



ICTP | International Centre for Theoretical Physics
SAIFR | South American Institute for Fundamental Research

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Gravitational Waves meet Amplitudes in the Southern Hemisphere

Towards 5PN in compact binary dynamics

Stefano Foffa



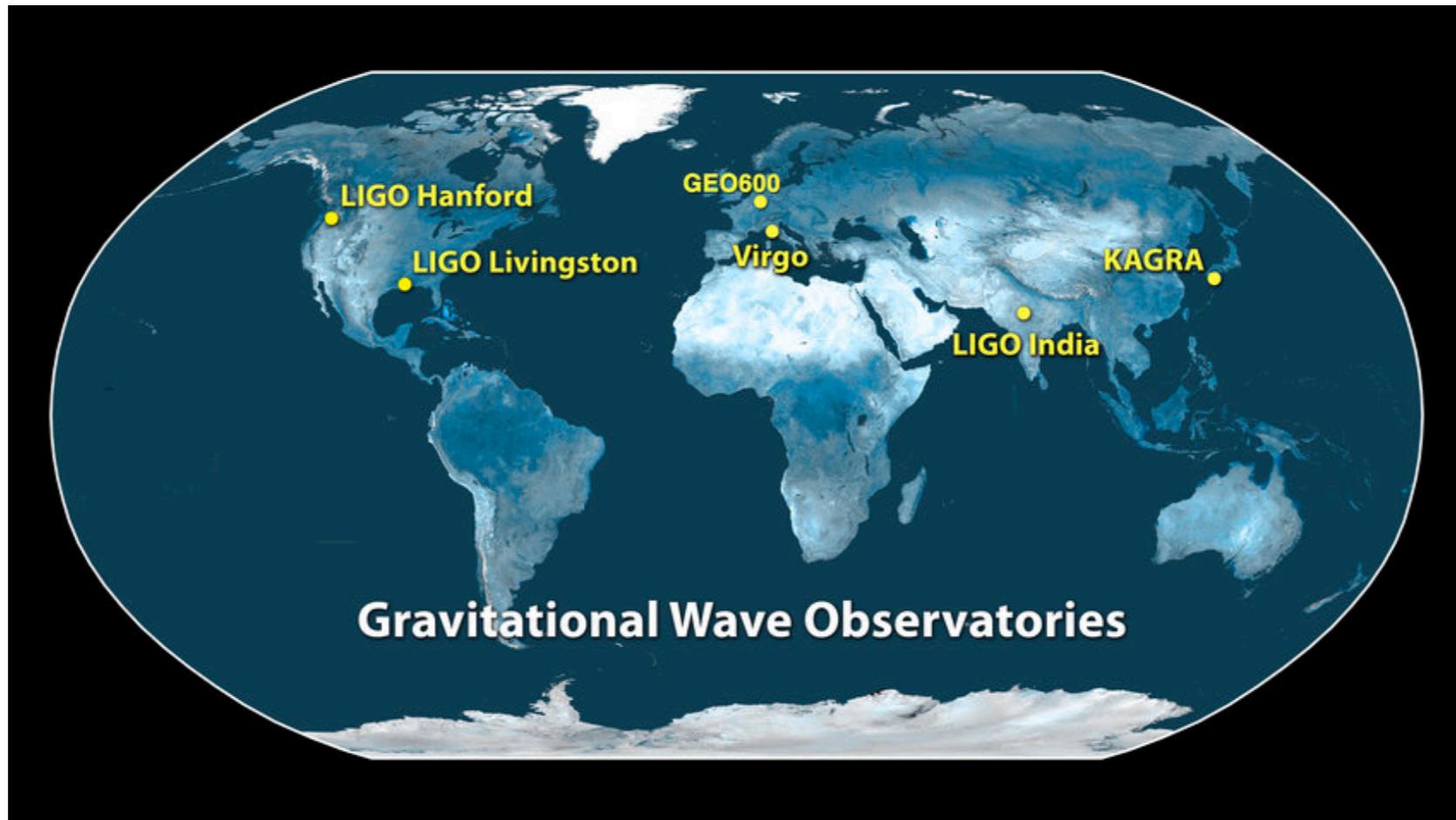
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Département de physique théorique



Southern emisphere GW observatories are less necessary than in the electromagnetic case



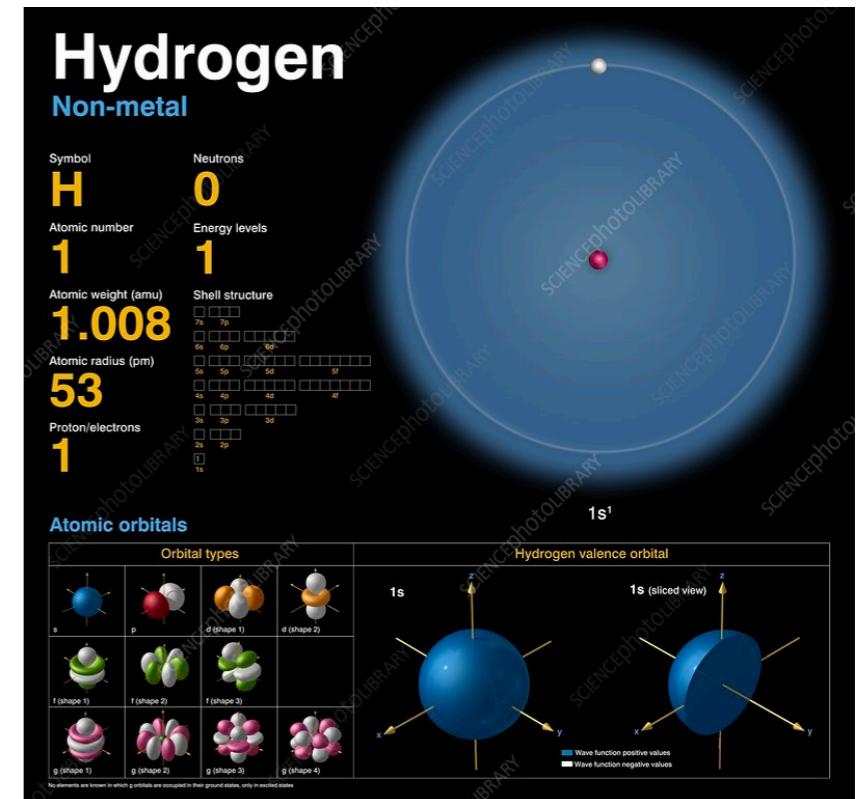
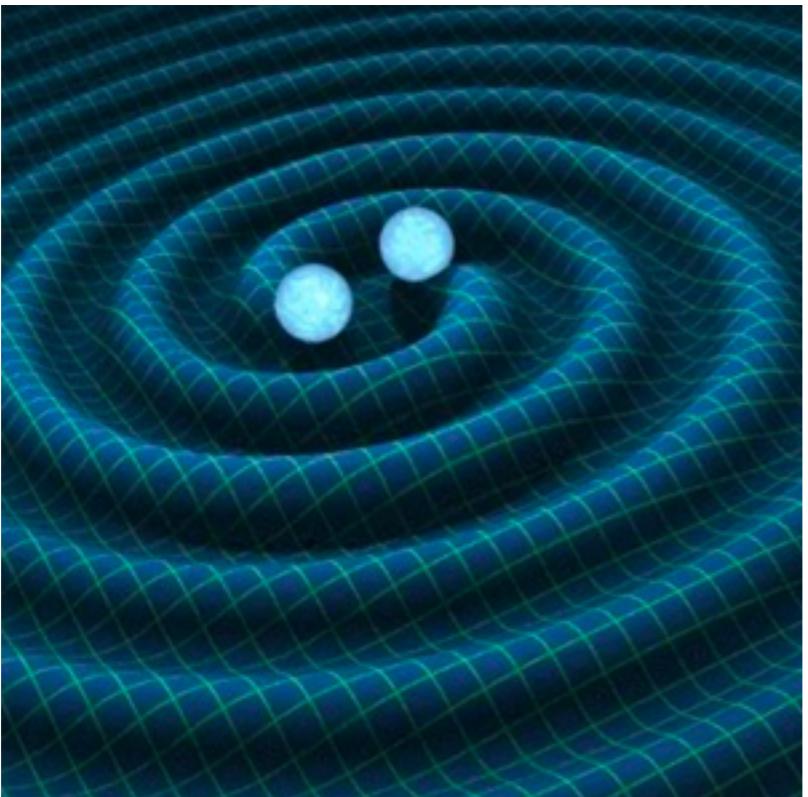
It is even more important to have programs like this to grow a strong and global community of GW/GR enthusiasts

