

Heavy quark masses from heavy quarkonium

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Motivation

m_Q . Fundamental parameter of the Standard Model.

Heavy Quarkonium (near threshold) physics very sensitive to its value.

Usual definitions:

- ▶ $m_{\overline{\text{MS}}}$ → short distance mass.
- ▶ m_{OS} → natural definition for heavy quark physics near threshold.

Heavy Quarkonium (non-perturbative = Coulomb resummation) dynamics.

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Nonrelativistic systems near threshold

$$v = \sqrt{1 - \frac{4m_Q^2}{s}} \ll 1$$

$$m_Q \gg m_Q v \gg m_Q v^2 \quad (\text{EFT}'s)$$

$\frac{\alpha_s}{v} \sim 1 \rightarrow$ Coulomb resummation \rightarrow Schrödinger equation (pNRQCD)

$$\text{LO} \sim \sum_{n=0}^{\infty} c_n \frac{\alpha_s^n}{v^n}$$

$$\text{NLO} \sim \sum_{n=0}^{\infty} c_n \frac{\alpha_s^n}{v^n} \times (\alpha_s, v)$$

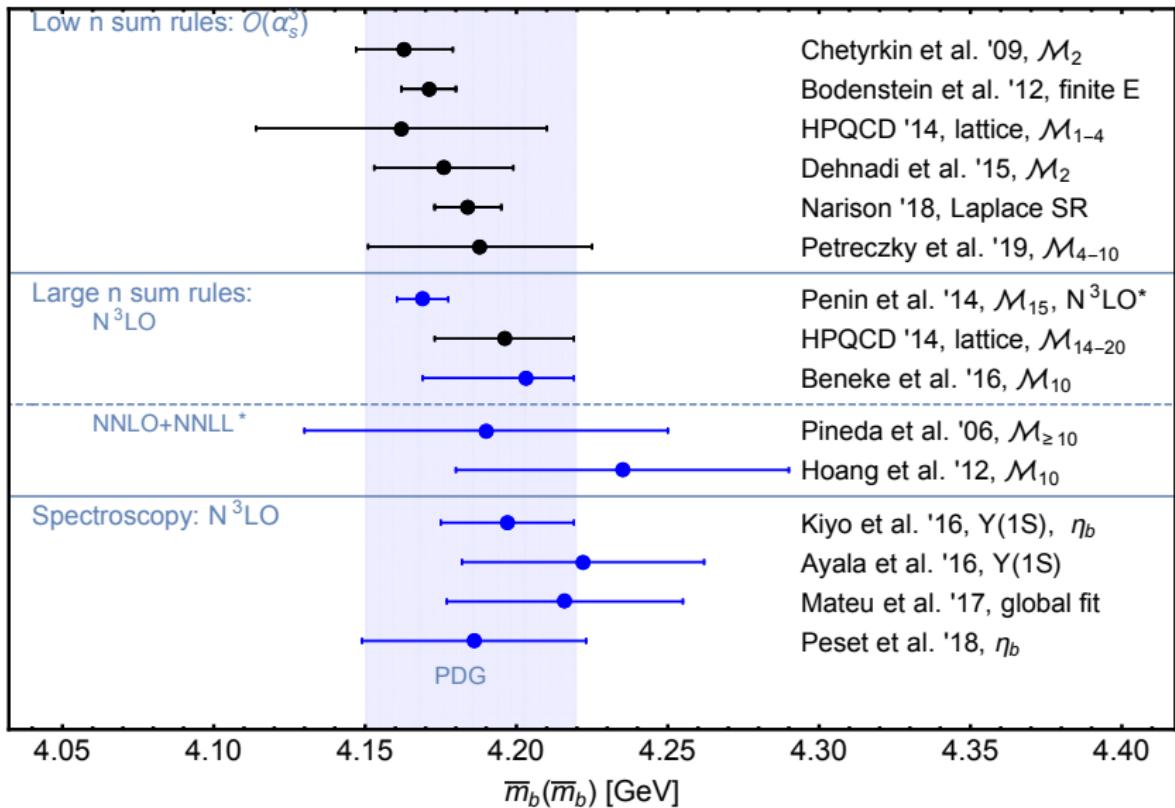
$$\text{NLL} \sim \sum_{n=0}^{\infty} c_n \frac{\alpha_s^n}{v^n} \times (\alpha_s, v) \sum_{k=0}^{\infty} d_k \alpha_s^k \ln^k(\alpha, v)$$

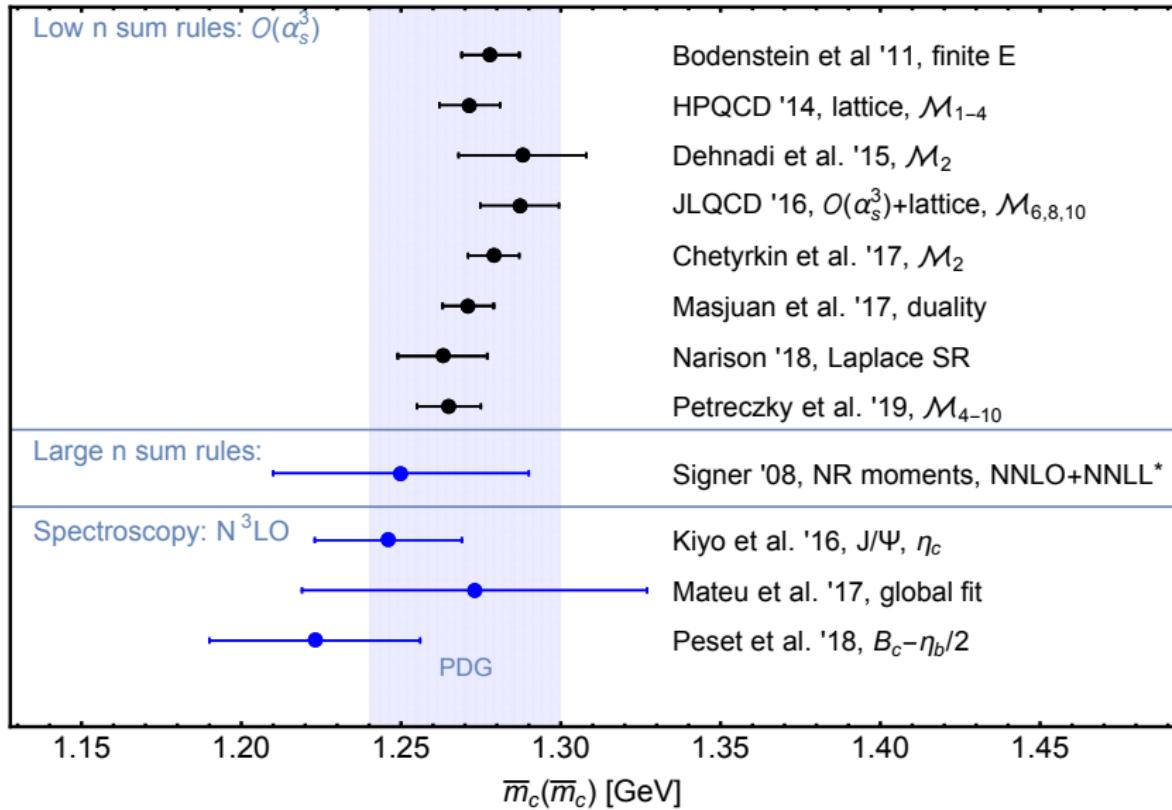
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Heavy Quarkonium (non-perturbative = Coulomb resummation) dynamics:

