





Lattice **QCD** with an inhomogeneous magnetic field background

Workshop on Electromagnetic Effects in Strongly Interacting Matter - ICTP-SAIFR, SP

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Strongly magnetized physical systems	Lattice QCD and magnetic fields	Lattice simulations	Summary & Conclusions	References

OUTLINE

- 1. Strongly magnetized physical systems
- 2. Lattice QCD and magnetic fields
- 3. Lattice simulations
- 4. Summary & Conclusions

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STRONGLY MAGNETIZED PHYSICAL SYSTEMS

Early universe $\sqrt{eB} \sim 1.5 \text{ GeV}$



Neutron stars $\sqrt{eB} \sim 1 \text{ MeV}$





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STRONGLY MAGNETIZED PHYSICAL SYSTEMS





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MAGNETIC FIELDS IN HEAVY-ION COLLISIONS

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MAGNETIC FIELDS IN HEAVY-ION COLLISIONS



Figure 2: Spatial distributions of the electric (right) and magnetic (left) fields in the transverse plane for an impact parameter b = 10 fm \mathscr{P} Deng and Huang 2012.

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MAGNETIC FIELDS IN HEAVY-ION COLLISIONS



Figure 2: Spatial distributions of the electric (right) and magnetic (left) fields in the transverse plane for an impact parameter b = 10 fm *P* Deng and Huang 2012.

Caveats:

- highly non-homogeneous background.
- A real E leads to sign problem.
- No Minkoswki time evolution from Euclidean simulations.

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Caveats:

What can we do?

- (highly non-homogeneous background.)
- A /veal/ E /veads /to/sign/problem/.
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Caveats:

- highly non-homogeneous background.
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- No Minkoswki time evolution from Euclidean simulations.

What can we do?

B(x) as background in lattice **QCD**!

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Lattice QCD and magnetic fields

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$$\langle \mathcal{O} \rangle = \frac{1}{\mathcal{Z}} \int \mathcal{D}\bar{\psi} \mathcal{D}\psi \mathcal{D}A \ \mathcal{O}e^{-S[\bar{\psi},\psi,A]}$$

Lattice QCD and magnetic fields

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$$\langle \mathcal{O} \rangle = \frac{1}{\mathcal{Z}} \int \mathcal{D}\bar{\psi}\mathcal{D}\psi\mathcal{D}A \ \mathcal{O}e^{-S[\bar{\psi},\psi,A]} \longrightarrow \frac{1}{\mathcal{Z}} \int \mathcal{D}A \det\left[\mathcal{D}(A) + m\right] \mathcal{O}e^{-S_g[A]}$$

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• quarks
$$\psi(x) \ x \in \mathbb{R} \longrightarrow \psi(n) \ n \in \mathbb{Z}$$



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- gluons $A_{\mu} \longrightarrow U_{\mu} = e^{iag A^b_{\mu} T_b} \in SU(3)$
- (anti-)periodic BC



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LATTICE **QCD** IN A NUTSHELL

$$\langle \mathcal{O} \rangle = \frac{1}{\mathcal{Z}} \int \mathcal{D}\bar{\psi}\mathcal{D}\psi\mathcal{D}A \ \mathcal{O}e^{-S[\bar{\psi},\psi,A]} \longrightarrow \frac{1}{\mathcal{Z}} \int \mathcal{D}A \det\left[\mathcal{D}(A) + m\right] \mathcal{O}e^{-S_g[A]}$$



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(anti-)periodic BC



1. Generate samples $\{\mathcal{O}_1, \mathcal{O}_2, ..., \mathcal{O}_N\}$ with a probability $\det [\mathcal{D}(A) + m] e^{-S_g}$ using Monte Carlo steps.

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LATTICE QCD IN A NUTSHELL

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• magnetic field
$$B \longrightarrow u_{\mu} = e^{iaqA_{\mu}} \in U(1)$$
 (BACKGROUND!)

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MAGNETIC FIELD ON THE LATTICE

Consider a uniform eld in the z directions:

B = B≵

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Stoke's theorem must hold:

inner area: A dx = SB



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 $e^{iqBS} = e^{iqB(L_x L_y S)}$



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inner area: A dx = SB I outer area: A dx = $(L_xL_y S)B$

$$e^{iqBS} = e^{iqB(L_xL_y S)}$$



$$qB = \frac{2 N_b}{L_x L_y}; \quad N_b 2 Z$$

The magnetic ux is quantized inside a box!