...before we start:

Compact binaries

are the new

hydrogen atom



Simple and rich of features at the same time

Instrumental to improve our understanding of certain dynamical regimes/configurations

ok let's be pragmatic

we need accurate WF's
(especially the **phase**)



balance equation:

conservative
dynamics

circular orbit

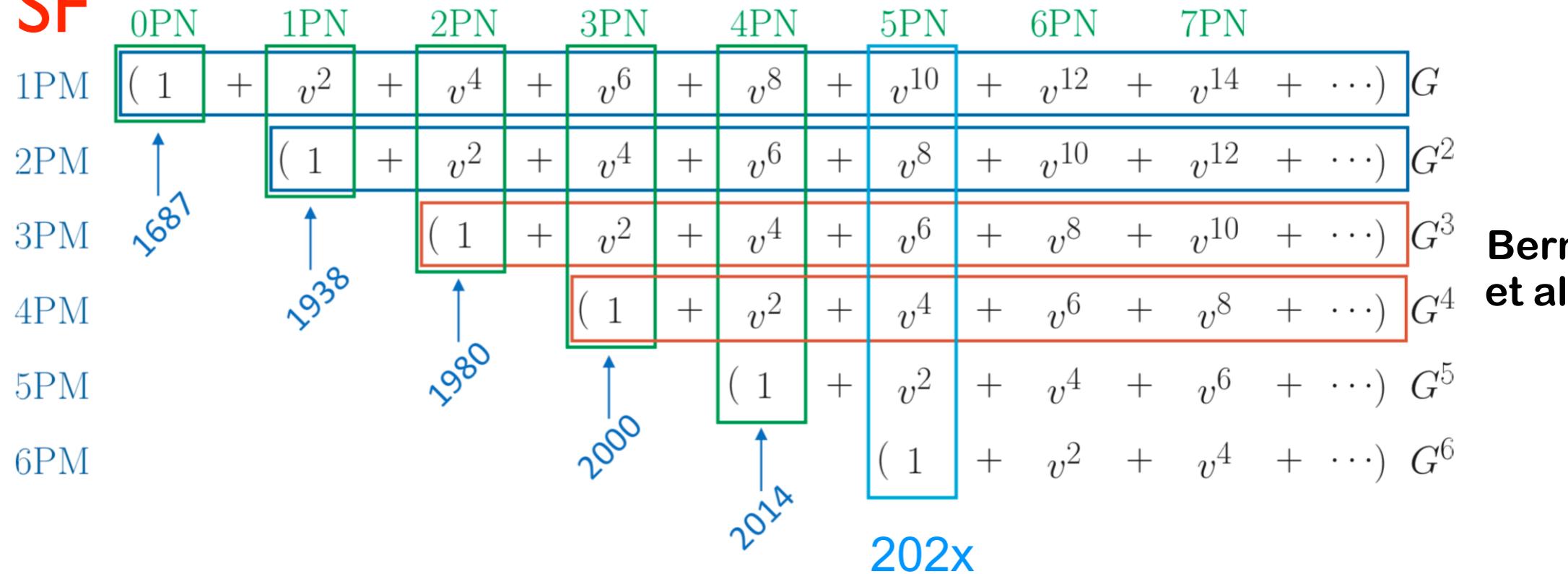
radiative
sector

$$\dot{E}(\omega) = P_{GW}(\omega)$$

(Schott terms can be put on either side!)

PN is the phenomenologically relevant perturbative scheme

PN, PM, SF



Energy:

Bern
et al.

0SF ISF 2SF 3SF 4SF

0PN	1	x^0
1PN	$1 + \nu$	x
2PN	$1 + \nu + \nu^2$	$x^2 \quad x \equiv \left(\frac{GM\omega}{c^3}\right)^{2/3}$
3PN	$1 + \nu + \nu^2 + \nu^3$	x^3
4PN	$1 + \nu + \nu^2 + \nu^3 + \nu^4$	x^4

(even more striking for the scattering angle)

