

Building a DM candidate

DM Flash-School
19/11/2018

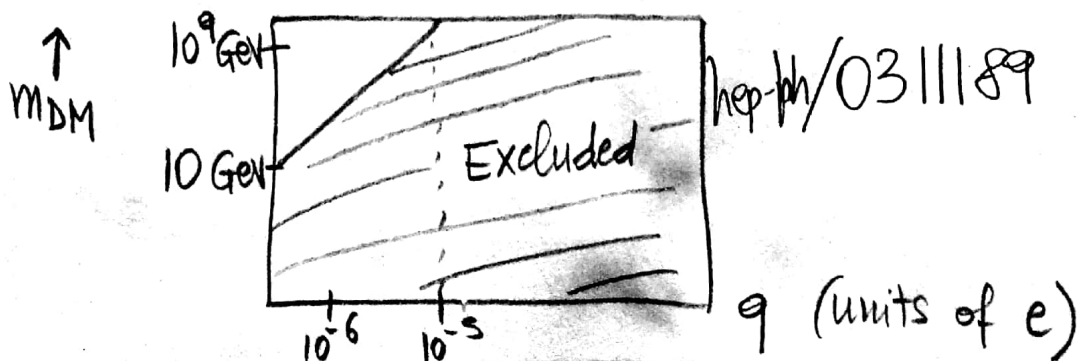
What we need to keep in mind to build a DM candidate?

- Lot of negatives
- No idea of what DM is (many candidates)

(Important) Properties

1. Massive (not radiation)
2. Nearly pressureless ($w=0$) @ $T \sim \text{keV} \Rightarrow \text{COLD}$
since then
3. Neutral (electrically)

→ need to ensure that the EM contact between DM & plasma do not delay pert. growth until recombination



Another possibility $\rightarrow$ maybe DM is neutral bound state as hydrogen (H-like)

Heavy hydrogen searches

$$\begin{array}{ccc} \text{frac} \lesssim 4 \times 10^{-14} & \text{for} & 5 m_p \lesssim M_{DM} \lesssim 1600 m_p \\ \uparrow & & \underbrace{\hspace{1cm}} \quad \underbrace{\hspace{1cm}} \\ \text{of H in} & & 5 \text{ GeV} \quad \quad \quad 1.6 \text{ TeV} \\ \text{heavy H form} & & \end{array}$$

Aside: can DM be colored (= strongly interacting)?

1801.01135 shows that it is possible if heavy enough

($M_{DM} \sim 12 \text{ TeV}$)
if $\chi = \text{octet of } SU(3)_c$

4. Interactions $\rightarrow$ for sure with gravity $\checkmark$
(exp. obs.)

$\rightarrow$ maybe with SM
(necessity if thermally produced)

$\rightarrow$ maybe with itself (in addition to grav.)

$$\frac{\sigma_{DM-DM}}{M_{DM}} \lesssim 1 \text{ cm}^2/\text{g} \quad [\text{Bullet cluster}]$$

5. Stable

How much? At least over the age of universe

1606.02073 DM $\rightarrow$ inv. particles

$$f_{\text{DM}} \cdot T_{\text{DM}} \lesssim 0.086 / T_{\text{universe}}$$

$\hookrightarrow$ fraction of decaying DM

That's about what we know

How can we implement this in a model?

Framework $\rightarrow$ QFT

DM $\rightarrow$ associated w/ a field
(scalar, fermion, vector...)
whose excitations we
interpret as a DM particle

Key object to do computations $\rightarrow \mathcal{L}$

SM $\rightarrow$ non-abelian gauge theory (massless vectors)
with SSB (some vectors become massive)

Theory completely determined once we give its
- symmetries $SU(3)_C \times SU(2)_L \times U(1)_Y$

- particle content

$$SM \begin{cases} q = (3, 2, \frac{1}{6}) \\ l = (1, 2, \frac{1}{2}) \\ u^c = (\bar{3}, 1, -\frac{2}{3}) \\ d^c = (\bar{3}, 1, -\frac{1}{3}) \\ e^c = (1, 1, +1) \end{cases}$$

Stable particles in SM $\rightarrow p, e, \nu$

Only DM candidate $\rightarrow \nu$ [inside $l = \begin{pmatrix} \nu \\ e \end{pmatrix}$]

BUT $\Omega_\nu \simeq 0.005$ vs. $\Omega_{DM} \simeq 0.12$
+ ν is hot candidate (rel. @ decoupling)

$\Rightarrow$ we need to add something new to the SM
[evidence for physics Beyond SM]

How do we implement the properties listed before?