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UNIFORM MAGNETIC FIELD ON THE LATTICE

$$B = r \qquad A$$
$$A_y = Bx \qquad A_x = A_z = A_t = 0$$

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UNIFORM MAGNETIC FIELD ON THE LATTICE

$$B = r \qquad A$$

$$A_y = Bx \qquad A_x = A_z = A_t = 0$$

$$u_y = e^{jaqBx} \qquad u_x = u_z = u_t = 1$$

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UNIFORM MAGNETIC FIELD ON THE LATTICE

$$\begin{split} B &= r \qquad A \\ A_y &= Bx \qquad A_x = A_z = A_t = 0 \\ u_y &= e^{iaqBx} \qquad u_x = u_z = u_t = 1 \\ u_y(L_x) &= e^{ia\,2\,Nb=L_y} \ \textbf{6} \ u_y(0) \end{split}$$

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UNIFORM MAGNETIC FIELD ON THE LATTICE

$$A_y = Bx \quad A_x = A_z = A_t = C$$

$$u_y = e^{iaqBx}$$
 $u_x = u_z = u_t = 1$

 $u_y(L_x) = e^{ia 2 Nb=L_y} \in u_y(0)$

We can perform gauge transformations on the links

$$u^{0}(x) = (x)u (x + a^{y})^{y}$$

a is the lattice spacing.

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We can perform gauge transformations on the links

$$u^{0}(x) = (x)u (x + a^{y})^{y}$$

a is the lattice spacing.

$$u_{x} = \begin{pmatrix} e^{iqBL_{x}y} & \text{if } x = L_{x} & a \\ 1 & \text{if } x \in L_{x} & a \\ u_{y} = e^{iaqBx} & 0 & x & L_{x} & a \\ u_{z} = 1 & & \\ u_{t} = 1 & & \end{pmatrix}$$

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INHOMOGENEOUS MAGNETIC FIELD ON THE LATTICE

$$\mathsf{B} = \frac{\mathsf{B}}{\cosh \frac{\mathsf{x} - \mathsf{L}_{\mathsf{x}} = 2}{2}} \mathsf{Z}$$

Pro le motivated by heavy-ion collision scenarios / Deng and Huang 2012, / Cao 2018.

$$qB = \frac{N_b}{L_y \tanh \frac{L_x}{2}} \qquad N_b 2 Z$$

$$u_{x} = \begin{pmatrix} e^{-2iqB y \tanh\left(\frac{L_{x}}{2}\right)} & \text{if } x = L_{x} & a \\ 1 & \text{if } x \in L_{x} & a \end{pmatrix}$$
$$u_{y} = e^{iaqB} \left[\tanh\left(\frac{x - L_{x} = 2}{2}\right) + \tanh\left(\frac{L_{x}}{2}\right) \right]; \quad 0 \quad x \quad L_{x} \quad a$$
$$u_{z} = u_{t} = 1$$

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THE SIMULATION SET UP

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THE SIMULATION SET UP

• $N_f = 2 + 1$ improved staggered fermions with physical masses;

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THE SIMULATION SET UP

- $N_f = 2 + 1$ improved staggered fermions with physical masses;
- Lattices: 16³ 6 24³ 8 28³ 10 36³ 12 ! continuum limit (lattice spacing ! 0, V = const.);
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- Magnetic eld

$$B = \frac{B}{\cosh \frac{x - L_x = 2}{2}} 2 eB = \frac{3 N_b}{L_y \tanh \frac{L_x}{2}} 0.6 \text{ fm}$$

strength 0 GeV $p = 1.2 \text{ GeV}$ magnetars, HIC and early
universe.

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• Temperature range 68 MeV T 300 MeV (crossover transition at T_c 155 MeV).

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• Local chiral condensates (u and d quarks!)

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• Local chiral condensates (u and d quarks!)

renormalization
$$(x; T; B) = \frac{m_{ud}}{m^4} (x; T; B) (x; T; 0)$$

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• Local chiral condensates (u and d quarks!)

$$\stackrel{\text{renormalization}}{!} (x; T; B) = \frac{m_{ud}}{m^4} (x; T; B) (x; T; 0)$$
$$(x; T; 0) (L=2; T; B)$$

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$$\stackrel{\text{renormalization}}{!} (x; T; B) = \frac{m_{ud}}{m^4} (x; T; B) (x; T; 0)$$
$$(x; T; 0) (L=2; T; B)$$

