Testing a Modified Gravity Theory in the Milky Way (arXiv:1810.07200 PRD in press)

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- Brief Review of Modified Gravity Theories
- Moffat's Modified Gravity Theory and the rotation curve of the Milky Way
- Data Sets and Morphologies
- Results
- Conclusions

# Review of Modified Gravity Theories

 MOdified Newtonian Dynamics (MOND, M.Milgrom 1983) The acceleration of a point particle:

$$\mathbf{a} = rac{MG}{R^2} 
u \left( \mathbf{q}, rac{MG}{R^2 \mathbf{a}_0} 
ight)$$

q is an adimensional parameter depending on the orbit's shape and es  $a_0$  is a free parameter. For  $a >> a_0$  the theory behaves as Newtonian Mechanics, while for  $a \le a_0$  the theory becomes scale invariant and  $a \simeq = \eta(q) \frac{(MGa_0)^{1/2}}{R}$ 

• Tensor-Vector-Scalar theory of gravity (TeVeS, J.D.Bekenstein 2004)

$$S = S_{\hat{g}} + S_A + S_{\phi} + S_m$$

Gravity is described by a metric  $g_{\mu\nu}$  as in RG plus a vector field  $A_{\mu}$  and a scalar field  $\phi$ .

# Problems for MOND and TEVES

- MOND
  - Different values of a<sub>0</sub> are needed to explain rotation curves of galaxies (Randriamampandry & Carignan MNRAS 439, 2132 (2014)).
  - The theory can not explain the gravitational lensing effect (Clowe et al ApJL, 648, L109 (2006)).
  - The theory is not able to explain the observed matter power spectrum (Dodelson IJMP D 20, 2749 (2011)).
- TeVeS
  - The theory is not able to explain the observed matter power spectrum (Dodelson IJMP D 20, 2749 (2011)).
  - It is not possible to reconcile gas profile and strong-lensing measurements in well known cluster systems (Nieuwenhuizen et al MNRAS 476, 3393 (2018)).
- Open Discussion
  - According to Clowe et al 2006, TeVeS and MOG have difficulties to explain the bullet cluster, while Brownstein & Moffat 2007 claim the opposite.

MOG action The MOG weak field approximation lpha and  $\mu$  estimates

# Moffat's MOG

The MOdified Gravity theory (also named as Scalar-Tensor-Vector Gravity) is a covariant modification of General Relativity.

- MOG was proposed by J. Moffat in 2006.
- Two scalar fields and one vector field are added to RG.
- It has been used to describe observations of the Solar System (Moffat IJMP D16, 2075(2008)) and rotation curves of spiral galaxies (Moffat & Rahvar MNRAS 436, 1439 (2013)), without the need of dark matter.
- There are claims that MOG can fit both Bullet and the Train Wreck merging clusters (Brownstein & Moffat, MNRAS 382, 29 (2007); Israel & Moffat Galaxies 6, 41 (2018)).
- There are also articles that apply the MOG to gravitational waves, black holes, binary pulsars, lensing, globular clusters, motion of satellite galaxies, gravitational stability of galactic disks, N-body simulations, galactic sun's motion. cosmological data

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# MOG action

$$S_{\mathrm{MOG}} = S_{\mathrm{G}} + S_{\phi} + S_{\mathrm{S}} + S_{\mathrm{M}}.$$

$$egin{aligned} S_{\mathrm{G}}&=-rac{1}{16\pi}\intrac{1}{G}\left(R+2\Lambda
ight)\sqrt{-g}d^{4}x,\ S_{\phi}&=-rac{1}{4\pi}\int\omega\left[rac{1}{4}B^{\mu
u}B_{\mu
u}-rac{1}{2}\mu^{2}\phi_{\mu}\phi^{\mu}+V_{\phi}(\phi_{\mu}\phi^{\mu})
ight]\sqrt{-g}d^{4}x, \end{aligned}$$