2PN PN iteration

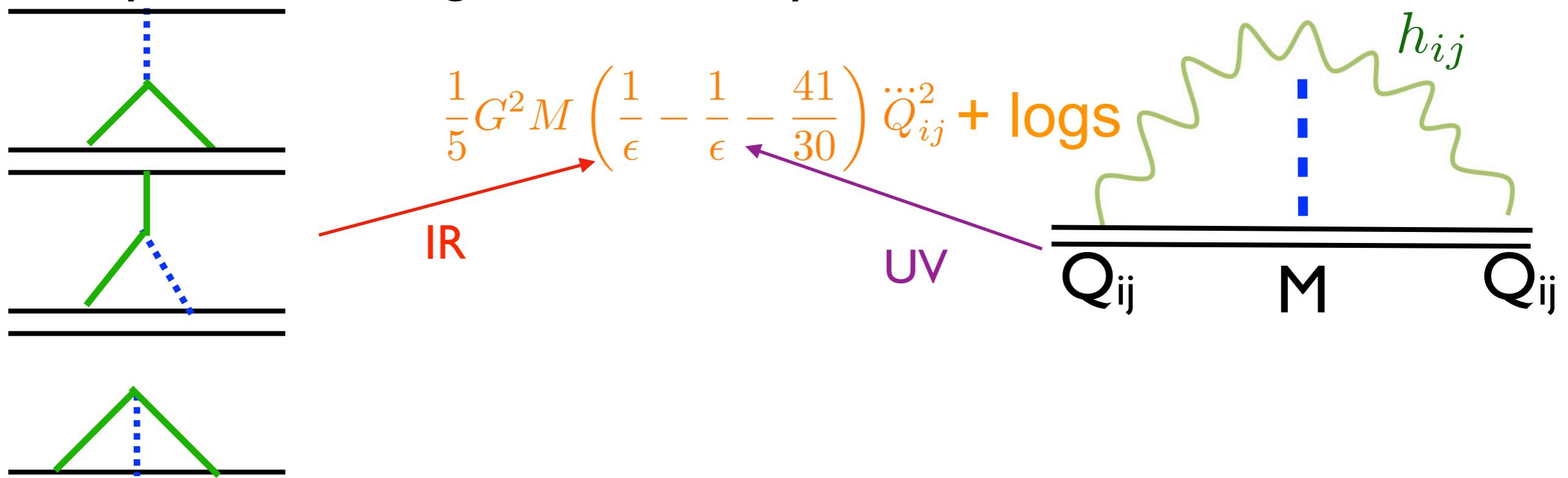
Problems in inverting Poisson operators

3PN ADM, PN iteration, surface integrals

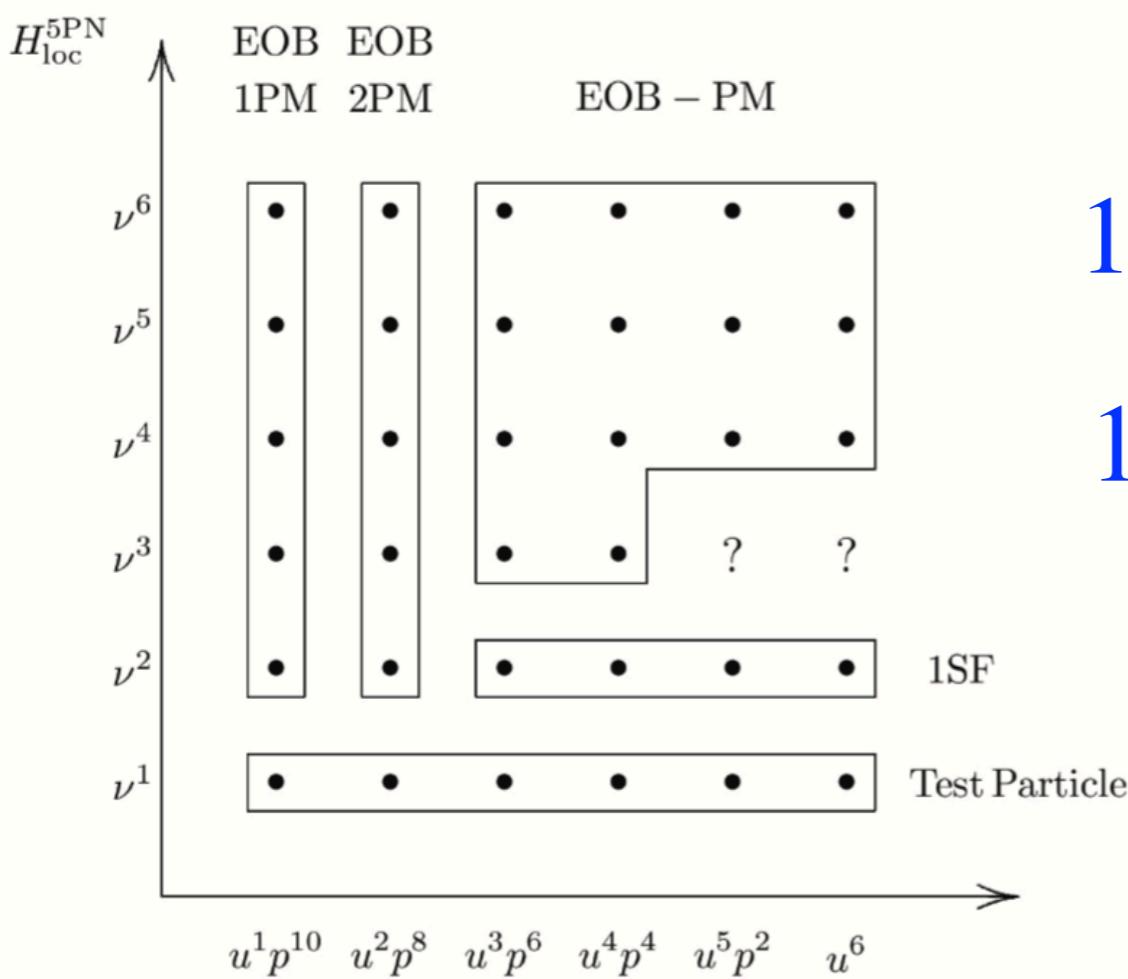
First divergencies: Hadamard and dimensional regularisations

4PN ADM+SF, Fokker lagrangian, EFT

Physical IR divergencies killed by tail contributions



First appearance of highly nontrivial 4-loop integrals:
resort to particle physics technology



test particle limit

1SF redshift z translated into H_{EOB}

1PM and 2PM amplitude results

scattering angle properties

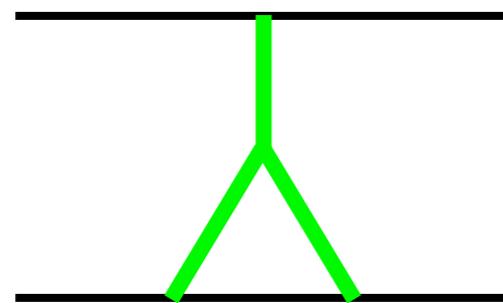
from 1909.02375

34 out of 36 coefficients of local c.o.m. Hamiltonian

(and 147 out of 151 6PN local terms)

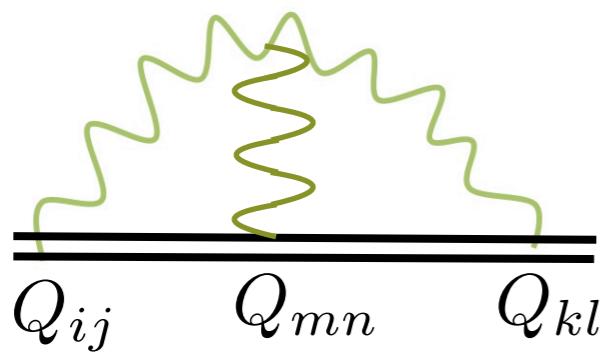
5PN, ab initio

Potential modes



(including finite size effects)

