



Quasinormal modes for Kerr black hole via Painlevé transcendents

João Paulo Cavalcante * and Bruno Carneiro da Cunha†

Department of Physics, Federal University of Pernambuco, 50670-901, Recife, Brazil



Introduction

The Teukolsky Master equation (TME) governs linear perturbations of the Kerr metric [2], where, for vacuum perturbations, one has

$$\frac{1}{\sin \theta} \frac{d}{d\theta} \left[\sin \theta \frac{dS}{d\theta} \right] + \left[a^2 \omega^2 \cos^2 \theta - 2a\omega s \cos \theta - \frac{(m + s \cos \theta)^2}{\sin^2 \theta} + s + \lambda \right] S(\theta) = 0,$$

$$\Delta^{-s} \frac{d}{dr} \left(\Delta^{s+1} \frac{dR(r)}{dr} \right) + \left(\frac{K^2(r) - 2is(r-M)K(r)}{\Delta} + 4is\omega r - s\lambda_{\ell, m} - a^2 \omega^2 + 2am\omega \right) R(r) = 0,$$

where $\Delta = r^2 - 2Mr + a^2 = (r - r_+)(r - r_-)$, $K(r) = (r^2 + a^2)\omega - am$. M and $J = aM$ are the mass and angular momentum of the black hole. The spin-weight field s can assume the values $0, \pm 1$, and ± 2 .

Riemann-Hilbert map

The Riemann-Hilbert map, between t_0, c_0 and σ, η , is made possible by the isomonodromic τ function, which has a natural expansion in terms of monodromy data [3]. Thus, the RHm is expressed in terms of τ_V by

$$\tau_V(\vec{\theta}; \sigma, \eta; t_0) = 0, \quad t_0 \frac{d}{dt_0} \log \tau_V(\vec{\theta}_-; \sigma - 1, \eta; t_0) - \frac{\theta_0(\theta_t - 1)}{2} = t_0 c_{t_0} \quad (3)$$

where $\vec{\theta} = \{\theta_0, \theta_t, \theta_s\}$ are the parameters in the CHE associated to the local monodromy of solutions and $\vec{\theta}_- = \{\theta_0, \theta_t - 1, \theta_s + 1\}$. In turn, the Riemann-Hilbert map associated to the DCHE is given by

$$\tau_{III}(\vec{\theta}; \sigma, \eta; u_0) = 0, \quad u_0 \frac{d}{du_0} \log \tau_{III}(\vec{\theta}_-; \sigma - 1, \eta; u_0) - \frac{(\theta_0 - 1)^2}{8} - \frac{1}{2} = u_0 k_0, \quad (4)$$

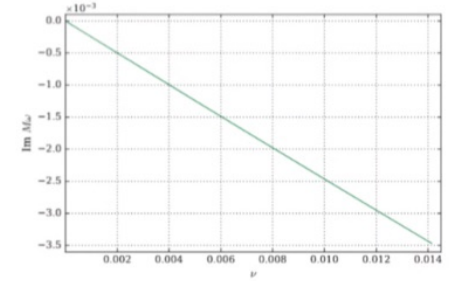
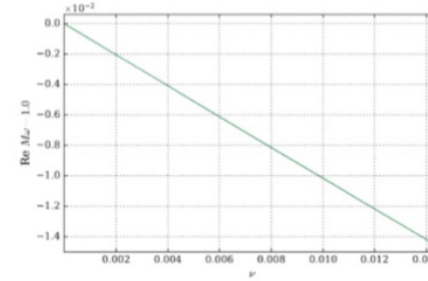
where $\vec{\theta} = \{\theta_0, \theta_s\}$ are the parameters in the DCHE associated to the local monodromy of solutions and $\vec{\theta}_- = \{\theta_0 - 1, \theta_s + 1\}$.

The function τ_V and τ_{III} can be expressed in terms of Fredholm determinant [4] or via Nekrasov partition function [5], while the parameters σ and η are functions of the monodromy parameters of the equations (CHE and DCHE).

Numerical Results

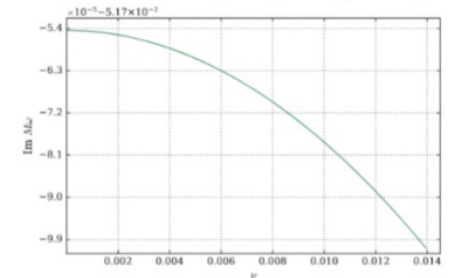
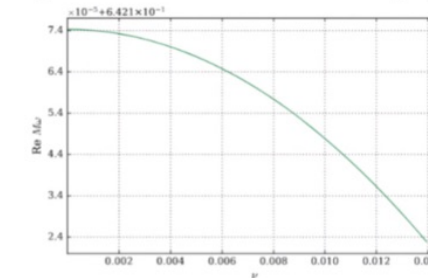
Extremal Limit $a \rightarrow M$: A. We observed numerically that for the modes $l = m$, with $m \neq 0$, the eigenfrequencies tend to $m/(2M)$. In this case, the Riemann Hilbert map (3) actually solved the QNMs for $a/M \in [0, 1]$.

Near-extremal behavior for the fundamental quasi-normal frequency for $s = -2, l = m = 2$, where $a/M = \cos(\nu)$.



B. $M\omega$ does not go to $m/2$. All modes with $l \neq m$, including those with negative m , will not tend to $M\omega = m/2$ in the extremal limit. In this situation the modes for $a = M$ are calculated using the Riemann-Hilbert map (4).

The near-extremal behavior for the fundamental quasi-normal frequency for $s = -1, l = 2$ and $m = 1$, where the mode calculated using τ_V converges to the frequency for τ_{III} as ν goes to 0.



Bibliography

- [1] Carneiro da Cunha, Bruno and Cavalcante, João Paulo, Confluent conformal blocks and the Teukolsky master equation, Phys. Rev. D, 102, 10, 2020.
- [2] Teukolsky, Saul A., Perturbations of a rotating black hole. 1. Fundamental equations for gravitational electromagnetic and neutrino field perturbations, Astrophys. J., vol 185, 1973
- [3] M. Jimbo, T. Miwa, and A. K. Ueno, Monodromy Preserving Deformation of Linear Ordinary Differential Equations with Rational Coefficients, I, Physica D2 (1981) 306-352.
- [4] da Cunha, Bruno Carneiro and Cavalcante, João Paulo, Teukolsky master equation and Painlevé transcendents: Numerics and extremal limit, Phys. Rev. D, Vol 104, 2021.
- [5] O. Gamayun, N. Iorgov, and O. Lisovyy, How instanton combinatorics solves Painlevé VI, V and III, J.Phys. A46.