- ▶ Heavy Quarkonium mass
- ▶ Non-relativistic sum rules
- ▶ t - \bar{t} production near threshold

Highest sensitive to the heavy quark pole mass





Spectrum

$$M_n = 2m - \frac{m\alpha^2}{4n^2} \left(1 + \frac{\alpha}{\pi} P_1(L_\nu) + \frac{\alpha^2}{\pi^2} P_2(L_\nu) + \frac{\alpha^3}{\pi^3} P_3(L_\nu) \right)$$

- ▶ $\mathcal{O}(m\alpha^5 \ln \frac{1}{\alpha})$ $m_1 = m_2$; Brambilla et al. (1999)
- ▶ $\mathcal{O}(m\alpha^5)$ $m_1 = m_2$ n=1; Penin, Steinhauser (2002)
- ▶ $\mathcal{O}(m\alpha^5)$ $m_1 = m_2$ S-wave; Beneke et al., Penin et al. (2005)
- ▶ $\mathcal{O}(m\alpha^5)$ $m_1 = m_2$ Kiyo, Sumino (2014)
- ▶ $\mathcal{O}(m\alpha^5)$ $m_1 \neq m_2$ Peset, Pineda, Stahlhofen (2015)

Fixed order computation → NNNLO

Resummation of logarithms → NNNLL (not yet complete)

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Pole mass and renormalons

$$m_{\text{OS}} = m_{\overline{\text{MS}}} + \sum_{n=0}^{\infty} r_n \alpha_s^{n+1},$$

Renormalon (OPE) analysis predicts $r_n \sim n!$. Therefore m_{OS} suffers from renormalon ambiguities:

$$\delta_{np}^{(\text{pert.})} m_{\text{OS}} = \delta_{np}^{(\text{pert.})} (m_{\overline{\text{MS}}} + r_0 \alpha_s + r_1 \alpha_s^2 + \dots) \sim \Lambda_{\text{QCD}}!$$

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$$m_{\text{OS}} = m_{\overline{\text{MS}}} + \int_0^\infty dt e^{-t/\alpha_s} B[m_{\text{OS}}](t), \quad B[m_{\text{OS}}](t) \equiv \sum_{n=0}^\infty r_n \frac{t^n}{n!}.$$

The behavior of the perturbative expansion at large orders is dictated by the closest singularity to the origin of its Borel transform ($u = \frac{\beta_0 t}{4\pi}$).

$$B[m_{\text{OS}}](t) = N_m \nu \frac{1}{(1 - 2u)^{1+b}} (1 + c_1(1 - 2u) + \dots) + (\text{analytic term}),$$

$$r_n \xrightarrow{n \rightarrow \infty} N_m \nu \left(\frac{\beta_0}{2\pi} \right)^n \frac{\Gamma(n+1+b)}{\Gamma(1+b)} \left(1 + \frac{b}{(n+b)} c_1 + \dots \right).$$

$$b = \frac{\beta_1}{2\beta_0^2}, \quad c_1 = \frac{1}{4b\beta_0^3} \left(\frac{\beta_1^2}{\beta_0} - \beta_2 \right), \quad \dots$$

Beneke, Pineda, ...

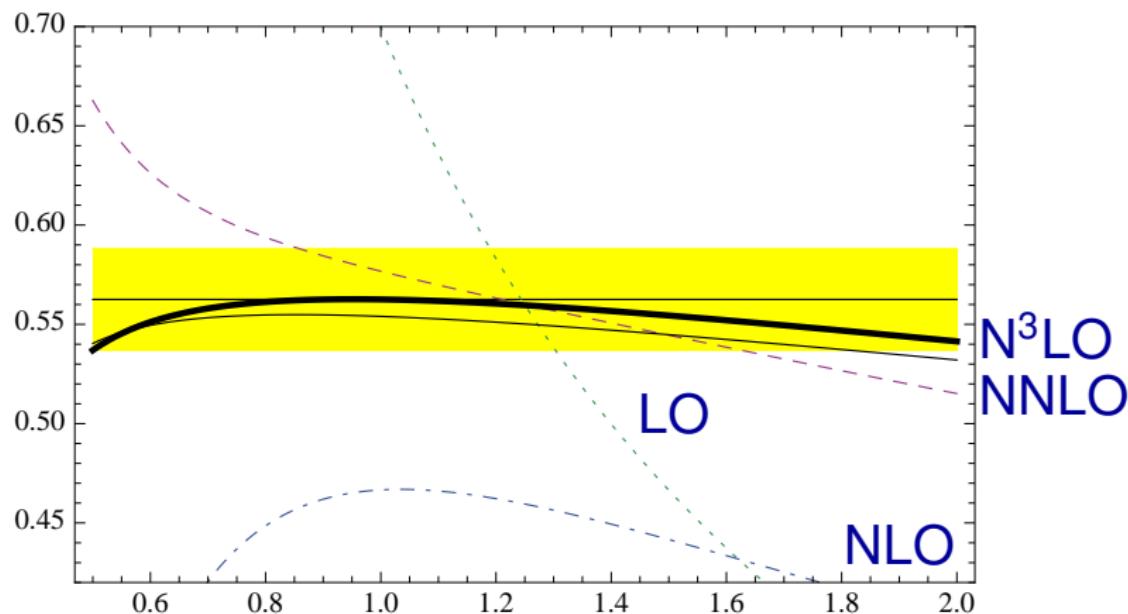


Figure: $2m + V$ renormalon free. $-N_V/2 = N_m$ for $n_l = 3$, as a function of $x \equiv \nu r$, obtained from $-(N_V/2)v_n/v_n^{\text{asym}}$. v_n^{asym} is truncated at $\mathcal{O}(1/n^3)$.

$$N_m(n_l = 0) = 0.600(29),$$

$$N_m(n_l = 3) = 0.563(26).$$

$$M_{\Upsilon(nS)} = 2m_{\text{OS}} + E_n^{\text{OS}} = 2m_{\text{OS}} + A_n^{1,\text{OS}} \alpha^2 + A_n^{2,\text{OS}} \alpha^3 + \dots$$

$M_{\Upsilon(nS)}$ is renormalon free $\rightarrow E_n^{\text{OS}}$ has a renormalon.

