

Beyond Λ CDM

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Focus on Gravity.

References:

P. G. Ferreira (1902.10503)

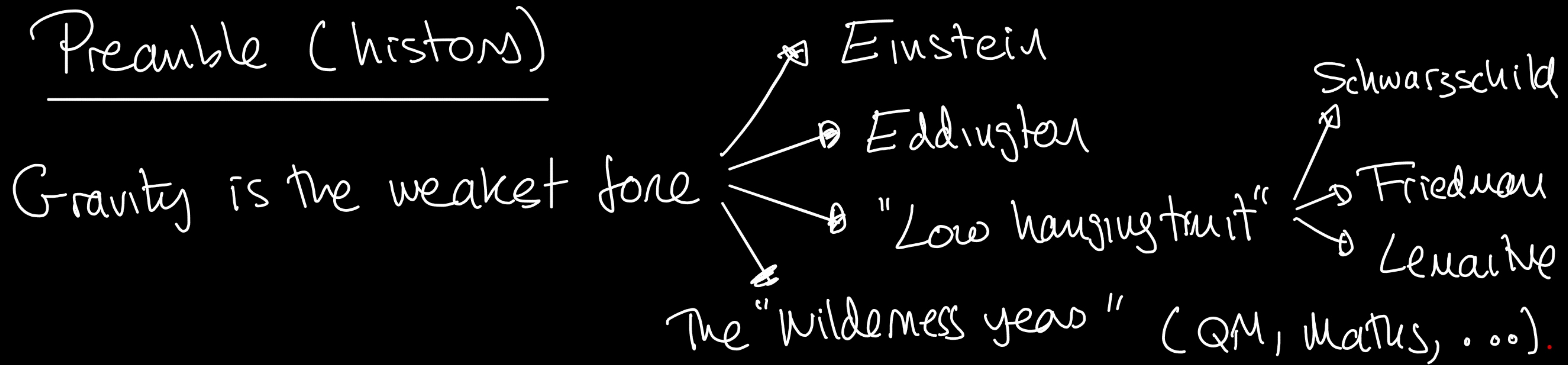
T. Baker et al (1908.03430)

