

# Entanglement Shadows

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Contains results from “Casting Shadows on Holographic Reconstruction”  
by B. Freivogel, R. Jefferson, L. Kabir, B. Mosk and I-S. Yang  
hep-th/1412.5175

# Outline

- 1 Holographic Entanglement Entropy
- 2 Entanglement Shadows
- 3 Outlook

# Entanglement Entropy

General and useful tool to measure correlations in quantum system which is dived in two parts

$$S = -\text{Tr} \rho_A \log \rho_A \quad \rho_A = \text{Tr}_B \rho_{AB}$$

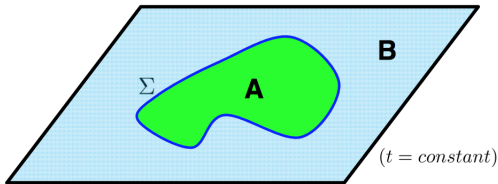


Image: R. Myers

# Holographic Entanglement Entropy

Ryu & Takayanagi '06

$$S(A) = \min_{\partial V = \Sigma} \frac{A_V}{4G}$$

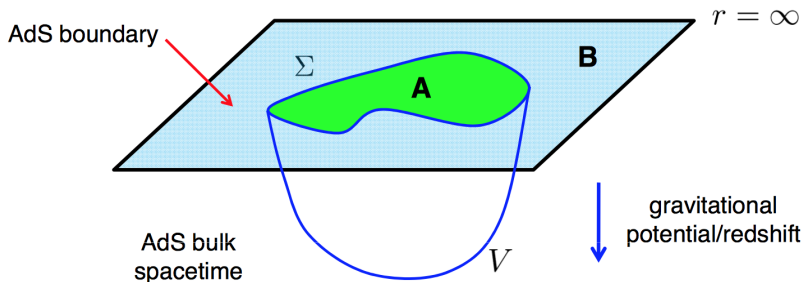
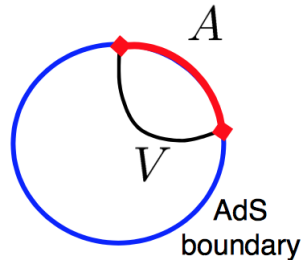
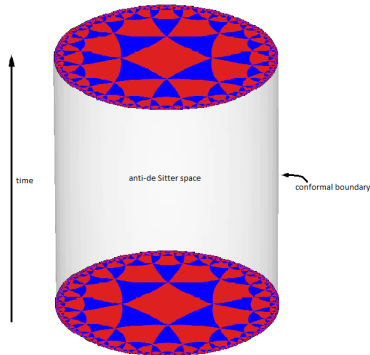


Image: R. Myers



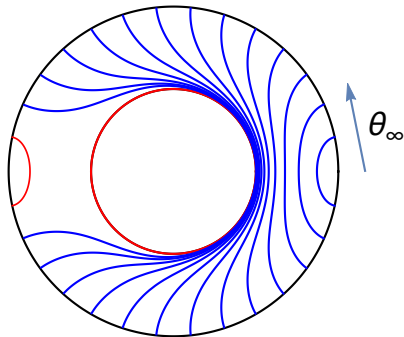
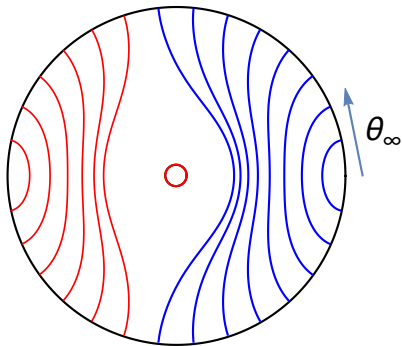
Reveals a link between geometry and quantum information

Reconstructing the bulk from the boundary:  
Gravitation from entanglement

[T Faulkner, M Guica, T Hartman, R Myers, M Van Raamsdonk, '13]

# Entanglement Shadows

Example: put a black hole in AdS and study the RT-surfaces for different boundary regions



# Holographic Shadows

What about other probes?

	Minimal Area	Wilson Loop	Causal
$d = 2, r_H \ll l_{\text{AdS}}$	$\sim l_{\text{AdS}}$	$\sim l_{\text{AdS}}$	$\sim l_{\text{AdS}}$
$d = 2, r_H \gg l_{\text{AdS}}$	$\sim e^{-r_H/l_{\text{AdS}}}$	$\sim r_H$	$\sim e^{-r_H/l_{\text{AdS}}}$
$d > 2, r_H \ll l_{\text{AdS}}$	$\sim r_H$	$\sim (r_H l_{\text{AdS}})^{1/2}$	$\sim e^{-l_{\text{AdS}}/r_H}$
$d > 2, r_H \gg l_{\text{AdS}}$	$\sim e^{-r_H/l_{\text{AdS}}}$	$\sim r_H$	$\sim e^{-r_H/l_{\text{AdS}}}$

**Table :** Shadow summary for various probes in  $SAdS_{d+1}$ .

# Outlook

Shadows are very generic.



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Shadows are very generic.

And they leave a lot of questions:

- how does bulk field reconstruction work in the shadow?
- can Einstein equations still be reconstructed?
- how are bulk dof's in the shadow region encoded in the CFT?

## Summary

Holographic entanglement entropy is a wonderful probe connecting geometry and entanglement, although shadow regions form a serious obstacle to reconstruct (local) bulk physics from the CFT.

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Thanks for your attention!

Questions?



## Backup Slide: Wilson loops

Wilson loops can encounter a deconfining phase transition, which make them have a larger shadow

Only for a small BTZ black hole, this transition will happen after  $\theta_\infty > \frac{\pi}{2}$ , hence it will never encounter this deconfining transition.