$$S_{\rm S} = -\int \frac{1}{G} \left[ \frac{1}{2} g^{\alpha\beta} \left( \frac{\nabla_{\alpha} G \nabla_{\beta} G}{G^2} + \frac{\nabla_{\alpha} \mu \nabla_{\beta} \mu}{\mu^2} \right) + \frac{V_{\rm G}(G)}{G^2} + \frac{V_{\mu}(\mu)}{\mu^2} \right] \sqrt{-g} d^4 x.$$

where  $B_{\mu\nu} = \partial_{\mu}\phi_{\nu} - \partial_{\nu}\phi_{\mu}$ ,  $\omega$  is an adimentional coupling constants, G represents the gravitational coupling strength, and  $\mu$  is the mass of the vector field  $\phi$ ,  $V_i$  are the self interaction potentials associated with each of the fields. For simplicity:  $V_{\phi}(\phi_{\mu}\phi^{\mu}) = V(G) = V(\mu) = 0$ .

MOG action The MOG weak field approximation  $\alpha$  and  $\mu$  estimates

### The MOG weak field approximation

The equations for  $\phi_{\mu}$ , G, $\mu$  and the metric are solved considering perturbations around a Minkovsky space-time

 $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}.$   $\phi_{\mu} = \phi_{\mu(0)} + \phi_{\mu(1)},$   $G = G_{(0)} + G_{(1)}$  $\mu = \mu_{(0)} + \mu_{(1)}$ 

In a Minkovsky space-time  $\phi_{\mu(0)} = 0$  and  $G_{(0)} = \text{constant}$ . Besides,  $\mu = \text{constant}$  is fixed. The energy-momentum tensor is expressed as:

$$T_{\mu\nu} = T_{\mu\nu(0)} + T_{\mu\nu(1)}$$

MOG action The MOG weak field approximation  $\alpha$  and  $\mu$  estimates

### The MOG weak field approximation

After a lot of algebra the following equation is obtained:

$$\nabla \underbrace{\left(\nabla \Phi_{\rm eff} - \kappa \omega \nabla \phi^0\right)}_{\nabla \Phi_{\rm N}} = 4\pi G_0 \rho.$$

$$\Phi_{
m eff}(ec{x}) = -G_{
m N} \int rac{
ho(ec{x}')}{|ec{x}-ec{x}'|} (1+lpha-lpha e^{-\mu|ec{x}-ec{x}'|}) d^3x'$$

where  $G = G_N(1 + \alpha)$ .

- For scales lower than  $\mu^{-1}$ , the repulsive force cancels a part of the atractive force and newtonian gravity is recovered .
- For scales larger than μ<sup>-1</sup>, the repulsive force becomes weaker and a newtonian force with a larger gravitational constant is obtained.

MOG action The MOG weak field approximation lpha and  $\mu$  estimates

Estimates can be obtained of  $\alpha$  and  $\mu$  from the solutions for spherical symmetry (Moffat & Toth, Class. Quant.Grav. 26, 085002 (2009)).



In all versions of the theory  $\alpha$  and  $\mu$  are taken as constants.

Baryonic Morphologies Data sets

### Galaxy Tour

The Milky Way is a complex system formed by stars, gas and dark matter gravitationally bound together

The galaxy has three main baryonic components: disk, bulge and gas



Baryonic Morphologies Data sets

# Observationally inferred morphologies

The gravitational potential of our galaxy receives contributions from baryons, and presumably from dark matter, separately

 $\phi_{\rm tot} = \phi_{\rm bulge} + \phi_{\rm disk} + \phi_{\rm gas} + \phi_?.$ 

	model	specification	data
	1	exponential E2	optical
	2	gaussian G2	optical
	3	gaussian plus nucleus	infrared
bulge	4	truncated power law	infrared
	5	power law plus long bar	optical   infrared
	6	truncated power law	optical   infrared
	1 thin plus thick		optical
	2	thin plus thick	optical
disc	3	thin plus thick plus halo	optical
	4	thin plus thick plus halo	optical
	5	single maximal disc	optical
gas	1	H <sub>2</sub> , HI, HII	optical   microwave   radio