M. Ishak et al (1905.09687)

C. Burrage & J. Sakstein (1709.09071)

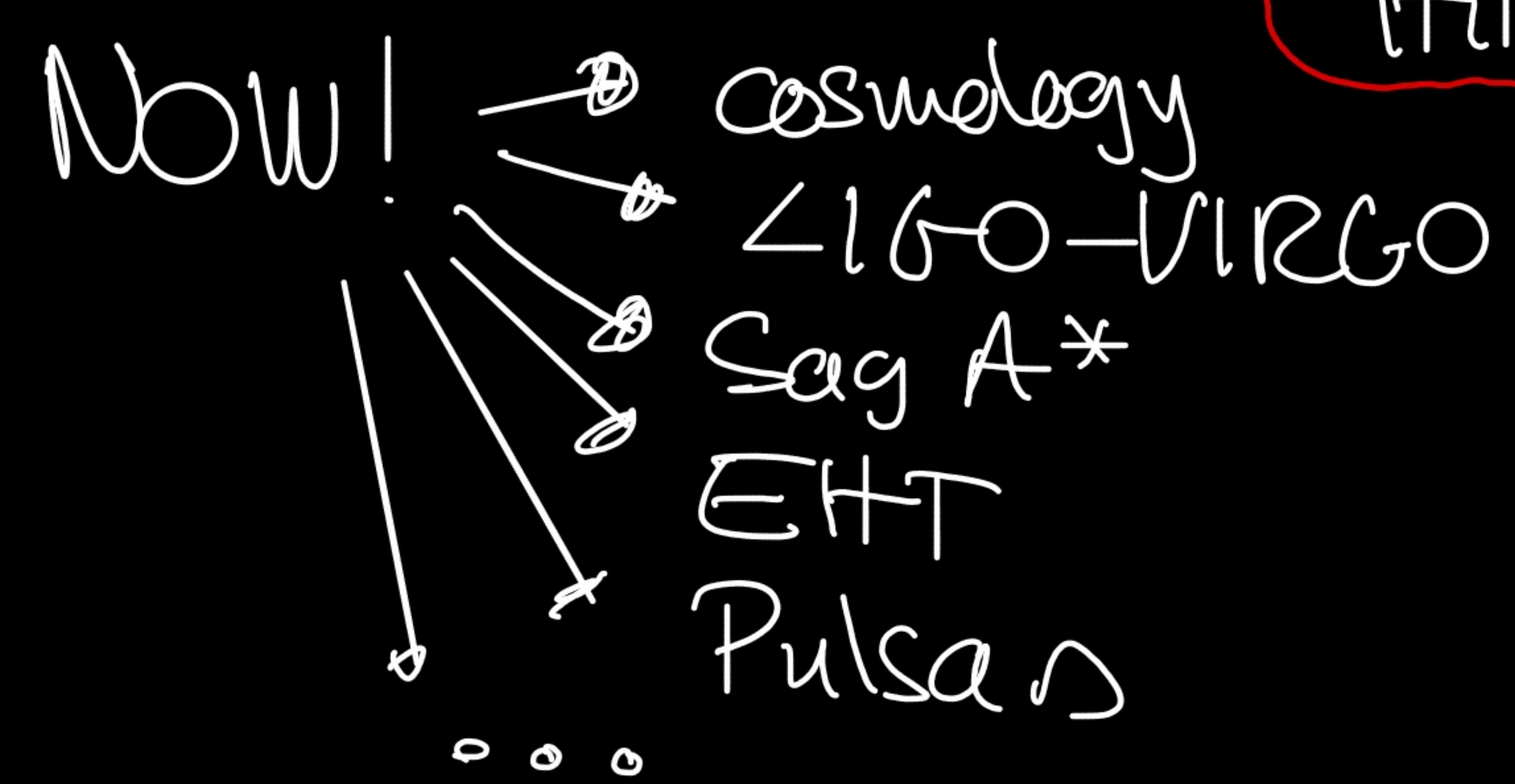
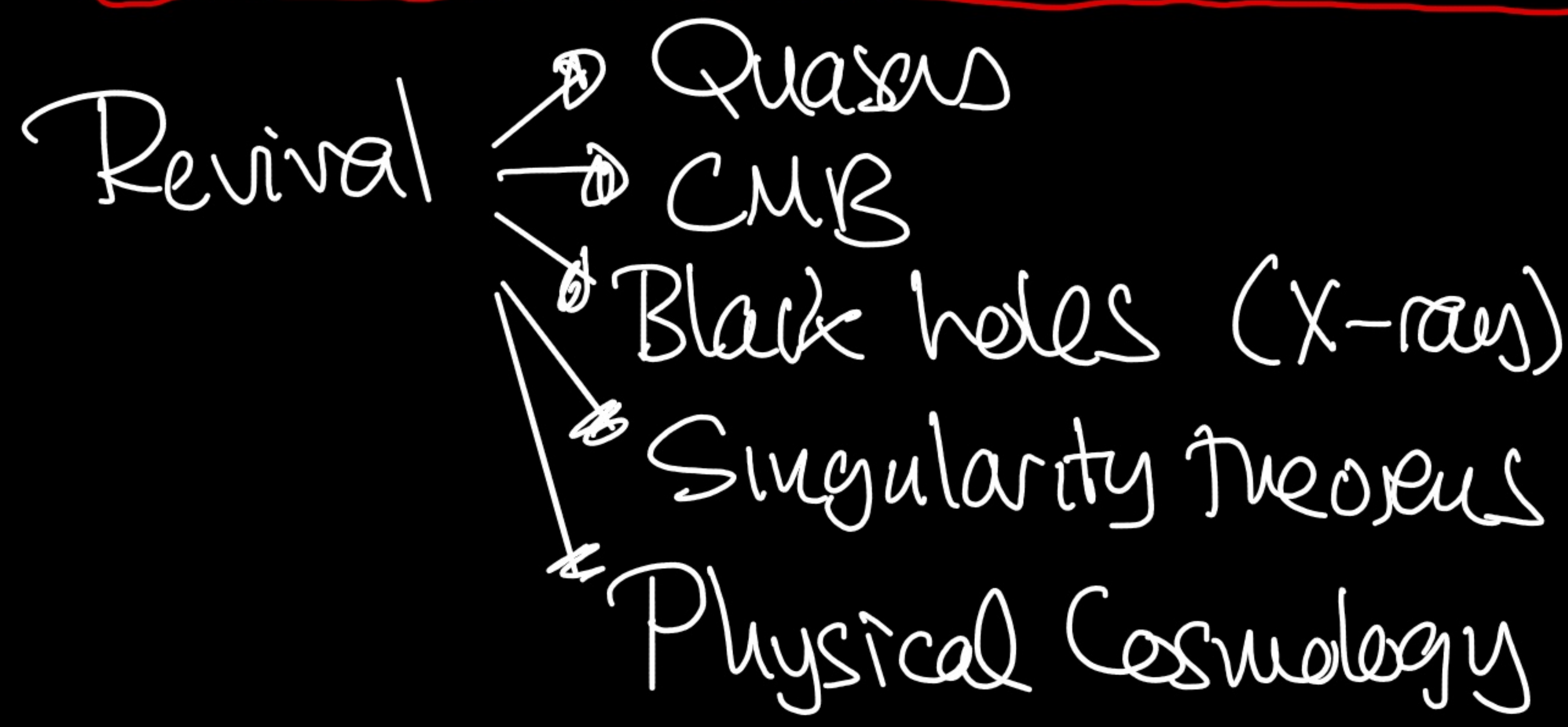
+ ...