Renormalon free definitions:

$$m_{\text{RS}}(\nu_f) = m_{\text{OS}} - \sum_{n=0(1)}^{\infty} N_m \nu_f \left(\frac{\beta_0}{2\pi} \right)^n \alpha_s^{n+1}(\nu_f) \sum_{k=0}^{\infty} c_k \frac{\Gamma(n+1+b-k)}{\Gamma(1+b-k)}.$$

Expansion in $\alpha_s(\nu)$

$$m_{\text{RS}}(\nu_f) = m_{\overline{\text{MS}}} + \sum_{n=0}^{\infty} r_n^{\text{RS}} \alpha_s^{n+1},$$

where $r_n^{\text{RS}} = r_n^{\text{RS}}(m_{\overline{\text{MS}}}, \nu, \nu_f)$ are of natural size. We now do not lose accuracy if we first obtain m_{RS} and later on $m_{\overline{\text{MS}}}$.

$$M_{\Upsilon(nS)} = 2m_{\text{RS}} + E_n^{\text{RS}} = 2m_{\text{RS}} + \sum_{m=1}^{\infty} A_n^{m,\text{RS}} \alpha_s^{m+1}.$$

Alternative: kinetic, PS, MSR, MRS, PV, ...

Issues

Bare data set: $\bar{b}b$, $\bar{c}c$, $\bar{b}c$,

Options:

Ground state only

Ground state +excitations

Pseudoscalar/vector

Bottom mass determinations: finite mass charm quark effects.

NP vs ultrasoft effects

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$n_f = 3$ light fermions in the $\overline{\text{MS}}$ -pole mass relation

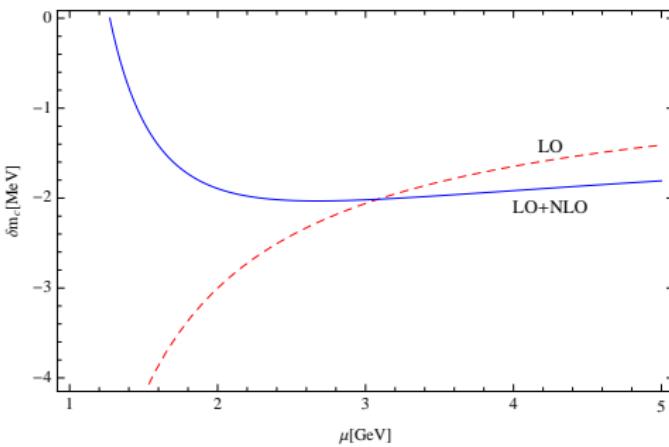
Light fermions: $n_f = 4 \rightarrow n_f = 3$

For soft scale Brambilla et al.:

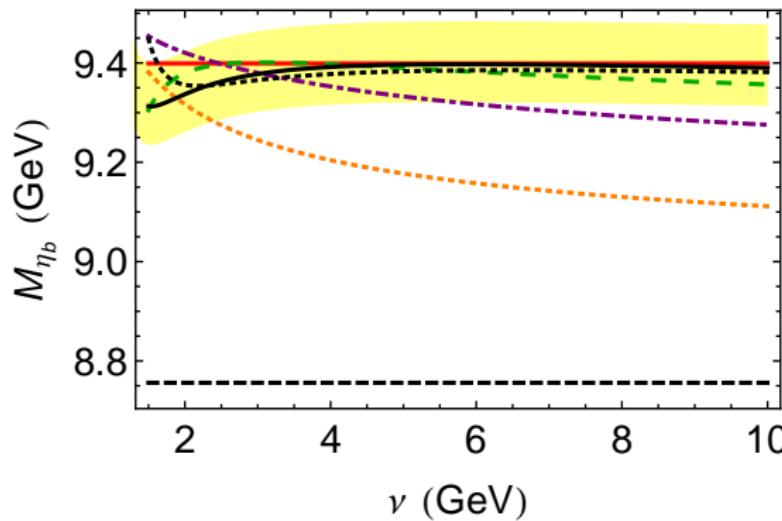
$$m_b \alpha_s < m_c$$

For hard scale Ayala et al.:

$$m_b e^{-n} < m_c$$

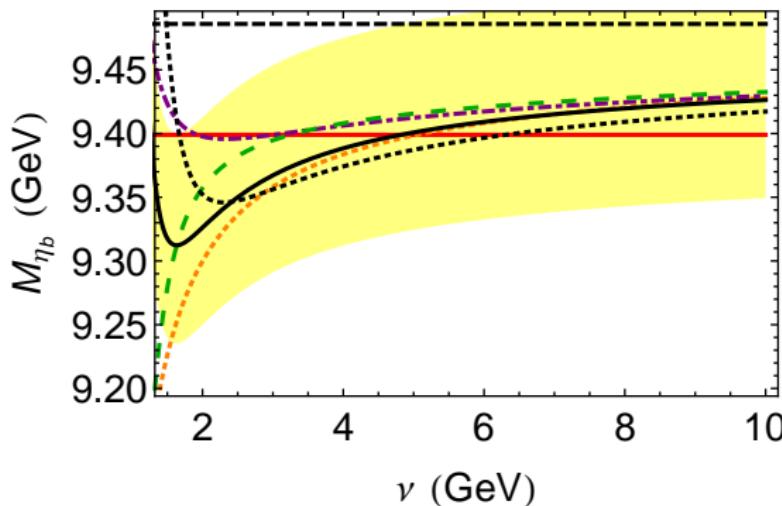


\overline{m}_b from M_{η_b}



RS scheme
Good convergence

\overline{m}_b from M_{η_b}



RS' determination (RS determination virtually identical); $\nu = 5$ GeV

$$\overline{m}_b(\overline{m}_b) = 4186(37) \text{ MeV}$$

Error dominated by ν dependence.
Ultrasoft effects are not important.