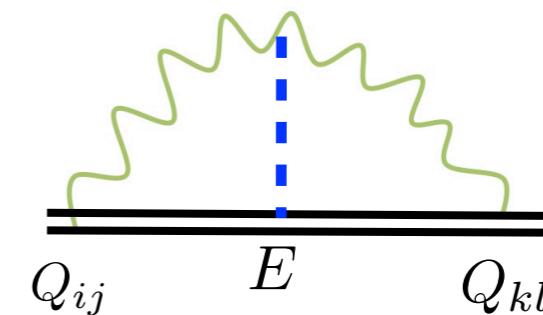
Memory



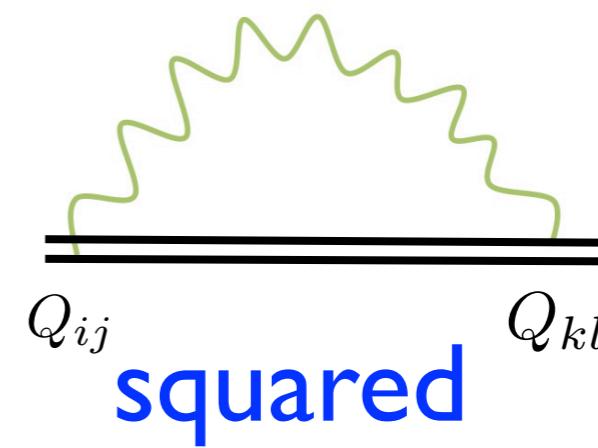
method of regions:

$$\frac{i}{\mathbf{k}^2 - k_0^2} \xrightarrow{\text{method of regions:}} \frac{i}{\mathbf{k}^2} \sum_{n \geq 0} \left(\frac{k_0^2}{\mathbf{k}^2} \right)^n$$

Tails



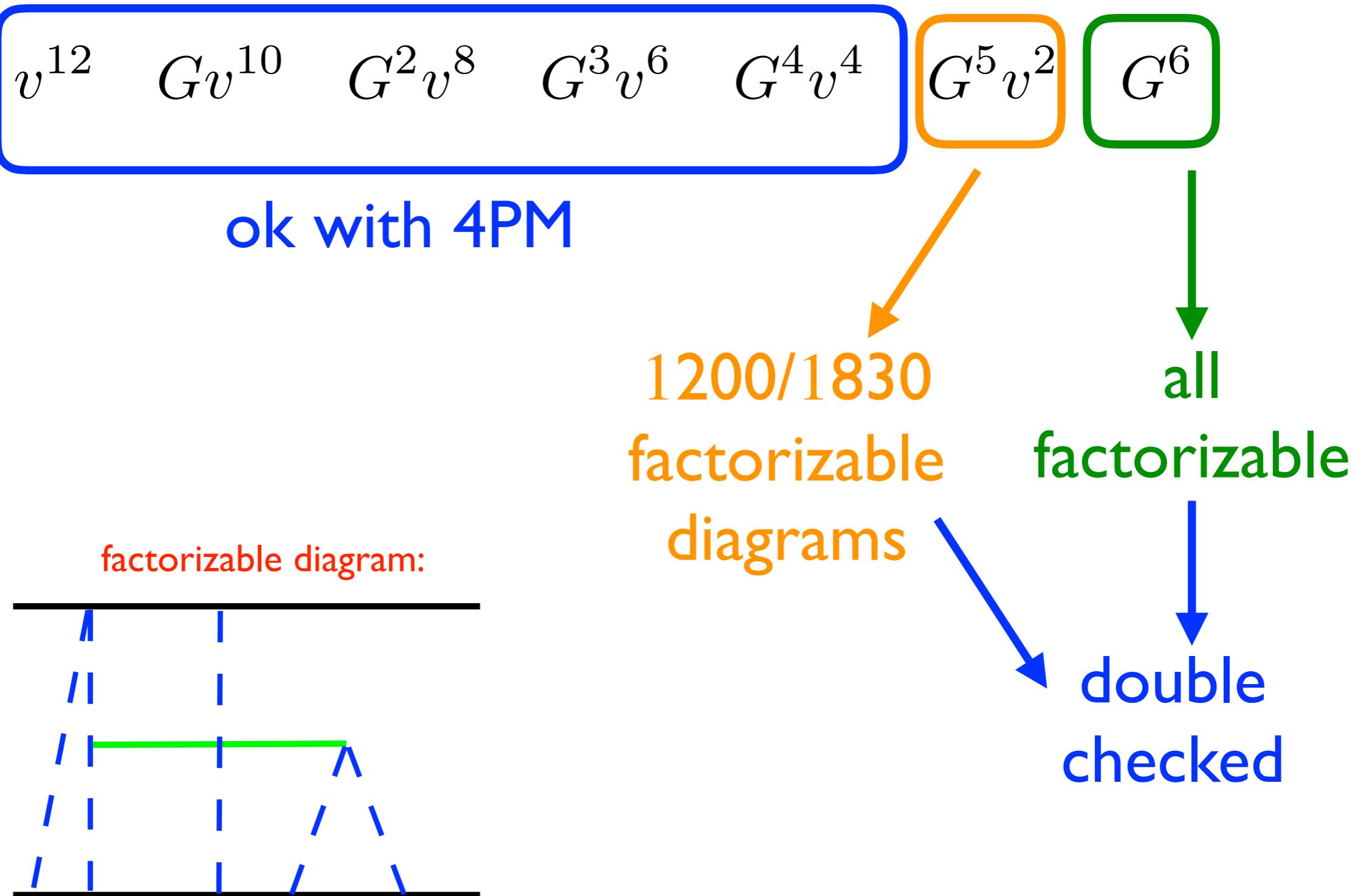
Radiation reaction



5PN, potential modes

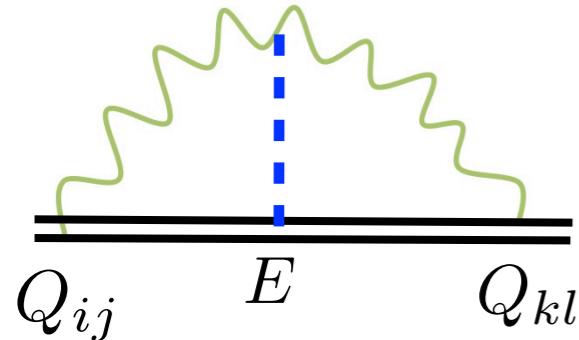
Blumlein et al.

