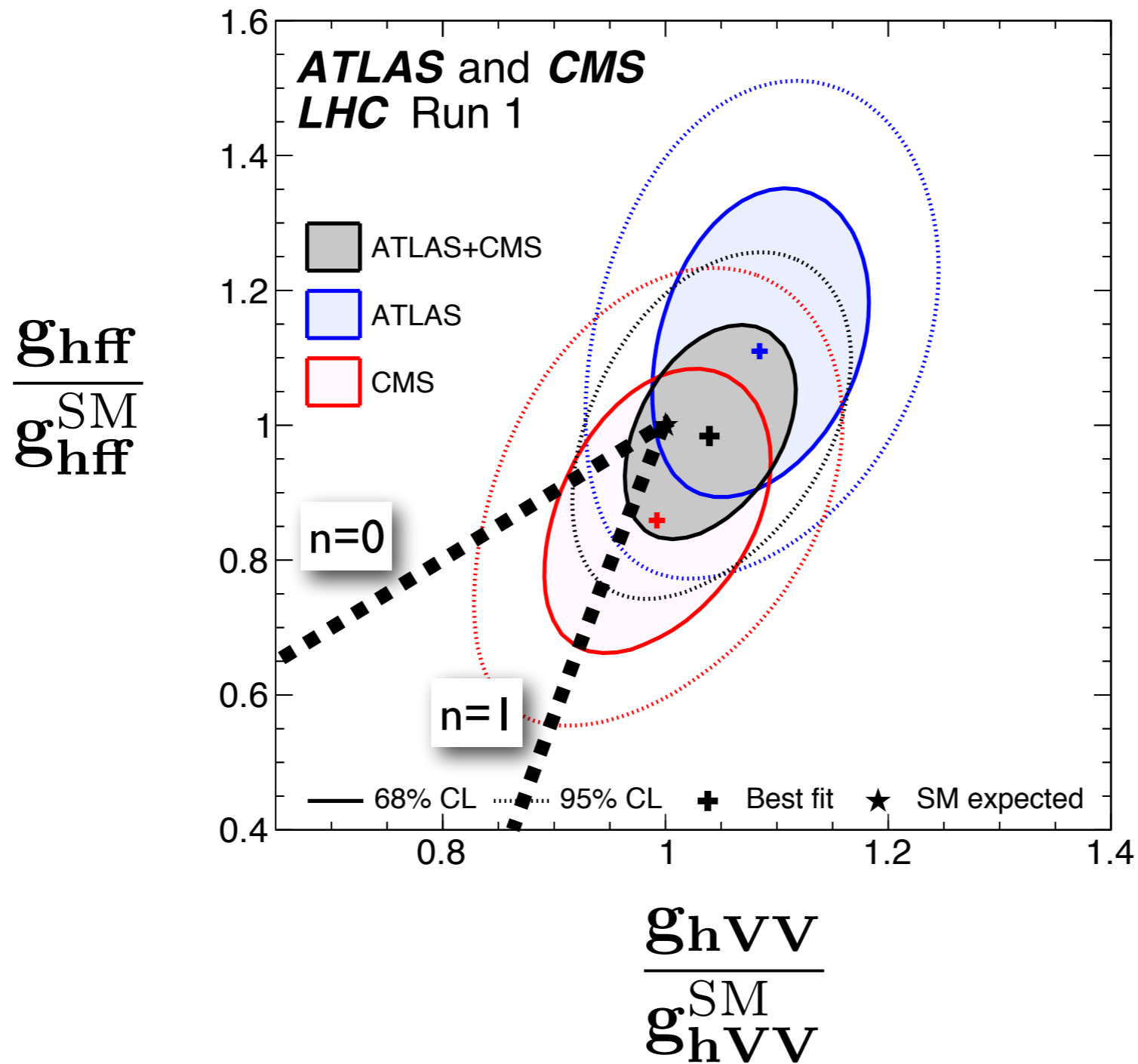
Signs of compositeness
in the Higgs (and top)



New resonances

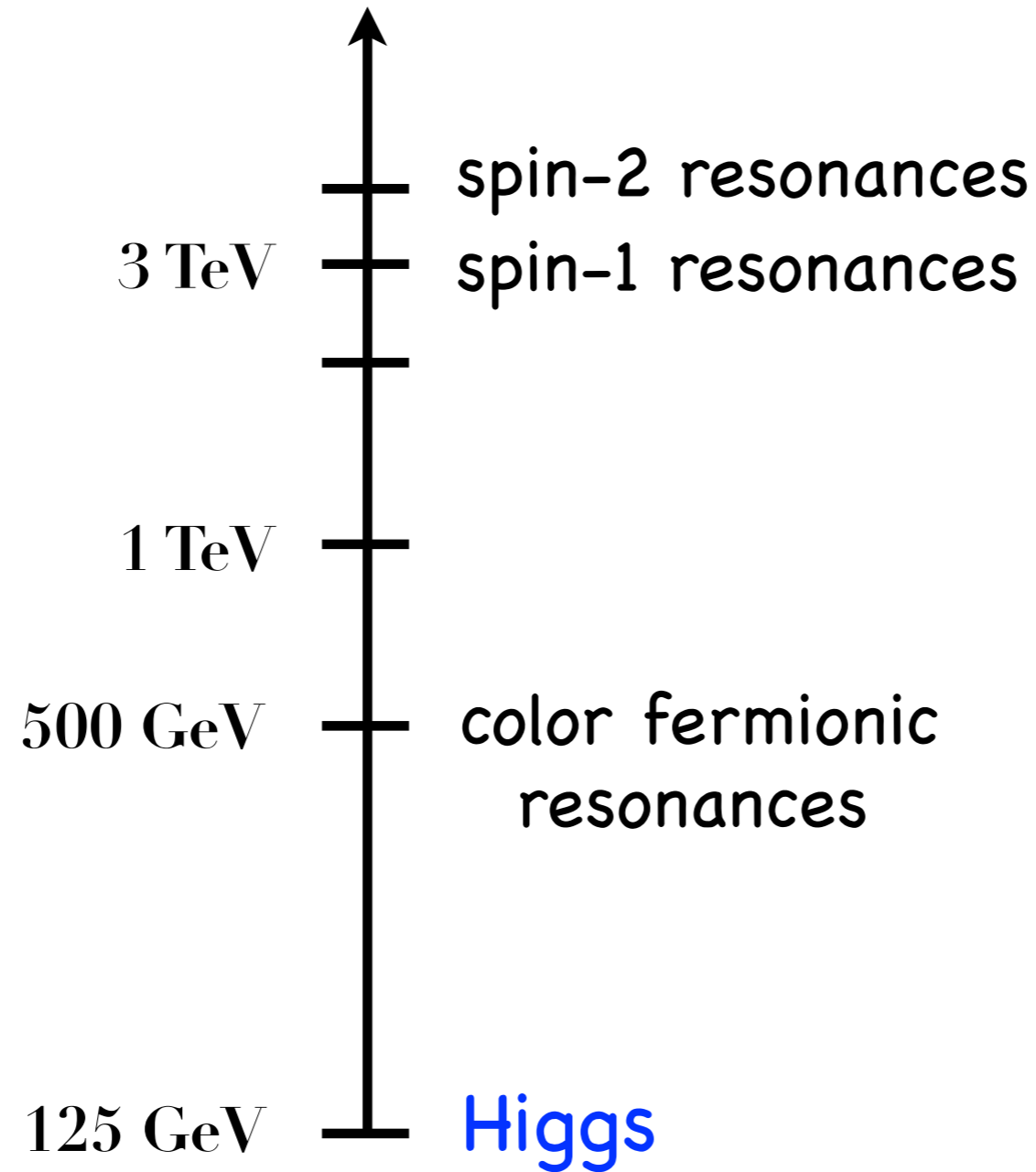


Signs of compositeness of the Higgs

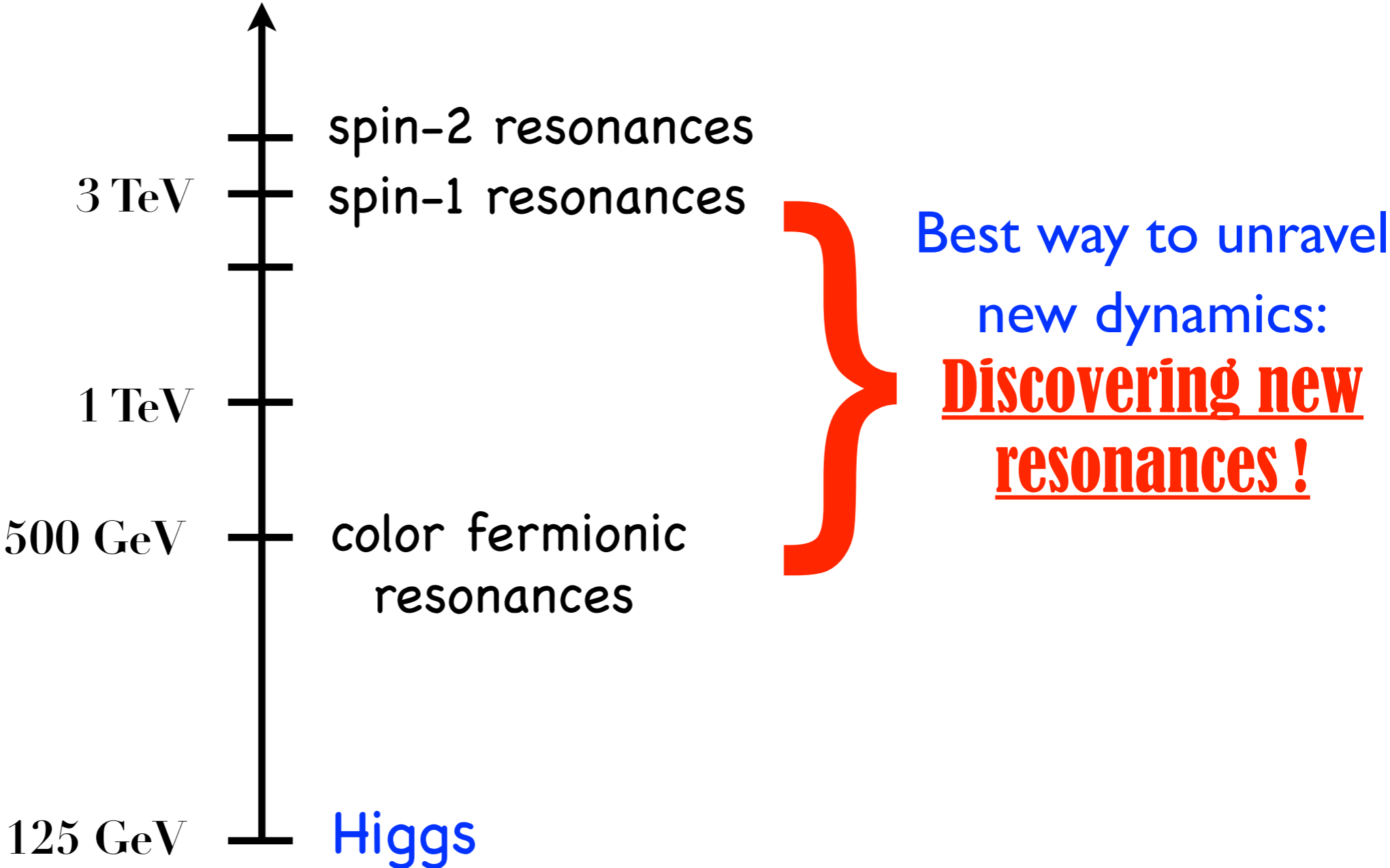


Entering the interesting region: bounds getting below 10%!

