

1. Let  $\xi^\alpha$  be a Killing vector. Show that the motion of a free particle conserves  $u_\alpha \xi^\alpha$ , using the geodesic equation in the form

$$u^\beta \nabla_\beta u^\alpha = 0.$$

For  $\xi^\alpha = t^\alpha$  and for  $\xi^\alpha = \phi^\alpha$ , what is the physical meaning of the conserved quantities? Check for a particle in flat space (in special relativity) that your identification is right.

2. Dragging of inertial frames. The metric of a rotating star or black hole has the form

$$ds^2 = -e^{2\nu}(dt - \omega d\phi)^2 + e^{2\psi} d\phi^2 + e^{2\mu}(d\varpi^2 + dz^2).$$

Show that a particle with zero angular momentum has angular velocity  $\omega$ . In particular, a particle dropped from rest at infinity, rotates with angular velocity  $\omega$  in the direction of the star's rotation.

3. Using the fact that the metric is asymptotically flat, show that the asymptotic form of the equation for  $\omega$  is

$$\nabla \cdot (r^2 \sin^2 \theta \nabla \omega) = 0,$$

and check that the equation is satisfied by the asymptotic form  $\omega = \frac{2J}{r^3}$ , with  $J$  a constant (we will find that  $J$  is the angular momentum of the spacetime). Consider a spherical shell rotating with angular velocity  $\Omega$ . For a nearly Newtonian shell,  $\omega \ll \Omega$ , and the metric coefficients in the equation for  $\omega$  are again those of flat space, so that the equation for  $\omega$  has the form

$$\nabla \cdot (r^2 \sin^2 \theta \nabla \omega) = -16\pi \rho r \sin^2 \theta \Omega,$$

Here  $\rho = \frac{m}{4\pi r_0^2} \delta(r - r_0)$  from the requirement  $\int \rho dV = m$ . Find the spherically symmetric solution for  $\omega$ .