Summing over Geometries in String Theory

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Based on [2008.07533] and [2102.12355]

Motivation

- Gravity should emerge from string theory at large distances.
- The gravitational path integral includes a sum over all distinct topologies with given boundary conditions, even those that do not contain saddles.

e.g. JT-gravity:



[Saad, Shenker, Stanford '19; ...]

String theory

- We usually formulate perturbative string theory on a fixed background and study perturbative string corrections around a given background.
- So how does the semiclassical sum over geometries emerges from the string path integral?

AdS/CFT correspondence

- In the AdS/CFT correspondence, we can be more precise because we know the answer for the string partition function: the dual CFT partition function.
- ► There are exact (no ensemble average!) AdS/CFT dualities:

$$\begin{split} \mathsf{AdS}_5 \times \mathsf{S}^5 & \Longleftrightarrow \, \mathcal{N} = 4 \,\, \mathsf{SYM} \,\,, \\ \mathsf{AdS}_3 \times \mathsf{S}^3 \times \mathbb{T}^4 & \Longleftrightarrow \, \mathsf{Sym}^N(\mathbb{T}^4) \,\,. \end{split}$$

How do we reproduce the exact boundary partition function from string theory?

A naive prescription

Inspired by semiclassical gravity, one might think

$$Z = \sum_{\text{bulk saddles } \mathcal{M}} \exp\left(\sum_{g=0}^{\infty} g_s^{2g-2} \int \mathscr{D}[\text{fields}] \, e^{-S_{\mathcal{M}}[\text{fields}]}\right) \\ + \text{ D-instantons } + ??? ,$$

where "fields" stands for all the worldsheet fields.

This is wrong in general! The sum over saddles overcounts configurations, since they can be represented by highly excited string states (gas of gravitons) on other geometries.

Tensionless AdS₃/CFT₂

There is a proposal for an exact AdS/CFT dual pair (in the sense that both the string side and the CFT side are under very good computational control): [LE, Gaberdiel, Gopakumar '18,...]

[see Matthias' talk]

 $\begin{array}{l} \mbox{Strings on } {\rm AdS}_3\times {\rm S}^3\times {\mathbb T}^4 \mbox{ with one unit of NS-NS flux} \\ \Longleftrightarrow {\rm Sym}^N({\mathbb T}^4) \ . \end{array}$

▶ We can study this theory on different backgrounds that are asymptotically $AdS_3 \times S^3 \times \mathbb{T}^4$ and address this question directly.

Result

There is no sum over geometry!

We simply have (when interpreted correctly)

$$Z_{\mathsf{CFT}} = \exp\left(\sum_{g=0}^{\infty} g_{\mathsf{s}}^{2g-2} \int \mathscr{D}[\mathsf{fields}] \, \mathrm{e}^{-S_{\mathcal{M}}[\mathsf{fields}]}\right) \; .$$

for any saddle geometry \mathcal{M} .

- All other semiclassical background geometries arise as highly excited string states.
- ▶ Perturbative string theory with asymptotic boundary conditions AdS₃ × S³ × T⁴ with unit of NS-NS flux is background independent.

Geometry in the tensionless limit

- In the tensionless limit, strings can become very large.
- On AdS₃, string states wind around the asymptotic boundary of AdS₃.
- ► Even the graviton is such a non-local string state ⇒ No local notion of spacetime geometry?
- Background geometry is still well-defined in the form of the worldheet theory.
- So 'summing over worldsheet theories' will be our proxy for 'summing over geometries'

t = 0 slice of AdS₃



Thermal partition function

- Let's consider the example of a single torus boundary.
- ► In 3d-gravity, we sum over all modular images of thermal AdS₃ (SL(2, Z) family of Euclidean black holes) [Maloney, Witten '07]



The string partition function on thermal \mbox{AdS}_3

- In the following, we explain how to compute the string partition function on thermal AdS₃.
- Other geometries will be similar.
- One can show that the perturbative string partition function is one-loop exact in g_s. [LE '21]
- So the partition function receives a sphere contribution and a torus contribution.
- ▶ sphere $\propto \frac{1}{g_s^2} \propto \frac{1}{G_N} \propto \text{on-shell gravity action} = \frac{c\pi}{6} \operatorname{Im} \tau_{\text{bdry}}$
- It is not known how to compute this directly in string theory and we will borrow the gravity result.