Preamble (history)



Feynman: "There exists one serious difficulty and that is the lack of experiments... The best viewpoint is to pretend that there are experiments and calculate. In this field we are not pushed by experiments but pulled by imagination"

1957 Chapel Hill



What is Gravity

Newton $\vec{x}(t)$

$$\frac{d^2 \vec{x}}{dt^2} = \vec{g} = -\vec{\nabla} \Phi$$

↑
grav. pot.

$$\nabla^2 \Phi = 4\pi G \rho$$

Einstein

$$\frac{d^2 x^\beta}{d\lambda^2} + \Gamma^\beta_{\mu\nu} \frac{dx^\mu}{d\lambda} \frac{dx^\nu}{d\lambda} = 0$$

(geodesic equ)

$$ds^2 = g_{\alpha\beta} dx^\alpha dx^\beta$$

(metric)

$$G^{\alpha\beta} = R^{\alpha\beta} - \frac{1}{2} R g^{\alpha\beta}$$
$$= 8\pi G T^{\alpha\beta} - \Lambda g^{\alpha\beta}$$

$g_{00} = -(1 + 2\Phi)$

Note that: ① $\nabla_\alpha G^{\alpha\beta} = 0$ and $\nabla_\alpha T^{\alpha\beta} = 0$

② $\Gamma \sim g^{-1} \partial g$ and $R \sim \partial \Gamma + \Gamma \Gamma$
(ie 2nd order)

What is Gravity

Einstein-Hilbert Action

$$S_{EH} = \frac{M_{Pl}^2}{2} \int d^4x \sqrt{-g} (R - 2\Lambda) + \int d^4x \sqrt{-g} \mathcal{L}(g, \text{matter})$$

where $M_{Pl}^2 = \frac{1}{8\pi G}$

Vary $\delta g^{\alpha\beta}$

$$\delta S = \frac{M_{Pl}^2}{2} \int d^4x \sqrt{-g} \left[(G_{\alpha\beta} + \Lambda g_{\alpha\beta}) \delta g^{\alpha\beta} + \underbrace{(g_{\alpha\beta} \square - \nabla_\alpha \nabla_\beta)}_{\text{surface term.}} \right] \delta g^{\alpha\beta} - \int d^4x \sqrt{-g} T_{\alpha\beta} \delta g^{\alpha\beta}$$

E.o.m. don't depend on boundary but need to impose $(\delta g + \delta \dot{g})$

Note: can add boundary term

$$+ M_{Pl}^2 \int_{\partial M} \sqrt{-\gamma} K$$

metric or boundary.

Extrinsic curvature

What is Gravity

Tests:

• gravitational redshift $\frac{\nu_B}{\nu_A} = \frac{dZ_A}{dZ_B} = \sqrt{\frac{g_{00}(A)}{g_{00}(B)}} = 1 - (\Phi_B - \Phi_A)$

(Pound Rebka '60 - new $\sim 0.01\%$)

• orbits (Mercury)