$\mathcal{O}(10^4)$ diagrams in EFT calculation



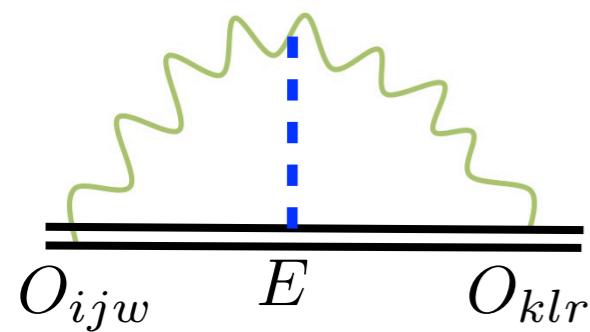
5PN tails

(tails computed for all multipoles)



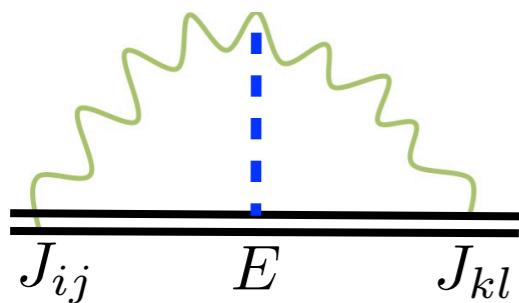
$$-\frac{1}{5} G^2 E \left(\frac{1}{\epsilon_{UV}} - \frac{41}{30} \right) \ddot{\dot{Q}}_{ij}^2$$

+ nonlocal terms



$$-\frac{1}{189} G^2 M \left(\frac{1}{\epsilon_{UV}} - \frac{82}{35} \right) \ddot{\ddot{O}}_{ijk}^2$$

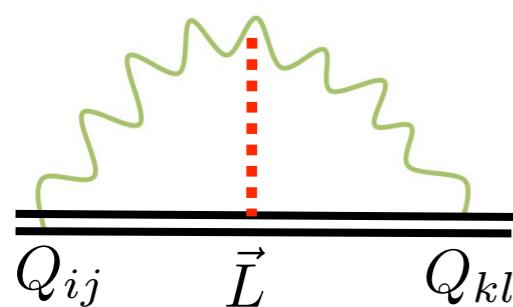
+ nonlocal terms



$$-\frac{16}{135} G^2 M \left(\frac{1}{\epsilon_{UV}} - \frac{49}{20} \right) \ddot{j}_{i|jk} [\ddot{j}_{i|jk} + \ddot{j}_{j|ik}]$$

+ nonlocal terms

ok with 1SF scattering angle χ^{4PM}



~~$$\frac{8}{15} \frac{1}{30} G^2 \epsilon_{ijk} L_k \ddot{\ddot{Q}}_{il} \ddot{\ddot{Q}}_{jl}$$~~



$\chi^{4PM} \mathcal{O}(\nu^2)$

Dissipative contributions: in-in formalism

(irrelevant for potential modes *and tails*)

$$e^{iS_{eff}[x]} = \int \mathcal{D}[h] e^{iS_{tot}[x,h]}$$

$$S_{tot}[x, h] \rightarrow S_{tot}[x_1, h_1] - S_{tot}[x_2, h_2]$$

$$x_- = x_1 - x_2$$

$$x_+ = \frac{1}{2}(x_1 + x_2)$$

$$\langle h_- h_+ \rangle = G_{ret}$$

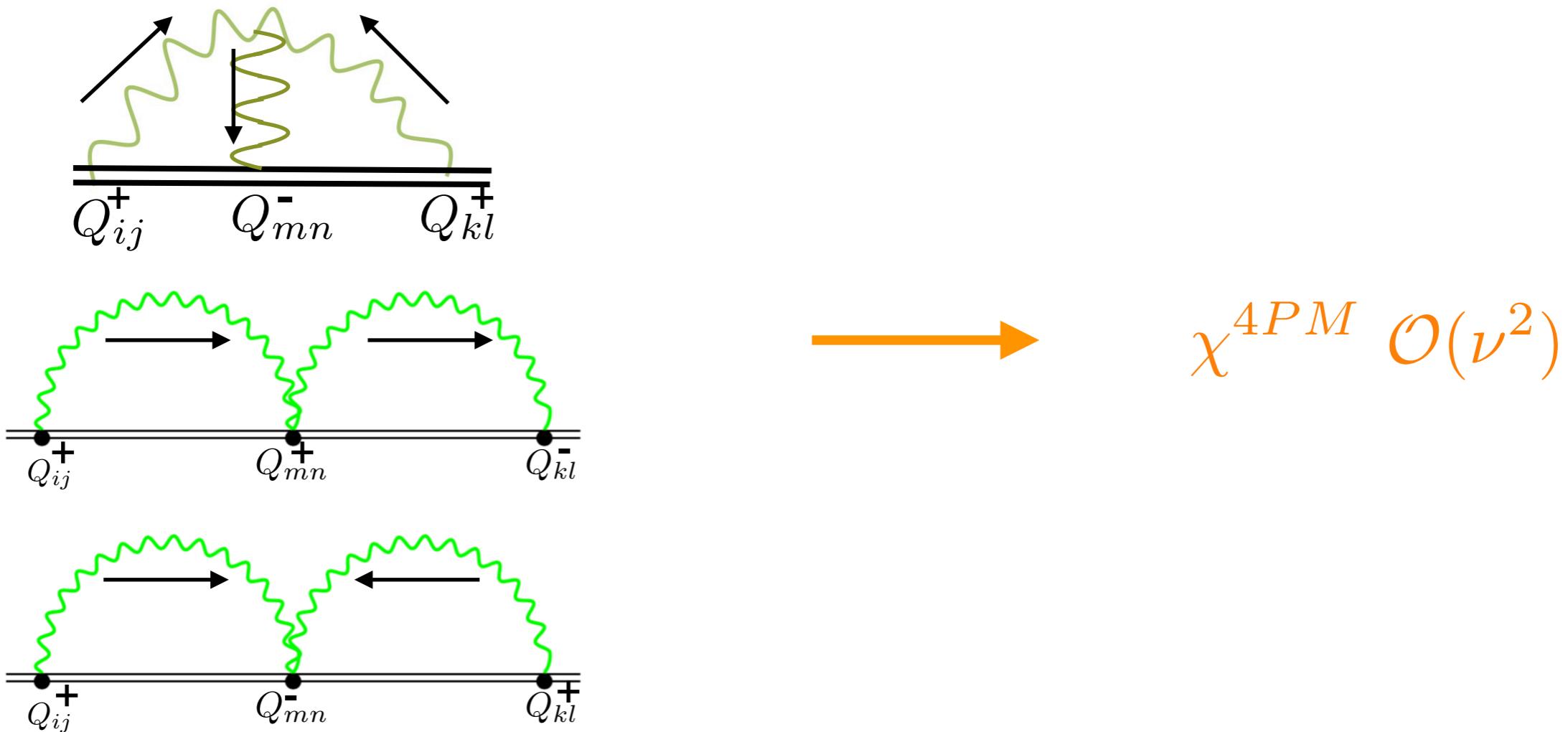
$$\left. \frac{\delta S_{eff}[x_\pm]}{\delta x_-} \right|_{x_- = 0} = 0$$