Grand canonical ensemble

- In the correspondence with the symmetric orbifold Sym^N(T⁴), N is the number of strings in the background.
- Since perturbative strings wind AdS₃ asymptotically, they contribute to this number and hence perturbative string theory computes the grand canonical partition function

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[Kim, Porrati '15]
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$$\begin{aligned} \mathfrak{Z}_{\mathsf{Sym}(\mathbb{T}^4)} &= \sum_{N=0}^{\infty} p^N Z_{\mathsf{Sym}^N(\mathbb{T}^4)} \\ &= \exp\left(\sum_{a,d=1}^{\infty} \sum_{b=1}^d \frac{p^{ad}}{ad} Z_{\mathbb{T}^4}\left(\frac{a\tau_{\mathsf{bdry}} + b}{d}\right)\right) \end{aligned}$$

This expresses the partition function as a sum over holomorphic covering surfaces.

Torus partition function

- Thus, we need to compute the string torus partition function.
- ► Thermal AdS₃ is obtained via a Z-orbifold of global AdS₃.

 \Rightarrow The worldsheet theory is the $\mathbb{Z}\text{-orbifold}$ of the worldsheet theory of global AdS_3.



We use the hybrid formalism to describe the worldsheet theory of global AdS₃.
[Berkovits, Vafa, Witten '99]

 $\mathsf{PSU}(1,1|2)_1/\mathbb{Z}\times\mathsf{top.}$ twisted $\mathbb{T}^4\times\mathsf{ghosts}$

[see Matthias' lectures at pre-strings]

Torus partition function

The worldsheet sphere + torus partition function takes the form

$$\frac{\operatorname{Im} \tau_{\mathsf{bdry}}}{2} \sum_{a,b,c,d \in \mathbb{Z}} p^{\det \binom{a}{c} \frac{b}{d}} \delta^2 \big(\tau_{\mathsf{bdry}}(c\tau + d) - a\tau - b \big) Z^{\mathbb{T}^4} \begin{bmatrix} \frac{b}{2} \\ \frac{a}{2} \end{bmatrix} (\tau) \ .$$

▶ ∑_{a,b,c,d} is the sum over worldsheet instanton sectors.
 ▶ The partition function localizes in the moduli space of Riemann surface to tori that cover the boundary torus holomorphically. [Matthias' talk, Rajesh @ Strings 2020]

The string partition function

- For the string partition function, one integrates over the moduli space of tori.
- This is straightforward because of the presence of the δ-function:

$$\begin{split} \mathfrak{Z}_{\mathsf{thermal}} \operatorname{AdS}_3 &= \exp\left(\sum_{a,d=1}^{\infty} \sum_{b=1}^{d} \frac{p^{ad}}{ad} Z_{\mathbb{T}^4}\left(\frac{a\tau_{\mathsf{bdry}} + b}{d}\right)\right) \\ &\stackrel{!}{=} \mathfrak{Z}_{\mathsf{Sym}(\mathbb{T}^4)} \ . \end{split}$$

One can repeat the same exercise for other backgrounds [LE '20]

$$\mathfrak{Z}_{\mathsf{thermal}} \ \mathsf{AdS}_3 = \mathfrak{Z}_{\mathsf{Euclidean}} \ \mathsf{BTZ} = \mathfrak{Z}_{\mathsf{conical}} \ \mathsf{defect} = \mathfrak{Z}_{\mathsf{Sym}(\mathbb{T}^4)} \ .$$

These equalities hold (at least) to all orders in string loops.



Long strings on thermal AdS₃ can be equivalent to a Euclidean BTZ black hole (String/BH correspondence).

[Susskind '94; Horowitz, Polchinski '96; Giveon, Kutasov, Rabinovici, Sever '05]

A possible resolution of the factorization problem

 One can also consider the tensionless string on a Euclidean wormhole geometry



worldsheet contributions here cancel

One can argue that the string partition function only receives contributions from worldsheets cover either the left or right boundary.

A possible resolution of the factorization problem (cont'd)

The string partition function on the wormhole factorizes:

$$\mathfrak{Z}_{\mathsf{wormhole}} = \mathfrak{Z}_{\mathsf{L}} \times \mathfrak{Z}_{\mathsf{R}}$$
 .

- Since the worldsheets cannot enter the wormhole, the wormhole just looks like a disconnected geometry for the string.
- So there is no factorization problem.

Lessons

- The tensionless string is useful to explore quantum gravity at very short distances.
- The worldsheet theory simplifies enormously and is perturbatively under very good control.
- The natural ensemble of the duality is the grand canonical ensemble.
- The tensionless string is background independent and one does *not* need to sum over geometries.
- There does not seem to be a factorization problem for the tensionless string.