$$\Delta\theta \sim 43'' \text{ per century}$$

• Lenses

$$\Delta\theta_{GR} = 2\Delta\theta_N = \frac{4GM}{R}$$

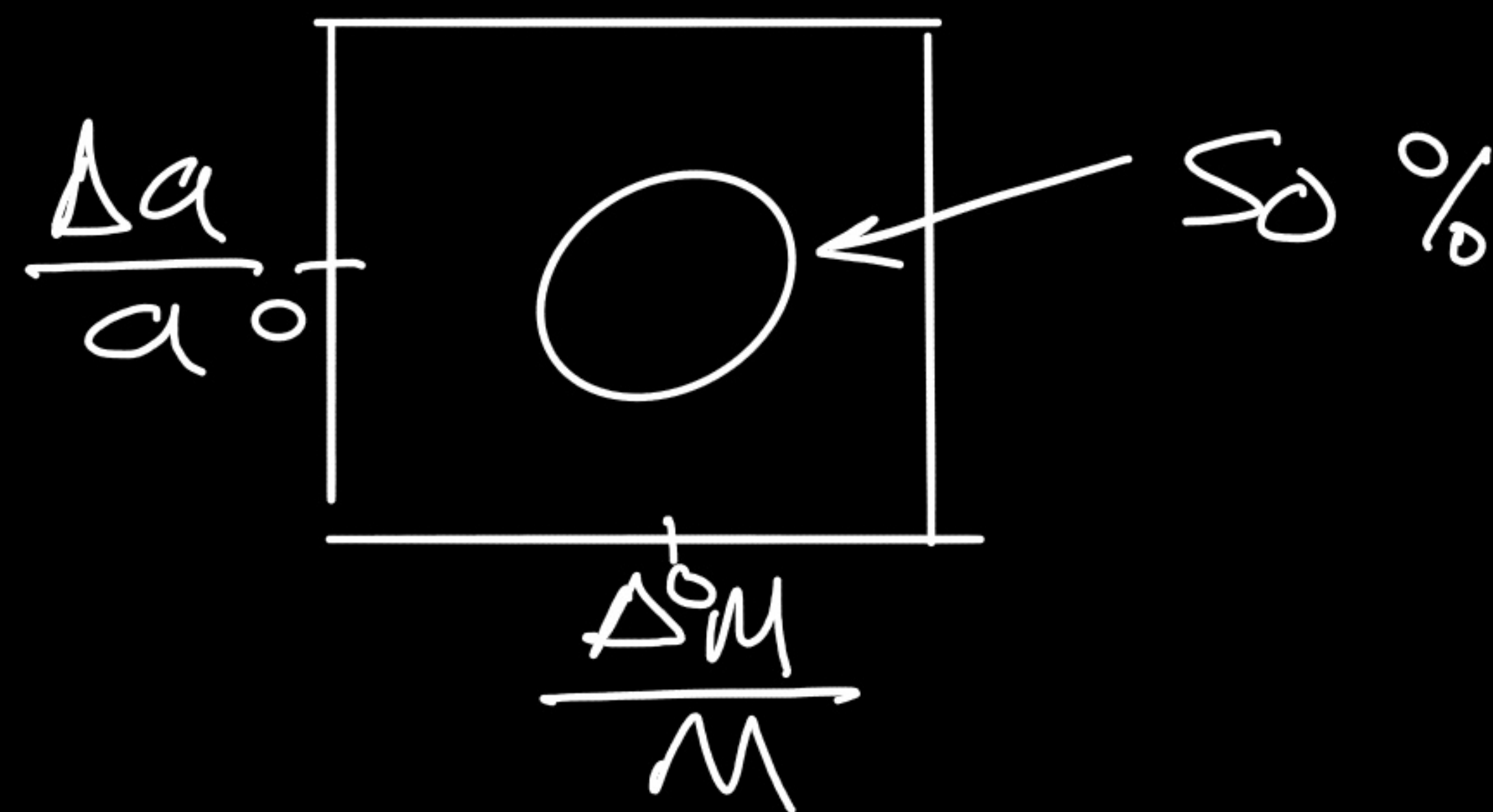
• Strong lenses
• weak lenses

• orbits (Shapiro)

$$c\Delta t \approx \frac{2GM}{c^2} \ln \left[\sqrt{\frac{r^2}{R^2} - 1} + \frac{r}{R} \right]$$

• binary pulsars $\frac{\dot{P}}{P}$

• Gravitational Waves



Why GR?

Lovejoy's Theorem (1971) — EH action is the only local, 4D, 2nd order equations of motion for metric.

Note: none of these are essential

(Lovejoy, J. Math. Phys. 12, 874, 1971).

Feynman / Weinberg / Deser (1960s) — EH action is the only 'consistent' theory of a spin 2 field.

Feynman, Lectures on Gravitation

Weinberg, P. Rev., 138, 878, 1965

Deser, GRG, 1, 9, 1970

But see: Padmanabhan gr-qc/0409089

Why GR?

