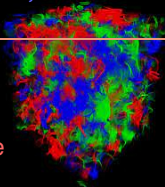




# Lattice QCD with an inhomogeneous magnetic field background

Workshop on Electromagnetic Effects in Strongly Interacting Matter - ICTP-SAIFR, SP



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# OUTLINE

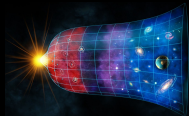
1. Strongly magnetized physical systems
2. Lattice QCD and magnetic fields
3. Lattice simulations
4. Summary & Conclusions

# **Strongly magnetized physical systems**

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# STRONGLY MAGNETIZED PHYSICAL SYSTEMS

Early universe  
 $\rho \frac{eB}{eB}$  1.5 GeV



Heavy-ion collision  
 $\rho \frac{eB}{eB}$  0.5 GeV

Neutron stars  
 $\rho \frac{eB}{eB}$  1 MeV

QCD vacuum

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$\mu_B$  Early universe  $\mu_B$  Heavy-ion collision  
 $\frac{\mu_B}{eB}$  1:5 GeV  $\frac{\mu_B}{eB}$  0:5 GeV

$\mu_B$  Neutron stars QCD vacuum  
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# MAGNETIC FIELDS IN HEAVY-ION COLLISIONS

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Figure 2: Spatial distributions of the electric (right) and magnetic (left) fields in the transverse plane for an impact parameter  $b = 10$  fm  Deng and Huang 2012.

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- highly non-homogeneous background.
- A real E leads to sign problem.
- No Minkowski time evolution from Euclidean simulations.




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What can we do?

$B(x)$  as  
background in  
lattice QCD!

# Lattice QCD and magnetic fields

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$$\langle h O_i \rangle = \frac{1}{Z} \int \mathcal{D}D \mathcal{D}A \mathcal{D}e^{S[D;A]} \dots$$

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$$\langle h | O | i \rangle = \frac{1}{Z} \int \mathcal{D}D \mathcal{D}A \text{Oe}^{S[D;A]} \quad ; \quad \frac{1}{Z} \int \mathcal{D}A \det \mathcal{D}(A) + m \text{Oe}^{S_g[A]}$$

# LATTICE QCD IN A NUTSHELL

$$\langle h | O_i | h \rangle = \frac{1}{Z} \int \mathcal{D}U \mathcal{D}\psi \mathcal{D}\bar{\psi} e^{-S[U; \psi, \bar{\psi}]} \frac{1}{Z} \int \mathcal{D}U \det \mathcal{D}(A) e^{-S_g[A] + m \bar{\psi} \psi}$$

quarks  $(x) \times 2 R!$   $(n) n 2 Z$



# LATTICE QCD IN A NUTSHELL

$$\langle \mathcal{O} \rangle = \frac{1}{Z} \int \mathcal{D}A \mathcal{D}\psi \mathcal{D}\bar{\psi} e^{-S[\psi, \bar{\psi}; A]} \mathcal{O} = \frac{1}{Z} \int \mathcal{D}A \det \mathcal{D}(A) e^{-S_g[A]} \mathcal{O}$$

quarks  $\psi(x) \in \mathbb{R}^2$   $(n) \in \mathbb{Z}$   
 gluons  $A_\mu \in \mathbb{U} = e^{iagA_\mu^b T_b} \in \text{SU}(3)$

## LATTICE QCD IN A NUTSHELL

$$\langle \text{Tr} O \rangle = \frac{1}{Z} \int \mathcal{D}U \mathcal{D}A \det \mathcal{D}(A) e^{S_f[A, U]} \frac{1}{Z} \int \mathcal{D}A \det \mathcal{D}(A) e^{S_g[A]}$$

quarks  $(x) \times 2 R!$   $(n) n \times 2 Z$   
 gluons  $A!$   $U = e^{iagA^b T_b} \in SU(3)$   
 (anti-)periodic BC

# LATTICE QCD IN A NUTSHELL

$$Z = \int \prod_{\text{links}} dA \int \prod_{\text{quarks}} d\psi d\bar{\psi} e^{-S_g[A] - S_f[\psi, \bar{\psi}; A]}$$

quarks  $\psi(x) \in \mathbb{C}^2 \otimes \mathbb{C}^3$   
 gluons  $A_\mu \in \mathfrak{su}(3)$   
 (anti-)periodic BC

1. Generate samples  $\{O_1; O_2; \dots; O_N\}$  with a probability  $\det \mathcal{D}(A) + m e^{-S_g}$  using Monte Carlo steps.

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magnetic field  $B_\mu$   $u = e^{iaqA_\mu} \in U(1)$  (BACKGROUND!)

# MAGNETIC FIELD ON THE LATTICE

Consider a uniform field in the  $z$  directions:

$$\mathbf{B} = B \hat{z}$$

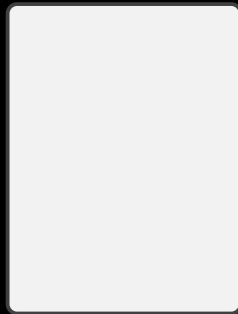
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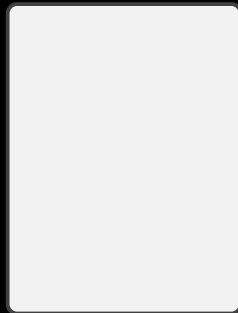
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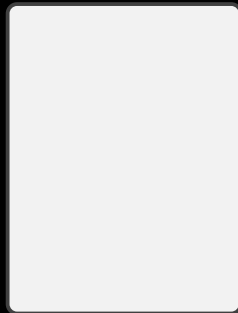
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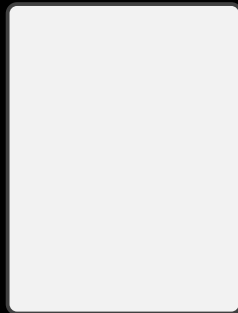
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$$qB = \frac{2\pi N_b}{L_x L_y}; \quad N_b \in \mathbb{Z}$$

The magnetic flux is quantized inside a box!