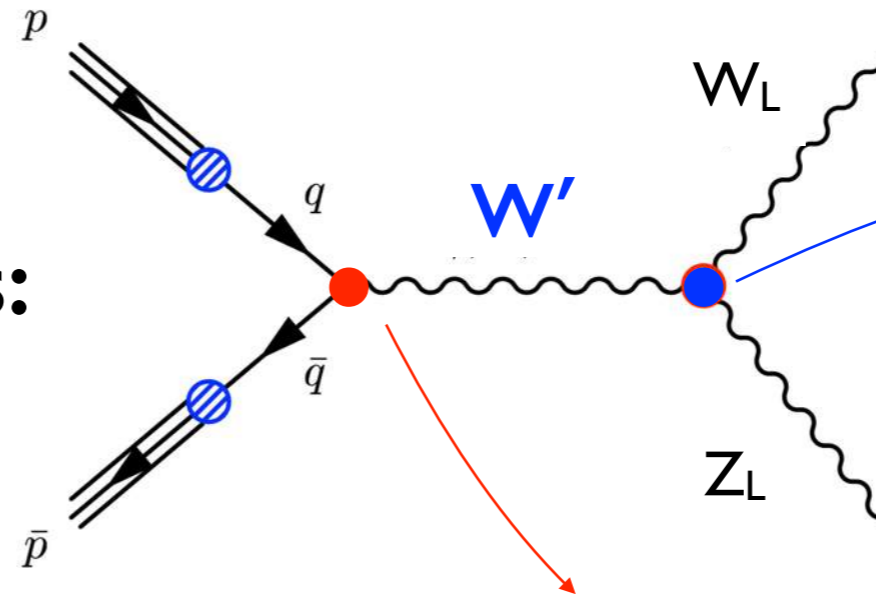
Expected spectrum in Composite Higgs Scenarios



Expected spectrum in Composite Higgs Scenarios



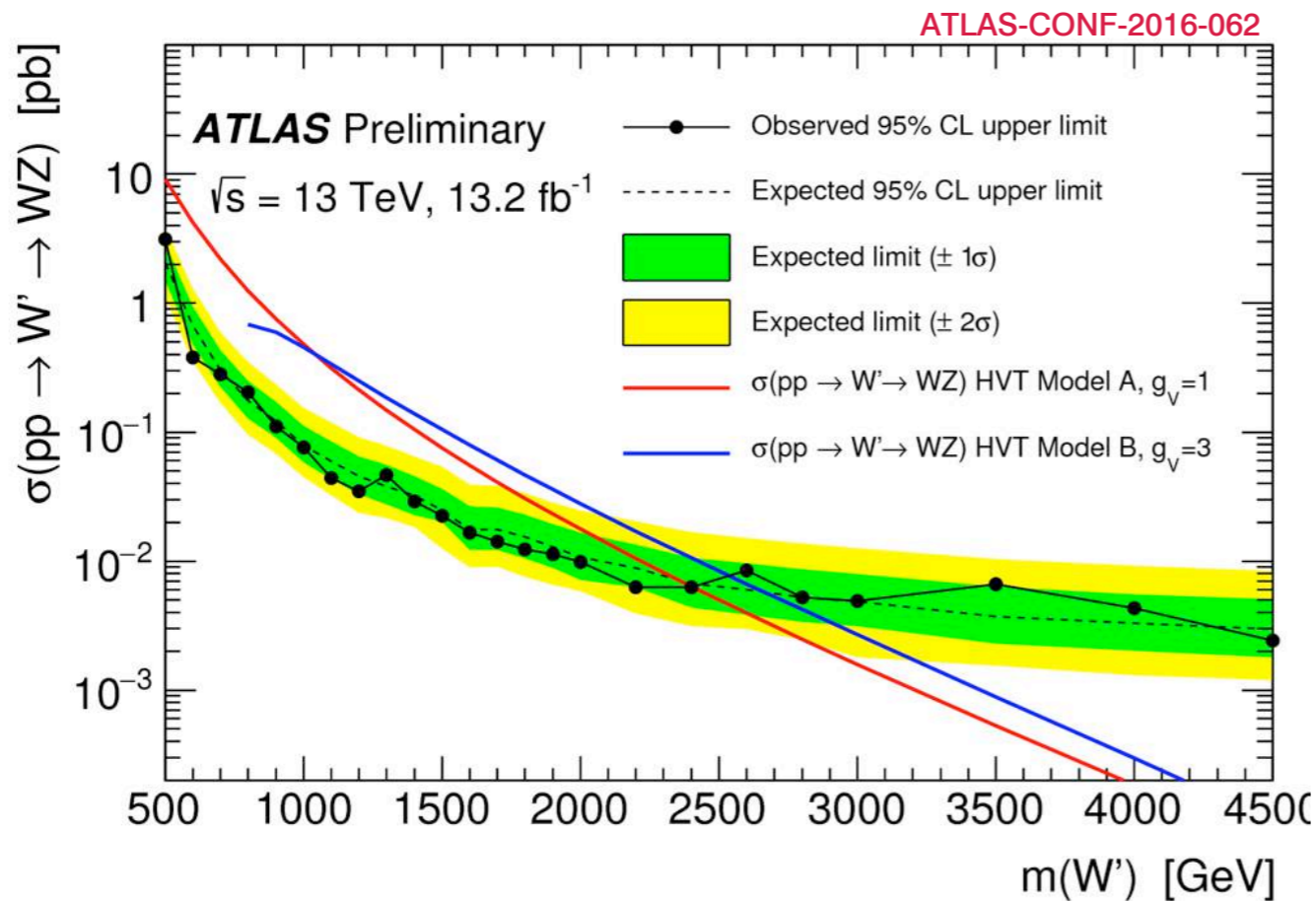
Spin-1 resonance searches:



enhanced by large couplings from the composite sector

through mixing with the SM W:
suppressed by large couplings from the composite sector

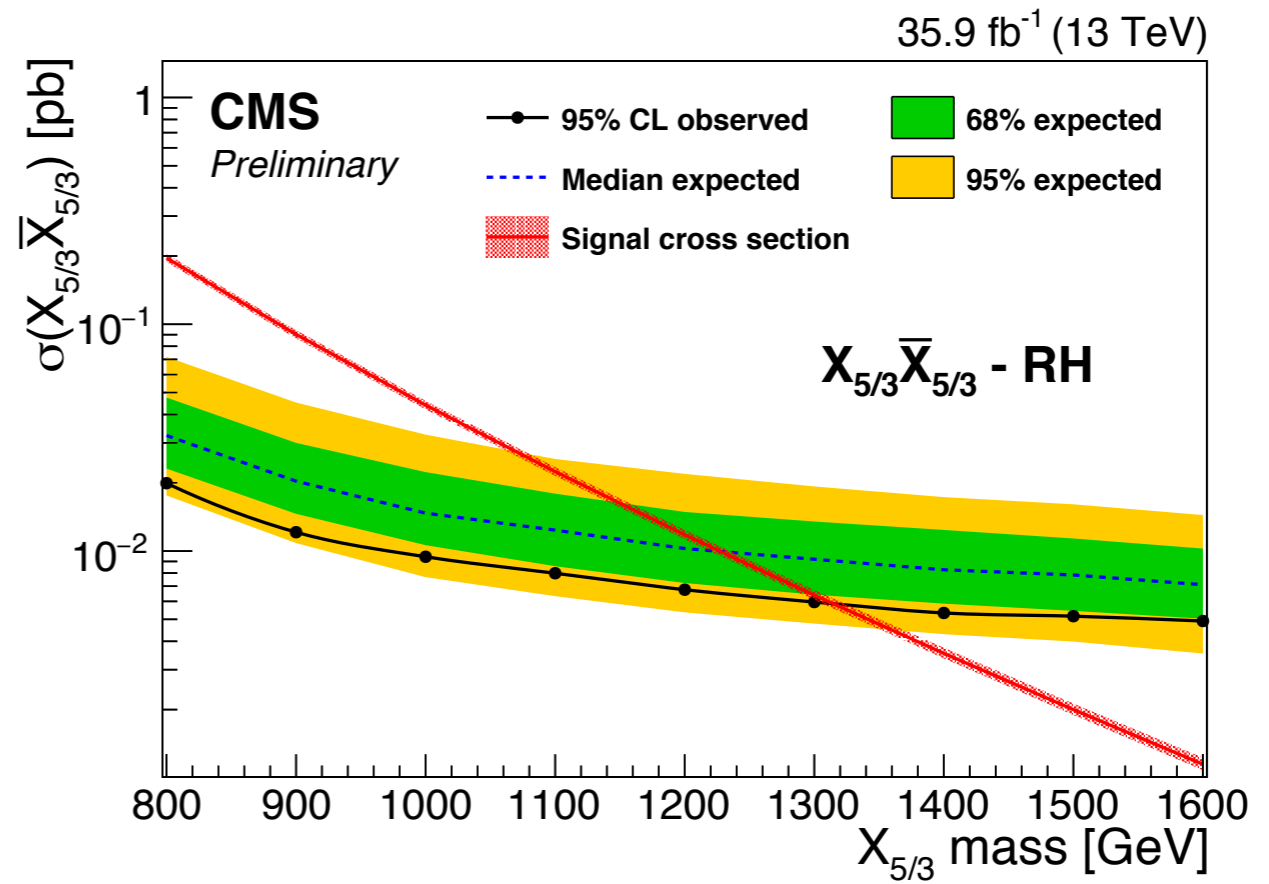
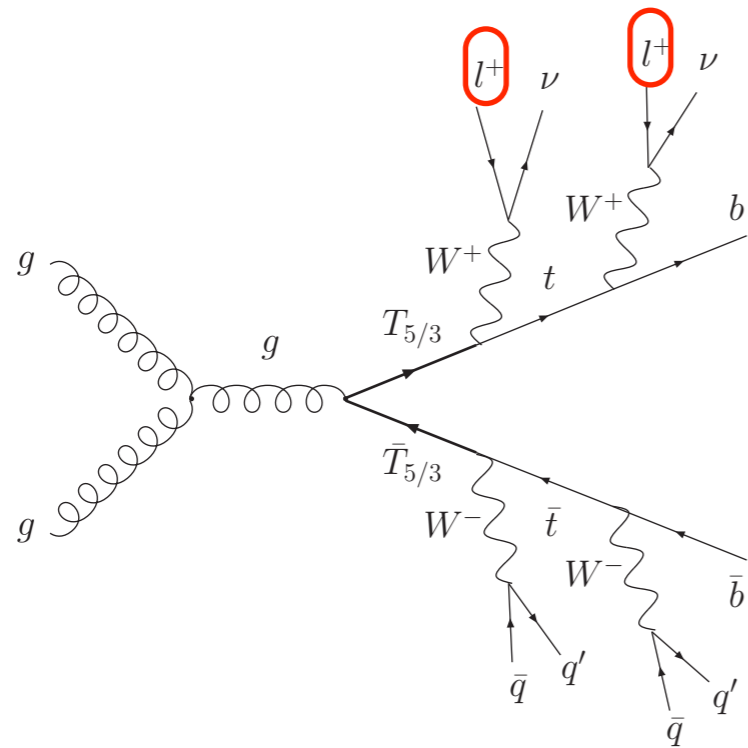
$W' \rightarrow \ell \nu qq$



$$m(W') \gtrsim 2.5 \text{ TeV}$$

scratching
the interesting regions!

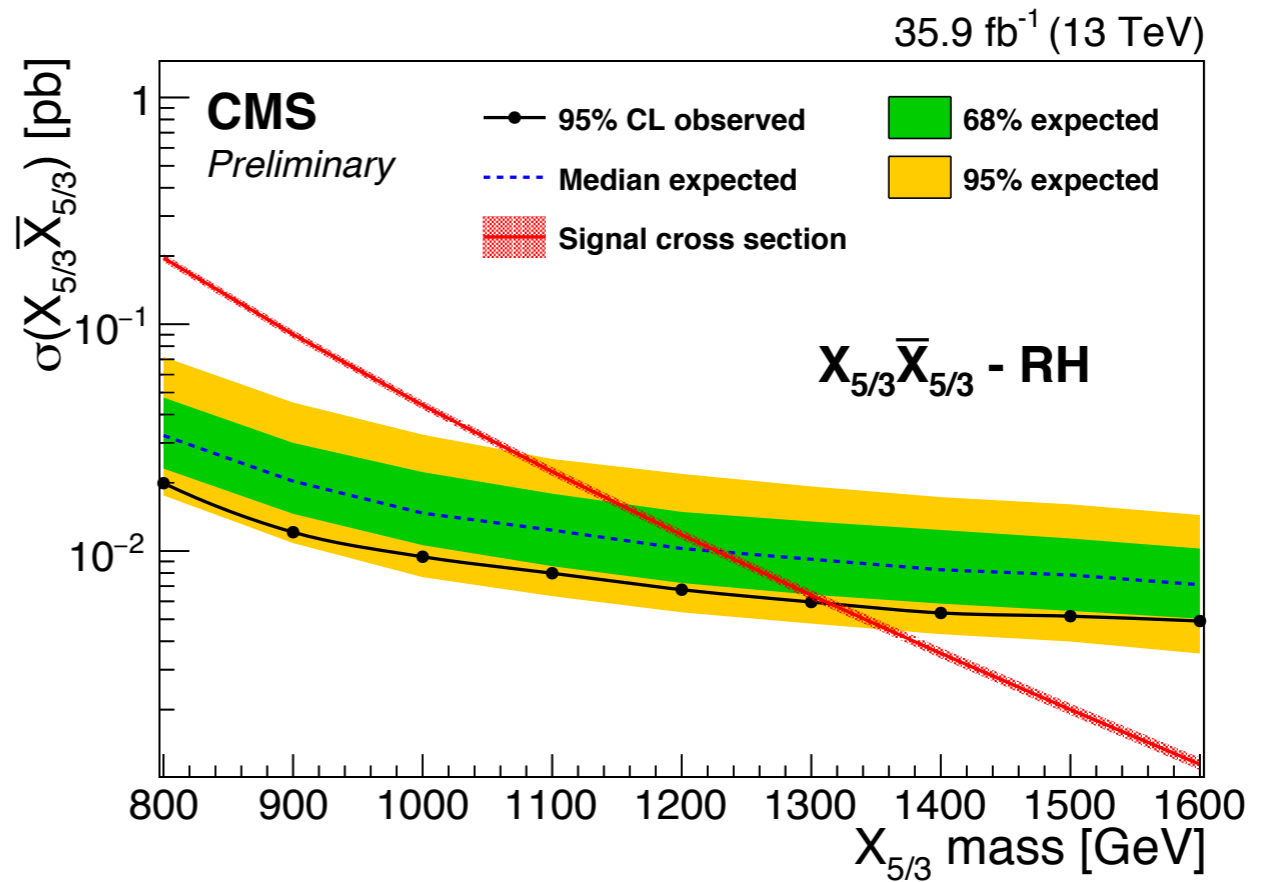
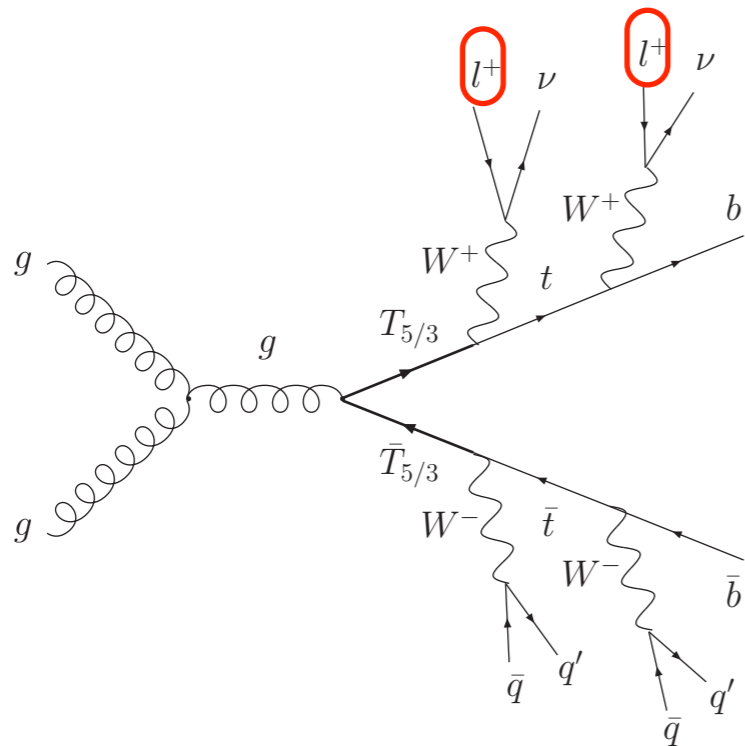
Colored fermion resonances at LHC 13 TeV



