

Status and Results from the

XENON1T

Dark Matter Experiment

Adam Brown



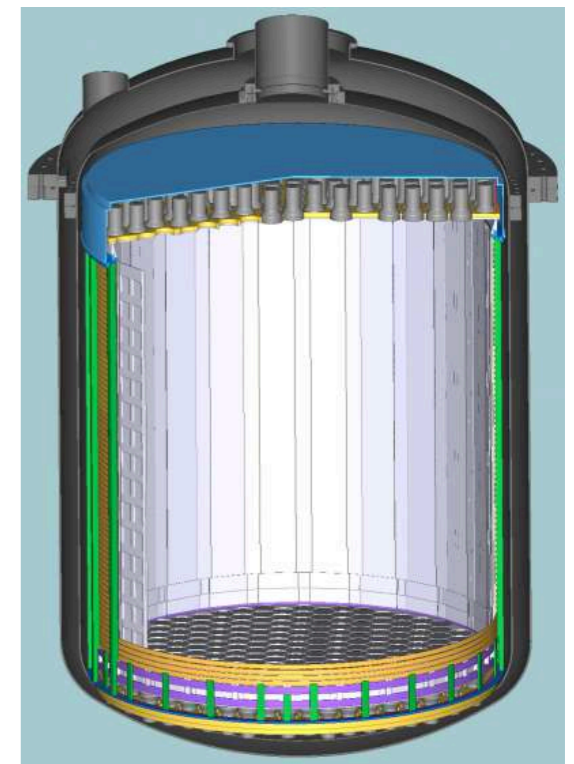
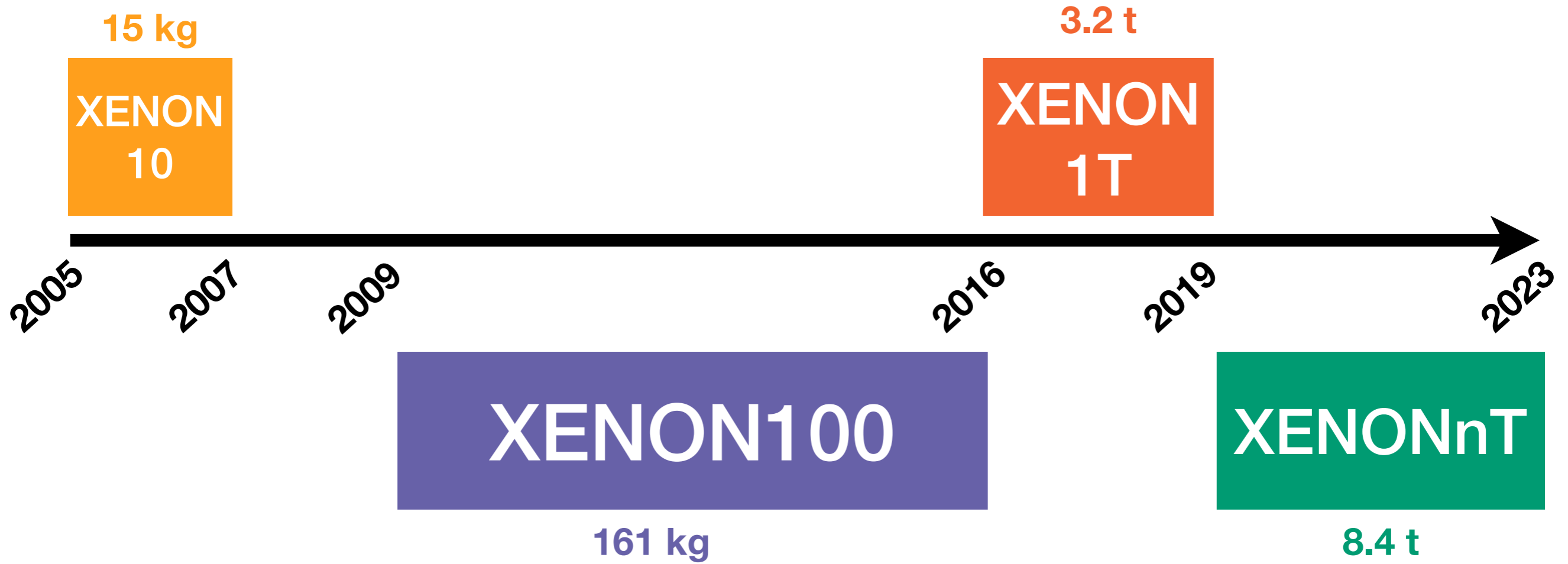
**University of
Zurich**^{UZH}