Ultrasoft and NP effects

$$\delta E \Big|_{\mathcal{O}(\alpha_s^5)} = A m \alpha_s^5 \left(\ln \frac{m \alpha_s}{\nu_{us}} + K \right) \Big|_{\text{soft}} + \delta E_{us}(\nu_{us})$$

$$\delta E_{us}(\nu_{us}) \simeq \frac{T_F}{3N_c} \int_0^\infty dt \langle n | \mathbf{r} e^{-t(H_0 - E_n)} \mathbf{r} | n \rangle \langle g \mathbf{E}^a(t) \phi(t, 0)_{ab}^{\text{adj}} g \mathbf{E}^b(0) \rangle (\nu_{us}),$$

where $H_0 \equiv \frac{\mathbf{p}^2}{m} + \frac{1}{2N_c} \frac{\alpha_s}{r}$ and $E_n \equiv -m C_f^2 \alpha_s^2 / (4n^2)$.

If $\Lambda_{\text{QCD}} \ll m \alpha_s^2$

$$\delta E_{us}(\nu_{us}) = \delta E_{us, \text{pert.}} + \delta E_{us, \text{nopert.}},$$

$$\delta E_{us, \text{pert.}} = A m \alpha_s^5 \left(\ln \frac{\nu_{us}}{m \alpha_s^2} + K' \right), \quad \delta E_{us, \text{nopert.}} = \sum_{n=0}^{\infty} C_n O_n,$$

where $C_n \sim 1/(m^{3+2n} \alpha_s^{4+4n})$ and $O_n \sim \Lambda_{\text{QCD}}^{4+2n}$.

Voloshin, Leutwyler, Pineda, Rauh

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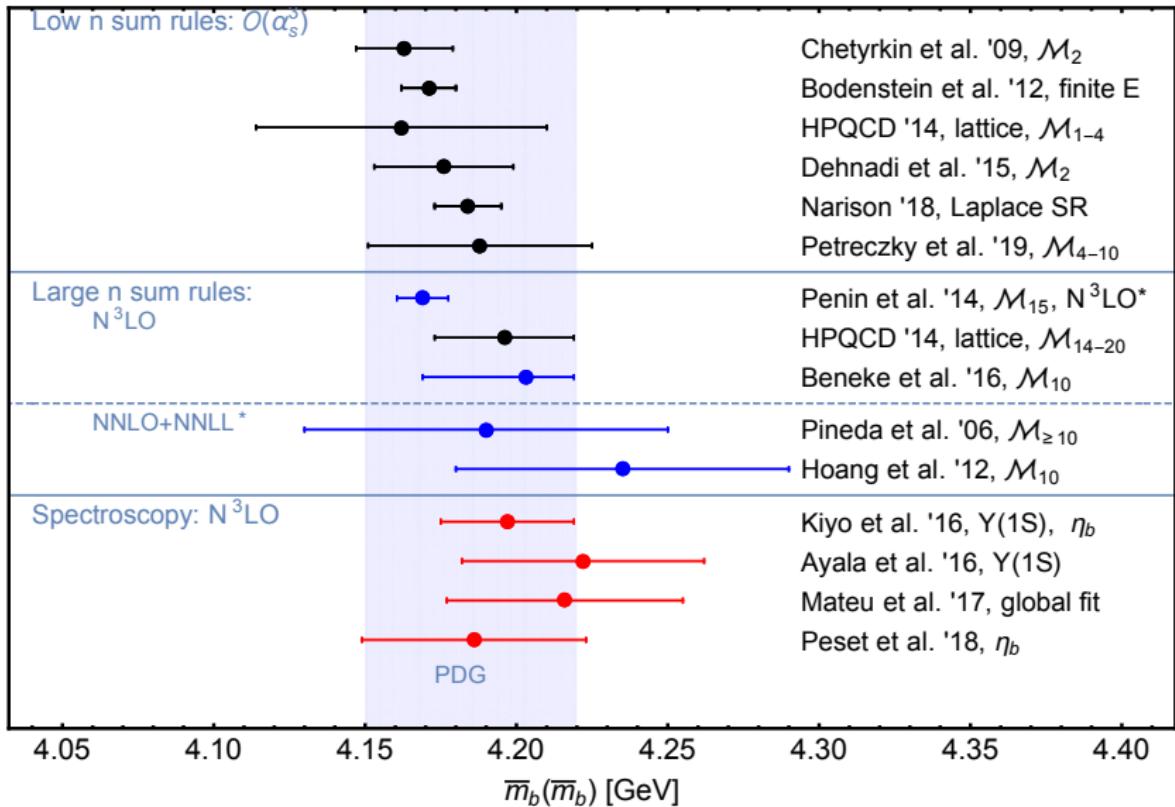
$$\delta E \Big|_{\mathcal{O}(\alpha_s^5)} = A m \alpha_s^5 \left(\ln \frac{m \alpha_s}{\nu_{us}} + K \right) \Big|_{\text{soft}} + \delta E_{us}(\nu_{us})$$

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If $\Lambda_{\text{QCD}} \gg m \alpha_s^2$