if $S_{eff}[x_\pm] = S_{eff}[x_1] - S_{eff}[x_2]$

the dynamics is conservative
and one can avoid using Keldish variables

Memory and double emission

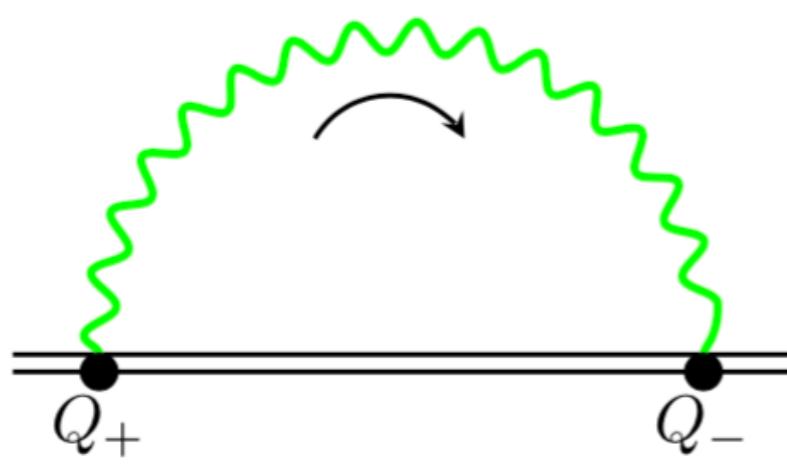
5PN, finite and local-in-time



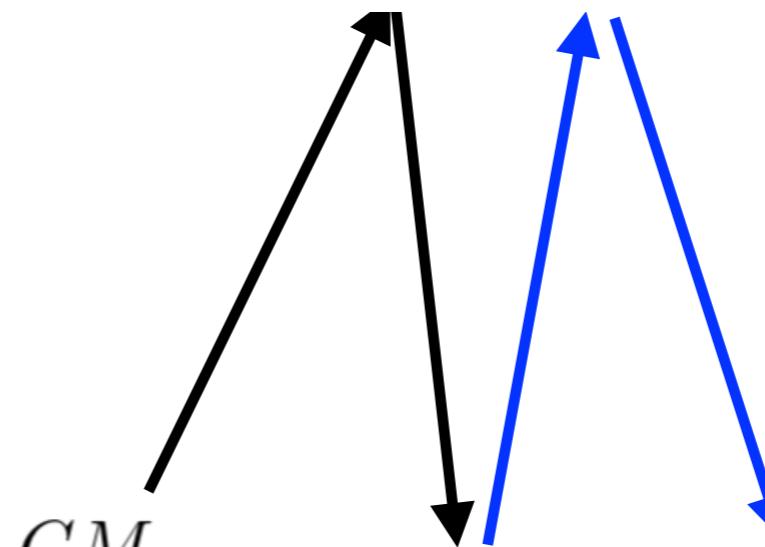
To be recomputed taking into account of
Ward correction

Simple self-energy diagram

at 2.5PN does not contribute to the conservative dynamics, gives only dissipative effects



$$\mathbf{a}_{rr}^i \simeq \mathbf{a}_{BT}^i = -\frac{2}{5}GQ_{ij}^{(5)}x^j$$



**Burke
Thorne
2.5PN**

5PN

But
 (radiation reaction)²

$$\mathbf{a} = -\frac{GM}{r^3}\mathbf{r} + \dots + \mathbf{a}_{rr} + \dots + \mathbf{a}_{BT^2}$$

N

affects 5PN

$$\mathbf{a}_{BT^2} = \frac{G^4 M^4 \nu^2}{r^6} \left[(c_1 v^4 + c_2 v^2 v_n^2 + c_3 v_n^4) \mathbf{r} + (c_4 v^2 + c_5 v_n^2) v_r \mathbf{v} \right] + \mathcal{O}(G^5)$$

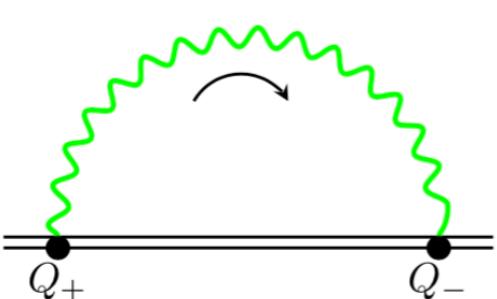
Conservative and dissipative terms

balance equation

$$\dot{E} = \mu \mathbf{a} \cdot \mathbf{v}$$

ambiguity due to Schott terms

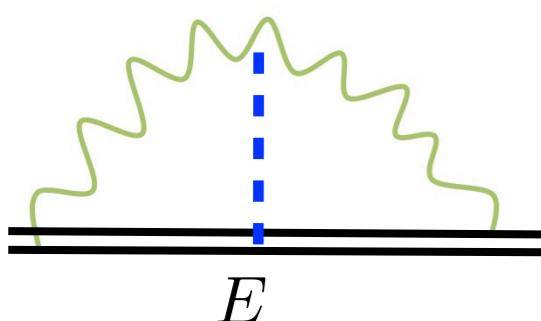
e.g.



$$\dot{E} = \mu \mathbf{a}^{BT} \cdot \mathbf{v} \simeq -\frac{G}{5} Q_{ij}^{(5)} \dot{Q}_{ij} = -\frac{G}{5} \ddot{Q}_{ij}^2 + \dot{V}_{Schott}$$

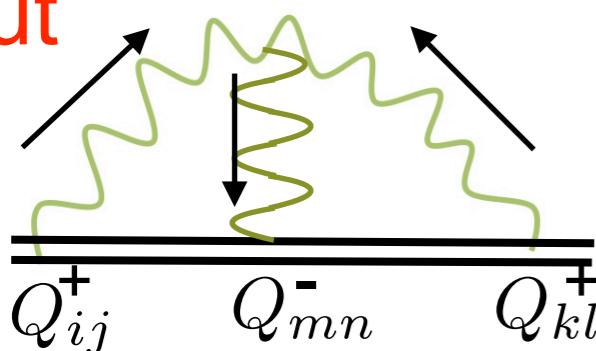
$$V_{Schott} = -\frac{G}{5} \left[Q_{ij}^{(4)} \dot{Q}_{ij} - Q_{ij}^{(3)} \ddot{Q}_{ij} \right]$$

half-integer PN, do not get mixed with conservative part



dissipative term is nonlocal, still do not get mixed

But



5PN local dissipative, mixes with conservative potential contributions

and the same for

$$\mathbf{a}_{BT^2}$$

5PN dynamics missing elements

Tuttifrutti: two 2FS coefficient and dissipative?

