

# Exercises IV Joint ICTP School on Cosmology

## lecture 04

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**1 - Inflation models (II) (4pt).** Consider the inflation model specified by  $V(\phi) = \frac{1}{2}m^2\phi^2$ .

- (a) Compute the amplitude of the scalar perturbations  $\Delta_{\mathcal{R}}^2$  at  $N_{cmb} = 60$  e-folds before the end of inflation (2 pt).
- (b) Measurements of the CMB anisotropies yield  $\Delta_{\mathcal{R}}^2 \sim 10^{-9}$ . Use this to determine  $m$  (1pt).
- (c) Compute the scalar spectran tilt  $n_s$  and the tensor-to-scalar ratio  $r$ . Compare your result to Fig. 8 of the Planck publication arxiv:1807.06211. (1pt)

**2 - Continuity equation (10 pt).** Consider the  $\nu = 0$  component of energy momentum conservation,

$$0 = \nabla_{\mu} T_{\nu}^{\mu} = \partial_{\mu} T_{\nu}^{\mu} + \Gamma_{\mu\alpha}^{\mu} T_{\nu}^{\alpha} - \Gamma_{\mu\nu}^{\alpha} T_{\alpha}^{\mu}, \quad (1)$$

up to first order in perturbations of the metric and the stress energy tensor. Work in the approximation of a flat universe ( $k = 0$ ) and consider the Universe filled with a single fluid with  $p = \omega\rho$ . In Newtonian gauge, the Christoffel symbols read

$$\begin{aligned} \Gamma_{00}^0 &= \mathcal{H} + \Psi', & \Gamma_{i0}^0 &= \partial_i \Psi, & \Gamma_{00}^i &= \delta^{ij} \partial_j \Psi, \\ \Gamma_{ij}^0 &= \mathcal{H} \delta_{ij} - [\Phi' + 2\mathcal{H}(\Psi + \Phi)] \delta_{ij}, & \Gamma_{j0}^i &= (\mathcal{H} - \Phi') \delta_j^i, & \Gamma_{jk}^i &= -2\delta_{(j}^i \partial_{k)} \Phi + \delta_{jk} \delta^{il} \partial_l \Phi, \end{aligned}$$

with the comoving Hubble parameter  $\mathcal{H} = aH$  and prime denoting the derivative with respect to conformal time.  $\mu, \nu = 0..3$  are Lorentz indices,  $i, j = 1..3$  are spatial indices.

- (a) Write Eq. (1) in terms of  $T_0^0$ ,  $T_0^i$  and  $T_j^i$  and drop all terms with are second order (or higher) in the perturbations (2 pt)
- (b) Substituting the relevant Christoffel symbols, show that (2pt)

$$\partial_0 T_0^0 + \partial_i T_0^i + (3\mathcal{H} - 3\Phi') T_0^0 - 3(\mathcal{H} - \Phi') T_i^i = 0 \quad (2)$$

- (c) Expand  $T_0^0$ ,  $T_0^i$  and  $T_j^i$  to first order (see lecture). Inserting this into Eq. (2) show that (5 pt)

$$\bar{\rho}' = -3\mathcal{H}(\bar{\rho} + \bar{p}) \quad (3)$$

$$\delta\rho' = -3\mathcal{H}(\delta\rho + \delta p) + (\bar{\rho} + \bar{p})(3\Phi' - \partial_i v^i) \quad (4)$$

- (d) Show that this implies (1pt)

$$\left(\frac{\delta\rho}{\bar{\rho}}\right)' = -3\mathcal{H}\left(\frac{\delta p}{\bar{\rho}} - \frac{\bar{p}}{\bar{\rho}^2}\delta\rho\right) + \left(1 + \frac{\bar{p}}{\bar{\rho}}\right)(3\Phi' - \partial_i v^i) \quad (5)$$