Sketch of field theory argument

$$h_{\mu\nu} \longrightarrow h_{\mu\nu} + \partial_\mu \xi_\nu + \partial_\nu \xi_\mu$$

↑
spin 2
field

(gauge symmetry)

Where could this symmetry come from?

$$\tilde{g}_{\alpha\beta} \longrightarrow \frac{\partial x^\mu}{\partial \tilde{x}^\alpha} \frac{\partial x^\nu}{\partial \tilde{x}^\beta} g_{\mu\nu}$$

Expand:

$$g_{\alpha\beta} = \eta_{\alpha\beta} + h_{\alpha\beta}$$

↑
Minkowski

$$\tilde{x}^\mu = x^\mu + \xi^\mu(\tilde{x}^\alpha)$$

So gauge symmetry is linearized diffeo. invariance

Why GR?

Massless Fierz-Pauli action

$$S_{\text{FP}} = \frac{M_{\text{Pl}}^2}{2} \int d^4x \left[-\frac{1}{2} \partial_\mu h_{\nu\sigma} \partial^\mu h^{\nu\sigma} + \partial_\mu h_{\nu\sigma} \partial^\nu h^{\mu\sigma} - \partial_\mu h \partial^\mu h + \frac{1}{2} \partial_\mu h \partial^\mu h \right]$$

Vary with $\delta h^{\alpha\beta}$ we get linearized GR.

Pick a gauge ($\partial^\mu h_{\mu\nu} = \partial^\nu h_{\mu\mu} = 0$) $\square h_{\mu\nu} = 0$

But no source...

Add $S_\phi = \frac{1}{2} \int d^4x \partial_\mu \phi \partial^\mu \phi$ to get $\begin{cases} \square h_{\mu\nu} = 0 \\ \square \phi = 0 \end{cases}$ uncoupled...

We want $\square h_{\mu\nu} \sim T_{\mu\nu}^\phi$

where
$$T_{\mu\nu}^\phi = -\partial_\mu \phi \partial_\nu \phi + \frac{1}{2} \eta_{\mu\nu} (\partial\phi)^2$$

(Energy momentum tensor).

Why GR?

Add a coupling $S_1 = S_{FP} + S_\phi + \frac{\kappa}{2} \int d^4x h^{\mu\nu} T_{\mu\nu}$

Now $\square h_{\mu\nu} = \kappa T_{\mu\nu}$

$$\square \phi = \kappa \partial_\mu \left[h^{\mu\nu} \partial_\nu \phi - \frac{1}{2} h \partial_\mu \phi \right]$$

Now

$$0 \Rightarrow \partial^\nu \square h_{\mu\nu} = \kappa \partial^\mu T_{\mu\nu}(\phi) = \kappa \square \phi \partial_\nu \phi \neq 0$$

← inconsistent →

Add graviton self energies

$$S_2 = S_1 + \frac{\kappa}{2} \int d^4x h^{\mu\nu} T_{\mu\nu}^h$$

quadratic in ∂h and h .

Doesn't work: $\square h_{\mu\nu} \neq \kappa T_{\mu\nu}^\phi + \kappa T_{\mu\nu}^h$

Note: keep on adding $O(3), O(4), \dots$ only works when

$O(\infty) \iff$ EH action!

Why GR?

What about higher derivatives?

Sakharov 1967: $S[\phi, g_{\alpha\beta}] = \int d^4x \sqrt{-g} g^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi$

(note \rightarrow no kinetic term for $g_{\alpha\beta}$...)

Integrate out ϕ

$$\int \mathcal{D}\phi e^{\frac{iS[\phi, g]}{\hbar}} \rightarrow e^{\frac{iS_{\text{eff}}}{\hbar}}$$

Now $S_{\text{eff}} \sim \sqrt{-g} \left[\Lambda + \frac{M_{\text{pl}}^2}{2} R + a_1 R_{\mu\nu} R^{\mu\nu} + a_2 R^2 + \dots \right]$

Naively we expect corrections.

Highly suppressed

E.g: $\frac{M_{\text{pl}}^2}{2} R + a_2 R^2 = \frac{M_{\text{pl}}^2}{2} R \left[1 + a_2 \frac{R}{M_{\text{pl}}^2} \right]$

We need $R \sim M_{\text{pl}}^2$
but, typically $R \ll M_{\text{pl}}^2$.

Note: Ostrogradsky too...

(see Woodard [astro-ph/0601672](https://arxiv.org/abs/astro-ph/0601672))

Why GR?

- ⊙ GR leading theory for gravity
- ⊙ Beautifully confirmed by experiment
- ⊙ Theoretical prejudice (Lovejoy's Theorem)
- ⊙ Theoretical prejudice (Field Theory of sp_{u-2} particle)
- ⊙ Higher derivatives: suppressed or unstable

But

- ⊙ Lovejoy gives us a way of exploring other options...