on behalf of

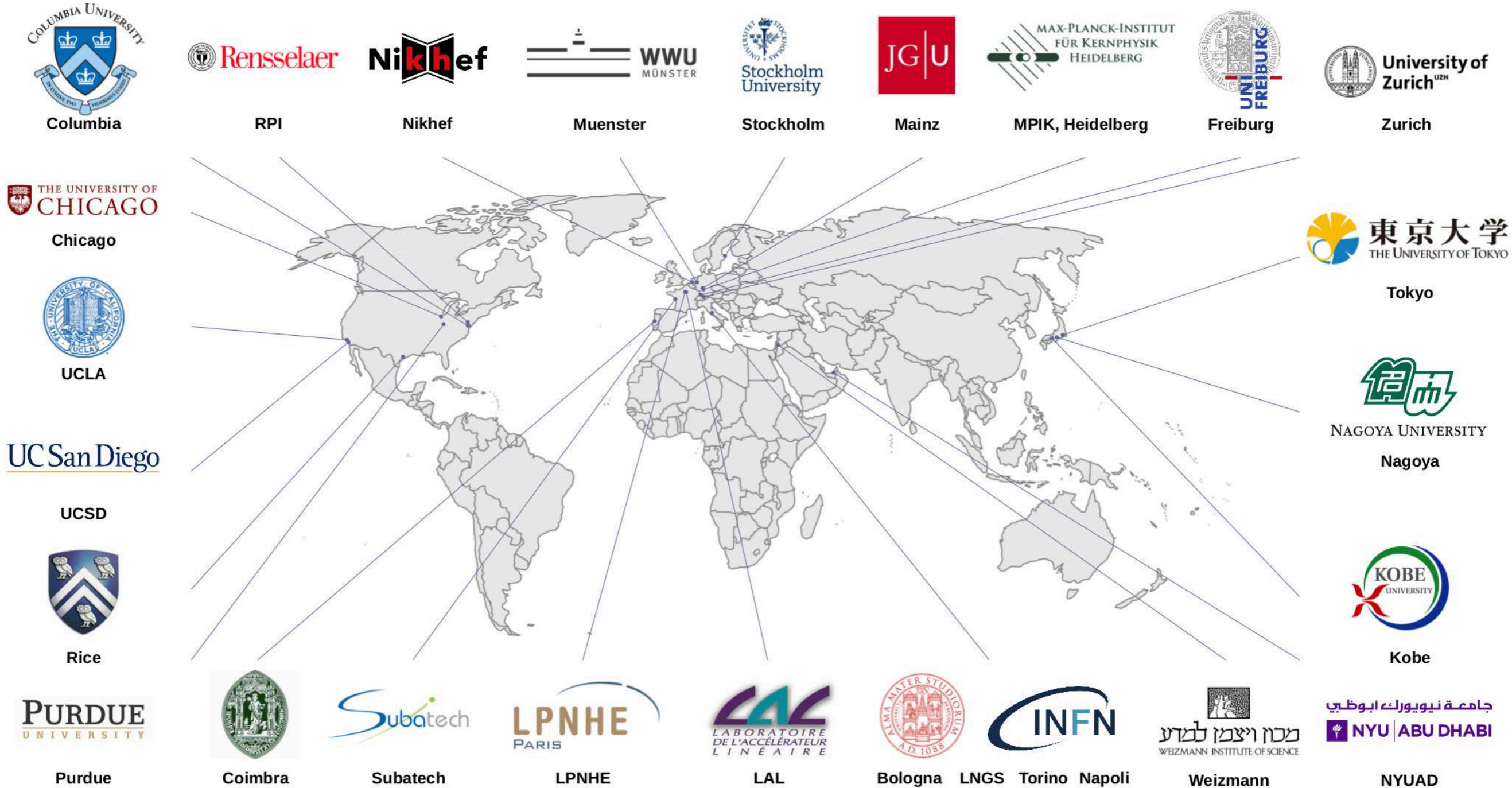


the XENON
Collaboration

Evolution of the XENON TPCs



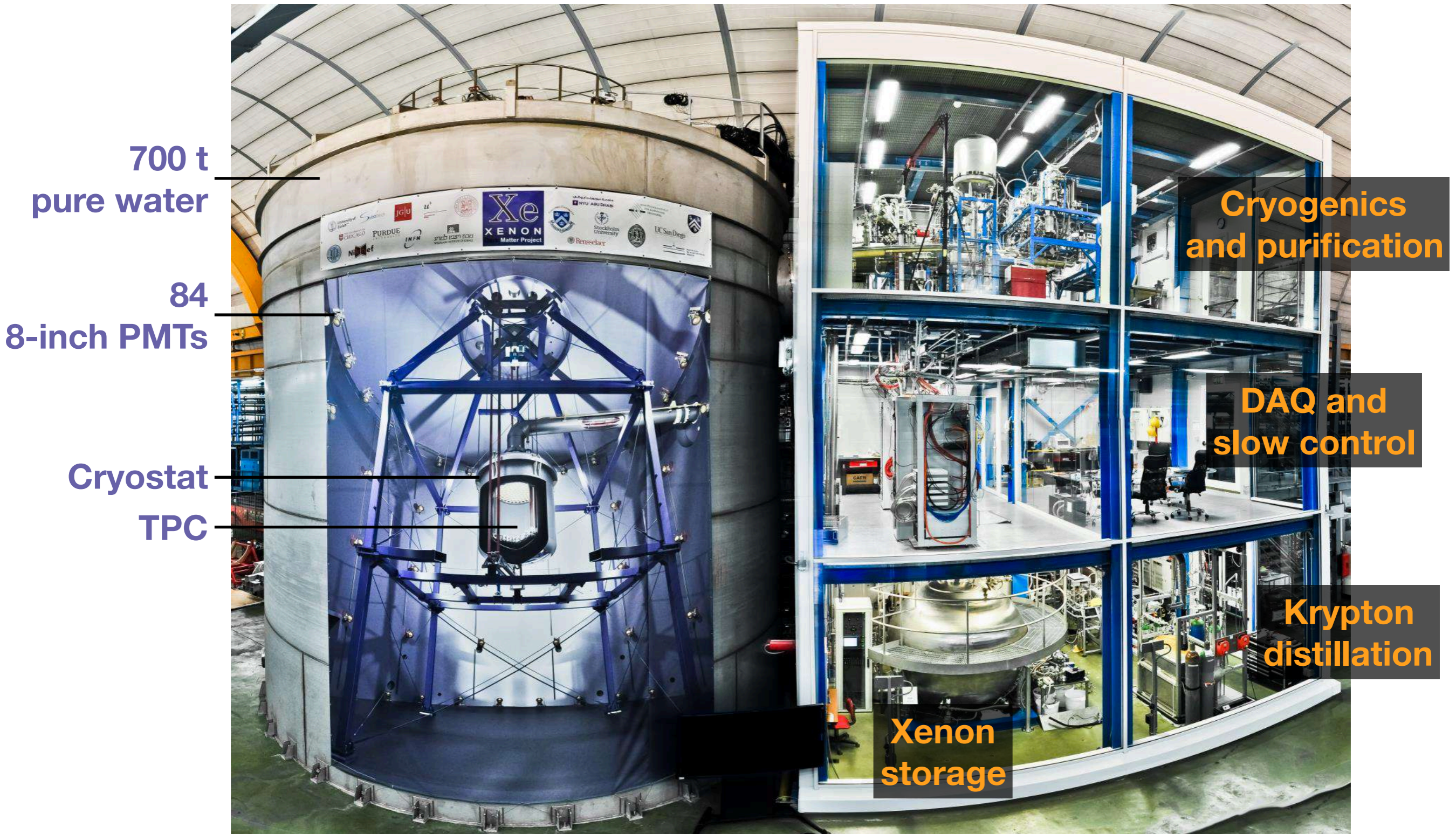
The XENON collaboration



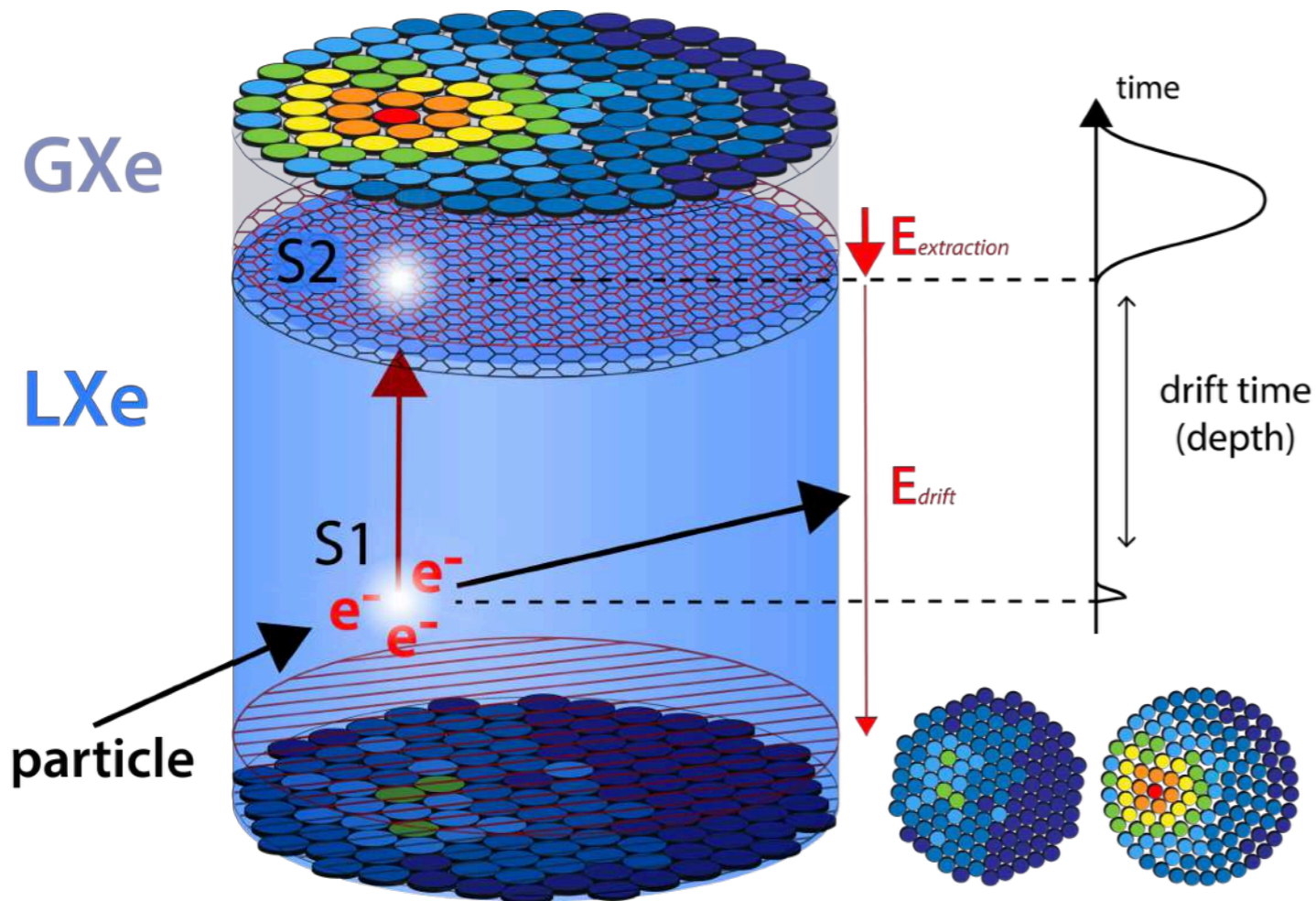
160 scientists

27 institutes

11 countries



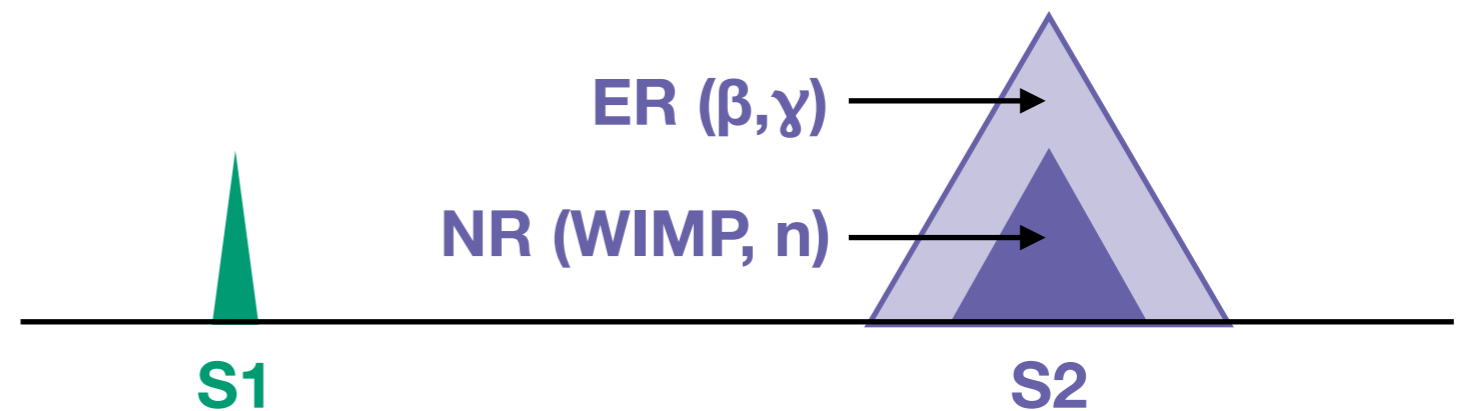
Dual-phase time projection chamber



- Particle interacts in liquid xenon
- ▶ S1: prompt scintillation signal
 - ▶ S2: electrons drifted to gas produce proportional scintillation signal
 - ▶ Heat

Position reconstruction

- ▶ (x, y) from S2 pattern on top PMTs
- ▶ z (depth) from S1 – S2 delay



Particle type identification

- ▶ S2:S1 ratio different for electronic recoils (β, γ) and nuclear recoils (WIMP, n)



TPC

holding 2 t
liquid xenon

~ 1 m diameter
~ 1 m length

Highly reflective
PTFE walls

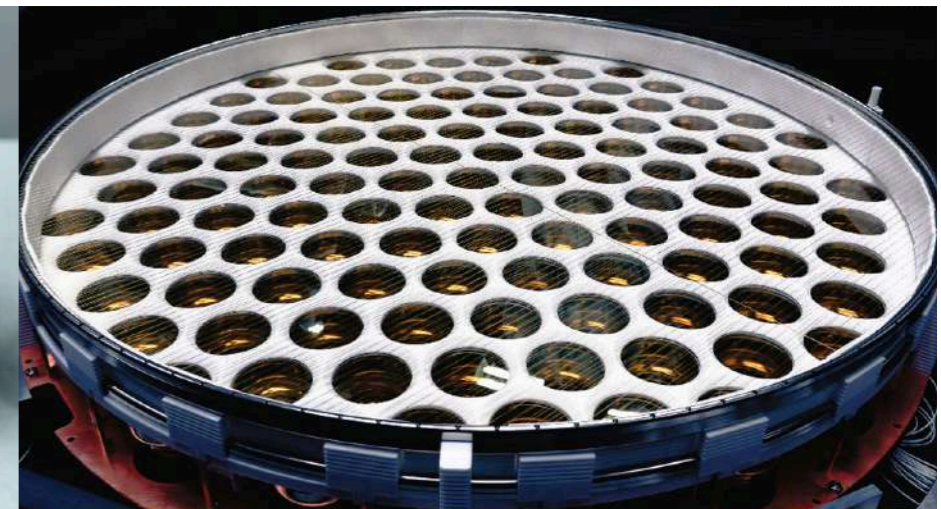
74 copper field
shaping rings

Five high-transparency
electrodes



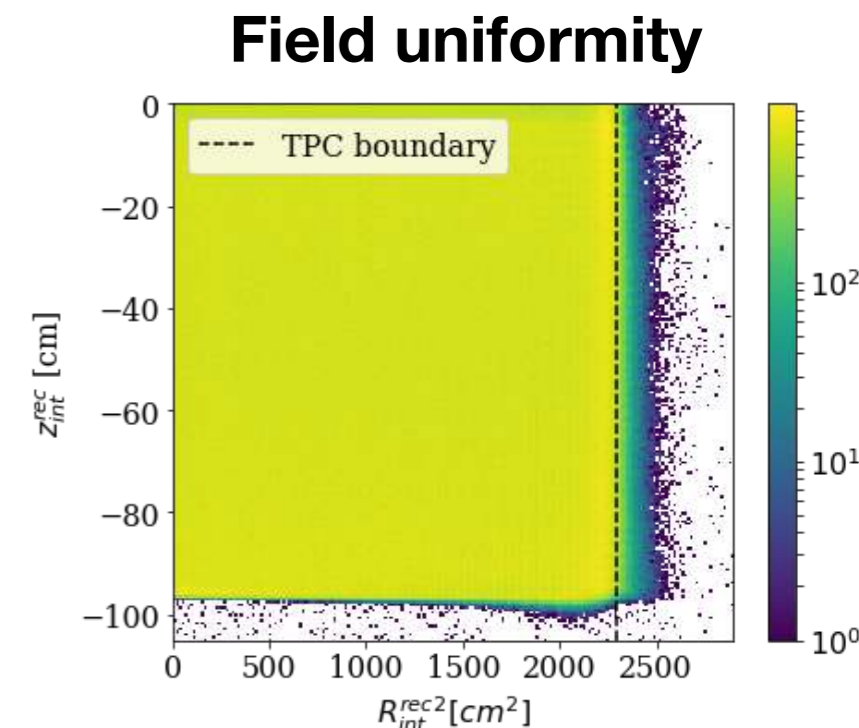
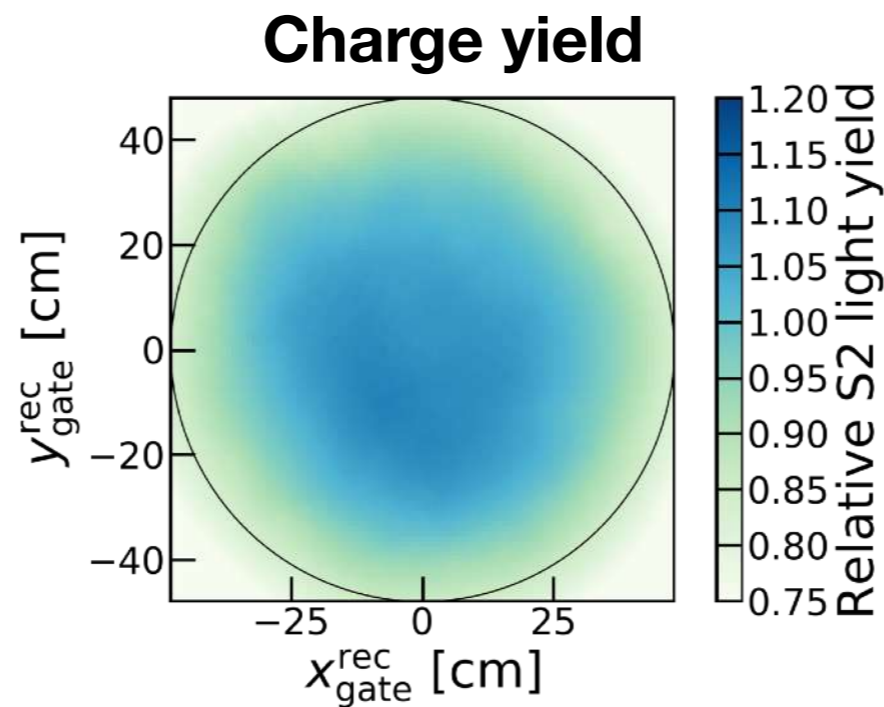
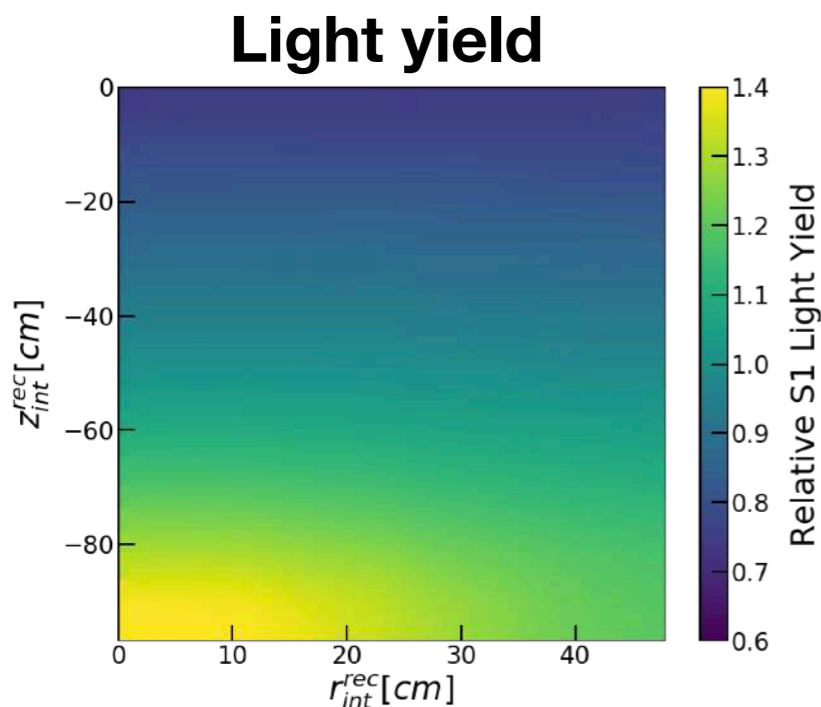
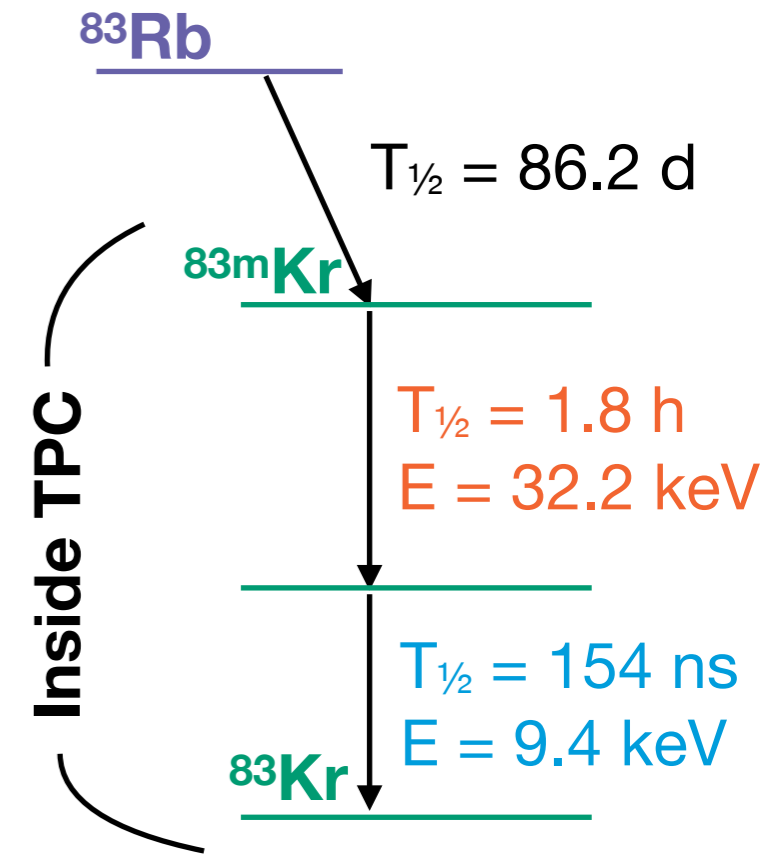
Hamamatsu R11410-21 **PMTs**