Local Polyakov loop

$$P = \frac{1}{L_x L_y} X_{y;z} \operatorname{ReTr}_{n}^{Y} U_t(x;y;z;n)$$

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$$\stackrel{\text{renormalization}}{!} (x;T;B) = \frac{m_{ud}}{m^4} (x;T;B) (x;T;0)$$
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$$P = \frac{1}{L_x L_y} X_{y;z} \operatorname{ReTr}_{n}^{Y} U_t(x;y;z;n) \qquad \stackrel{\text{renormalization}}{\stackrel{P(x;T;B)}{:}} \frac{P(x;T;B)}{P(x;T;0)}$$

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• Local chiral condensates (u and d quarks!)

$$\stackrel{\text{renormalization}}{!} (x;T;B) = \frac{m_{ud}}{m^4} (x;T;B) (x;T;0) (x;T;0) (L=2;T;B)$$

Local Polyakov loop

$$P = \frac{1}{L_x L_y} X_{y;z} \operatorname{Re} \operatorname{Tr} Y_{n} U_t(x;y;z;n) \qquad \stackrel{\text{renormalization}}{!} \frac{P(x;T;B)}{P(x;T;0)}$$
$$P(x;T;0) \qquad P(L=2;T;B)$$

• Local electric currents (u, d and s quarks!)

$$hJ_i(x) i = e - \frac{2}{3}u^i u - \frac{1}{3}d^i d - \frac{1}{3}s^i s$$

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CHIRAL CONDENSATE

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CHIRAL CONDENSATE

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CHIRAL CONDENSATE



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CHIRAL CONDENSATE IN THE T-B PLANE

What happens to the peak of the condensate as a function of T and B?

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CHIRAL CONDENSATE IN THE T-B PLANE

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CHIRAL CONDENSATE IN THE T-B PLANE

What happens to the peak of the condensate as a function of T and B?



Valence effect vs Sea effect = (inverse) magnetic catalysis

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The Polyakov loop is typically broader than the chiral condensate.

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The Polyakov loop is typically broader than the chiral condensate.

$$C(x,x') = rac{1}{m_\pi^3} \left\langle \ ar{\psi}\psi(x) P(x') \
ight
angle_c$$



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The Polyakov loop is typically broader than the chiral condensate.

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The interaction of the condensate with *P* causes the dips! (Local inverse magnetic catalysis)

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INHOMOGENEOUS VS UNIFORM CASE

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$\mathbf{J}\sim \boldsymbol{\nabla}\times \mathbf{B}$

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$$\mathbf{J} \sim \mathbf{\nabla} \times \mathbf{B} \quad \longrightarrow \quad J_y \sim \frac{\partial B_z}{\partial x} = -\frac{2B}{\epsilon \cosh\left(\frac{x - L_x/2}{\epsilon}\right)^2} \tanh\left(\frac{x - L_x/2}{\epsilon}\right)$$

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Figure 6: Lattice electric currents for RHIC-like ($\sqrt{eB} = 0.1$ GeV) and LHC-like ($\sqrt{eB} = 0.5$ GeV) magnetic fields, respectively.

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(BARE) MAGNETIC SUSCEPTIBILITY

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(BARE) MAGNETIC SUSCEPTIBILITY

$$\frac{1}{\mu_0}\mathbf{B} = \mathbf{H} + \mathbf{M} \quad \longrightarrow \quad \mathbf{J}_{tot} = \mathbf{J}_f + \mathbf{J}_m \quad \longrightarrow \quad \mathbf{J}_m = \boldsymbol{\nabla} \times \mathbf{M}$$

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- Linear response term: $\mathbf{M}\approx \chi_m \mathbf{H}$

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(BARE) MAGNETIC SUSCEPTIBILITY

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- Linear response term: $\mathbf{M} \approx \chi_m \mathbf{H}$
- $\frac{\chi_m}{1+\chi_m} \nabla \times \mathbf{B} = \mathbf{J}_m$

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(BARE) MAGNETIC SUSCEPTIBILITY

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• Linear response term: $\mathbf{M} \approx \chi_m \mathbf{H}$

•
$$\frac{\chi_m}{1+\chi_m} \nabla \times \mathbf{B} = \mathbf{J}_m$$


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The divergence is independent of *T*: $\chi_m^r(T) \equiv \chi_m(T) - \chi_m(0)$

Strongly magnetized physical systems	Lattice QCD and magnetic fields	Lattice simulations	Summary & Conclusions	References
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Great agreement with the current-current method! @ Bali, Gergely Endrődi,

and Piemonte 2020

eferences

Summary & Conclusions

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- The knowledge of these processes is important to capture the correct physics in heavy-ion collision studies (QCD models, hydrodynamics, etc.);

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Strongly magnetized physical systems	Lattice QCD and magnetic fields	Lattice simulations	Summary & Conclusions	References

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