1. - we need to be able to write a mass term

$\rightarrow$ not difficult

$$\mathcal{L} \supset -m^2 \chi^\dagger \chi \quad (\chi \text{ scalar})$$

$$\mathcal{L} \supset -m \chi \chi \quad (\chi \text{ fermion, Majorana mass})$$

$$\mathcal{L} \supset -m \chi^c \chi \quad (\chi \text{ fermion, Dirac mass})$$

2 → Neutral $Q = T_L^3 + Y$

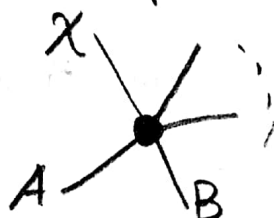
• If X is $SU(2)_L$ singlet ($T_L^3 = 0$)
 $\Rightarrow Y = 0$

• If X is $SU(2)_L$ multiplet ($T_L^3 \neq 0$)
 $\Rightarrow$ need to ensure a neutral component inside the multiplet

3 — Stable. Let's make X COMPLETELY STABLE

Aim: avoid decay $X \rightarrow A + B + \dots$

Decay happens when $m_X > m_A + m_B + \dots$ (*)
and comes from an interaction



Eq. (*) automatically not satisfied if (at least)
one between $A, B, \dots$ is X itself

because $m_X \geq m_X + m_B + \dots \Rightarrow 0 \geq m_B + \dots$
absurd!

$\Rightarrow$ If we guarantee that all interaction vertices contain at least 2 X , then $X =$ stable

Super easy way $\rightarrow$ ^{introduce} new parity, X odd

$$\begin{cases} X \rightarrow -X \\ SM \rightarrow SM \end{cases}$$

$$\mathcal{L} \supset \lambda X SM_1 SM_2 \dots \xrightarrow{\text{parity}} -\lambda X SM_1 SM_2 \dots$$

NOT INVARIANT

$$\mathcal{L} \supset \lambda XX SM_1 SM_2 \dots \xrightarrow{\text{parity}} \lambda XX SM_1 SM_2 \dots$$

INVARIANT
CONTAINS 2 X
 $\Rightarrow X$ STABLE

Other way $\rightarrow$ charge X under a continuous group (i.e. $U(1)$) avoids interactions containing one X only.

Possible caveat: it is believed that gravity does not respect global, continuous symm.

But it can arise as accidental symm (like B or L numb. in SM)

What you need to be able to compute?

1. Production mechanism (ensuring $\Omega_{DM} \simeq 0.12$)
2. Exp. bounds on cross section w/ nucleons
3. Exp. bounds on annihilation cross-section

Production mechanism

Thermal



DM was in eq. with SM thermal bath

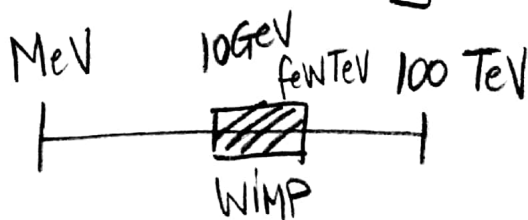


FREEZE OUT
(see KOLB-TURNER or DODELSON)



$\langle \sigma v \rangle_{\chi\chi \rightarrow \text{SM SM}}$

relevant quantity



Thermalish



intermediate
(not quite in eq.,
or in eq. only for
a while, or trying
to reach eq. ...)



- INCOMPLETE THERMALIZATION

- FREEZE-IN

- 3 → 2 PROCESSES
(SIMP DM)
(CANNIBAL DM)

- FORBIDDEN DM



$\langle \sigma v \rangle$ relevant quantity

Non-thermal



DM production not dictated by thermal contact with SM bath



- ASYMMETRIC DM
(1308.0338)

- COHERENT OSCILL.



usually nothing to do with

$\langle \sigma v \rangle$

Relevant quantities

$\langle \sigma v \rangle$

- production (except non-thermal)
- CMB bound
- ID

σ_n ($\chi_n \rightarrow \chi_n$)
- DD

Since DM non-rel

@ decoupling

we can expand in $v \ll 1$

(equivalent to partial wave expansion)

$$\langle \sigma v \rangle = \sigma_0 + \sigma_2 v^2 + \sigma_4 v^4 + \dots$$

S-wave p-wave d-wave