248 3-inch PMTs in XENON1T
low radioactivity & VUV-sensitive
QE ~ 35% at 175 nm



$^{83\text{m}}\text{Kr}$ as calibration source

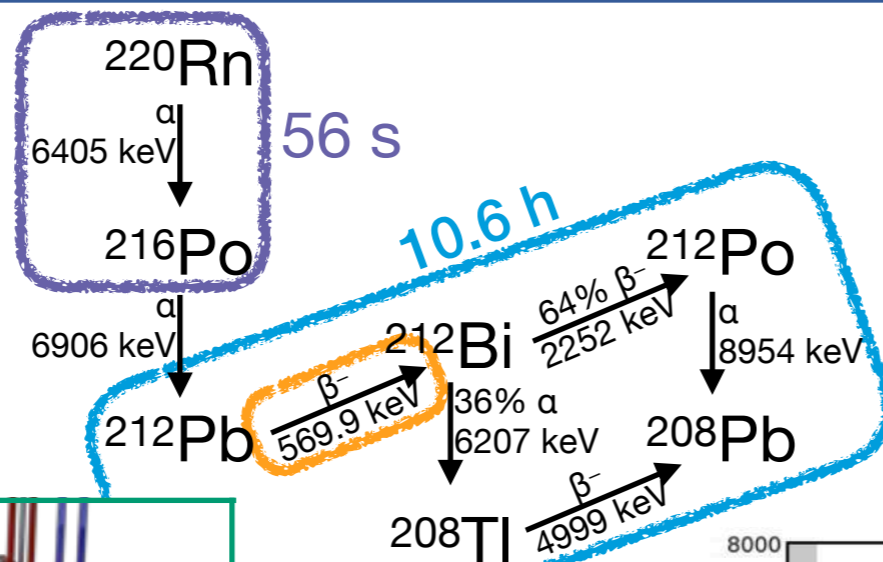
- ▶ Injected into LXe
- ▶ Double emission allows very pure identification
- ▶ Used to correct spatially varying detector response:
 - S1: light collection efficiency (x, y, z)
 - S2: electron lifetime (z), charge amplification (x, y)
 - Position: due to non-uniform field (x, y, z)



Calibration of ER / NR bands

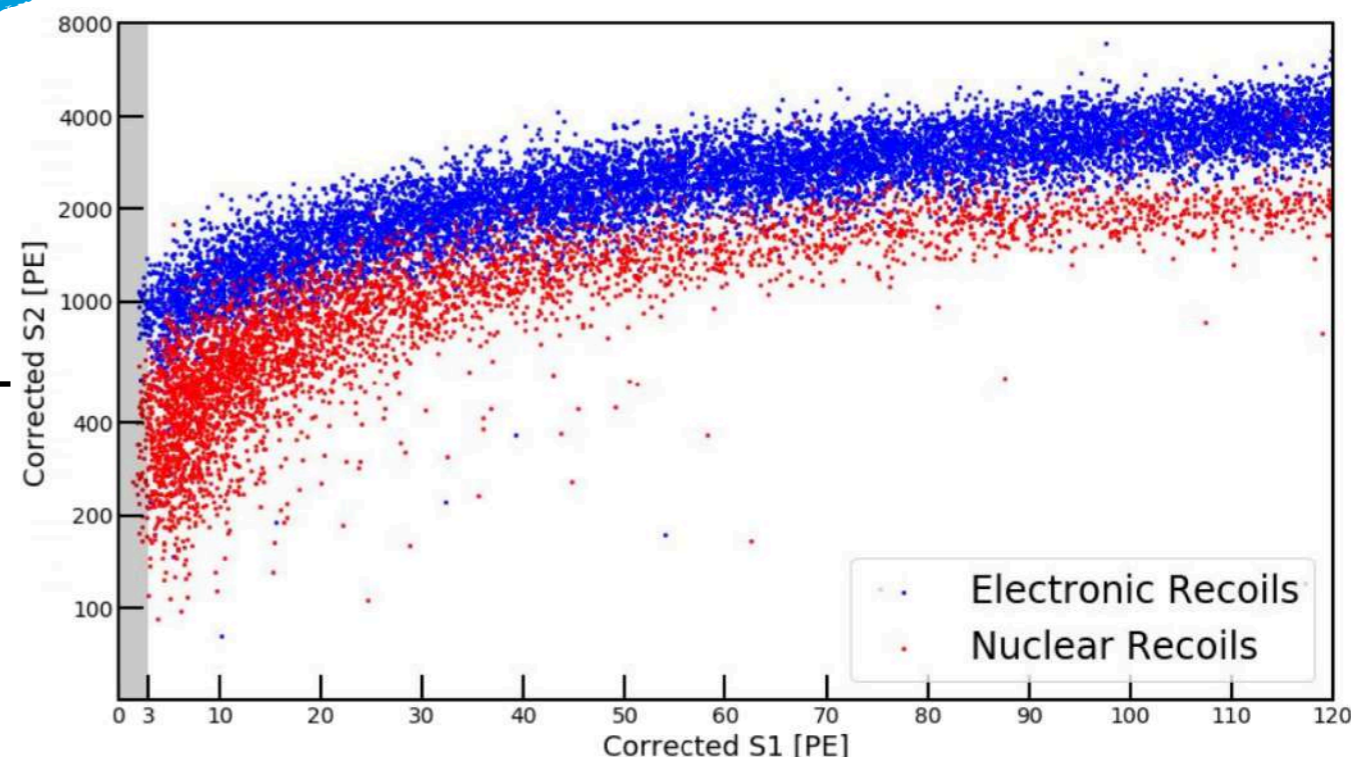
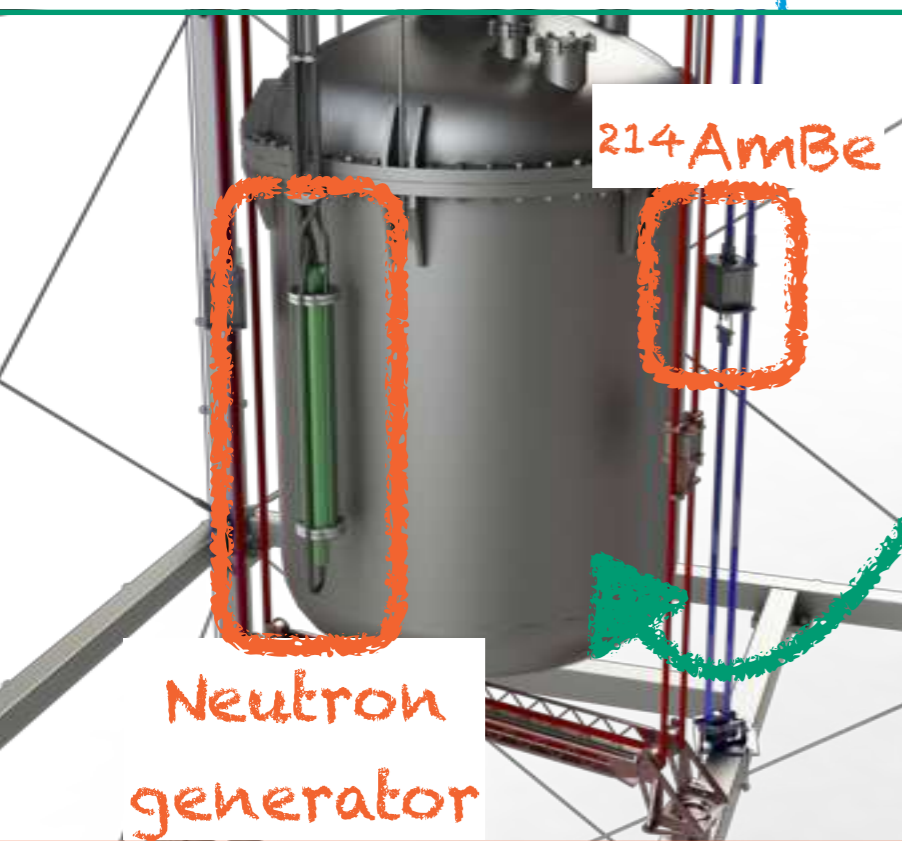
JINST 11:P04004 (2016)

^{220}Rn injected into liquid xenon
Decays away in few days



Electronic recoils

^{212}Pb β^- decay:
continuous low E spectrum



External sources mounted on belt system
D-D fusion generator

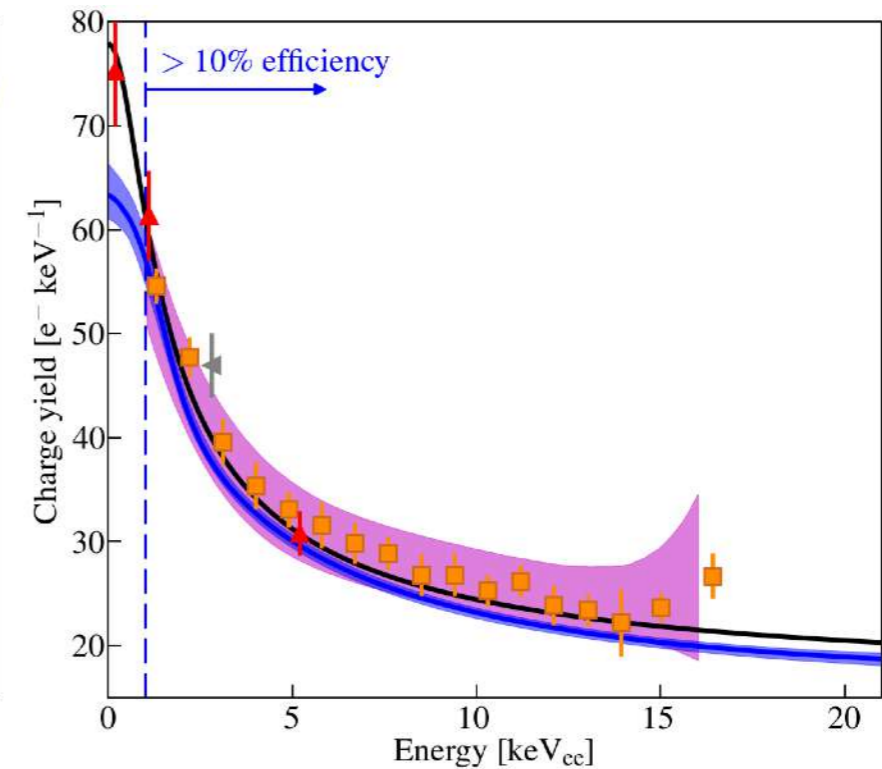
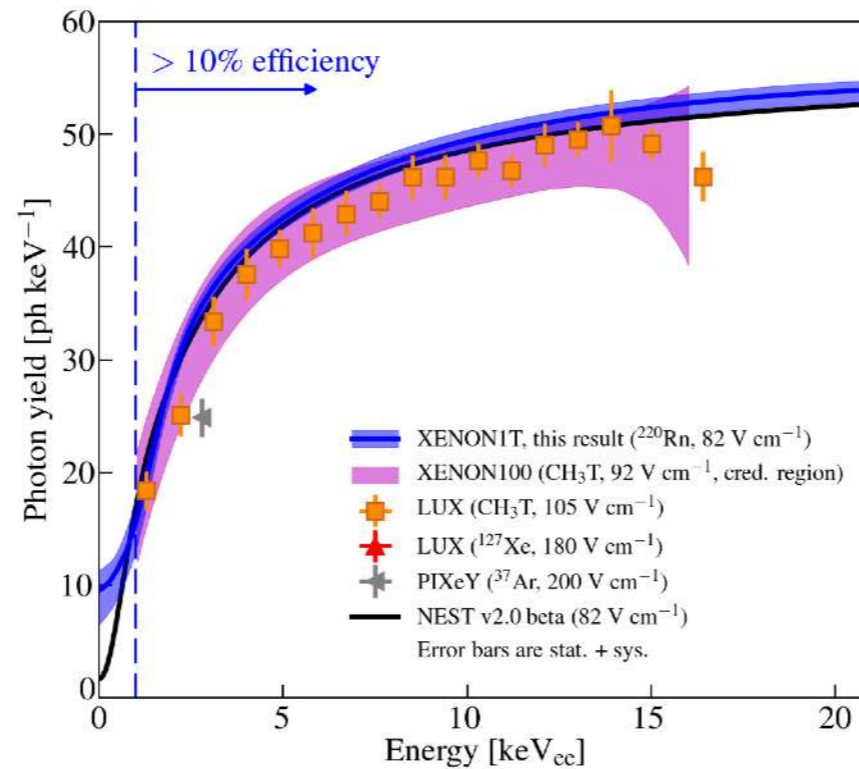
High neutron flux (2000 n/s)
Reduce calibration weeks \rightarrow days
compared to AmBe source

