LAGO: the Latin American Giant Observatory, current status, projects and future perspectives.

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Latin America Giant Observatory





The Latin American Giant Observatory

LAGO is an extended astroparticle observatory at continental scale: from México to Antarctica

- Scientific goals
 - Astroparticle physics to study the extreme universe
 - Transient and long term space weather phenomena trough Solar modulation of Cosmic Rays
 - Measurement of background radiation at ground level
- Academic goals
 - Train Latin American students in High Energy and Astroparticle physics
 - Build a Latin American network of Astroparticle and Cosmic Rays researchers

The Latin American Giant Observatory

Autonomous, reliable, simple, cheap and smart (based on SBC and COTS) WCD with a single PMT (usually provided by LAGO in most of the participating countries)





New own designed electronic based on SteamLab RedPitaya H. Arnaldi et al, <u>IEEE2020</u>

SaaS (Sensors as a Service) Concept H. Asorey et al, <u>PoS(ICRC2015)</u>

LAGO Capabilities: Multi-spectral analysis

- Simultaneous measurements of secondaries at ground level
- Intensive simulation and data analysis frameworks

Connections CR Flux Astrophysics transients Modulated flux ··· ··· Modulated flux Geomagnetic field Primaries ··· ··· Primaries Atmospheric conditions Secondary particles ··· ··· Secondary particles Detector response Signals

Synergy

Flux variation of signals at detector level⇔Transients



<u>ARTI</u>, the LAGO simulation framework



A novel cloud-based framework for standardized simulations in the Latin American Giant Observatory, A. J. Rubio- Montero, et. al., WSC '21

<u>ARTI</u>, the LAGO simulation framework



Detailed flux of of secondary particles at detector level for all LAGO sites and other locations around the World. The EOSC-Synergy cloud services implementation for the Latin American Giant Observatory (LAGO), H. Asorey et al, PoS(ICRC2021)261 A novel cloud-based framework for standardized simulations in the Latin American Giant Observatory, A. J. Rubio- Montero, et. al., WSC '21

LAGO Universities

Ajuste exponencial para la diferencia de tiempo entre pulsos para un detector Cherenkov de Agua (LAGO COLLABORATION)



Experimental, astro-ph & hep-ph courses availables

Muon decay: electroweak theory, python, data analysis, simulations, detector physics, statistics, ...

H. Asorey et al, Rev. Bras. Ensino Fís. 40 (4)

Yearly LAGO workshop and AP&HE physics schools (combined mode since 2012!) More than 400 participating students in total



Last in person meeting - Dec. 2019 - @ CAC - Bs As

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H. Asorey et al, Rev. Bras. Ensino Fís. 40 (4)

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Next edition Feb. 2022 @ Tucuman, Argentina



LAGO Universities

The Latin American Giant Observatory (LAGO)

LAGO

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Home	About	Publications	Activities	News	Contact			
_AGO Thesis								
Title					Author [Director]	Thesis Type	School	Year
<u>Caracteriz</u> simulaciór	<u>ación de per</u> 1 de la colabo	files atmosféricos pración LAGO	<u>para la cadena</u>	<u>de</u>	Grisales- Casadiegos, J. []	Undergraduate thesis	Escuela de Física, Universidad Industrial de Santander, Bucaramanga, Colombia	2019
<u>Estimaciór</u> fondo de r	<u>n de la respu</u> ayos cósmic	lesta de un detecto os en Bucaraman	or Cherenkov d ga(956 m s.n.m	<u>e agua al</u>))	Jaimes-Motta, A. []	Undergraduate thesis	Escuela de Física, Universidad Industrial de Santander, Bucaramanga, Colombia	2018
Procedimiento de instalación, calibración y sincronización del arreglo de detectores cherenkov de agua (guane), para la detección y estudio de rayos cósmicos en Bucaramanga				<u>ón del</u> la a	Hernández-Barajas SP, León-Carreño YF. []	Undergraduate thesis	Escuela de Ingeniería Eléctrica, Electrónica y de Telecomunicaciones, Universidad Industrial de Santander, Bucaramanga,	2018