First important
constraint
from LHC:

$$M(X_{5/3}) \gtrsim 1.3 \text{ TeV}$$

Colored fermion resonances at LHC 13 TeV



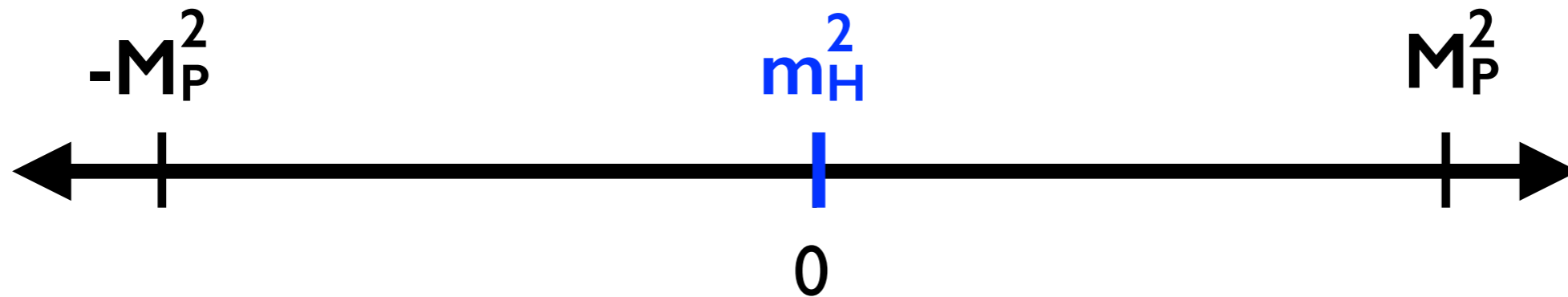
First important
constraint
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$$M(X_{5/3}) \gtrsim 1.3 \text{ TeV}$$

The situation starts being worrisome..
but not yet desperate

||

Symmetry approach: $m_H \rightarrow 0$ a special point?



Why special? More symmetry?

Not in the SM but possible if extended

Supersymmetry: boson ↔ fermion

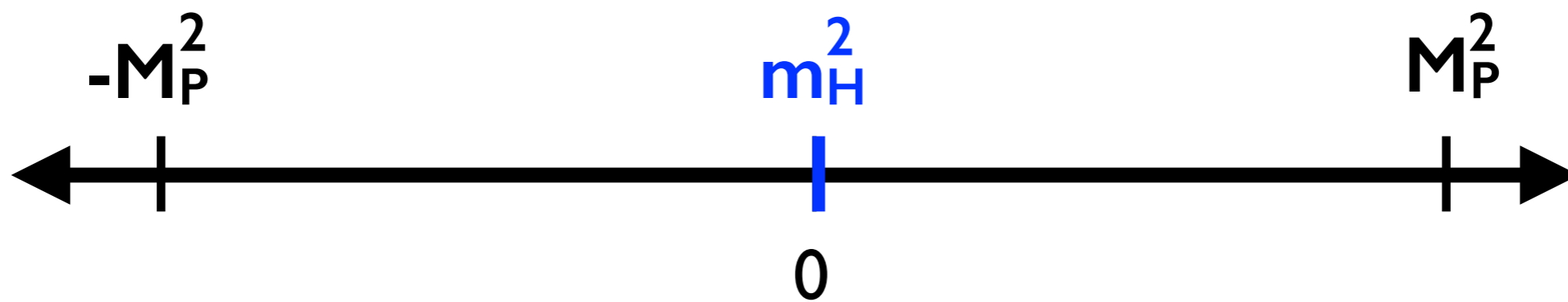
$$m_{\text{boson}} \leftrightarrow m_{\text{fermion}}$$

$m_{\text{fermion}} = 0$ of a fermion can be guaranteed:

	Massless	Massive
Fermion (charged)	2 dof Ψ_L	4 dof Ψ_L, Ψ_R



Supersymmetry guarantees $m_{\text{boson}}=0$



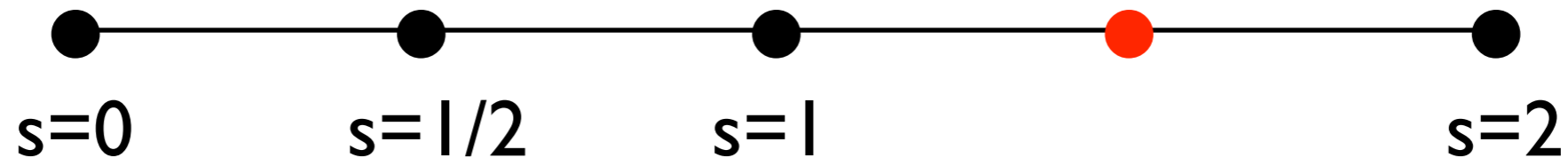
Supersymmetric point!

Lesson from the present SM of particles

Symmetries must be *accidental*
or needed for consistency
of the model

SM+Gravity:

Consistent theory of $s=0$, $s=1/2$, $s=1$ and $s=2$ particles

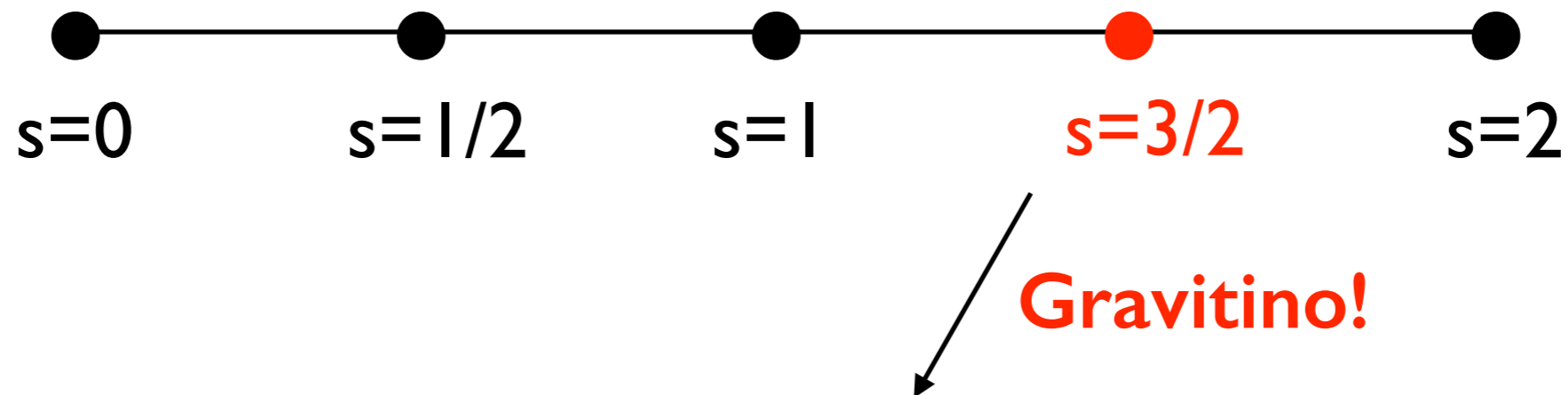


Lesson from the present SM of particles

Symmetries must be *accidental*
or needed for consistency
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SM+Gravity \longrightarrow Add spin=3/2 particle

Consistent theory of $s=0$, $s=1/2$, $s=1$ and $s=2$ particles



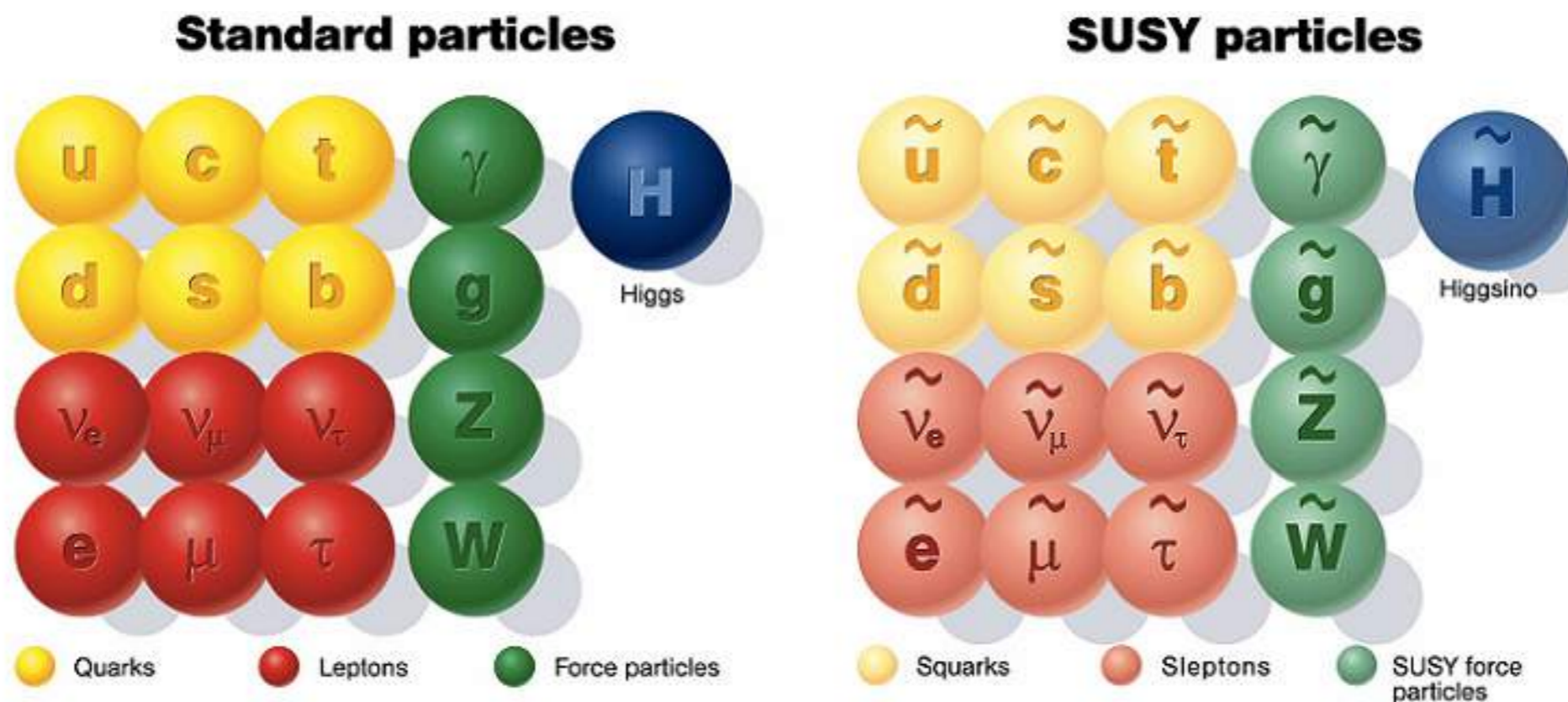
Consistent theory of spin-3/2 must have **supersymmetry!**