Nuclear recoils

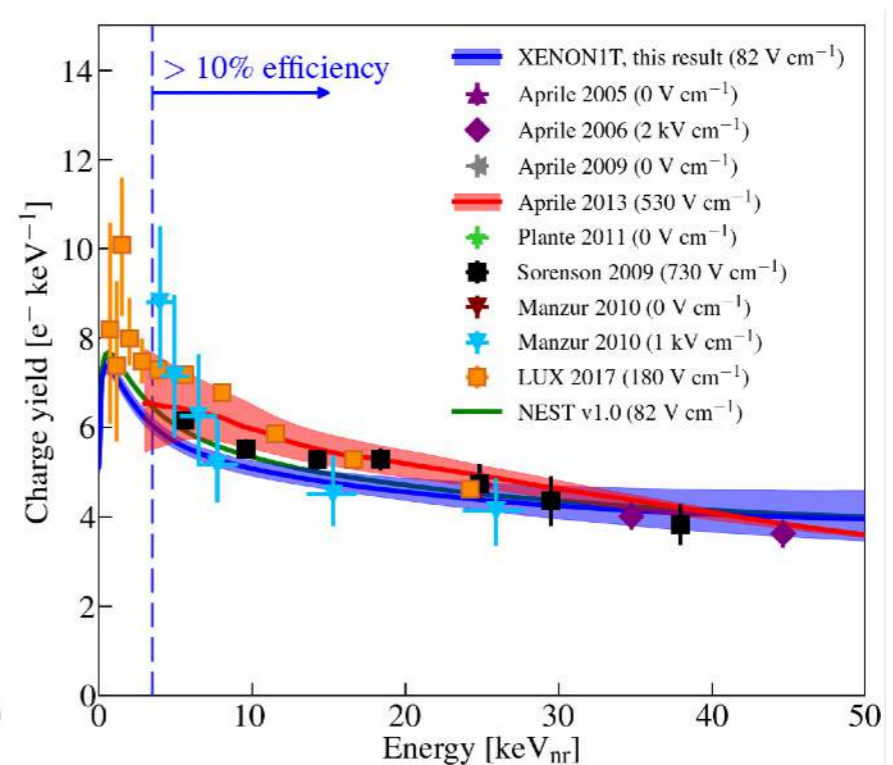
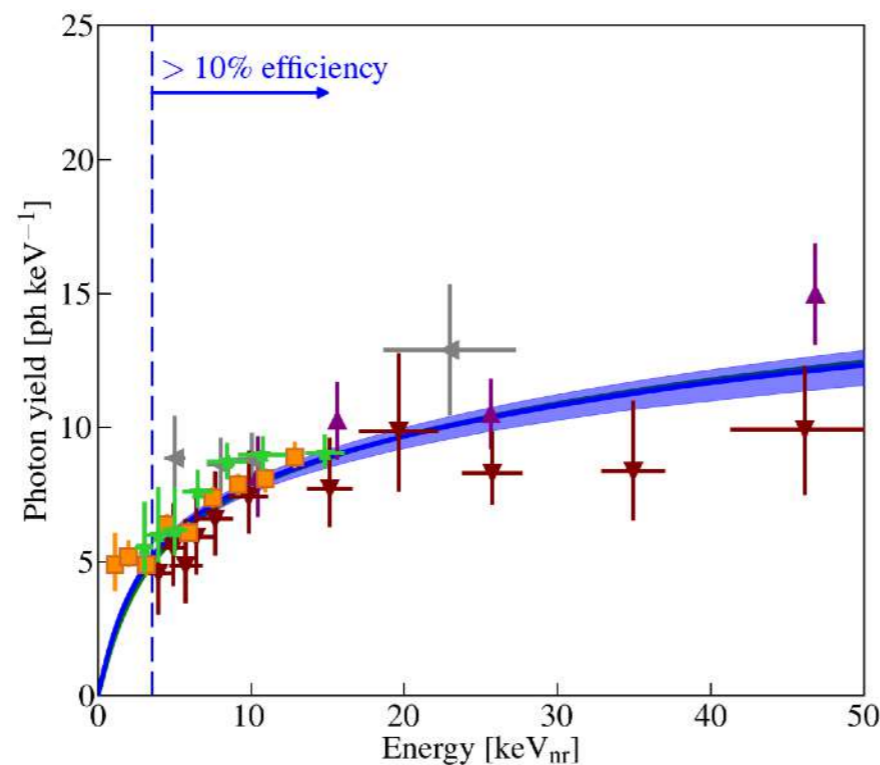
Calibration of ER / NR bands

Models for detector response physically motivated from LXe microphysics
Fit with the calibration data

ER



NR



4.1% Materials

- ▶ HPGe γ screening
- ▶ Material selection
- ▶ Suppressed by fiducialisation

1.4% ^{136}Xe

85.4% ^{222}Rn

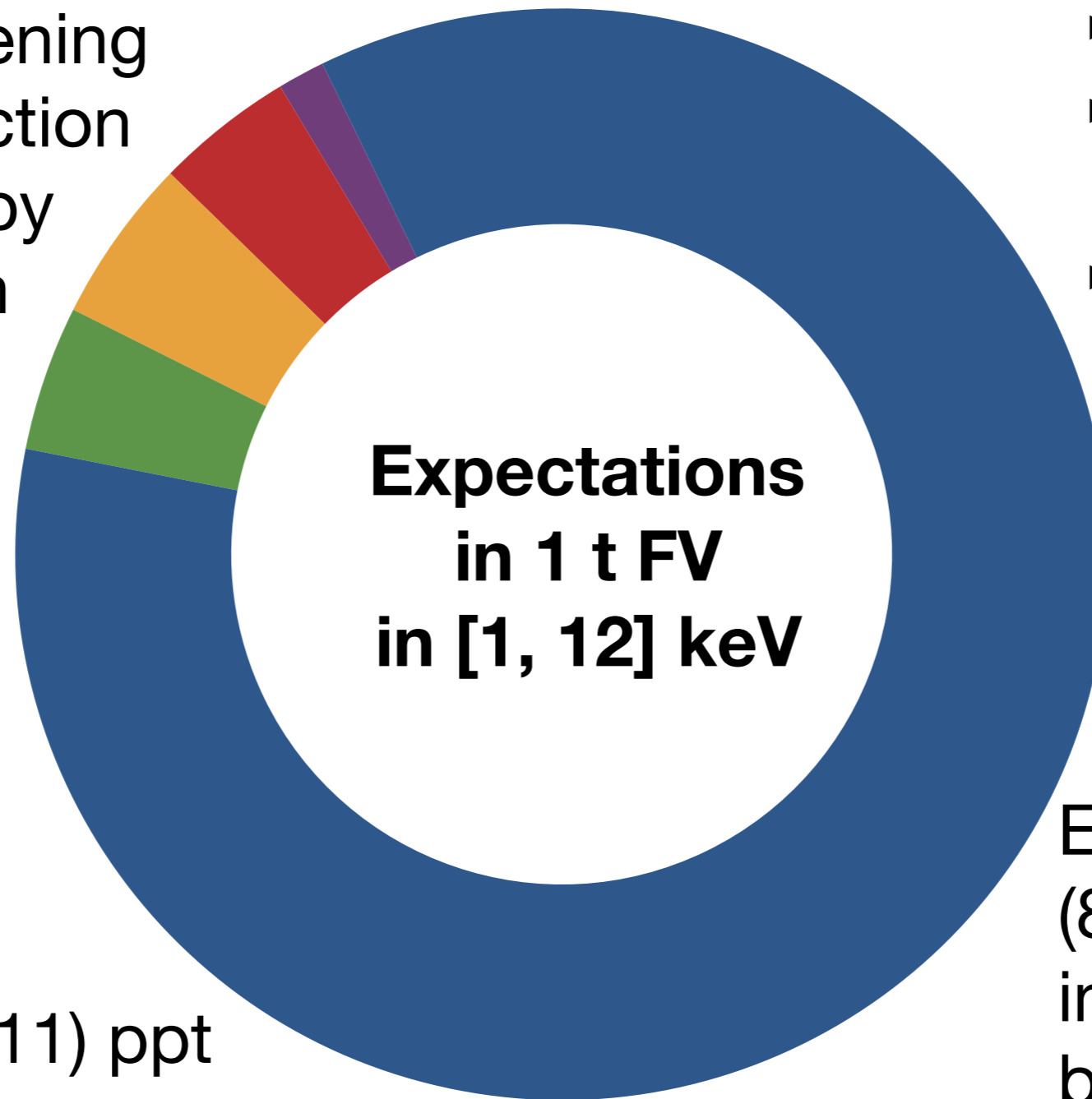
- ▶ $\sim 10 \mu\text{Bq/kg}$
- ▶ Control surface emanation
- ▶ Further reduction by online distillation (more later)

4.9% Solar ν

4.3% ^{85}Kr

- ▶ Cryogenic distillation
- ▶ natKr (0.66 ± 0.11) ppt

EPJ C 77:275 (2017)



ER Rate
(82 ± 5) $\text{ev}/(\text{keV t y})$
in 1.3 t
below $25 \text{ keV}_{\text{ee}}$

Lowest ER background ever achieved in DM detector

< 0.01 eV Cosmogenic n

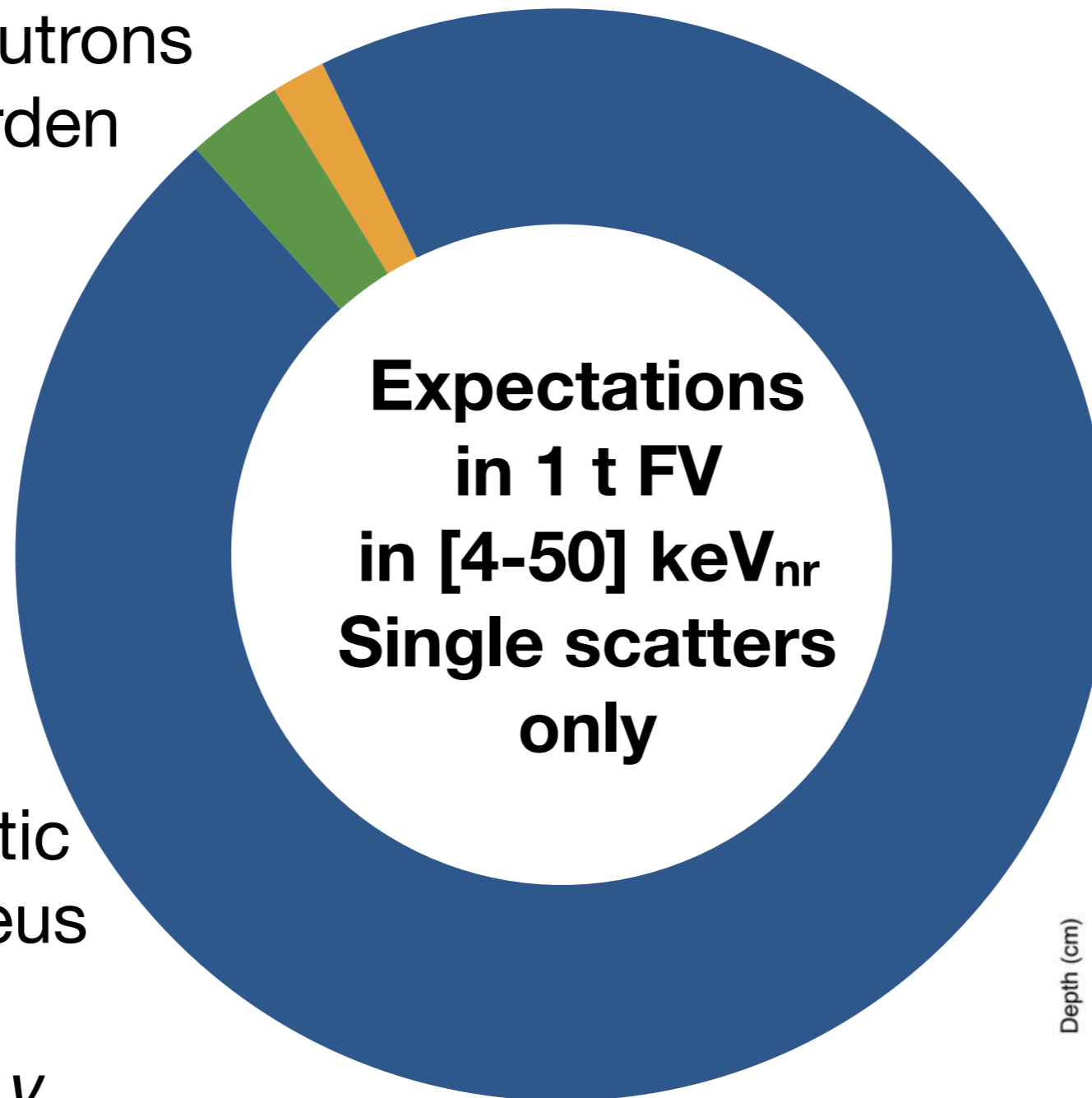
- ▶ μ -induced neutrons
- ▶ Rock overburden
- ▶ Muon veto