PM: 5PM and 6PM but no new 5-loop integrals at 5PN and memory?

EFT: memory scattering angle test

- New features:
- memory, nonlinear radiation reaction
 - mixing of conservative and dissipative terms

What to expect at 6PN?

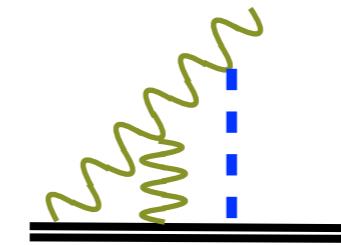
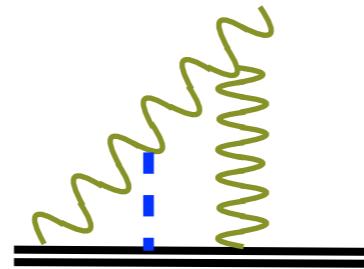
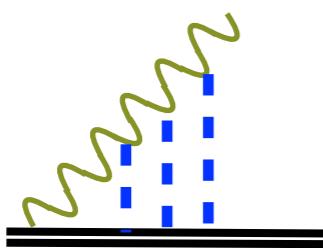
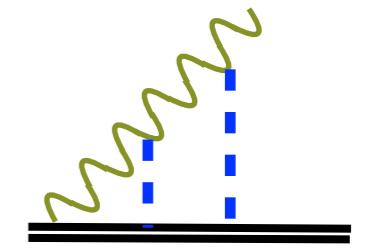
partial Tuttifrutti results
potential done with both EFT and PM up to G^4
5 and 6-loop integrals at $\mathcal{O}(G^6)$

GW radiation

Blanchet et al.

done in MPM formalism up to 4.5PN

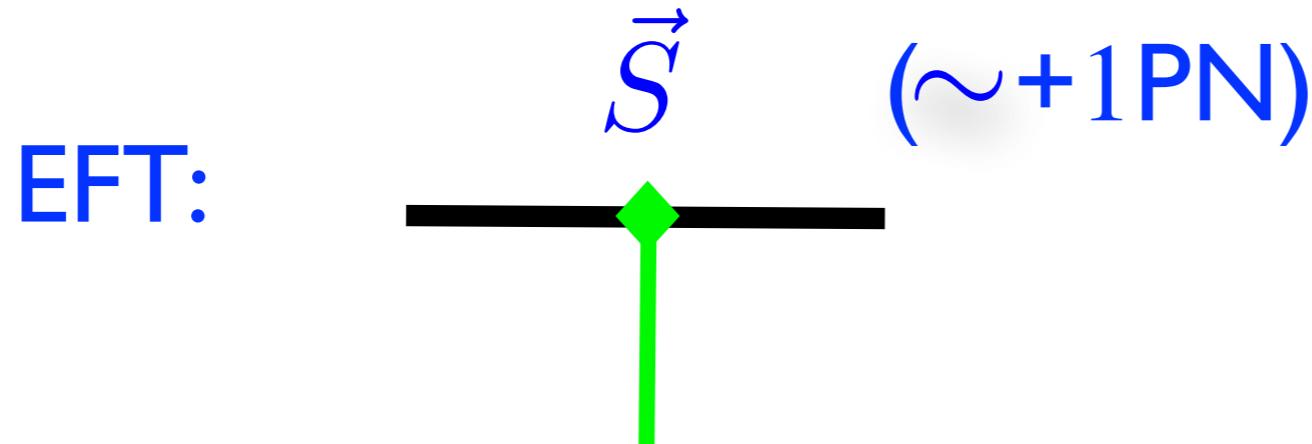
- General solution in far zone (**radiative moments**)
- Matching with near zone (**source moments**)
- Knowledge of 4PN source dynamics
- IR and UV divergencies involved at 3PN (tails-of-tails)
- Subtle gauge issues
- Tail-of-tail-of-tail, memory-of-tail and tail-of memory



GW phase known at 4.5PN
in principle extendable up to 5PN or even 5.5PN

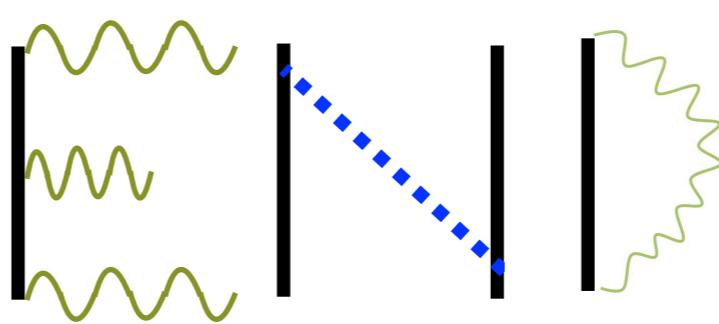
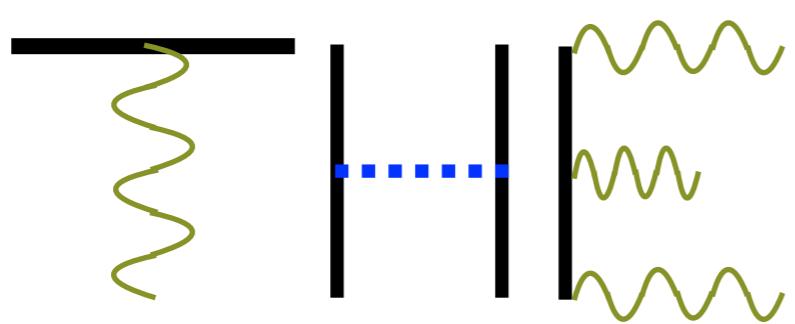
demonstrative 2PN calculation in EFT