Imposing supersymmetry to the SM \Rightarrow **MSSM**

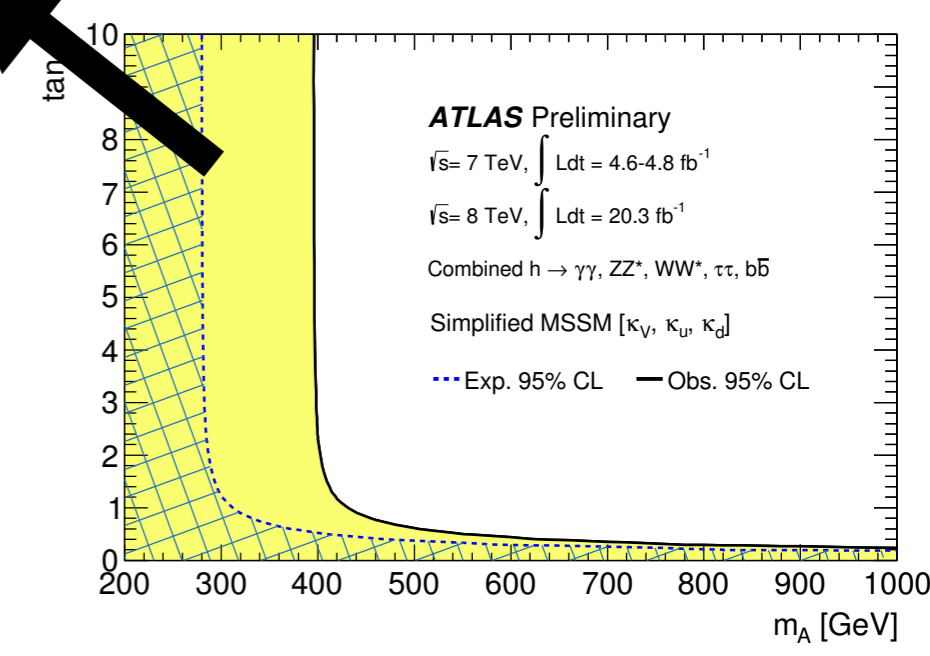
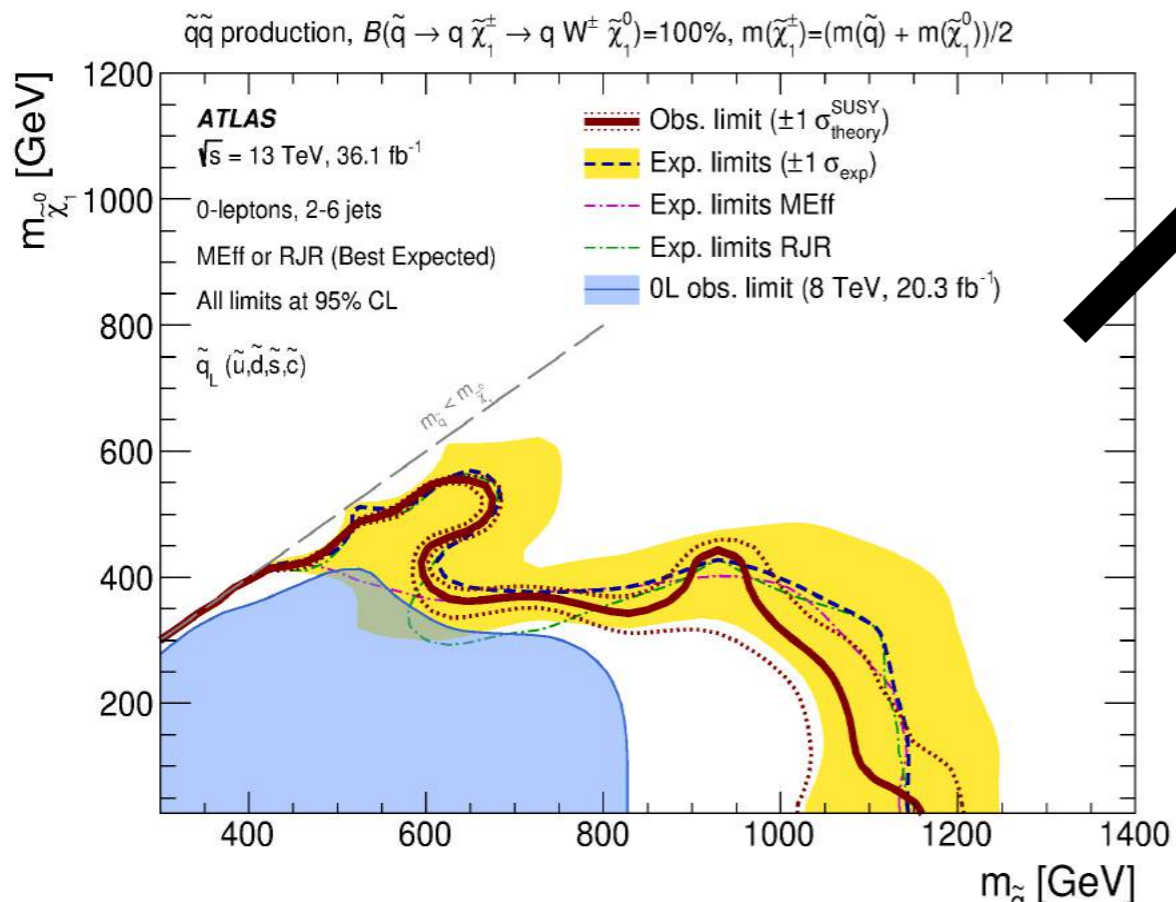
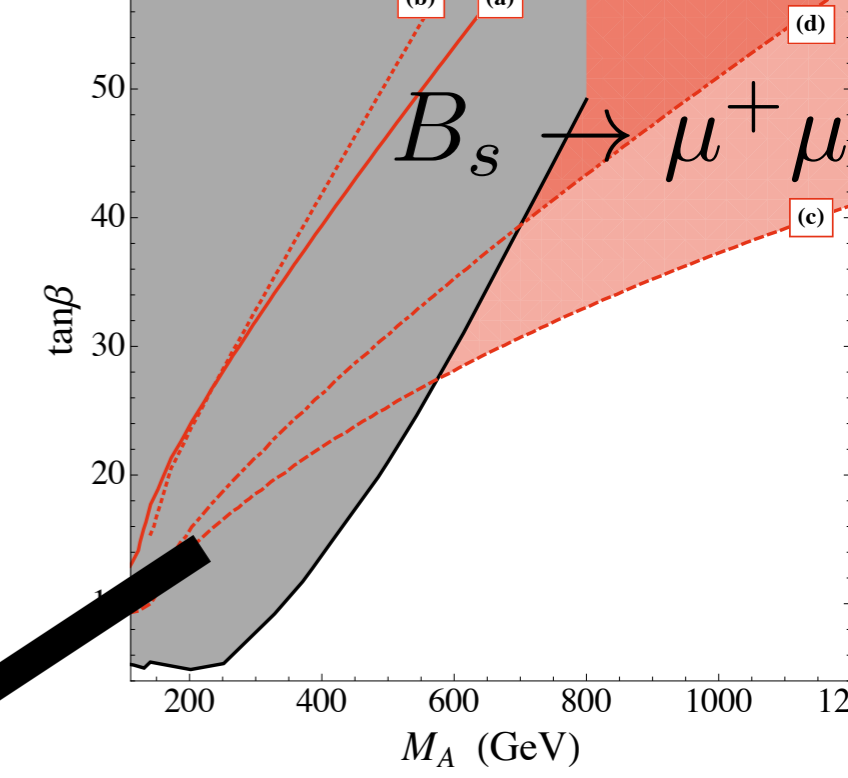
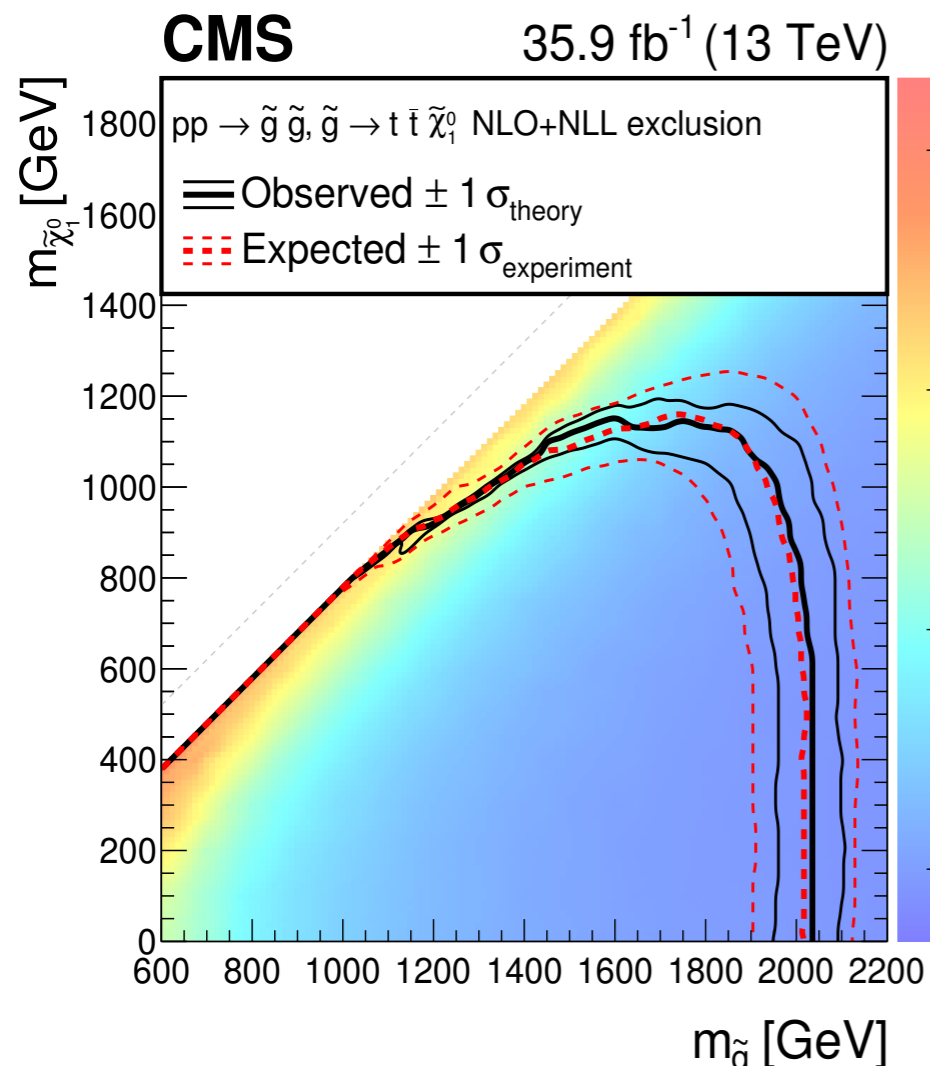
The spectrum is doubled:

SM fermion \Rightarrow New scalar

SM boson \Rightarrow New majorana fermion



we must break supersymmetry to give them mass



Higgs coupling measurements

New-Physics at the TeV

Pros

Hierarchy problem

Cons

No new particles seen,
no new flavor-violations seen,
no deviations on Higgs couplings seen,
no deviations on Z/W couplings seen,
no WIMP detected,
no EDMs seen,

New-Physics at the TeV

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Putting Time into the game

explaining fine-tunings by *accidents* in history



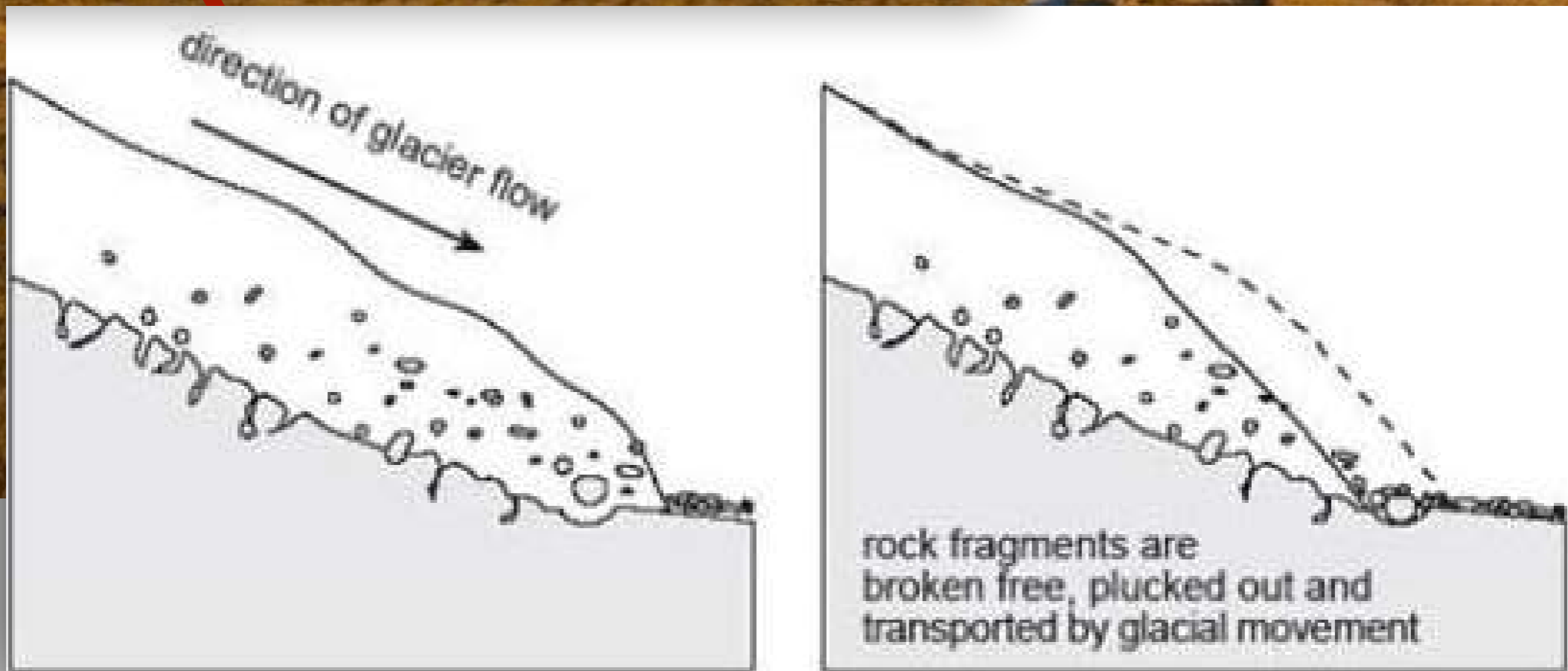
Canadian Prairies in Alberta



Canadian Prairies in Alberta

**A hierarchy
problem**

Solution to the hierarchy problem:



“Relaxation” mechanism

P.W. Graham, D.E. Kaplan, S.Rajendran
arXiv:1504.07551

Higgs-mass parameter \longrightarrow Field-dependent Higgs mass

$$m_H^2 |H|^2$$

$$m_H^2(\phi) |H|^2$$

ϕ could evolve
to a value where

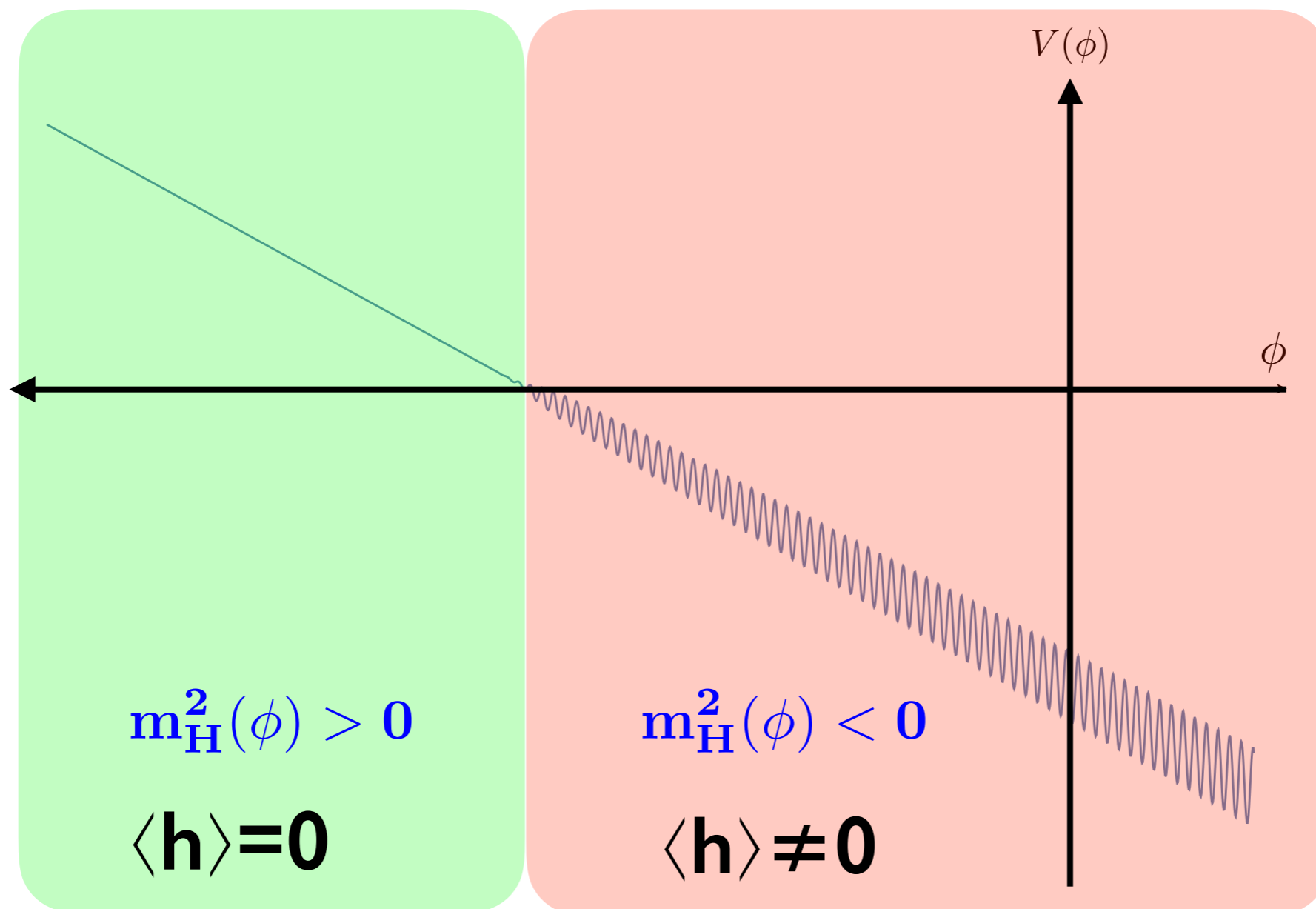
$$m_H^2(\phi) \ll M_P^2$$

Axion-like ϕ & Higgs h potential:

“technically natural”

$$V(\phi, h) = \Lambda^3 g \phi - \frac{1}{2} \Lambda^2 \left(1 - \frac{g\phi}{\Lambda} \right) h^2 + \epsilon \Lambda_c^4 \left(\frac{h}{\Lambda_c} \right)^n \cos(\phi/f)$$

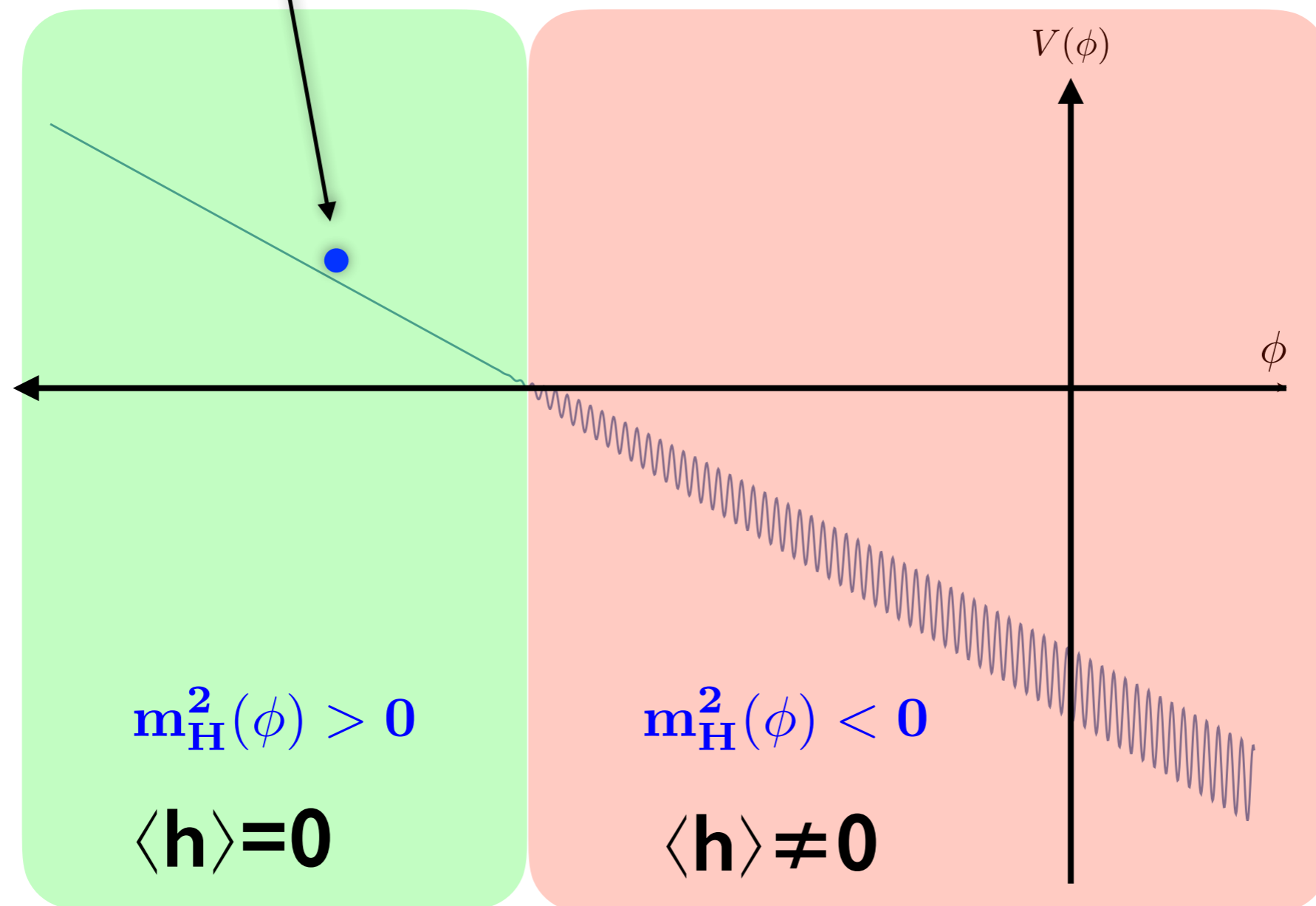
P.W. Graham, D.E. Kaplan, S.Rajendran
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Cosmological evolution can lead to a small EW scale

$$V(\phi, h) = \Lambda^3 g \phi - \frac{1}{2} \Lambda^2 \left(1 - \frac{g\phi}{\Lambda} \right) h^2 + \epsilon \Lambda_c^4 \left(\frac{h}{\Lambda_c} \right)^n \cos(\phi/f)$$

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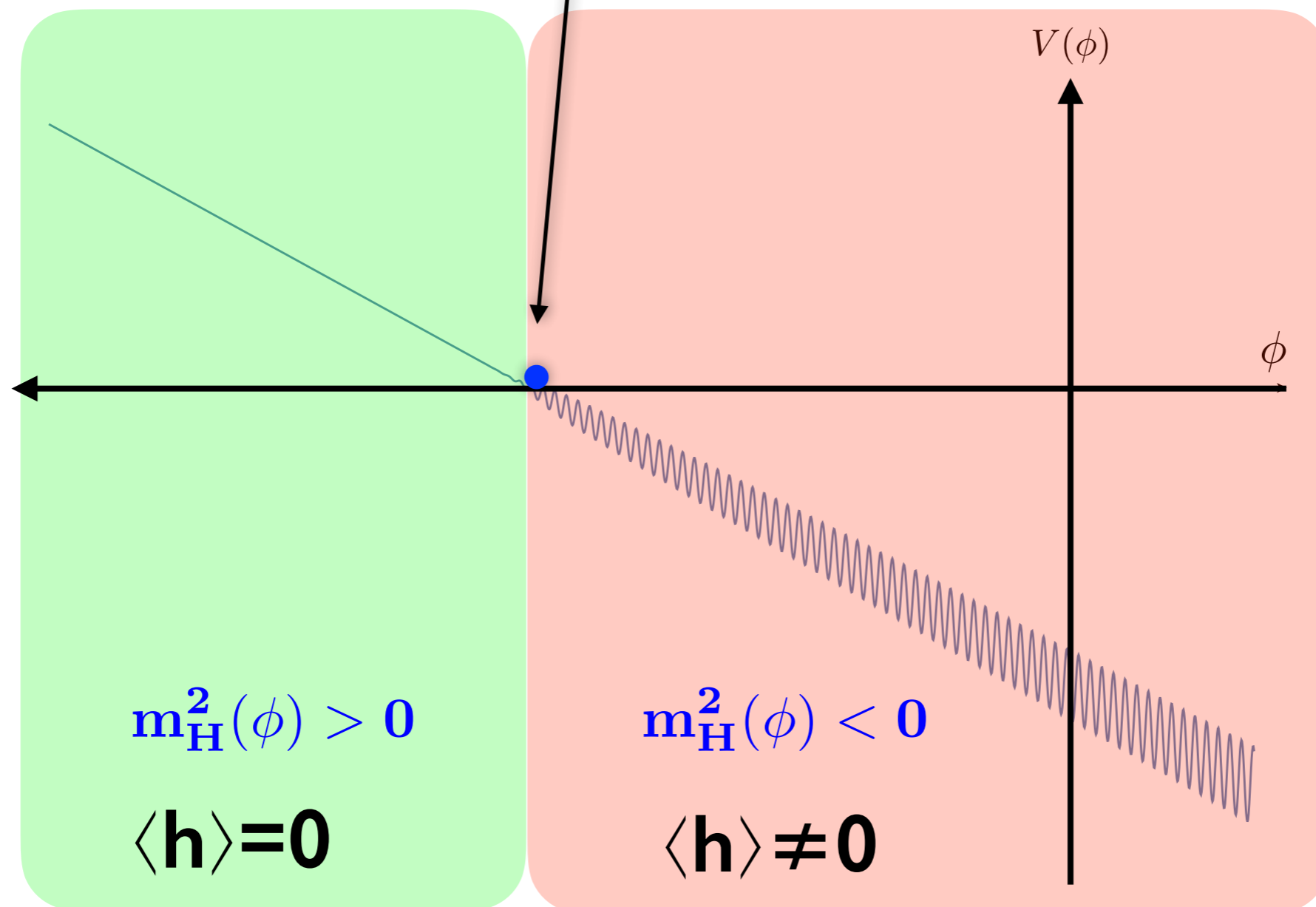