JINST 9:P11006 (2014)

0.6 eV Radiogenic n

- ▶ From (α , n) and spontaneous fission
- ▶ Material selection
- ▶ Mostly multiple scatter
- ▶ Fiducialisation

EPJ C 77:890 (2017)

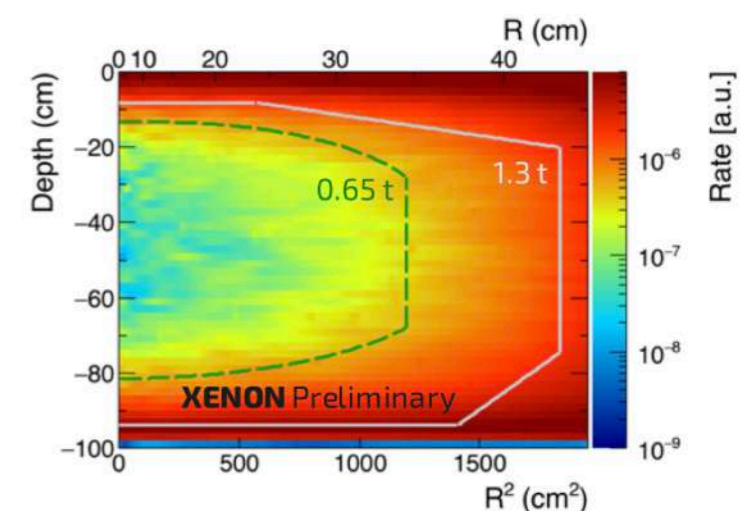


**Expectations
in 1 t FV
in [4-50] keV_{nr}
Single scatters
only**

0.02 eV CEVNS

- ▶ Coherent elastic neutrino-nucleus scattering
- ▶ From ^8B solar ν
- ▶ Irreducible, very low energy (< 1 keV)

JCAP 04:027 (2016)



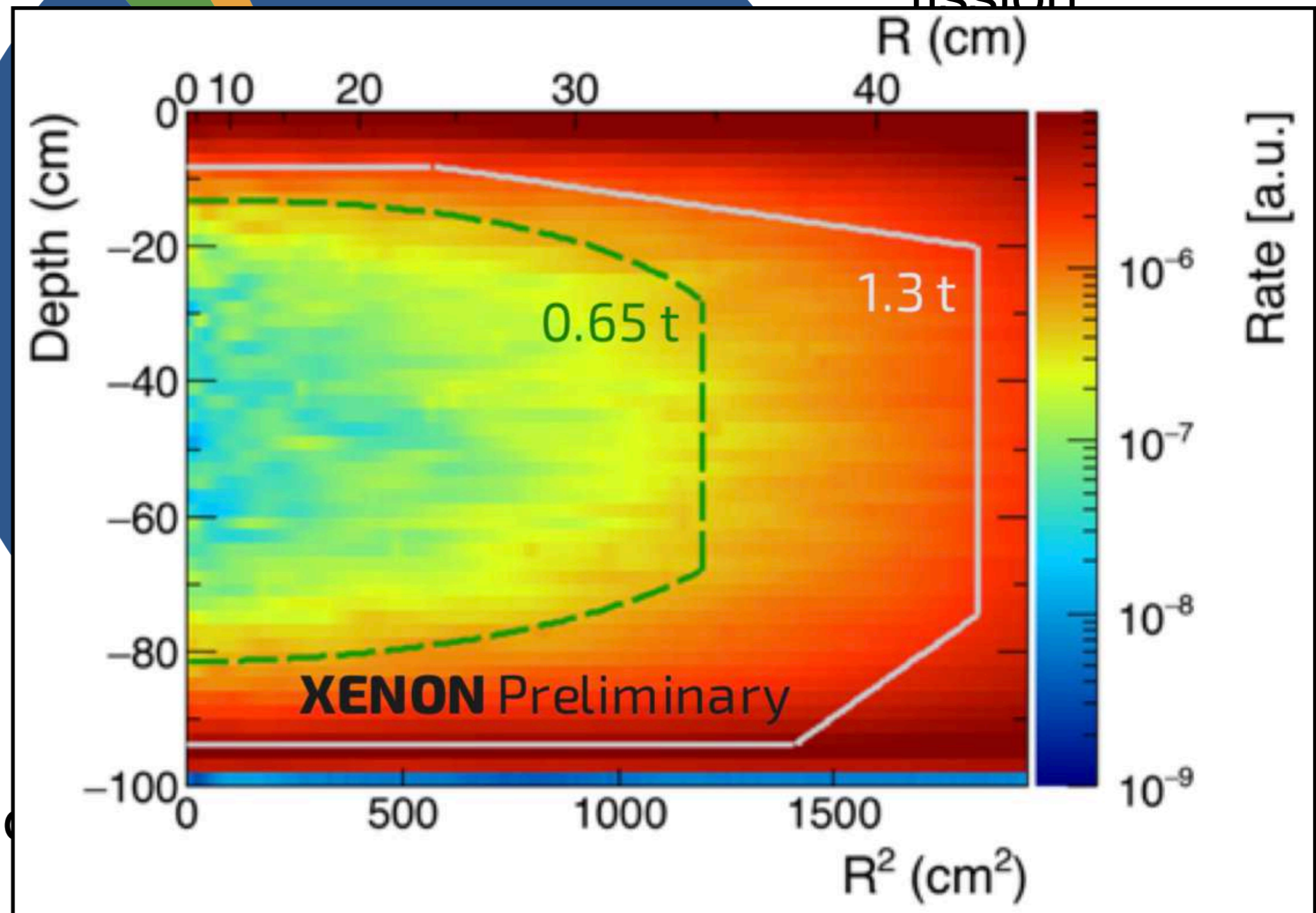
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JINST 9:P11006 (2014)

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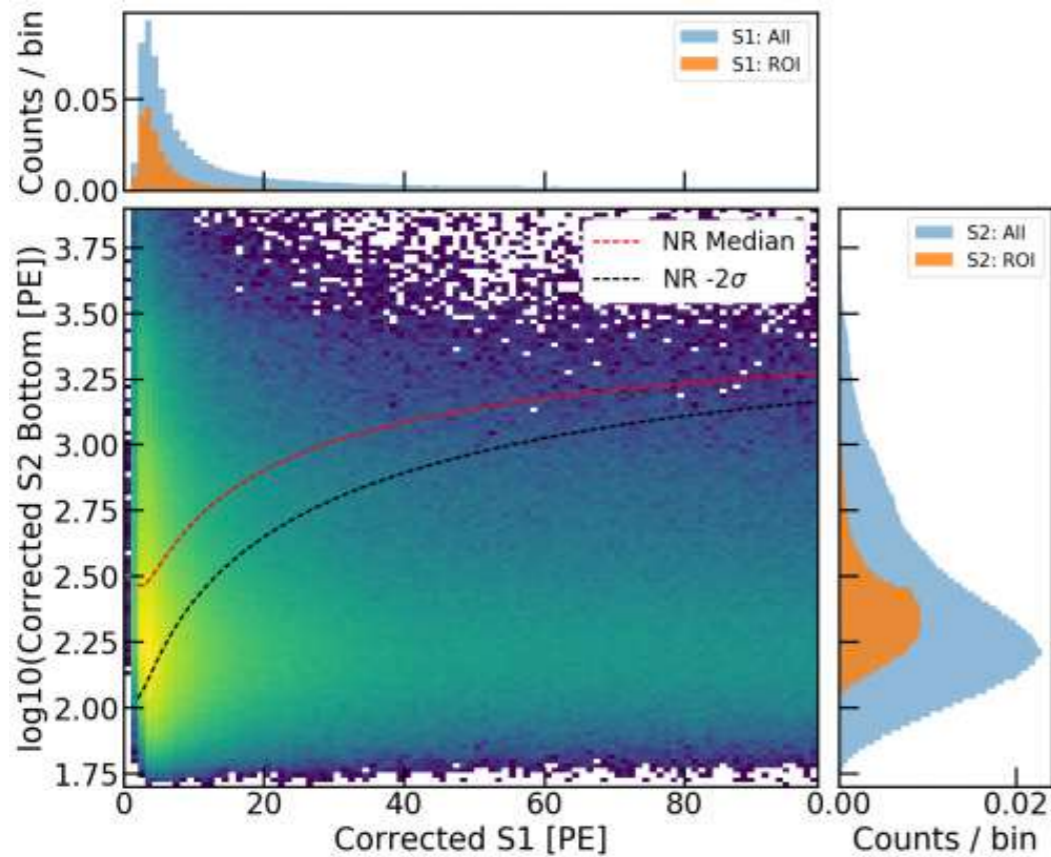


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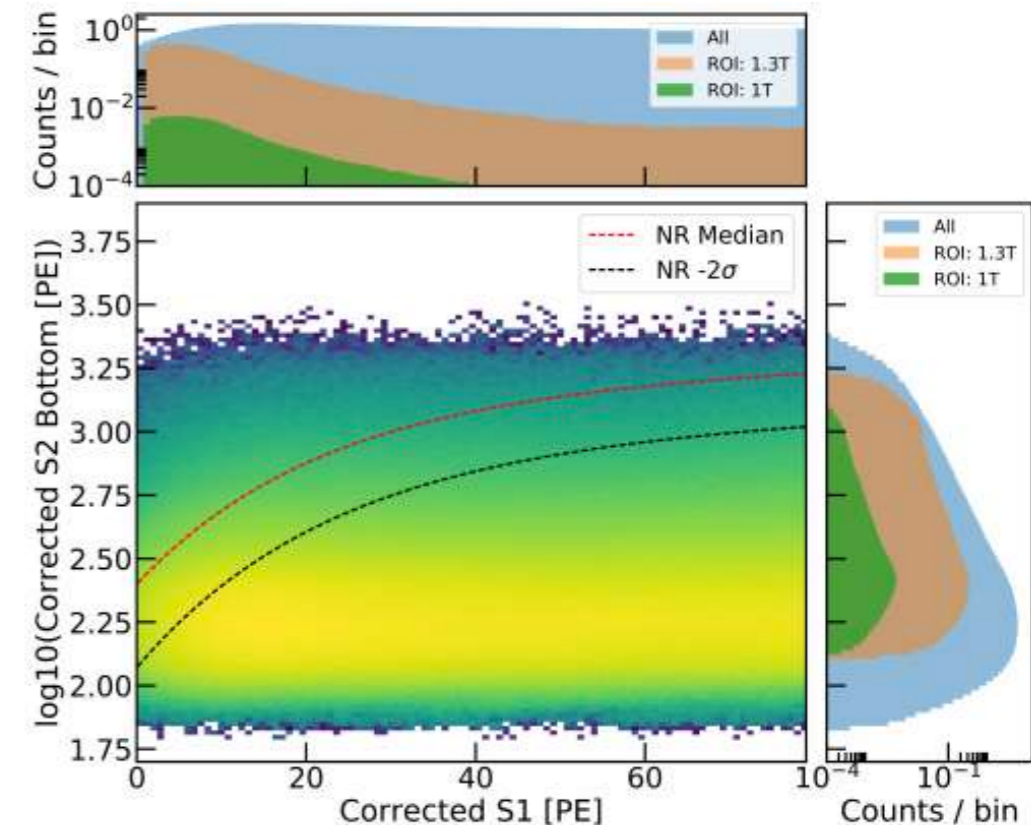
Accidental coincidence