ArXiv:1511.05571

Baryonic Morphologies Data sets

### Data Sets

- The Galkin compilation comprises the velocity measurements of 2701 objects at R > 2.5 kpc
  - 2095  $\rightarrow$  Gas (HI, HII, C0, giant molecular clouds)
  - 506  $\rightarrow$  Stars (open clusters, planetary nebulae, cepheids, carbon stars)
  - 100  $\rightarrow$  masers (molecular clouds, comets, planetary and stellar atmospheres)
- The Huang compilation comprises 43 data obtained from a binning of :
  - 16000  $\rightarrow$  red clump giants selected from LSS-GAC y SDSSIII/APOGEE surveys
  - $\bullet~5700 \rightarrow$  Halo K stars selected from the SDSS/SEGUE survey

Baryonic Morphologies Data sets

### Data Sets

### Huang compilation:

- 43 data
- r=[4.59,98.97] kpc

### Galkin compilation:

- 2701 objects
- r=[2.5,24.81] kpc



Baryonic Morphologies Data sets

#### • Photometric data are used to trace each barionic component

This technique has been used to show that an extra component is needed to explain the observed rotation curve of the MW (locco et al., Nature Physics, 2015).





### Parameters $\alpha$ and $\mu$

$$\Phi_{\rm MOG}(\vec{x}) = -G_{\rm N} \int \frac{\rho_b(\vec{x}') + \rho_d(\vec{x}') + \rho_g(\vec{x}')}{|\vec{x} - \vec{x}'|} (1 + \alpha - \alpha e^{-\mu|\vec{x} - \vec{x}'|}) d^3x'$$

- $(\alpha, \mu)^{SG} = (8.89, 4.2 \times 10^{-2})$ , best fit obtained by Moffat & Rahvar MNRAS 436, 1439 (2013) to fit spiral galaxies;
- $(\alpha, \mu)^{MW} = (15.01, 3.13 \times 10^{-2})$ , obtained by Moffat considering  $M_{Moffat}^{MW} = 4 \times 10^{10} M_{\odot}$ ;
- (α, μ)<sup>C</sup>, considering the mass of the Milky Way we obtain for each one of our observationally inferred morphologies.

$$\alpha = \frac{M}{(\sqrt{M} + E)^2} \left(\frac{G_{\infty}}{G_N} - 1\right) \qquad \qquad \mu = \frac{D}{\sqrt{M}}$$