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Higgs mass-squared
turns negative: $\langle h \rangle \neq 0$

P.W. Graham, D.E. Kaplan, S.Rajendran
arXiv:1504.07551

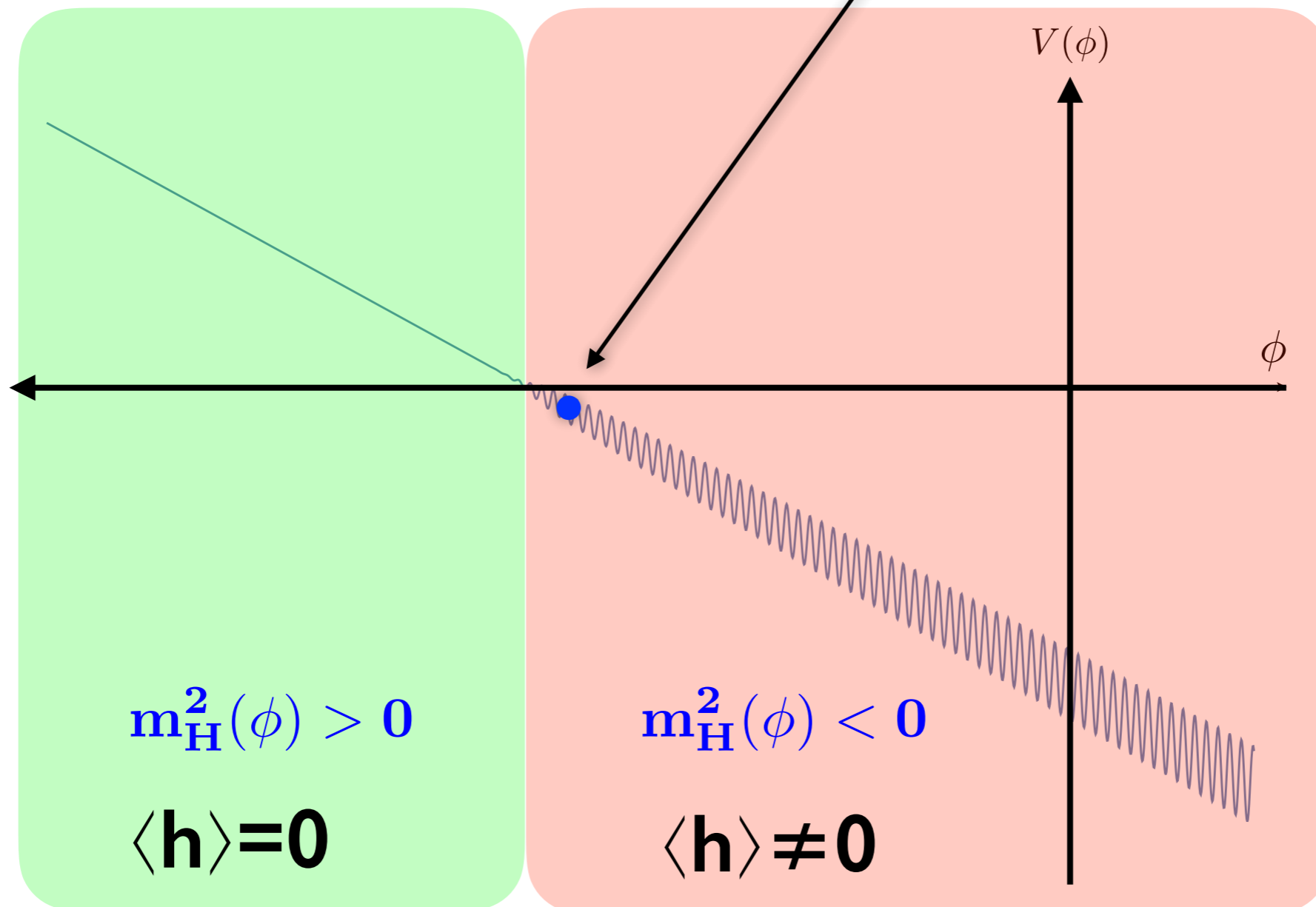


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becomes
more & more
important

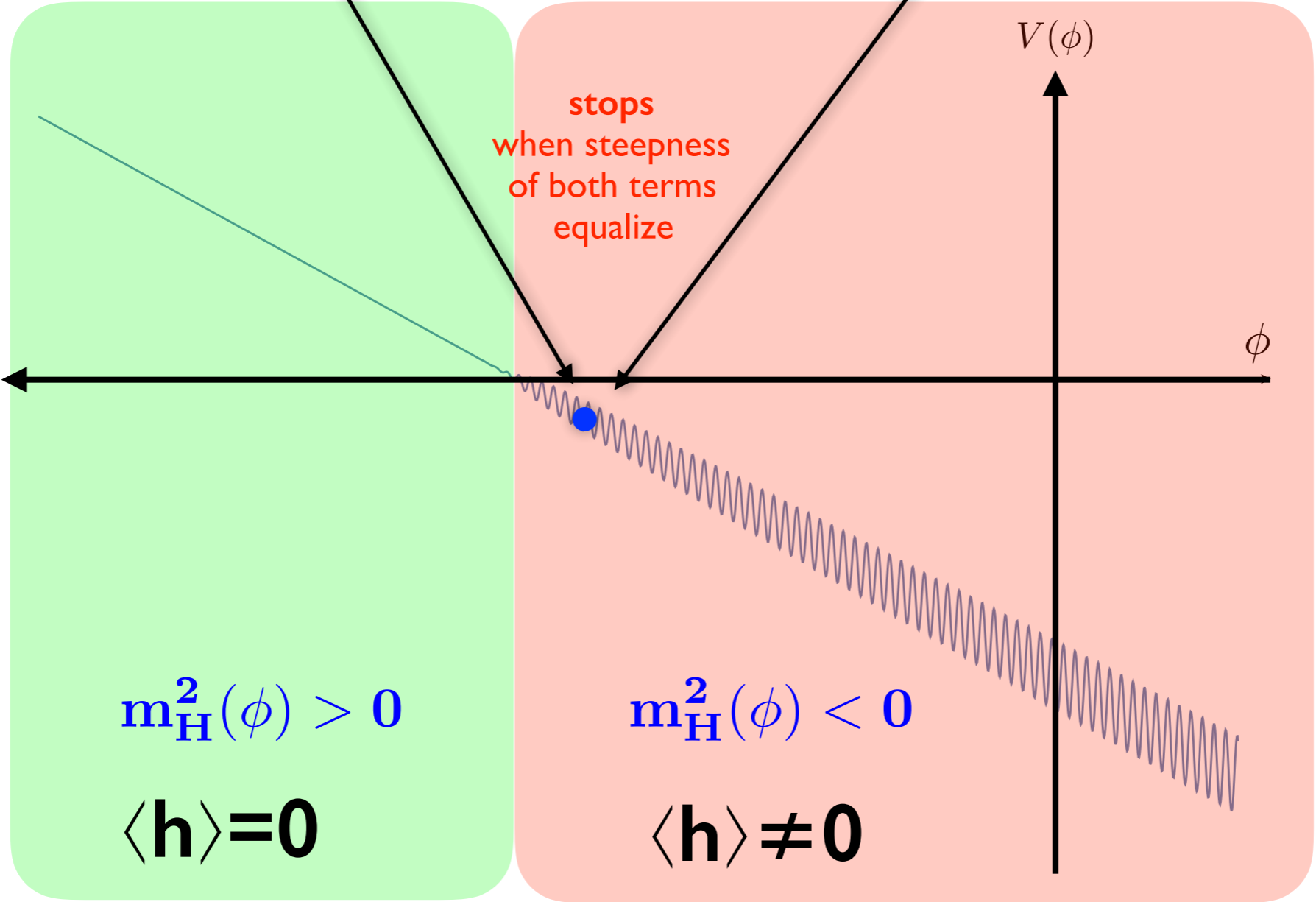
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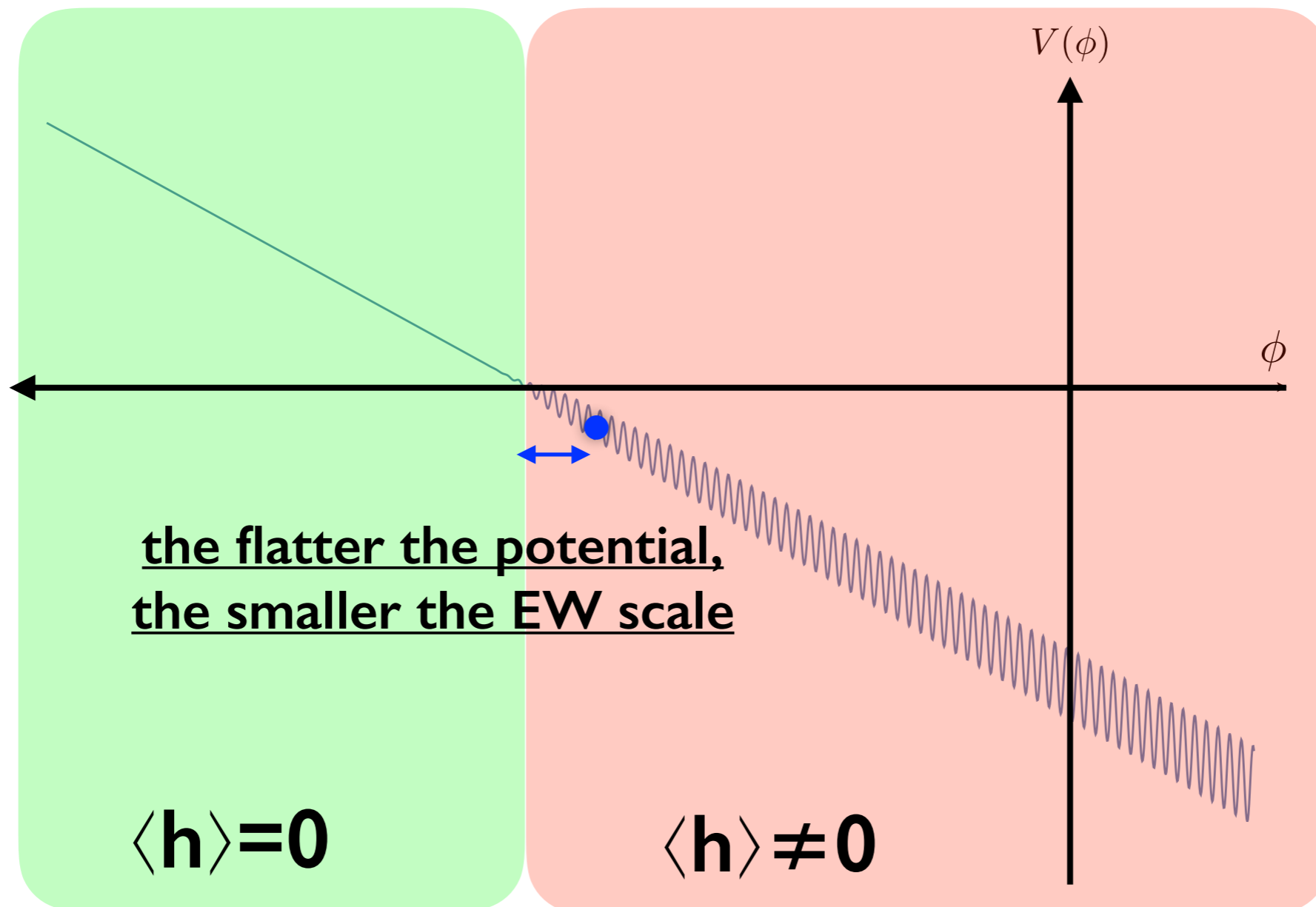
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Cosmological evolution can lead to a small EW scale

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P.W. Graham, D.E. Kaplan, S.Rajendran
arXiv:1504.07551



Tuning the initial conditions?



Tuning the initial conditions?



No, if slow rolling due to a friction:
possible in the **inflationary epoch!** (Hubble friction)

can be neglected \rightarrow
$$\ddot{\phi} + 3H_I \dot{\phi} = -\partial_{\phi} V(\phi)$$

Tuning the initial conditions?



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Long period of inflation needed,
in order for ϕ to “scan” large ranges of the Higgs mass

e-folds needed: $N_e \gtrsim \frac{H_I^2}{g^2 \Lambda^2} \sim 10^{40}$

Tuning the initial conditions?