- ▶ Random pairing of S1-like and S2-like lone signals



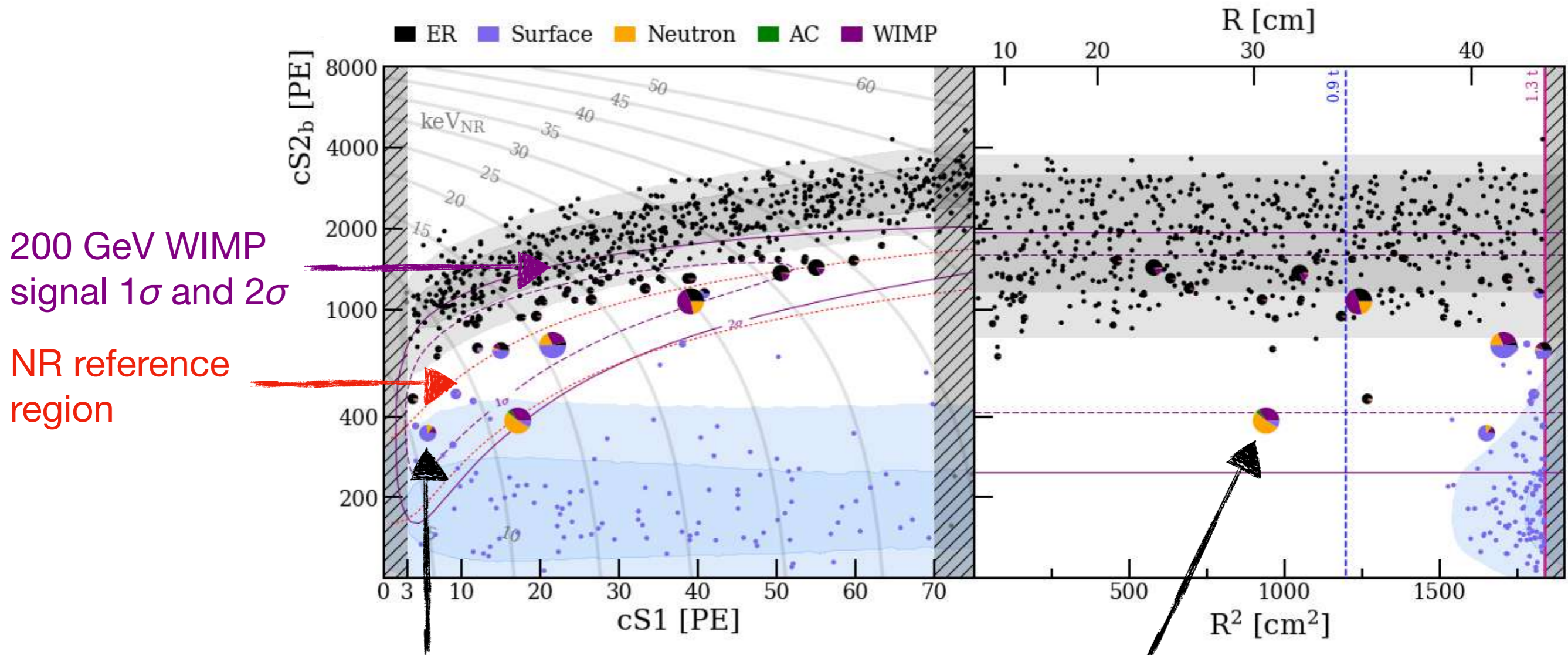
Surface events

- ▶ Events at outer edge of TPC
- ▶ Position reconstruction trickiest here
- ▶ Charge loss on walls



Dark matter search results

- ▶ Results interpreted with **profile likelihood analysis**
- ▶ 4 dimensions:
 - 3 unbinned**: $cS1$, $cS2_{\text{bottom}}$, R
 - 1 binned (binary)**: core 0.65 t volume with low radiogenic n rate



200 GeV WIMP
signal 1σ and 2σ

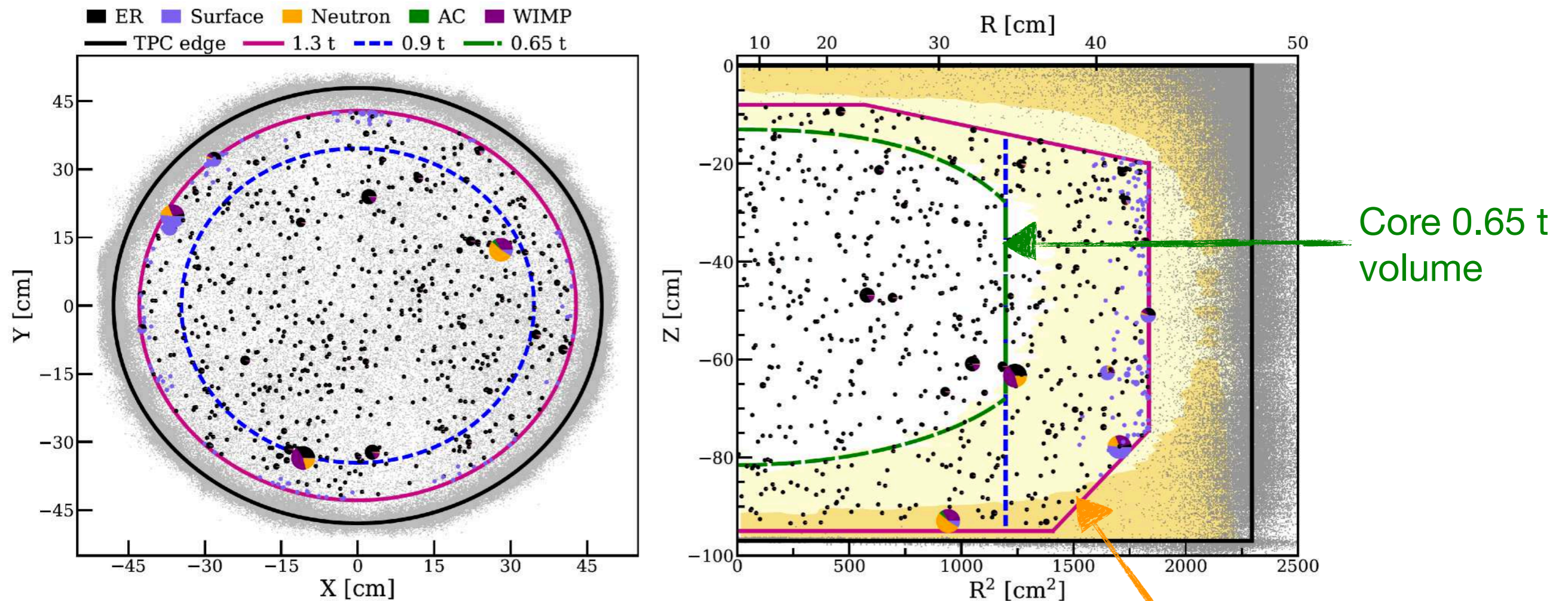
NR reference
region

Pie charts show relative contribution of
each component to that event's PDF

Large pie charts have
more WIMP probability

Dark matter search results

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1σ & 2σ of radiogenic neutron pdf

Sensitivity

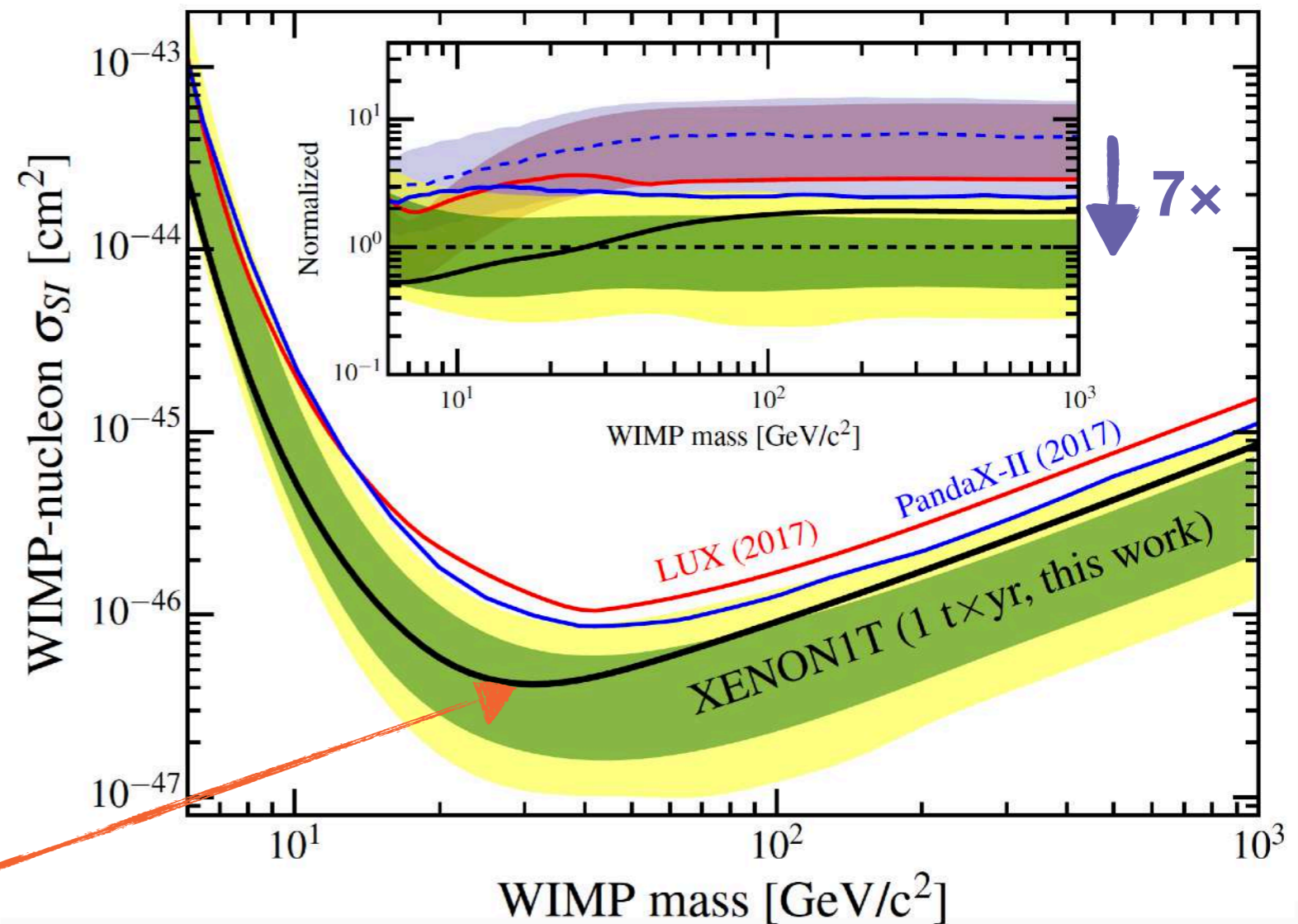
- ▶ 7-times improvement of previous-generation experiments

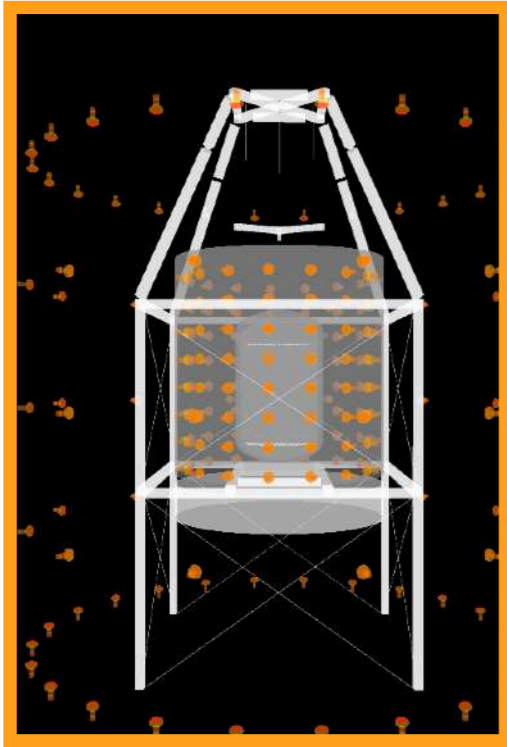
Limit

- ▶ Strongest exclusion limits for WIMPs above 6 GeV
- ▶ Under fluctuation < 8 GeV
Over-fluctuation for higher mass