26 thesis / 44 publications / 15 astroparticle schools in LA efficiency: (scientific production / investment) tends to infinity I. Sidelnik for LAGO, LAS4RI forum, 2020



70.4%

thesisType

Undergraduate Thesis

LAGO Virtual





Own designed hierarchical data analysis and virtualized docker-based tools

Measured: 2 TB/year-detector. 4 quality levels: L1: raw data, L2: preliminary, L3: Data Quality, L4: High Quality

Simulated: Up to 1 PB (estimated), EOSC-Synergy thematic service lead by CIEMAT: S0: raw data, S1: simulated and modulated particles at ground, S2: simulated signals at detector level.

Data is transferred to the central repository and is mirrored to several sites (+ each site has its own local data)

H. Asorey, Red Clara symposium, 2021



LACO SW

LAGO studies Earth-Sun

connection by measuring the time-evolving secondary signals from ground level. Atmospheric and geomagnetic conditions are continuously monitored

Antarctic dedicated SW sites



LAGO SW

3%

2%

1%

0%

-1%

-2%

Intensive data analysis to detect secular and transient Solar Wind and Magnetosphere interaction: Forbush decreases (iCMEs) and Solar daily modulation (R. Aguiar et al, PoS(ICRC2021)1267)



CME1 CME2 CME3 Kp = 4 Kp = 5 Kp = 8

CME4

Kp = 6

CME5

Kp = 3





Long term analysis for space climate at low rigidity sites: LAGO observations at the Antarctica Peninsula

Observations of the cosmic ray detector at the Argentine Marambio base in the Antarctic Peninsula

N. Santos et al, PoS(ICRC2021)304

Seismic activity could produce changes in local geoelectric and geomagnetic fields E and B

Study of the effect of seismically-induced geoelectric and geomagnetic fields on secondary particle detection at a LAGO site.

Diego Alberto Coloma Borja et al, PoS(ICRC2021)929



LAGO HE program

Small arrays of WCD at very high altitude sites (<4500 m asl)

FOV overlapped by design for simultaneous measurements

New high altitude sites projects!

- Mt. Chimborazo @ Ecuador (ESPOCH) [Just started]
- Atacama @ Chile (Apex site) [application for funding ongoing]
- Imata @ Perú [application for funding ongoing]
- San Antonio de los Cobres @ Argentina



LAGO Capabilities for detecting high energy component of GRBs and Gamma Sources

C. Sarmiento-Cano et al, PoS(ICRC2021)929

LAGO ML program and Michel spectrum



Molina el al- AGRANDA 2021 DATA ACQUISITION + DATA CLEANING + DATASET • 2 WCDs - 2 sites

- original data set > 2.5. millions samples (raw data
- Very aggressive > we keep 5% of the data.
- Cleaning of raw pulses: saturated, duplicated multiple peaks, negative amplitudes.

FEATURES

 Charge histogram, peak, charge/peak, pulse rise time, delta time (consecutive pulses), pulse width

DIMENSIONALITY

< PCA





----- T3 = 100, T3 = 65

----- T3 = 100, T3 = 85

----- T3 = 100. T3 = 95

T3 = 100 T3 = 100

3000 3500 400

Altura <= 80 cm

1500 2000 2500

Charge [ADC]

T3 = 100, T3 = 75





Machine Learning program at the LAGO Collaboration

M. Graciela Molina et al, AGRANDA meeting 2021

Michel spectrum calculation from water Cherenkov detector data of the LAGO Observatory

L. Otiniano et al

Conclusion

- LAGO major activities
- Long base WCD array from Mexico to Antarctica
- High and low altitude sites that allow us to perform space weather, high energy physics and background radiation measurements
- Complete simulation chain from the primary cosmic rays flux to signal at the WCD
- Multispectral analysis
- New sites @ latin america with projects and fund requested
- Local and regional integration of universities and institutes
- Student training in high energy physics @LA: schools and experiment @ different sites
- Very active LA community with several project funded and ongoing @ different institutions

LAGO constitutes a Latin American network of students and researchers in astroparticle and high energy physics