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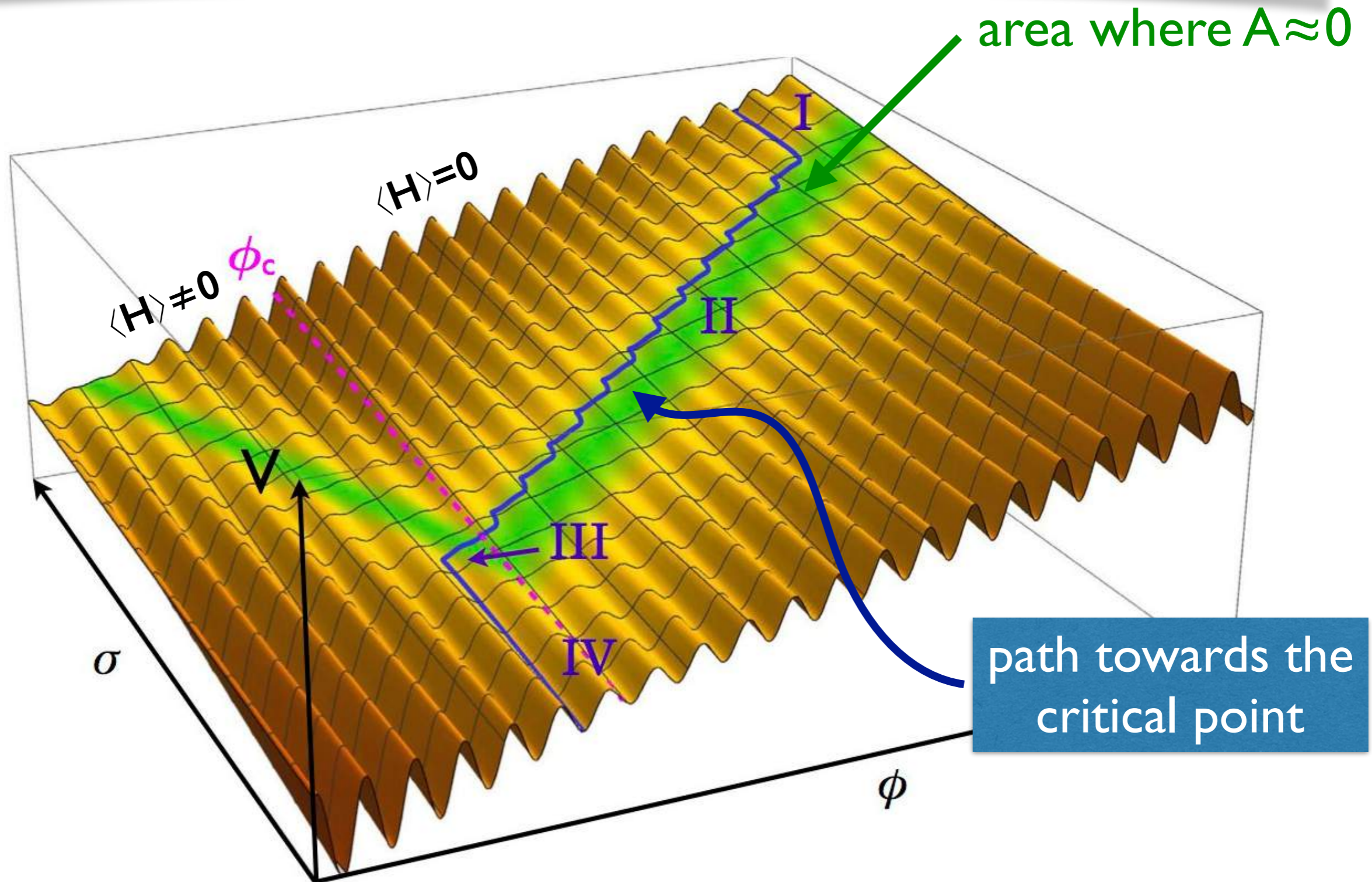
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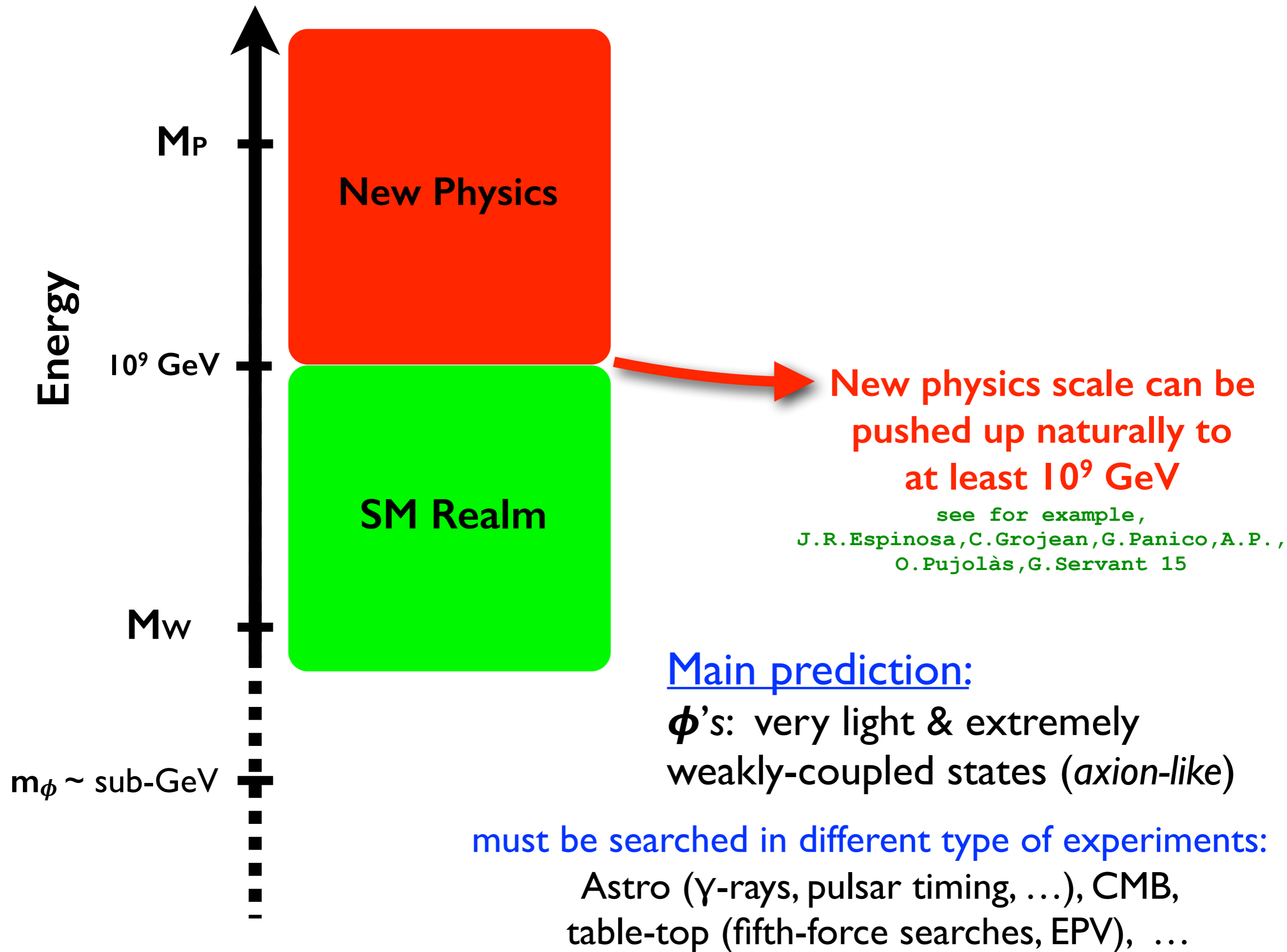
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ALPine Cosmology:

Phys.Rev.Lett. 115 (2015) 25, 251803

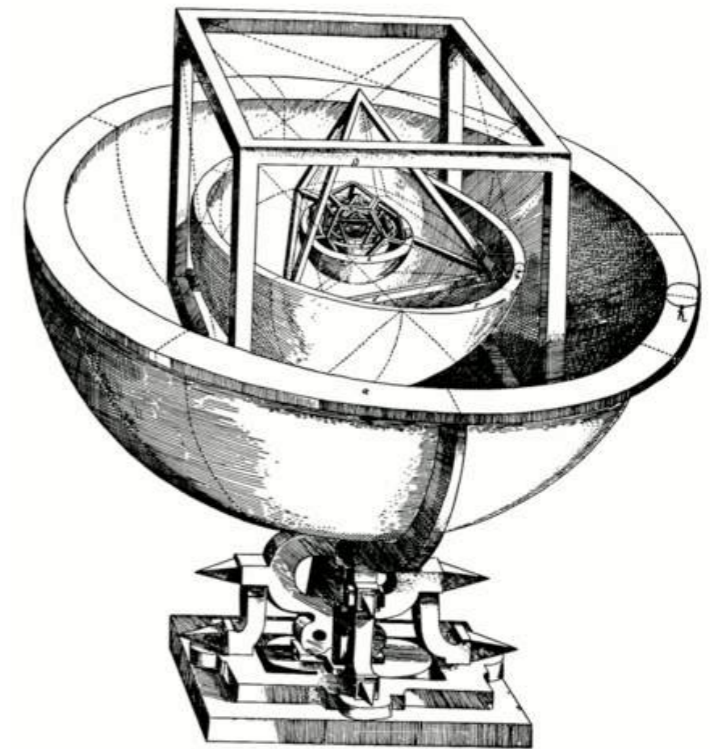
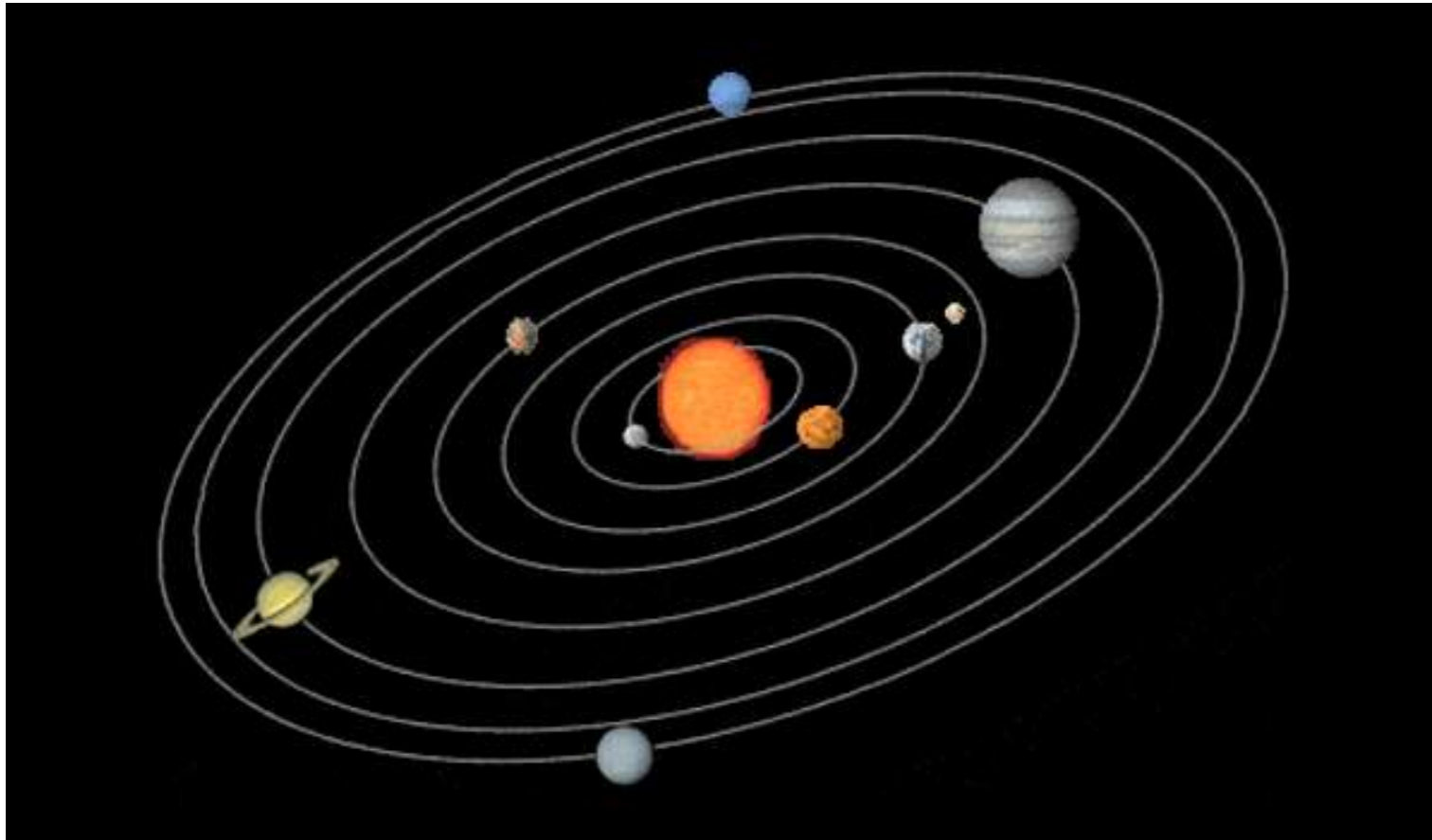
$$V(\phi, \sigma, H) = \Lambda^4 \left(\frac{g\phi}{\Lambda} + \frac{g_\sigma\sigma}{\Lambda} \right) + m^2(\phi)|H|^2 + A(\phi, \sigma, H) \cos(\phi/f)$$