 $\chi^2$  Test Best fit parameters

Conclusions

baryonic	Newton	MW	SG	С	$(\alpha, \mu)^{C}$	$M_{C}^{MW}$ [10 <sup>10</sup> M <sub>☉</sub> ]
morphology	$\tilde{\chi}^2$	$\tilde{\chi}^2$	$\tilde{\chi}^2$	$\tilde{\chi}^2$		
[disk] [bulge]	Huang - galkin	Huang - galkin	Huang - galkin	Huang - galkin		
$1 \ [44][40] \ G2$	31.83 - 10.69	4.50 - 4.25	4.68 - 4.25	8.59 - 5.96	$(15.79, 2.43 \times 10^{-2})$	$6.6^{+0.6}_{-0.4}$
$2 \ [44][40] \ E2$	30.80 - 9.89	4.11 - 3.83	4.25 - 3.83	8.00 - 5.39	$(15.80, 2.41\times 10^{-2})$	$6.7^{+0.7}_{-0.6}$
3 [44][45]	32.90 - 8.51	3.36 - 3.10	3.43 - 3.10	6.85 - 4.37	$(15.83, 2.39 \times 10^{-2})$	$6.8^{+0.7}_{-0.6}$
4 [44][46]	29.85 - 9.45	3.71 - 3.51	3.79 - 3.51	7.47 - 5.03	$(15.83, 2.39\times 10^{-2})$	$6.8^{+0.7}_{-0.6}$
5 [44][47]	35.73 - 11.40	4.93 - 4.66	5.16 - 4.66	9.21 - 6.51	$(15.77, 2.44 \times 10^{-2})$	$6.6 \pm 0.6$
6 [44][48]	28.67 - 13.65	6.17 - 6.00	6.48 - 6.00	13.00 - 8.43	$(15.74,2.47\times10^{-2})$	$6.4^{+0.6}_{-0.5}$
7 [39][40] G2	33.84 - 12.69	5.51 - 5.45	5.74 - 5.44	9.86 - 7.37	$(15.79, 2.42\times 10^{-2})$	$6.6\substack{+0.6\\-0.4}$
8 [39][40] <i>E2</i>	32.65 - 11.72	5.02 - 4.90	5.20 - 4.90	9.14 - 6.65	$(15.80, 2.41\times 10^{-2})$	$6.7^{\pm 0.7}_{-0.6}$
9 [39][45]	30.19 - 10.04	4.06 - 3.93	4.17 - 3.93	7.72 - 5.23	$(15.84, 2.38\times 10^{-2})$	$6.9^{+0.7}_{-0.6}$
10 [39][46]	31.62 - 11.22	4.54 - 4.50	4.66 - 4.50	8.53 - 6.22	$(15.83, 2.39\times 10^{-2})$	$6.9^{+0.7}_{-0.6}$
11 [39][47]	35.10 - 13.56	6.06 - 5.98	6.33 - 5.97	10.64 - 8.10	$(15.77, 2.44\times 10^{-2})$	$6.6\pm0.6$
12 [39][48]	38.46 - 16.32	7.66 - 7.74	8.03 - 7.74	15.79 - 10.60	$(15.73,2.47\times10^{-2})$	$6.4^{+0.6}_{-0.5}$
$13 \ [49][40] \ G2$	33.70 - 12.39	5.43 - 5.29	5.66 - 5.28	9.80 - 7.17	$(15.79, 2.42\times 10^{-2})$	$6.7^{+0.6}_{-0.4}$
$14\ [49][40]\ E2$	32.54 - 11.45	4.94 - 4.76	5.15 - 4.76	9.09 - 6.47	$(15.81,2.41\times10^{-2})$	$6.7^{+0.7}_{-0.6}$
15 [49][45]	30.14 - 9.82	4.02 - 3.83	4.14 - 3.83	7.71 - 5.11	$(15.84, 2.38\times 10^{-2})$	$6.9^{+0.7}_{-0.6}$
16 [49][46]	31.50 - 10.95	4.46 - 4.37	4.60 - 4.37	8.49 - 6.06	$(15.84, 2.38 \times 10^{-2})$	$6.9^{+0.7}_{-0.6}$

 $\tilde{\chi}^2_{5\sigma}=2.41$  for Huang,  $\tilde{\chi}^2_{5\sigma}=1.14$  for galkin.

