

From data to theory and back

Exercises

August 3rd, 2015

Exercise1 Linearized Riemann, Ricci and Einstein tensors

Using that the Christoffel symbols at linear level are

$$\Gamma_{\mu\nu}^{\alpha} = \frac{1}{2} (\partial_{\mu} h_{\nu}^{\alpha} + \partial_{\nu} h_{\mu}^{\alpha} - \partial^{\alpha} h_{\mu\nu})$$

derive eqs.

$$\begin{aligned} R_{\mu\nu\rho\sigma} &= \frac{1}{2} (\partial_{\nu} \partial_{\rho} h_{\mu\sigma} + \partial_{\mu} \partial_{\sigma} h_{\nu\rho} - \partial_{\mu} \partial_{\rho} h_{\nu\sigma} - \partial_{\nu} \partial_{\sigma} h_{\mu\rho}) , \\ R_{\mu\nu} &= \frac{1}{2} (\partial_{\rho} \partial_{\mu} h_{\nu}^{\rho} + \partial_{\rho} \partial_{\nu} h_{\mu}^{\rho} - \square h_{\mu\nu} - \partial_{\mu} \partial_{\nu} h) , \\ R &= \partial_{\mu} \partial_{\nu} h^{\mu\nu} - \square h , \\ G_{\mu\nu} &= \frac{1}{2} (\partial_{\rho} \partial_{\mu} h_{\nu}^{\rho} + \partial_{\rho} \partial_{\nu} h_{\mu}^{\rho} - \square h_{\mu\nu} - \partial_{\mu} \partial_{\nu} h - \eta_{\mu\nu} \partial_{\rho} \partial_{\rho} h^{\mu\nu} + \eta_{\mu\nu} \square h) , \end{aligned} \tag{1}$$

Exercise2 Retarded Green function I

Show that the two representation of the retarded Green function given by

$$\begin{aligned} G_{ret}(t, \mathbf{x}) &= -\delta(t-r) \frac{1}{4\pi r} , \\ G_{adv}(t, \mathbf{x}) &= -\delta(t+r) \frac{1}{4\pi r} , \end{aligned}$$

and

$$\begin{aligned} G_{ret}(t, x) &= -i\theta(t) (\Delta_{+}(t, x) - \Delta_{-}(t, x)) , \\ G_{adv}(t, x) &= i\theta(-t) (\Delta_{+}(t, x) - \Delta_{-}(t, x)) , \end{aligned}$$

where

$$\Delta_{\pm}(t, x) \equiv \int_{\mathbf{k}} e^{\mp ikt} \frac{e^{i\mathbf{k}\mathbf{x}}}{2k}$$

are equivalent. Hint: use that

$$\int_{-\infty}^{\infty} \frac{dk}{2\pi} e^{ikx} = \delta(x) ,$$

and that

$$\theta(t) \int_{-\infty}^{\infty} \frac{dk}{2\pi} e^{ik(t+r)} = 0 \quad \text{for } r \geq 0.$$

Exercise3 Retarded Green function II

Use the representation of the $G_{ret,adv}$ obtained in the previous exercise to show that

$$G_{ret}(t, \mathbf{x}) = - \int_{\mathbf{k}} \frac{d\omega}{2\pi} \frac{e^{-i\omega t + i\mathbf{k}\mathbf{x}}}{k^2 - (\omega + i\epsilon)^2},$$

$$G_{adv}(t, \mathbf{x}) = - \int_{\mathbf{k}} \frac{d\omega}{2\pi} \frac{e^{-i\omega t + i\mathbf{k}\mathbf{x}}}{k^2 - (\omega - i\epsilon)^2}.$$

where ϵ is an arbitrarily small positive quantity. Hint: use that

$$\theta(\pm t) = \mp \frac{1}{2\pi i} \int \frac{e^{-i\omega t}}{\omega \pm i\epsilon}.$$

Show that G_{ret} is real.

Exercise4 Feynman Green function I

Show that the G_F defined by

$$G_F(t, \mathbf{x}) = -i \int_{\mathbf{k}} \frac{d\omega}{2\pi} \frac{e^{-i\omega t + i\mathbf{k}\mathbf{x}}}{k^2 - \omega^2 - i\epsilon}$$

is equivalent to

$$G_F(t, \mathbf{x}) = \theta(t)\Delta_+(t, \mathbf{x}) + \theta(-t)\Delta_-(t, \mathbf{x}).$$

Derive the relationship

$$G_F(t, \mathbf{x}) = \frac{i}{2} (G_{adv}(t, \mathbf{x}) + G_{ret}(t, \mathbf{x})) + \frac{\Delta_+(t, \mathbf{x}) + \Delta_-(t, \mathbf{x})}{2}.$$

Exercise5 Feynman Green function II

By integrating over \mathbf{k} the G_F in the $\sim 1/(k^2 - \omega^2)$ representation, show that G_F implements boundary conditions giving rise to field h behaving as

$$h(t, \mathbf{x}) \sim \int d\omega e^{-i\omega t + i|\omega|r},$$

corresponding to out-going (in-going) wave for $\omega > (<)0$.