# UNIFORM MAGNETIC FIELD ON THE LATTICE

$$\mathbf{B} = r \hat{z} \quad \mathbf{A}$$

$$A_y = Bx \quad A_x = A_z = A_t = 0$$

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$$u_x = \begin{cases} e^{iqBL_x y} & \text{if } x = L_x \quad a \\ 1 & \text{if } x \notin L_x \quad a \end{cases}$$

$$u_y = e^{iaqBx} \quad 0 \leq x < L_x \quad a$$

$$u_z = 1$$

$$u_t = 1$$

## INHOMOGENEOUS MAGNETIC FIELD ON THE LATTICE

$$B = \frac{B}{\cosh \frac{x - L_x/2}{2} \hat{z}}$$

Proposed motivated by heavy-ion collision scenarios [Deng and Huang 2012](#),

[Cao 2018](#).

$$qB = \frac{N_b}{L_y \tanh \frac{L_x}{2}} \quad N_b \ll Z$$

$$u_x = \begin{cases} e^{2iqB y \tanh(\frac{L_x}{2})} & \text{if } x = L_x \\ 1 & \text{if } x \in L_x \end{cases} \quad a$$

$$u_y = e^{iaqB [\tanh(\frac{x - L_x/2}{2}) + \tanh(\frac{L_x}{2})]}; \quad 0 \leq x \leq L_x \leq a$$

$$u_z = u_t = 1$$



# Lattice simulations

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$$B = \frac{B}{\cosh \frac{x}{L_x=2}} \frac{1}{2} \hat{z}$$

$$eB = \frac{3 N_b}{L_y \tanh \frac{L_x}{2}} \quad 0.6 \text{ fm}$$

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strength 0 GeV  $\frac{p}{eB}$  1:2 GeV ! magnetars, HIC and early universe.

- Temperature range 68 MeV  $T$  300 MeV (crossover transition at  $T_c = 155 \text{ MeV}$ ).

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$$\langle P \rangle(x; T; 0) \quad \langle P \rangle(L=2; T; B)$$

- Local electric currents (**u**, **d** and **s** quarks!)

$$\langle J_i(x) \rangle = e \left( \frac{2}{3} \langle J_i^u \rangle + \frac{1}{3} \langle J_i^d \rangle + \frac{1}{3} \langle J_i^s \rangle \right)$$

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What happens to the peak of the condensate as a function of T and B?

- Magnetic catalysis T away from  $T_c$
- Inverse catalysis for T around  $T_c$

 Endrodi et al. 2019

Valence effect vs Sea effect = (inverse) magnetic catalysis

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# POLYAKOV LOOP

The Polyakov loop is typically broader than the chiral condensate.

$$C(x; x^0) = \frac{1}{m^3} \text{Tr} \left( \prod_{t=0}^{x^0-1} U(x, t) \right)$$

The interaction of the condensate with  $P$  causes the dips! (Local inverse magnetic catalysis)

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How different do the condensates behave in the two cases?

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$$J_r(B) = J_y \frac{\partial B}{\partial x} = \frac{2B}{\cosh \frac{x}{L_x}} \tanh \frac{x}{L_x}$$

Figure 6: Lattice electric currents for RHIC-like ( $\mu_B = 0.1$  GeV) and LHC-like ( $\mu_B = 0.5$  GeV) magnetic fields, respectively.

# (BARE) MAGNETIC SUSCEPTIBILITY

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$$\frac{1}{\mu_0} \mathbf{B} = \mathbf{H} + \mathbf{M} \quad ! \quad \mathbf{J}_{\text{tot}} = \mathbf{J}_f + \mathbf{J}_m \quad ! \quad \mathbf{J}_m = \gamma \mathbf{M}$$

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- $\frac{\chi_m}{1 + \chi_m} \mathbf{B} = \mathbf{J}_m$

The susceptibility contains an additive  
divergence  $\chi_m \sim \log(a)$

# (RENORMALIZED) MAGNETIC SUSCEPTIBILITY

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diamagnetism

# (RENORMALIZED) MAGNETIC SUSCEPTIBILITY

The divergence is independent of T:  $\chi_m(T) \sim \chi_m(0) + \frac{r_m}{T}$

- $r_m < 0$ :  
diamagnetism
- $r_m > 0$ :  
paramagnetism

# (RENORMALIZED) MAGNETIC SUSCEPTIBILITY

The divergence is independent of T:  $\chi_m(T) = \chi_m(0) + \chi_m^r(T)$

- $\chi_m^r < 0$ :  
diamagnetism
- $\chi_m^r > 0$ :  
paramagnetism


Material	$\chi_m$
Al	$+2.2 \cdot 10^{-5}$
Glass	$-1.13 \cdot 10^{-5}$

# (RENORMALIZED) MAGNETIC SUSCEPTIBILITY

The divergence is independent of T:  $\chi_m^r(T) = \chi_m(T) - \chi_m(0)$

- $\chi_m^r < 0$ :  
diamagnetism
- $\chi_m^r > 0$ :  
paramagnetism

Material	$\chi_m$
Al	$+2.2 \cdot 10^{-5}$
Glass	$-1.13 \cdot 10^{-5}$

Great agreement with the current-current method!  Bali, Gergely Endrodi,

and Piemonte 2020



## Summary & Conclusions

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- The knowledge of these processes is important to capture the correct physics in heavy-ion collision studies (QCD models, hydrodynamics, etc.);

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