$$\delta E_{us}(\nu_{us}) \sim \Lambda_{\text{QCD}}^3 \langle r^2 \rangle_n$$



\overline{m}_c from $M_{B_c} - M_{\eta_b}/2$ and M_{η_c}

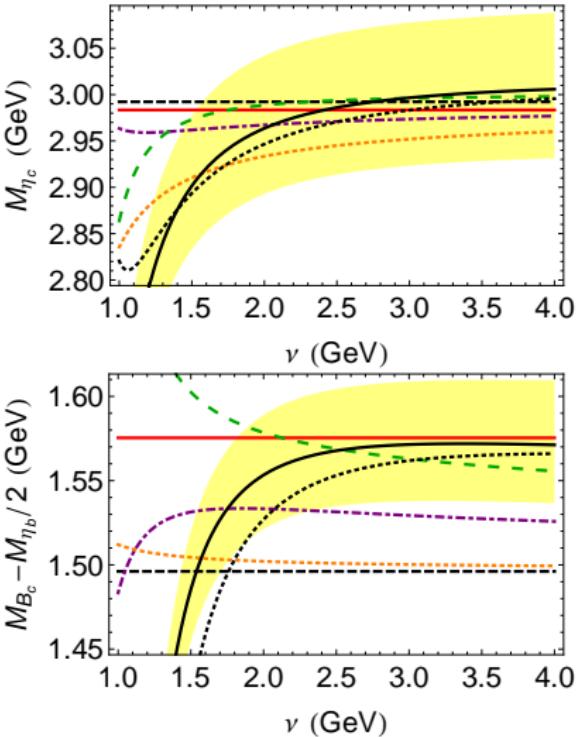
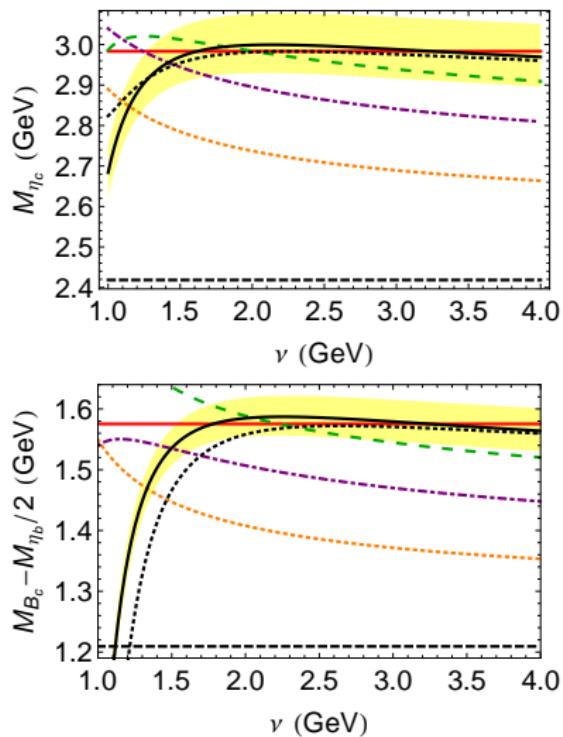


Figure: **Left panels:** RS scheme. **Right panels:** RS' scheme.

Determinations from M_{η_c}

$$\begin{aligned} m_{c,\text{RS}}(1\text{GeV}) &= 1202^{+15}_{-16}(\nu)^{-15}_{+11}(\nu_f)^{-10}_{+10}(\alpha_s)^{-34}_{+34}(N_m); \\ \Rightarrow \bar{m}_c &= 1217^{+12}_{-13}(\nu)^{-1}_{-6}(\nu_f)^{-8}_{+8}(\alpha_s)^{+10}_{-10}(N_m). \end{aligned}$$

$$\begin{aligned} m_{c,\text{RS'}}(1\text{ GeV}) &= 1495^{-11}_{+50}(\nu)^{-9}_{+20}(\nu_f)^{+4}_{-4}(\alpha_s)^{-20}_{+20}(N_m); \\ \Rightarrow \bar{m}_c &= 1222^{-9}_{+40}(\nu)^{-7}_{+16}(\nu_f)^{-7}_{+7}(\alpha_s)^{+11}_{-11}(N_m), \end{aligned}$$

Determinations from $M_{B_c} - M_{\eta_c}/2$

$$\begin{aligned} m_{c,\text{RS}}(1\text{GeV}) &= 1204^{+27}_{-8}(\nu)^{-26}_{+18}(\nu_f)^{-17}_{+16}(\alpha_s)^{-33}_{+33}(N_m)^{-1}_{+1}(\bar{m}_b); \\ \Rightarrow \bar{m}_c &= 1220^{+21}_{-6}(\nu)^{-7}_{-4}(\nu_f)^{-14}_{+13}(\alpha_s)^{+11}_{-11}(N_m)^{-1}_{+1}(\bar{m}_b). \end{aligned}$$

$$\begin{aligned} m_{c,\text{RS'}}(1\text{ GeV}) &= 1501^{+1}_{+23}(\nu)^{-14}_{-27}(\nu_f)^{-2}_{+2}(\alpha_s)^{-18}_{+18}(N_m)^{-1}_{+1}(\bar{m}_b); \\ \Rightarrow \bar{m}_c &= 1227^{-1}_{+18}(\nu)^{+11}_{-22}(\nu_f)^{-12}_{+12}(\alpha_s)^{+13}_{-13}(N_m)^{-1}_{+1}(\bar{m}_b) \end{aligned}$$

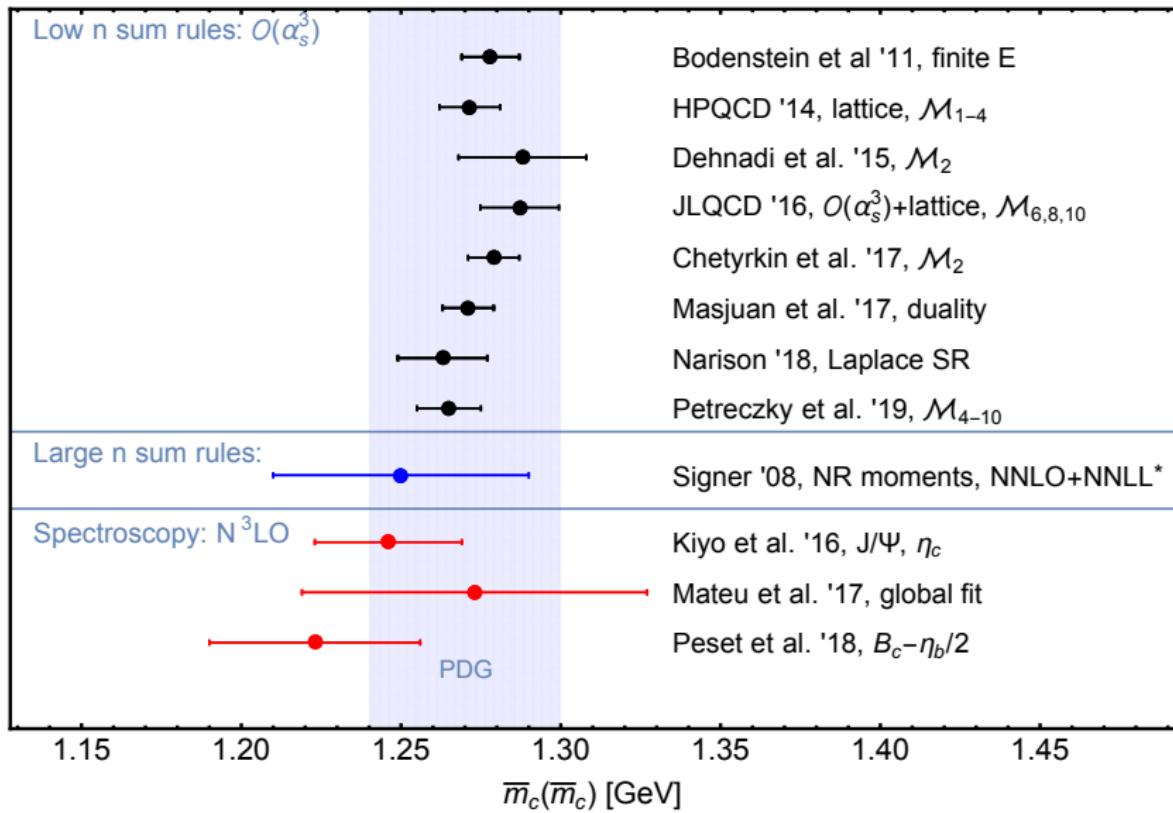
$$\bar{m}_c(\bar{m}_c) = 1223(33) \text{ MeV}$$

NP effects vs ultrasoft effects

$$\delta E \sim \Lambda_{\text{QCD}}^3 \left(\langle r^2 \rangle_{Bc} - \frac{1}{2} \langle r^2 \rangle_{\eta_b} \right)$$

and for the latter

$$\delta E \sim \Lambda_{\text{QCD}}^3 \langle r^2 \rangle_{\eta_c} .$$



Extra comments:

- ▶ Vector fits appear to give larger values for the masses than pseudoscalars
Possible explanation → hyperfine splitting sensitive to the resummation of logarithms.
- ▶ Resummation of logarithms will increase sensitivity to smaller scales.
- ▶ Experimental precision infinity → no need for excitations that may introduce systematics.
No problem if the theoretical errors are properly introduced.
They can be used to test assumptions.

Nonrelativistic Sum rules (b - \bar{b} , c - \bar{c})

$$J^\mu = \bar{Q} \gamma^\mu Q = B_1 \psi^\dagger \sigma \chi + \dots,$$

$$B_1 = 1 + a_1 \alpha_s + a_2 \alpha_s^2 + a_3 \alpha_s^3 + \dots$$

B_1 at NNLO: Hoang(QED); Beneke, Signer, Smirnov; Czarnecki, Melnikov

B_1 at NNNLO: Marquard, Piclum, Seidel, Steinhauser

B_1, B_0 at NLL: Pineda; Hoang, Stewart

B_1/B_0 at NNLL: Penin, Pineda, Smirnov, Steinhauser

B_1, B_0 at NNLL (partial): Pineda, Signer; Hoang, Stahlhofen

$$(q_\mu q_\nu - g_{\mu\nu}) \Pi(q^2) = i \int d^4x e^{iqx} \langle \text{vac} | J_\mu(x) J_\nu(0) | \text{vac} \rangle$$

$$\Pi(q^2) \sim B_1^2 \langle \mathbf{r} = \mathbf{0} | \frac{1}{E - H} | \mathbf{r} = \mathbf{0} \rangle$$

$$G(0, 0, E) = \sum_{m=0}^{\infty} \frac{|\phi_{0m}(0)|^2}{E_{0m} - E + i\epsilon - i\Gamma_t} + \frac{1}{\pi} \int_0^{\infty} dE' \frac{|\phi_{0E'}(0)|^2}{E_{0E'} - E + i\epsilon - i\Gamma_t}$$

Relation of the vacuum polarization with $\sigma_{t\bar{t}}$, non-relativistic sum rules and $\Gamma(V_Q(nS) \rightarrow e^+e^-)$

$$\Gamma(V \rightarrow e^+e^-) \sim \frac{1}{m^2} B_1^2 |\phi(\mathbf{0})|^2$$

$$\sigma_{t-\bar{t}} \sim B_1(\nu)^2 \text{Im} G(0, 0, \sqrt{s}) + \dots$$

$$M_n \equiv \frac{12\pi^2 e_b^2}{n!} \left(\frac{d}{dq^2} \right)^n \Pi(q^2) \Big|_{q^2=0} = \int_0^\infty \frac{ds}{s^{n+1}} R_{b\bar{b}}(s),$$

$$M_n = 48\pi e_b^2 N_c \int_{-\infty}^\infty \frac{dE}{(E + 2m_b)^{2n+3}} \left(B_1^2 - B_1 d_1 \frac{E}{3m_b} \right) \text{Im } G(0, 0, E)$$

Non-relativistic Sum rules: bottomonium

Pineda-Signer

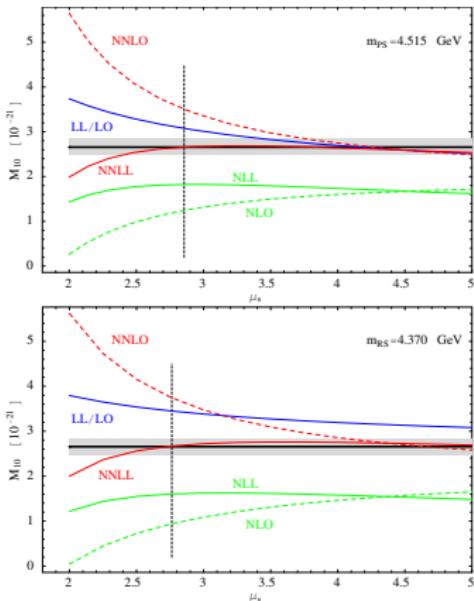


Figure: The moment M_{10} as a function of μ_s at LO/LL, NLO, NLL, NNLO and NNLL for m_{bPS} (2 GeV) = 4.515 GeV in the PS scheme (upper figure), and for m_{bRS} (2 GeV) = 4.370 GeV in the RS scheme (lower figure).

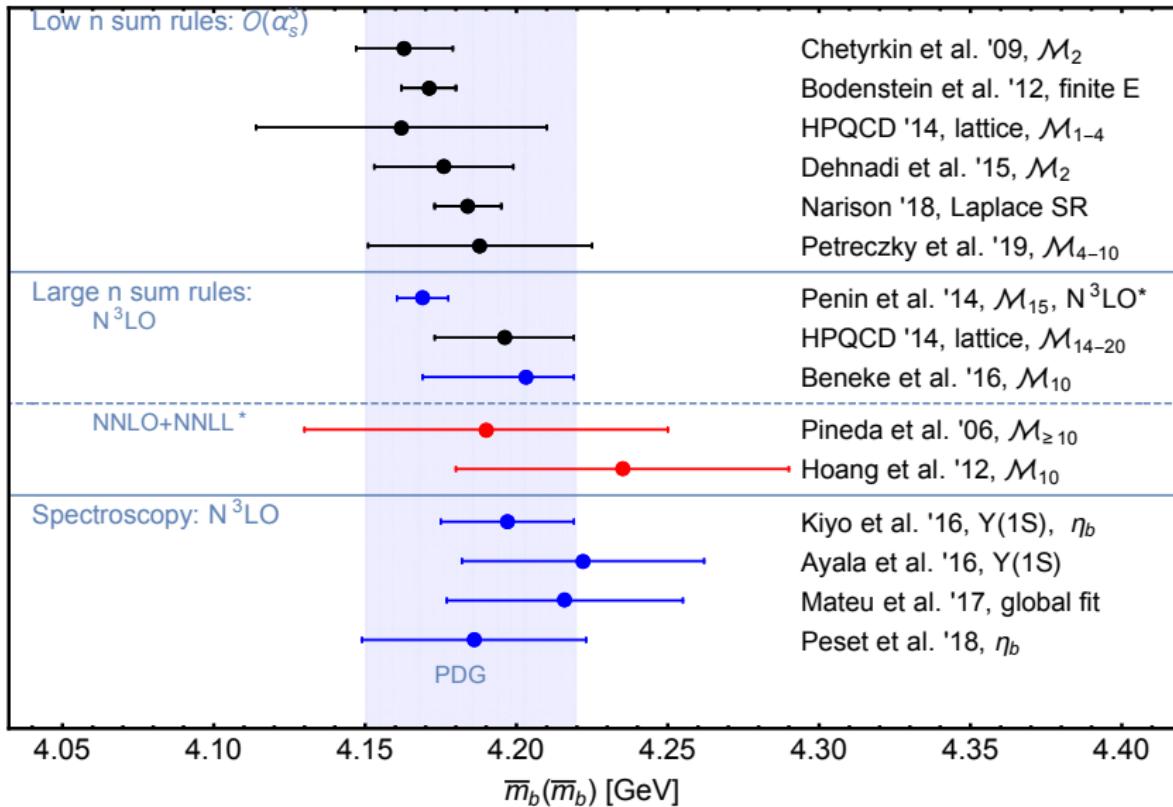
$$m_{b,\text{PS}}(2\text{GeV}) = 4.52 \pm 0.06 \text{ GeV},$$

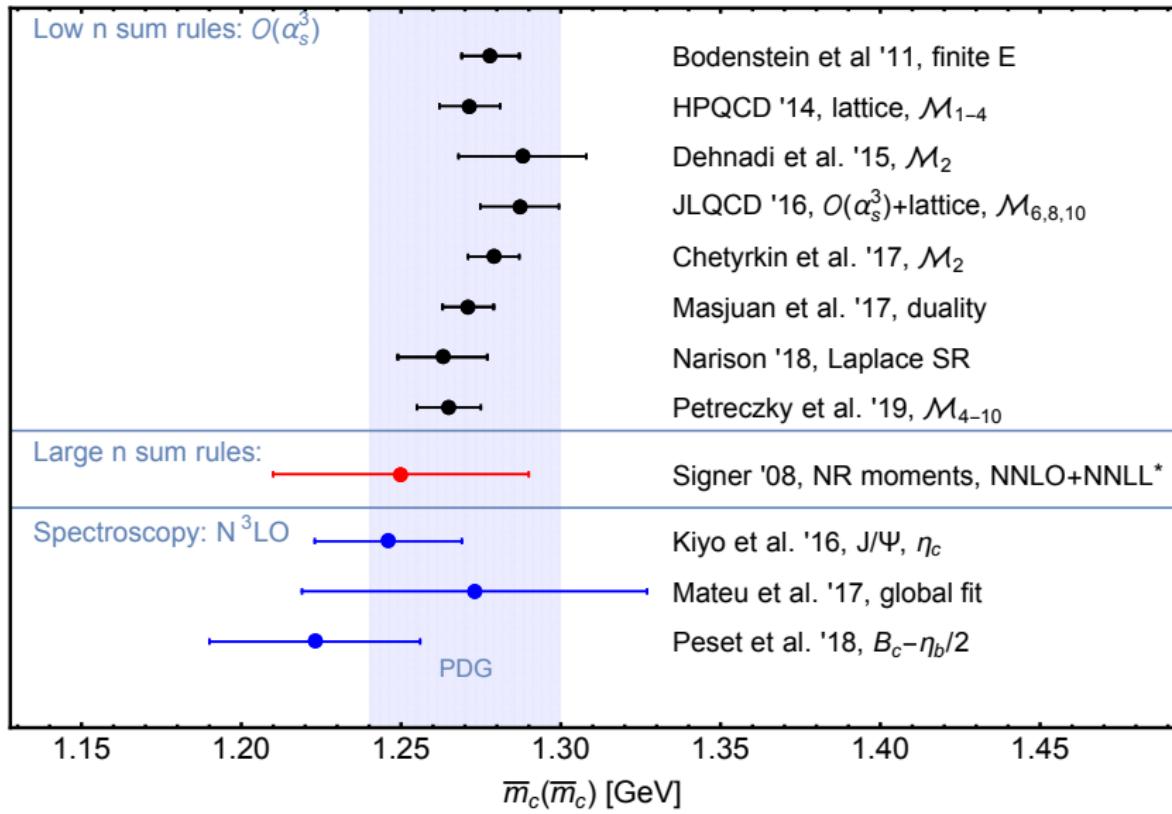
$$m_{b,\text{RS}}(2\text{GeV}) = 4.37 \pm 0.07 \text{ GeV}.$$

$$\overline{m}_b(\overline{m}_b) = 4.19 \pm 0.06 \text{ GeV}.$$

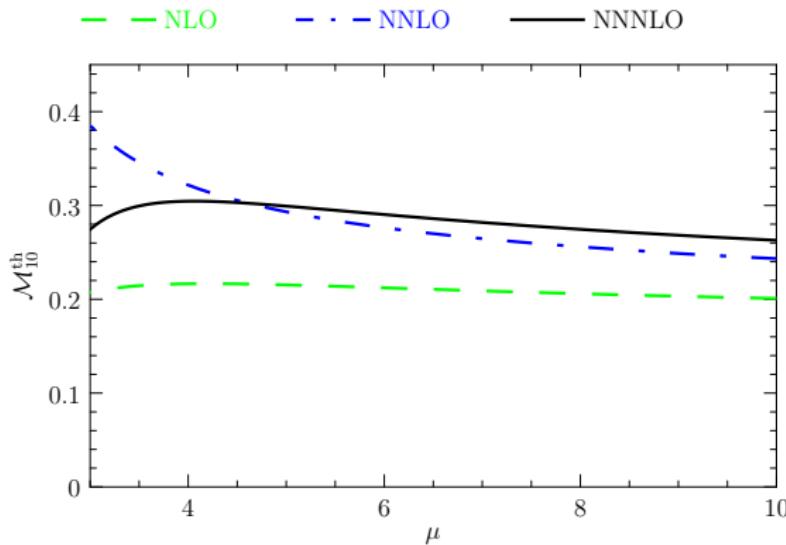
The perturbative series is **sign-alternating**. This is the opposite than for electromagnetic decays. The convergence of the perturbative series in sum rules is also better than for electromagnetic decays.

Low n sum rules very sensitive to the high energy behavior of R_{bb} . $n = 6$ Error 50 MeV, and bigger for smaller values of n .

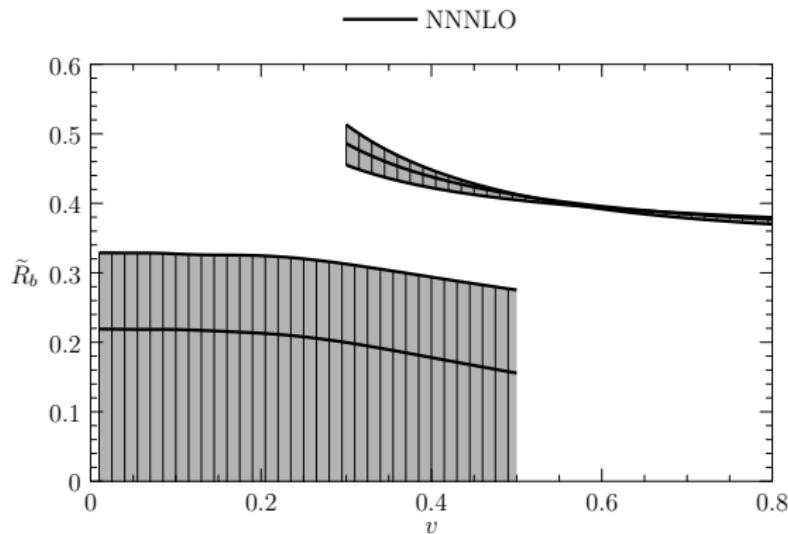


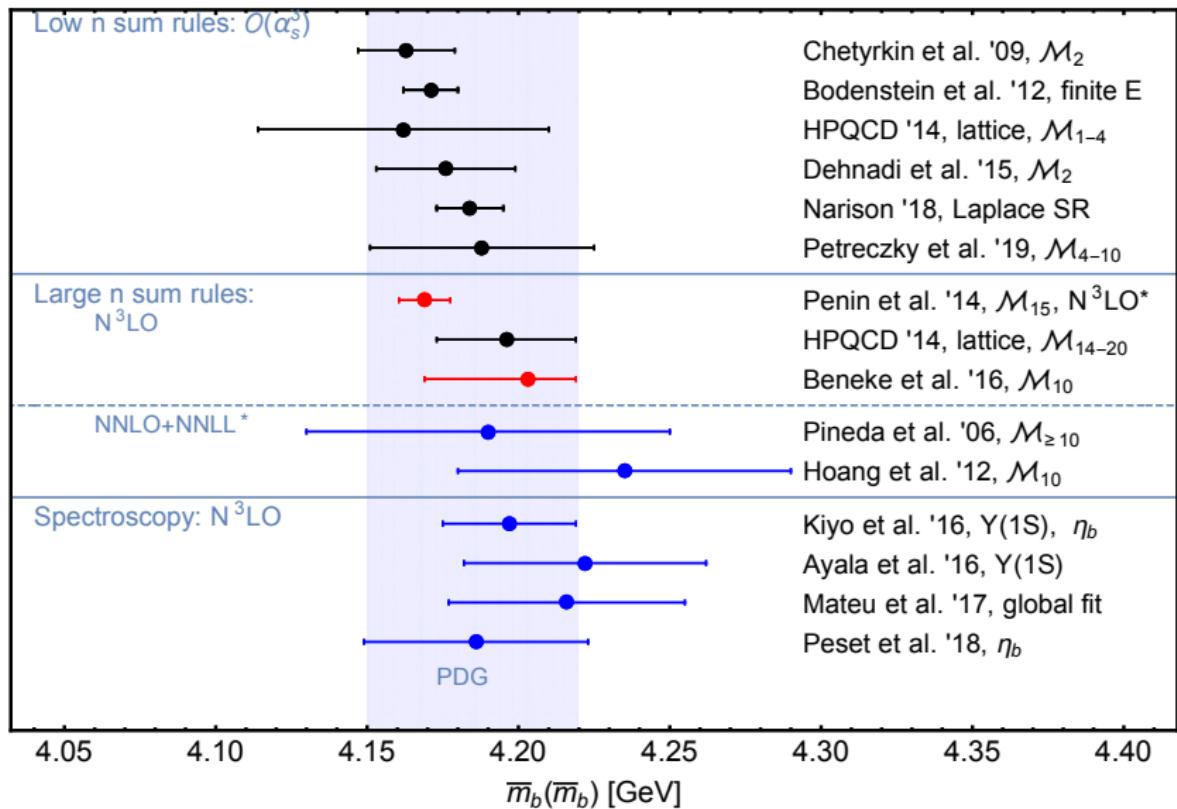


Martin Beneke, Andreas Maier, Jan Piclum, Thomas Rauh (NNNLO)



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CONCLUSIONS

Heavy quarkonium (near threshold) observables → ideal place to determine the heavy quark mass.

They show highest sensitivity to the heavy quark mass.

- ▶ Clean theoretical observables
- ▶ Different systematics than alternative determinations
- ▶ Similar numbers than alternative determinations
- ▶ Spectrum: Scale dependence increases from $NNLO \rightarrow N^3LO$

Possible issues:

mv^2 versus Λ_{QCD}

New effects show up at N^3LO

Perturbative ultrasoft effects small (with fixed order computations)

Renormalization group improvement?

No clear signal of the existence of $\mathcal{O}(\Lambda_{\text{QCD}}^3 r^2)$ effects.

Lattice simulations of the static potential also seem to support this.

A dedicated lattice simulation of δE_{us} needed.

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Nonperturbative effects are parametrically equal in spectroscopy and nonrelativistic sum rules. Overall coefficient significantly smaller in sum rules.

Non-relativistic sum rules

- ▶ 1) for which n is non-relativistic
- 2) for which n the perturbative expansion breaks down (smoking gun signal?)
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 - 1) Dependence with the continuum: Large experimental error
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