Exercise6 TT gauge

Show that the projectors defined by

$$\Lambda_{ij,kl}(\hat{n}) = \frac{1}{2} [P_{ik}P_{jl} + P_{il}P_{jk} - P_{ij}P_{kl}],$$

$$P_{ij}(\hat{n}) = \delta_{ij} - n_i n_j,$$

satisfies the relationships

$$P_{ij}P_{jk} = P_{ik}$$

$$\Lambda_{ij,kl}\Lambda_{kl,mn} = \Lambda_{ij,mn},$$

which characterize projectors operator.

Exercise7 Energy of circular orbits in a Schwarzschild metric

Consider the Schwarzschild metric

$$ds^2 = -\left(1 - \frac{2G_N M}{r}\right) dt^2 + \frac{dr^2}{\left(1 - \frac{2G_N M}{r}\right)} + r^2 d\Omega^2. \quad (2)$$

The dynamics of a point particle with mass m moving in such a background can be described by the action

$$S = -m \int d\tau = -m \int d\lambda \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda}}$$

for any coordinate λ parametrizing the particle world-line. Using $S = \int d\lambda L$, we can write

$$L = -m \left[\left(1 - \frac{2G_N M}{r}\right) \left(\frac{dt}{d\tau}\right)^2 - \frac{\left(\frac{dr}{d\tau}\right)^2}{\left(1 - \frac{2G_N M}{r}\right)} - r^2 \left(\frac{d\phi}{d\tau}\right)^2 \right]^{1/2}.$$

Verify that L has cyclic variables t and ϕ and derive the corresponding conserved momenta.

(Hint: use $g_{\mu\nu}(dx^\mu/d\tau)(dx^\nu/d\tau) = -1$. Result: $E = m(dt/d\tau)(1 - 2G_N M/r) \equiv m e$ and $L = m r^2(d\phi/d\tau) \equiv m l$).

By expressing $dt/d\tau$ and $d\phi/d\tau$ in terms of e and l , derive the relationship

$$e^2 = (1 - 2G_N M/r) \left(1 + l^2/r^2\right) + \left(\frac{dr}{d\tau}\right)^2.$$

From the circular orbit conditions ($\frac{de}{dr} = 0 = \frac{dr}{d\tau} = 0$), derive the relationship between l and r for circular orbits.

(Result: $l^2 = r^2/(\frac{r}{M} - 3)$).

Substitute into the energy function e and find the circular orbit energy

$$e(x) = \frac{1 - 2x}{\sqrt{1 - 3x}},$$

where $x \equiv (G_N M \dot{\phi})^{2/3}$ is an observable quantity as it is related to the GW frequency f_{GW} by $x = (G_N M \pi f_{GW})^{2/3}$.

(Hint: Use

$$\dot{\phi} = \frac{d\phi}{d\tau} \dot{\tau} = \frac{l}{r^2} \left[1 - \frac{2M}{r} - r^2 \dot{\phi}^2\right]^{1/2}$$

to find that on circular orbits $(M\dot{\phi})^2 = (G_N M/r)^3$, an overdot stands for derivative with respect to t .)

Derive the relationships for the Inner-most stable circular orbit

$$\begin{aligned} r_{ISCO} &= 6G_N M = 4.4\text{km} \left(\frac{M}{M_\odot} \right) \\ f_{ISCO} &= \frac{1}{6^{3/2}} \frac{1}{G_N M \pi} \simeq 8.8\text{kHz} \left(\frac{M}{M_\odot} \right)^{-1} \\ v_{ISCO} &= \frac{1}{\sqrt{6}} \simeq 0.41 \end{aligned}$$

Exercise8 Newtonian force exerced by GWs

Derive the equivalent Newtonian-like force

$$\ddot{\xi} = \frac{1}{2} \ddot{h}_{ij} \xi^j, \quad (3)$$

from the geodesic deviation equation

$$\frac{D^2 \xi^i}{d\tau^2} = -R^i{}_{0j0} \xi^j \left(\frac{dx^0}{d\tau} \right)^2.$$

Exercise9 World-line action

Derive the geodesic equation

$$\ddot{x}^\mu + \Gamma_{\rho\sigma}^\mu \dot{x}^\rho \dot{x}^\sigma = 0$$

from the world-line action

$$S_{wl} = \int dt d^3y \sqrt{-g_{\mu\nu} \dot{y}^\mu \dot{y}^\nu} \delta^{(3)}(y - x(t))$$