Minimum

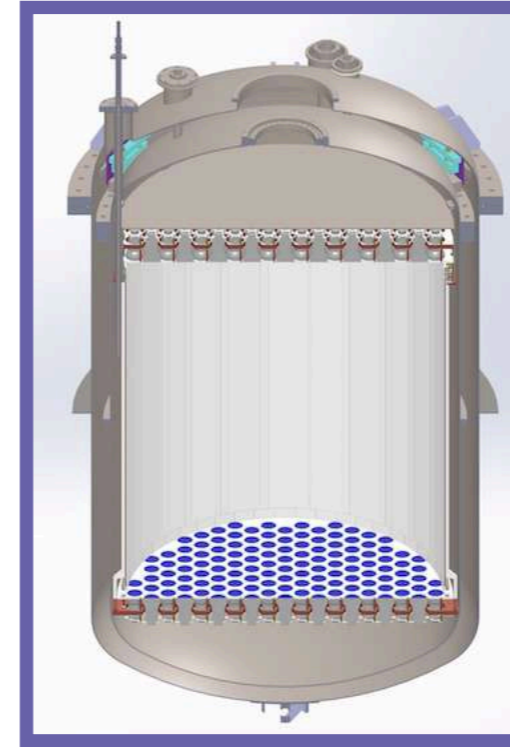
$$\sigma_{SI} < 4.1 \times 10^{-47} \text{ cm}^2 \text{ at } 30 \text{ GeV}$$





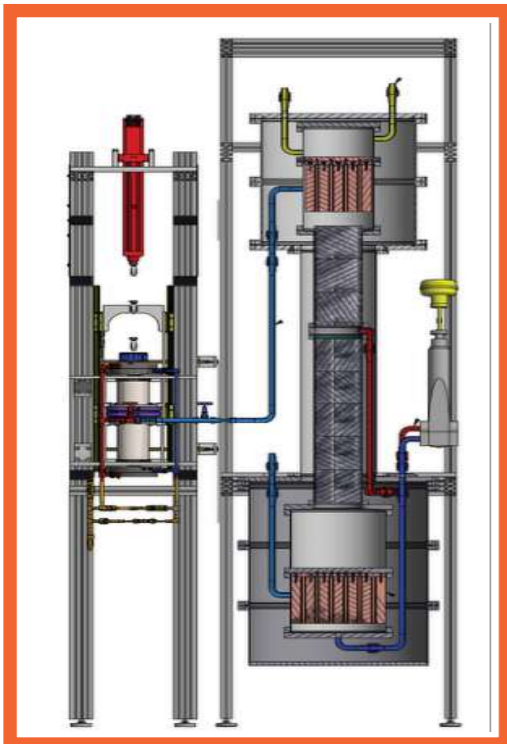
neutron veto

- ▶ Inner region
- ▶ optically separate
- ▶ extra PMTs
- ▶ Gd in the water tank
- ▶ 0.5% $\text{Gd}_2(\text{SO}_4)_3$



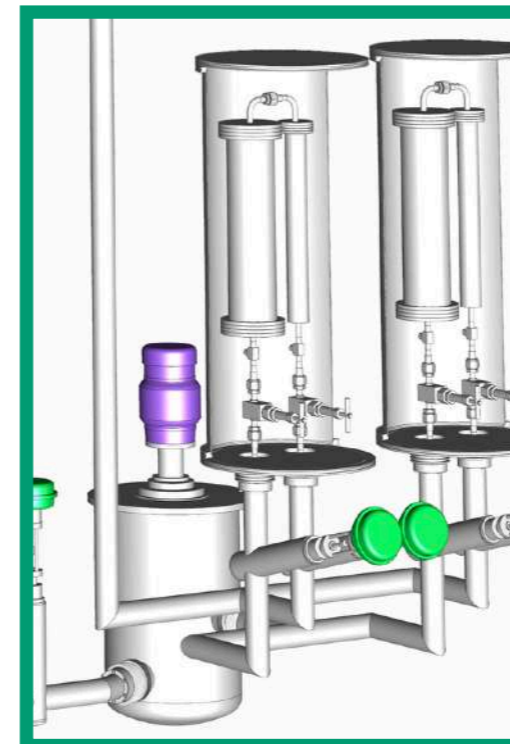
larger TPC

- ▶ Total 8.4 t LXe
- ▶ 5.9 t in TPC
- ▶ ~ 4 t fiducial



^{222}Rn distillation

- ▶ Reduce radon from pipes, cables, cryogenic system



LXe purification

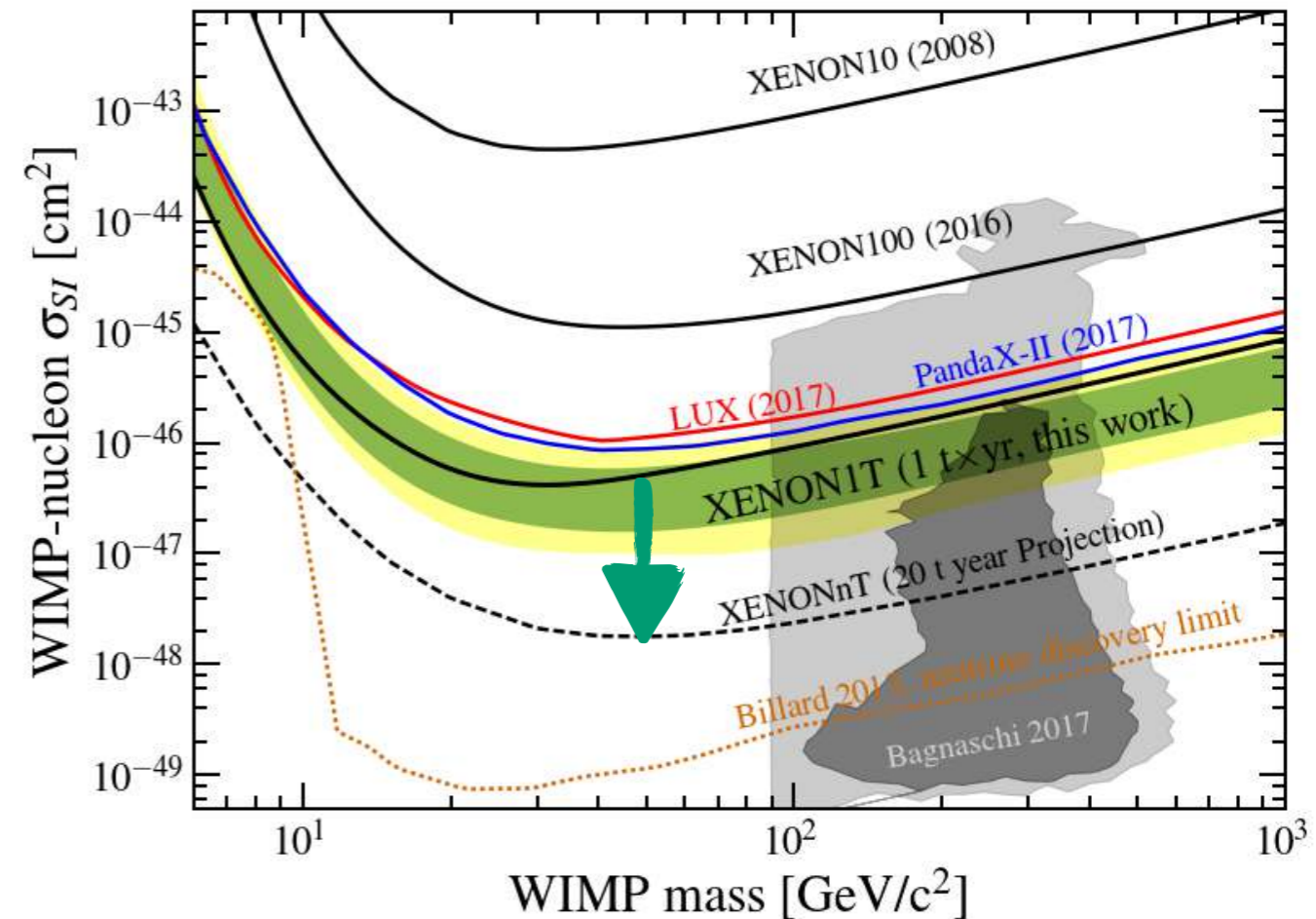
- ▶ Faster xenon cleaning
- ▶ 5L/min LXe (2500 slpm)
- ▶ Now: 120 slpm

XENON1T

- ▶ First multi-ton scale LXe TPC
- ▶ Stable operation > 1 year
- ▶ More results on the way: annual modulation, low-mass WIMP, $0\nu\beta\beta$, solar axions, dark photons

XENONnT

- ▶ 4–5× bigger fiducial mass
- ▶ 10× better sensitivity
- ▶ Preparations and tests underway
- ▶ Construction starting early next year
- ▶ Commissioning 2019



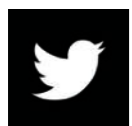
Stay tuned for more exciting results



contact@xenon1t.org



www.xenon1t.org



[@xenon1t](https://twitter.com/xenon1t)

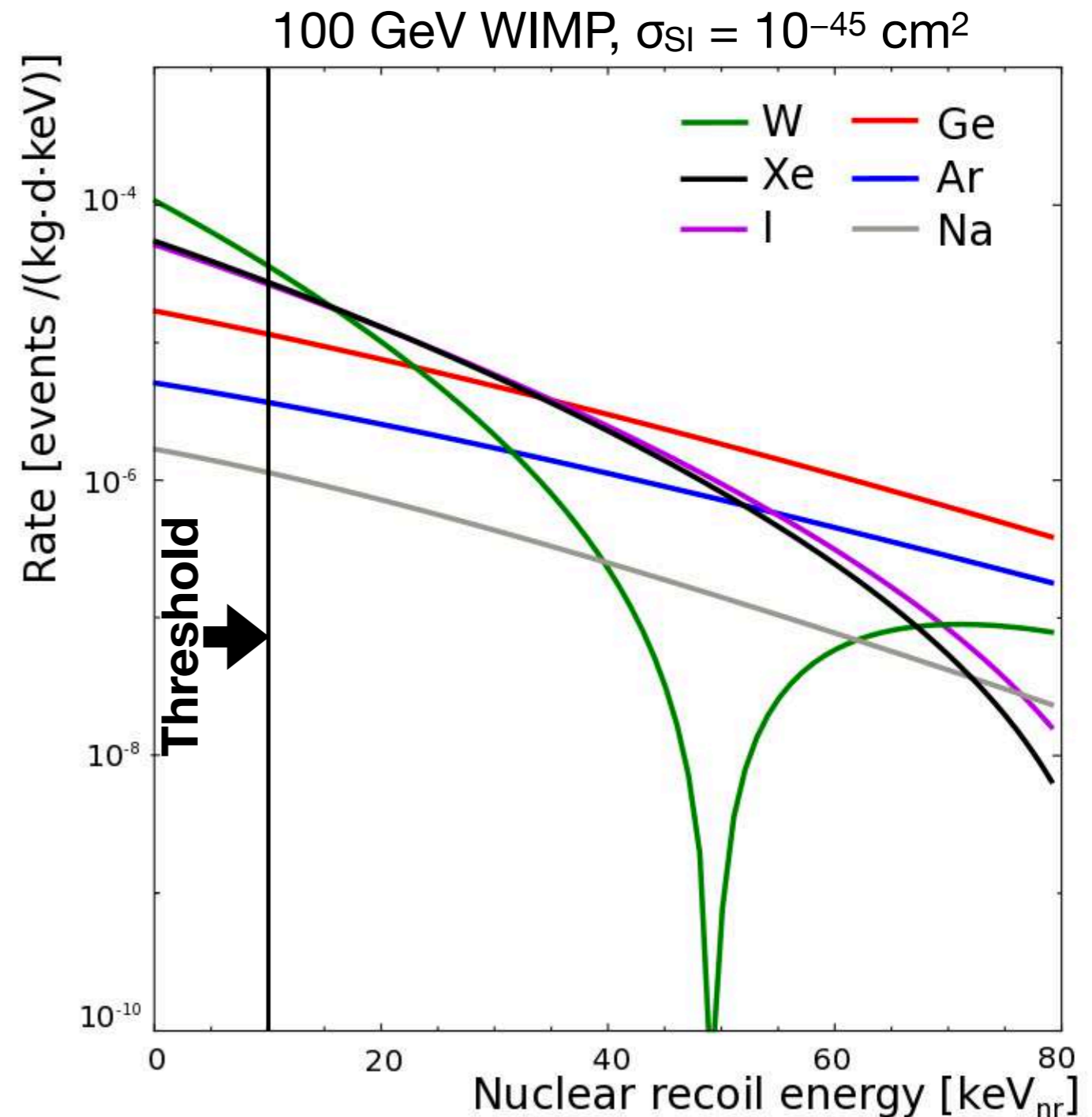


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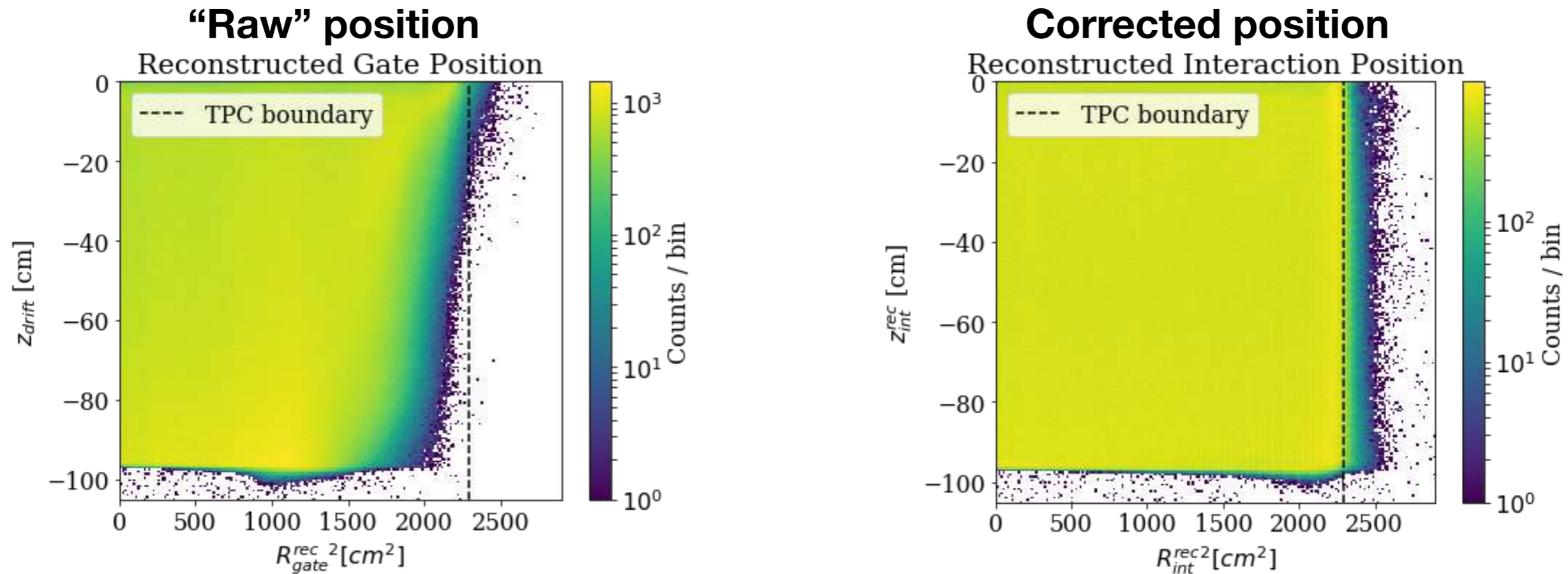
Backup slides