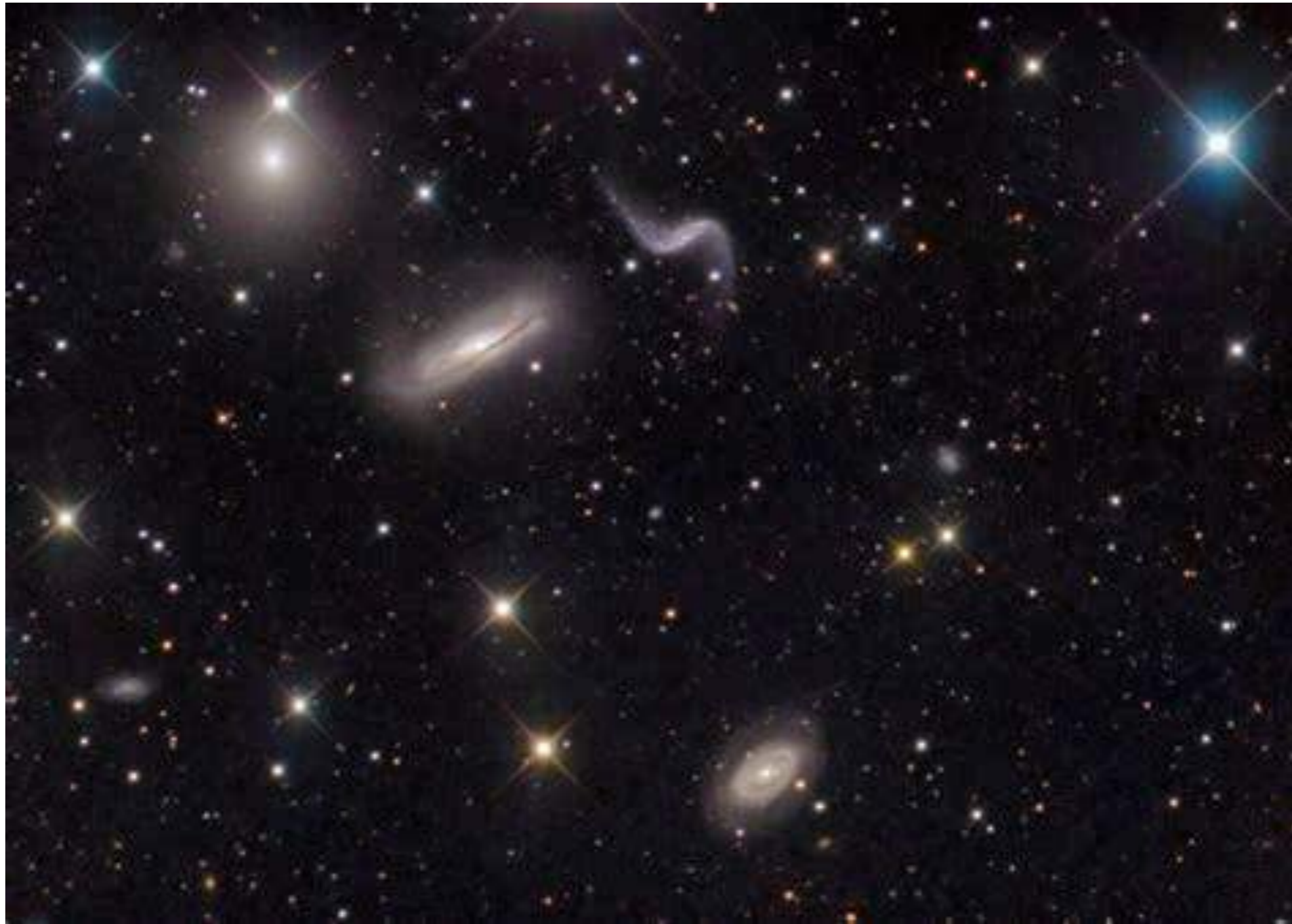
Alternative approach: *Lesson from history*

Orbit's planet: Fundamental scales?



Kepler's *Mysterium Cosmographicum*

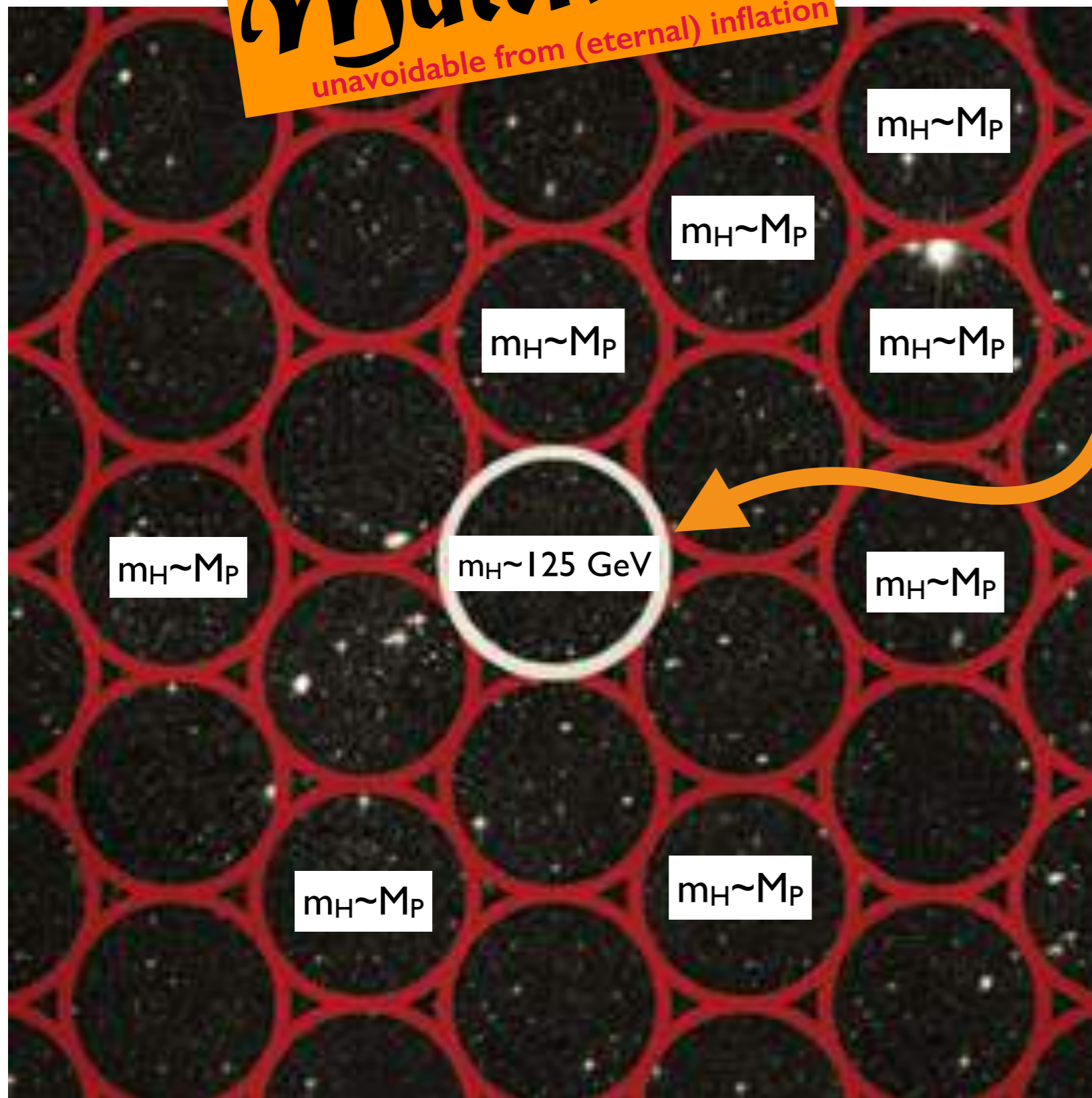
But many solar systems discovered



**Orbit's planet don't seem to be
fundamental entities**

Multiverse

unavoidable from (eternal) inflation



Our Universe is very delicate:
Change the SM parameters and could be uninhabitable

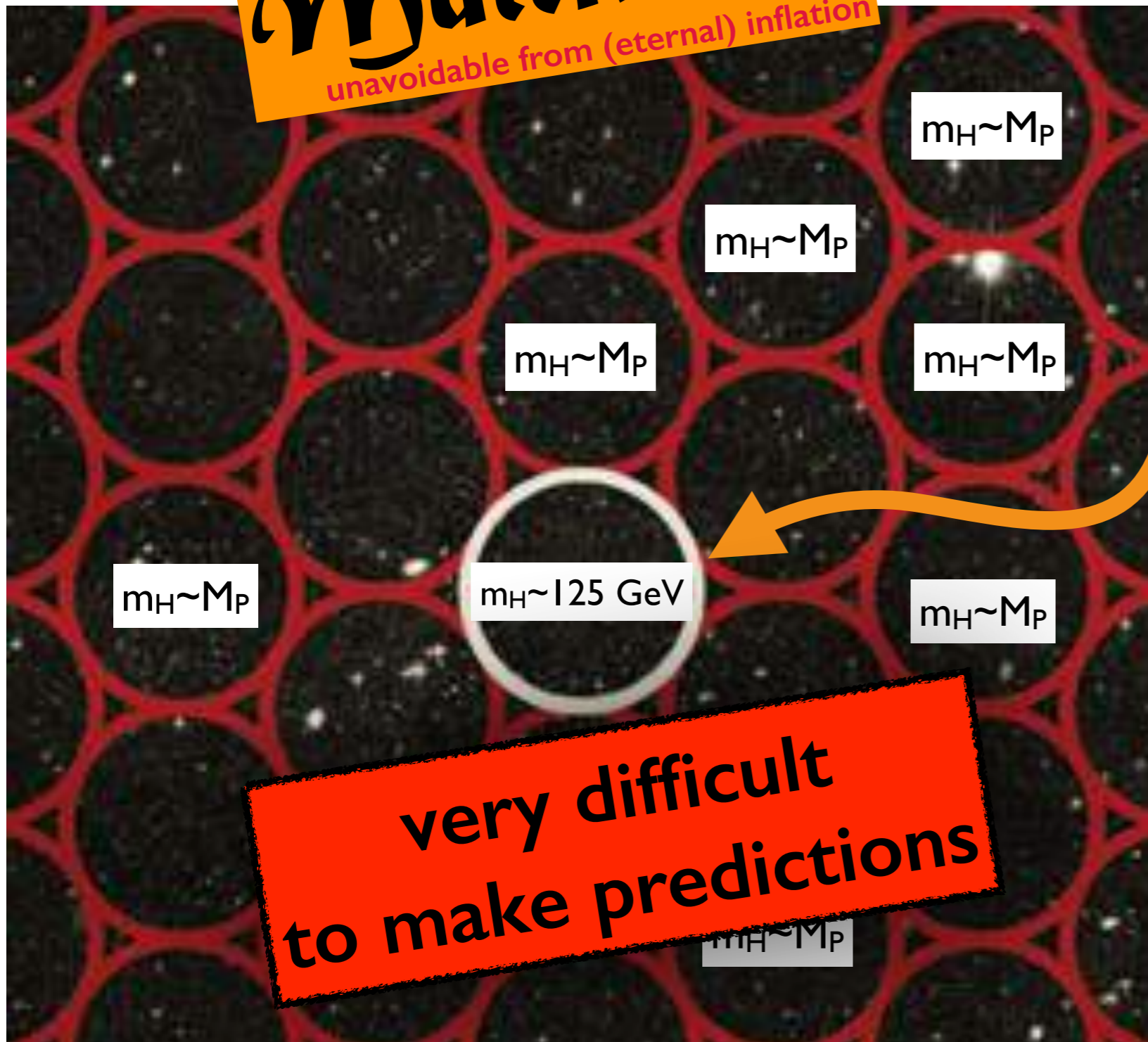
“Natural”, since only we can “live” in a Universe with these “fine-tuned” parameters



No new physics at the TeV!
(new physics in another universes)

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At present, the only scenario that could “explain” the present smallness of the cosmological constant!

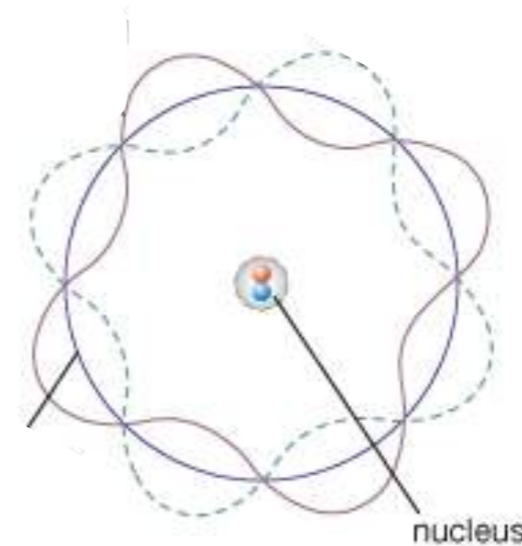
IV

“Quantum” constraints

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“Quantum” constraints

Quantum mechanics
restrict the atomic orbits:



Can some “quantum” condition restrict m_H to be small?

Weak Gravity Conjecture:

For a consistent theory of quantum gravity,
there must be a state of charge q and mass m
satisfying

Arkani-Hamed, Molt, Nicolis, Vafa 06

$$q > \frac{m}{M_{\text{P}}}$$

bound on mass scale

See Cheung, Remmen 14, for proposals along these lines

Conclusions

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➔ some proposals (*time* important player?):

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Let's enjoy a different
problem of scales:

