

Lecture III : Scalar Perturbations

Recap:

- Slowly rolling homogeneous scalar field $\rightarrow w \simeq -1$
 \rightarrow solves flatness & horizon problems
- SVT - decomposition:

$\delta g_{\mu\nu} \rightarrow 2$ scalars, 2 vectors, 1 tensor

$\delta\Phi \rightarrow 1$ scalar

coordinate transformation $\rightarrow 2$ scalars, 1 vector

- gauge dependence

$\Gamma \hat{=}$ choice of equal time hypersurface:

$$\delta X(t, \vec{x}) = \underbrace{X(t, \vec{x})}_{\substack{\text{locally} \\ \text{unambig.} \\ \text{defined}}} - \underbrace{\bar{X}(t)}_{\substack{\text{depends on choice} \\ \text{of equal time} \\ \text{hypersurface}}}$$

\downarrow
gauge dependent

\Rightarrow a) define gauge invariant scalar, e.g.

$$\mathcal{R} = \psi + \frac{H}{\dot{\phi}} \delta\phi \quad (\text{during slow roll inflation})$$

"comoving curvature perturbation"

= spatial curvature of constant- ϕ hypersurface

or b) choose a gauge

\rightarrow intermediate steps gauge dependent

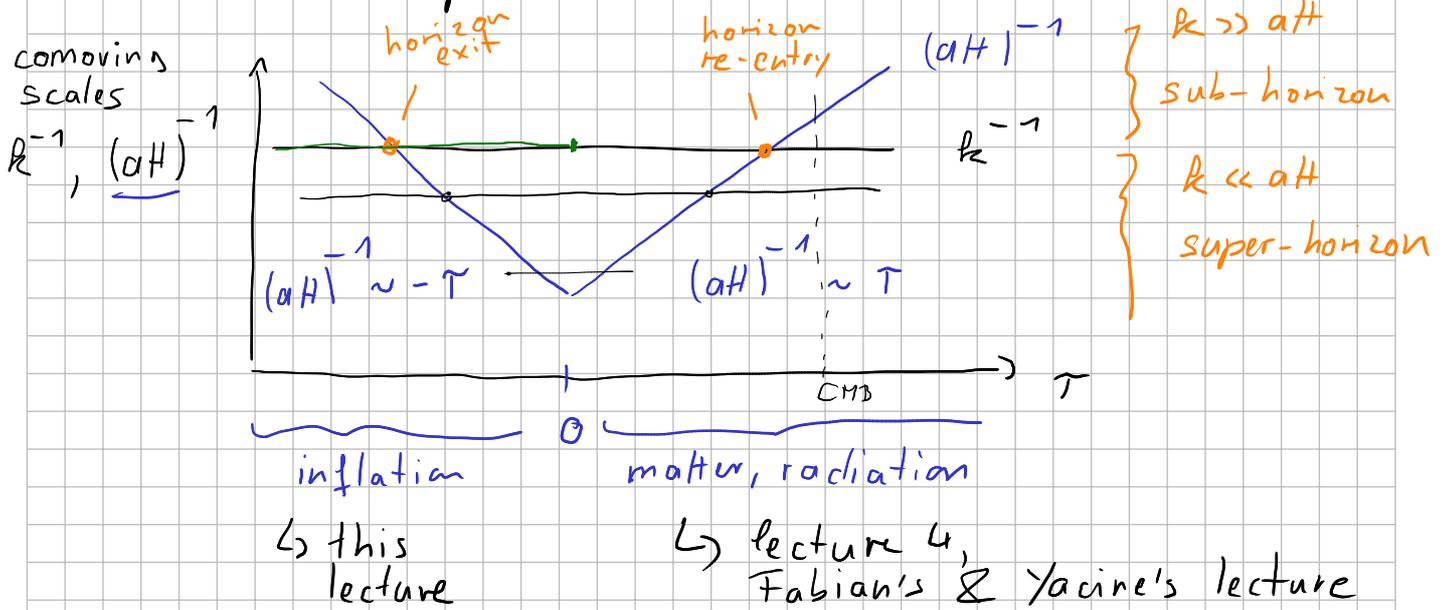
but physical observables are not

e.g. Newtonian gauge : $\mathcal{B} = \mathcal{F} = 0$

note: tensor sector has no gauge freedom.

0) scales and horizons

Consider perturbation with wavenumber k



Computational strategy:

- 1) expand action to 2nd order in perturbations
 \hookrightarrow derive eom
- 2) quantize & set initial conditions
- 3) compute power spectrum
(= statistical properties of perturbations)

1) EOM

gauge choice: $\delta\phi = 0$, $g_{ij} = a^2 [(1 - 2\mathcal{R})\delta_{ij} + h_{ij}]$

$$(S_\phi + S_{EH})^{(2)} = \dots = \frac{1}{2} \int d^4x a^3 \frac{\dot{\phi}^2}{H^2} [\dot{\mathcal{R}}^2 - a^{-2} (\partial_i \mathcal{R})^2]$$

$$z \equiv \frac{a\dot{\phi}}{H} \quad \underline{\quad} = \frac{1}{2} \int d\tau d^3x \left(\underbrace{a^2 z^2 \dot{\mathcal{R}}^2}_{(v' - z'R)^2} - \underbrace{z^2 (\partial_i \mathcal{R})^2}_{(\partial_i v)^2} \right)$$

$v \equiv z\mathcal{R}$
Mukhanov variable

$$= \frac{1}{2} \int d\tau d^3x \left[v'^2 - (\partial_i v)^2 - 2v'v \frac{z'}{z} - \left(v \frac{z'}{z} \right)^2 \right]$$

$$= \frac{1}{2} \int d\tau d^3x \left[v'^2 - (\partial_i v)^2 - \frac{d}{d\tau} \left[\frac{z'}{z} v^2 \right] + v^2 \frac{z''}{z} \right]$$

$$\text{p.I.} = \frac{1}{2} \int d\tau d^3x \left[v'^2 - (\partial_i v)^2 + \underbrace{\frac{z''}{z} v^2}_{\text{"mass term"}} \right]$$

Fourier expansion: $v(\tau, \vec{x}) = \int \frac{d^3k}{(2\pi)^3} V_{\vec{k}}(\tau) e^{i\vec{k}\cdot\vec{x}}$

$$\boxed{V_k'' + \left(k^2 - \frac{z''}{z} \right) V_k = 0} \quad (M) \quad \text{Mukhanov equation}$$

$\approx 2a^2 H^2$
slow-roll

\Rightarrow sub-horizon ($k \gg aH$) : free oscillator

super-horizon ($k \ll aH$) : z''/z important

2) Quantization and initial conditions

Quantization:

$$v \rightarrow \hat{v} = \int \frac{d^3k}{(2\pi)^3} \left[V_k(\tau) \hat{a}_{\vec{k}} e^{i\vec{k}\vec{x}} + V_k^*(\tau) \hat{a}_{\vec{k}}^\dagger e^{-i\vec{k}\vec{x}} \right]$$

$$\text{with } [\hat{a}_{\vec{k}}, \hat{a}_{\vec{k}'}^\dagger] = (2\pi)^3 \delta(\vec{k} - \vec{k}')$$

$$\text{HW} \hookrightarrow \frac{i}{\hbar} (V_k^* V_{k'} - V_{k'}^* V_k) = 1 \quad (*)$$

fixes one out of two initial conditions

Initial conditions:

$$\text{in the far past } k \gg aH \xrightarrow{(M)} V_k'' + k^2 V_k = 0$$

\hookrightarrow harmonic oscillator

\hookrightarrow unique state which minimizes energy

$$\lim_{\tau \rightarrow -\infty} V_k = \frac{1}{\sqrt{2k}} e^{-ik\tau} \quad \text{Bunch Davies vacuum}$$

only positive frequency mode due to (*)

\rightarrow initial conditions fully fixed

Solve (M) with these i.c. in de-Sitter ($H = \text{const}$):

$$\tau = -\frac{1}{aH} \rightarrow V_k'' + (k^2 - \frac{2}{\tau^2}) V_k = 0$$

$$\rightarrow V_k = \frac{e^{-ik\tau}}{\sqrt{2k}} \left(1 - \frac{i}{k\tau} \right) \quad (**)$$

τ super-horizon, $\frac{k}{aH} = -k\tau \ll 1$

$$V_k'' - \frac{2}{\tau^2} V_k = 0 \rightarrow V_k = A\tau^2 + B/\tau$$

$$R \propto \frac{V_k}{a} = \underbrace{-A H \tau^3}_{\tau \rightarrow 0 \text{ decays}} - \underbrace{BH}_{\text{constant}}$$

$\tau \rightarrow 0$
decays

constant

\rightarrow no oscillations, "freeze-out"

3) Power spectrum $\langle R_{\vec{k}} R_{\vec{k}'} \rangle$

- statistical homogeneity & isotropy (SHI)

$$\langle R(t, \vec{x}) R(t, \vec{x}') \rangle \stackrel{\text{SHI}}{=} \xi(t, |\vec{x} - \vec{x}'|)$$

\uparrow isotropy \uparrow homogeneity

in Fourier space

$$\langle R_{\vec{k}}(t), R_{\vec{k}'}(t) \rangle \stackrel{\text{SHI}}{=} (2\pi)^3 \delta(\vec{k} - \vec{k}') P_R(k)$$

\uparrow homogeneity \uparrow isotropy

we will also need

$$\int \frac{d^3k}{(2\pi)^3} P_R(k) = \int d \ln k \underbrace{\frac{k^3}{2\pi^2} P_R(k)}_{\equiv \Delta_R^2}$$

- power spectrum

$$R = \frac{v}{z} = \frac{vH}{a\dot{\phi}} \rightarrow \langle R_{\vec{k}} R_{\vec{k}'} \rangle = (2\pi)^3 \delta(\vec{k} - \vec{k}') \left(\frac{H}{\dot{\phi}}\right)^2 \frac{|V_k(\tau)|^2}{a^2}$$

$$\left| \frac{V_k(\tau)}{a} \right|^2 \stackrel{(\forall \nu)}{=} \frac{1}{a^2} \frac{1}{2k} \left(1 + \frac{1}{k^2 \tau^2}\right)$$

$$= \frac{H^2}{2k^3} \left(1 + \underbrace{k^2 \tau^2}\right)$$

$$= \text{const} \quad \hookrightarrow 0 \text{ on super-horizon scales}$$

$\Rightarrow \langle R^2 \rangle$ at horizon crossing, $k = a_* H_*$,

with $\phi = \phi_*$, remains constant after horizon crossing

$$\Rightarrow \langle R_{\vec{k}} R_{\vec{k}'} \rangle = (2\pi)^3 \delta(\vec{k} + \vec{k}') \left(\frac{H_*}{\dot{\phi}_*}\right)^2 \frac{H_*^2}{2k^2} \quad \forall \tau \text{ with } aH \gg k$$

$$\Rightarrow \Delta_R^2 = \frac{H_*^2}{(2\pi^2)} \frac{H_*^2}{\dot{\phi}_*^2}$$

$$H, \dot{\phi} \approx \text{const}$$

\Rightarrow scale invariant