Why xenon as a target?

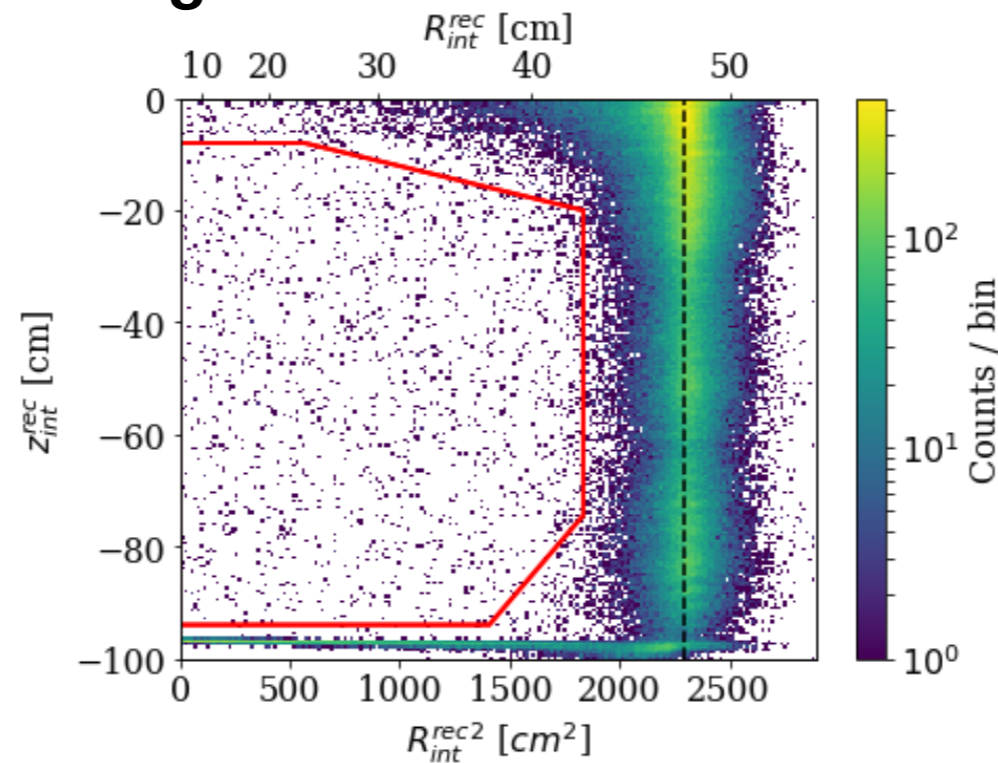
- ▶ High atomic mass
- ▶ Self-shielding
- ▶ ~ 50% odd-nucleon isotopes
- ▶ 178 nm scintillation light detectable directly — no wavelength shifters
- ▶ Radioactively pure: all isotopes either stable, short lifetime (≤ 36 d), or long lifetime (^{136}Xe : 2.2×10^{21} y)
- ▶ High charge & light yield
- ▶ High boiling point (~ -100 °C)

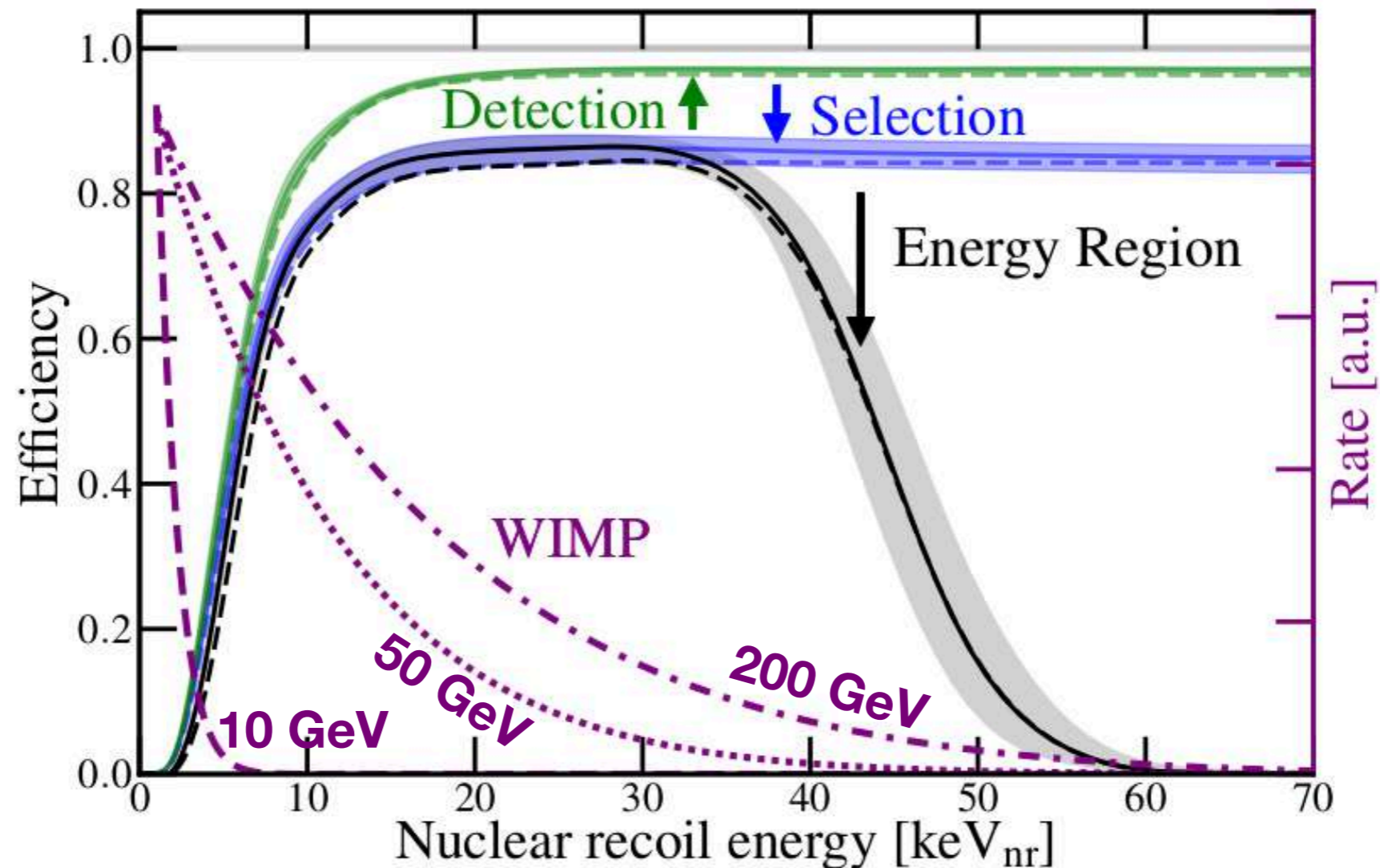


Field non-uniformity correction



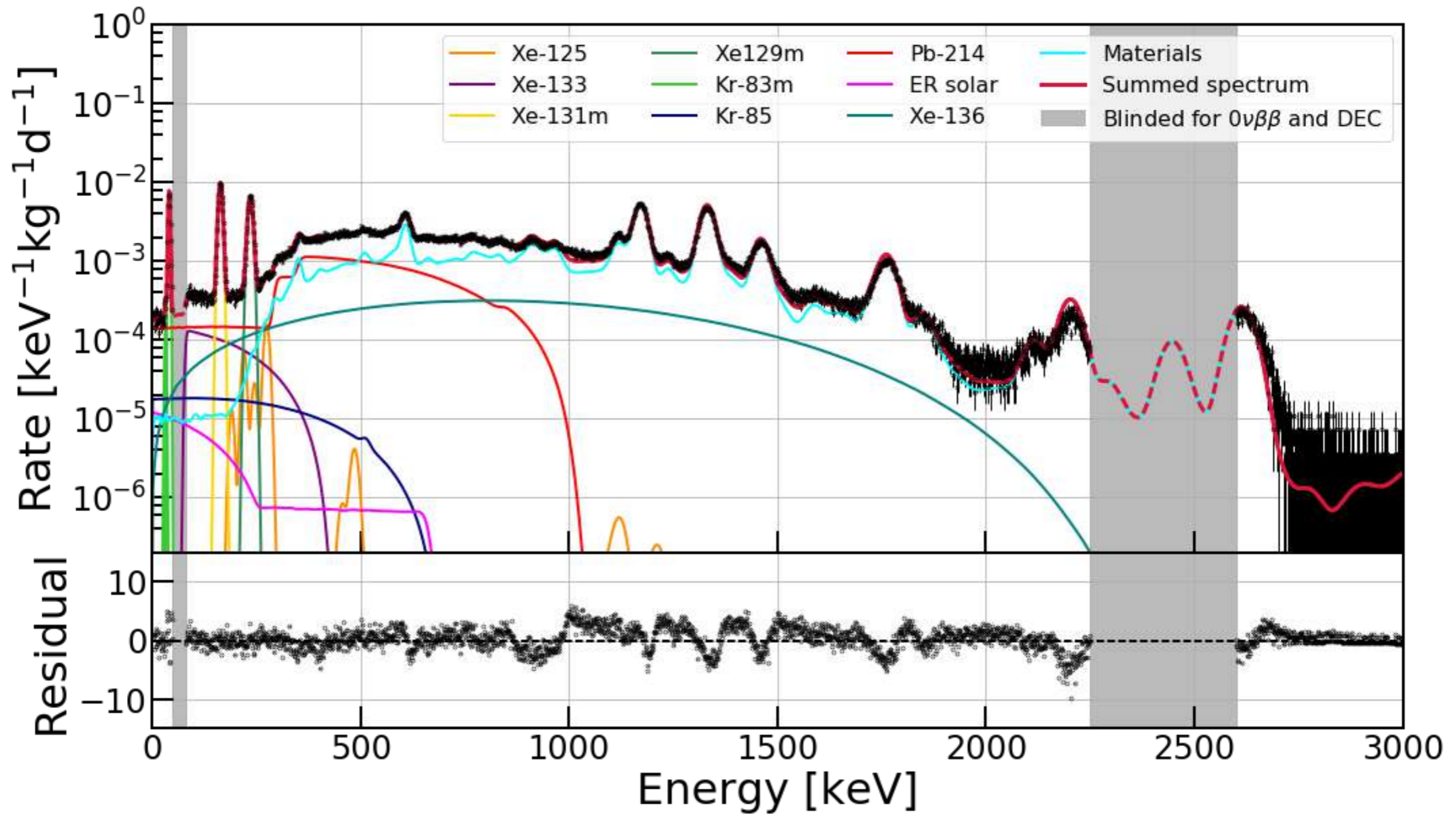
Background data – wall matches





- ▶ Dominant **detection efficiency** loss:
3 PMTs must contribute to S1 (so $S1 > 3$ pe)
- ▶ **Selection efficiencies** estimated from calibration or MC data
- ▶ Search region defined as 3–70 pe in cS1

In 1 t volume



XENON1T Cryogenic systems

LXe temperature stable at $-96.07\text{ }^{\circ}\text{C}$, RMS $0.04\text{ }^{\circ}\text{C}$
GXe pressure stable at 1.934 bar , RMS 0.001 bar