 $\chi^2$  Test Best fit parameters

17 [49][47]	34.93 - 13.23	5.96 - 5.80	6.24 - 5.79	10.56 - 7.86	$(15.78, 2.44 \times 10^{-2})$	$6.6\pm0.6$
18 [49][48]	38.18 - 15.89	7.5 - 7.48	7.87 - 7.47	15.49 - 10.27	$(15.74, 2.47 \times 10^{-2})$	$6.4^{+0.6}_{-0.5}$
$19 \ [50][40] \ G2$	32.81 - 11.45	5.22 - 4.91	5.18 - 4.90	8.46 - 5.96	$(15.91,2.32\times 10^{-2})$	$7.2^{+0.6}_{-0.5}$
20 [50][40] E2	31.79 - 10.66	4.79 - 4.48	4.76 - 4.47	7.86 - 5.35	$(15.92, 2.31 \times 10^{-2})$	$7.3^{+0.7}_{-0.6}$
21 [50][45]	33.86 - 9.26	3.99 - 3.71	3.99 - 3.70	6.69 - 4.35	$(15.95,2.29\times 10^{-2})$	$7.5\substack{+0.8 \\ -0.7}$
22 [50][46]	30.64 - 10.19	4.21 - 4.07	4.20 - 4.06	7.30 - 5.01	$(15.95, 2.29 \times 10^{-2})$	$7.5^{+0.7}_{-0.6}$
23 [50][47]	36.51 - 12.17	5.68 - 5.33	5.63 - 5.32	9.11 - 6.45	$(15.89, 2.34 \times 10^{-2})$	$7.2^{+0.7}_{-0.6}$
24 [50][48]	29.76 - 14.42	6.91 - 6.67	6.83 - 6.66	12.91 - 8.34	$(15.85, 2.37 \times 10^{-2})$	$7.0\pm0.6$
25 [37][40] G2	24.48 - 4.87	1.94 - 1.50	1.79 - 1.51	4.50 - 2.07	$(15.94, 2.30 \times 10^{-2})$	$7.4^{+0.7}_{-0.6}$
26 [37][40] E2	24.02 - 4.64	1.84 - 1.42	1.68 - 1.43	4.30 - 1.97	$(15.94, 2.29 \times 10^{-2})$	$7.4^{+0.8}_{-0.7}$
27 [37][45]	23.23 - 4.15	1.70 - 1.29	1.53 - 1.29	3.97 - 1.72	$(15.95, 2.29 \times 10^{-2})$	$7.5\substack{+0.8\\-0.7}$
28 [37][46]	22.9 - 4.47	1.58 - 1.26	1.32 - 1.27	3.82 - 1.84	$(15.98, 2.26 \times 10^{-2})$	$7.7\substack{+0.8\\-0.7}$
29 [37][47]	24.93 - 3.89	2.03 - 1.58	1.90 - 1.59	4.76 - 2.20	$(15.93, 2.30 \times 10^{-2})$	$7.4 \pm 0.7$
30 [37][48]	25.78 - 5.88	2.20 - 1.81	2.08 - 1.81	5.40 - 2.62	$(15.92, 2.31 \times 10^{-2})$	$7.3^{+0.7}$

 ${ ilde \chi}^2_{5\sigma}=2.41$  for Huang,  ${ ilde \chi}^2_{5\sigma}=1.14$  for galkin.



### Best fit values for $\alpha$ y $\mu$

- Representative Morphology
  - $\alpha \in$  [14.44, 16.43],  $\mu \in$  [2.29, 2.68] imes 10<sup>-2</sup> kpc<sup>-1</sup>
  - $(\alpha,\mu)_{\rm BF} = (16.4, 2.68 imes 10^{-2})$  with  $\tilde{\chi}^2_{Huang} = 8.60$

### Best Fit Morphology

- $\alpha \in$  [14.67, 16.59],  $\mu \in$  [2.13, 2.52]  $\times 10^{-2} \, {\rm kpc^{-1}}$
- $(\alpha,\mu)_{\rm BF} = (16.59, 2.52 \times 10^{-2})$  with  $\tilde{\chi}^2_{\it Huang} = 2.78$

$$\alpha = \frac{M}{(\sqrt{M} + E)^2} \left(\frac{G_{\infty}}{G_N} - 1\right) \qquad \qquad \mu = \frac{D}{\sqrt{M}}$$

Conclusions



# Representative morphology



Conclusions



# Best Fit morphology



# Summary and Conclusions

- We have performed a test of Moffat's MOG theory in the Milky Way using two recent compilations of data for the observed Rotation Curve, and adopting a complete set of observationally inferred morphologies for the stellar and gas components.
- We have also modified the key parameters of the theory, in order to match them to the baryonic mass of the Milky Way of each of our baryonic models.
- We conclude that modifying the gravitational potential according to the MOG theory, does not explain the discrepancy between the observed rotation curve, and that generated by the baryons only, in the Milky Way.

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- We conclude that modifying the gravitational potential according to the MOG theory, does not explain the discrepancy between the observed rotation curve, and that generated by the baryons only, in the Milky Way.

#### Gracias! Obrigado! Thank you!

### Best Fit morphology with only 12 data sets from galkin