Spin

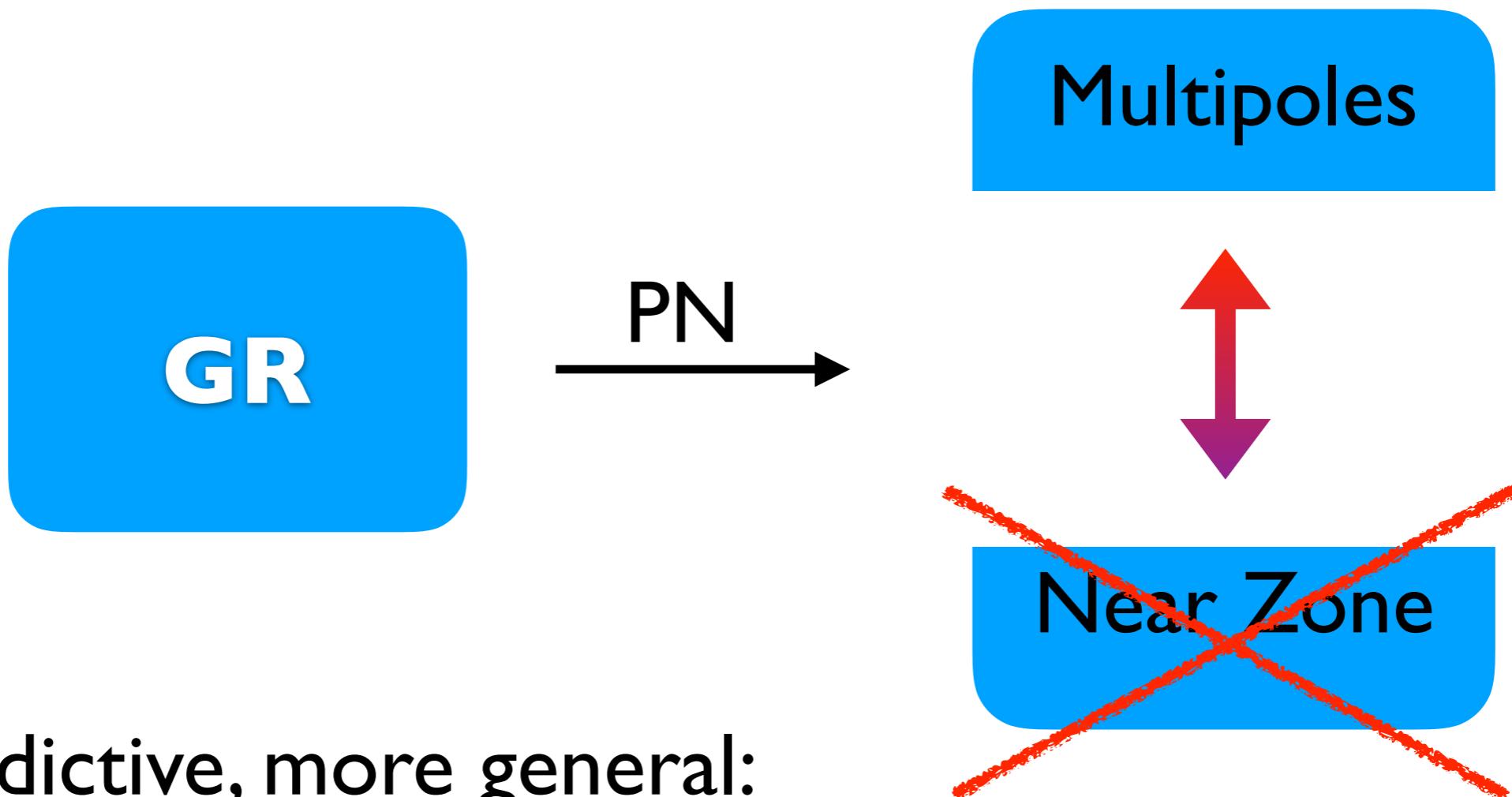


Φ_{GW}
aligned

	N	1PN	1.5PN	2PN	2.5PN	3PN	3.5PN	4PN	4.5PN	5PN
S^0	✓	✓	✓	✓	✓	✓	✓	✓	✓	??
S		✓	0	✓	✓	✓	✓	?	??	?
SS			✓	0	✓	✓	?	?	?	??
S^3						✓	0	?	?	?
S^4							?	0	?	?

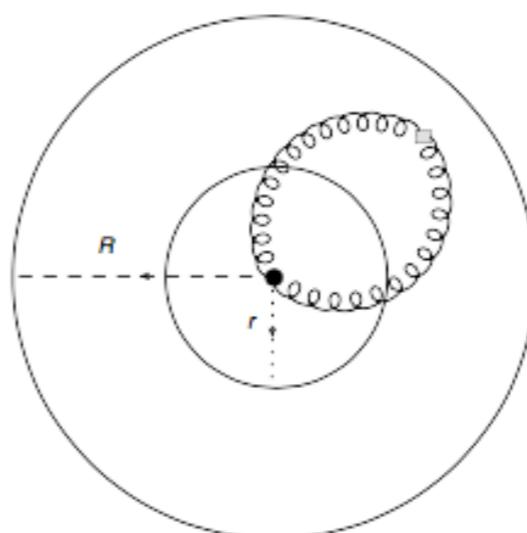


Renormalisation group analysis

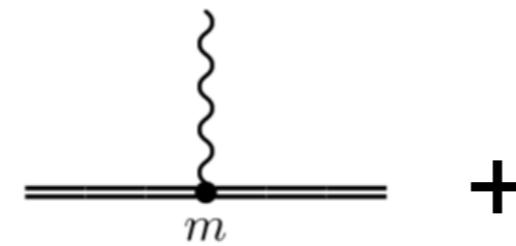
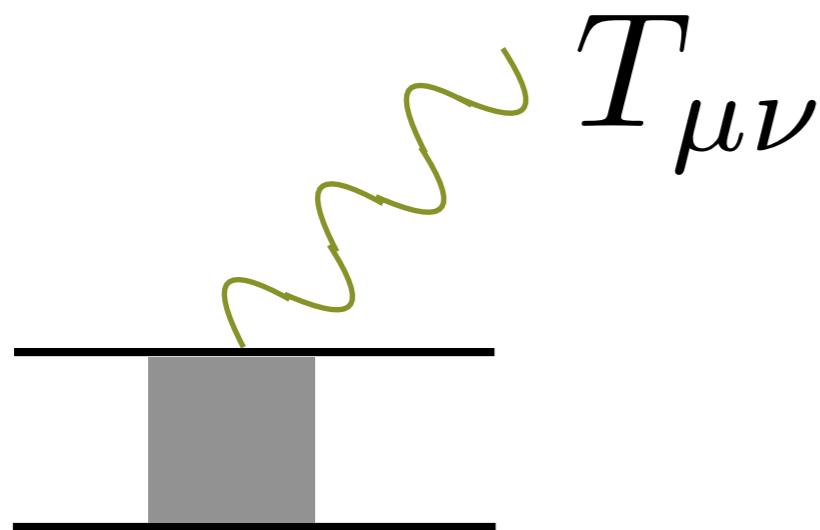


less predictive, more general:

UV divergencies \longrightarrow renormalization \longrightarrow universal IR log's

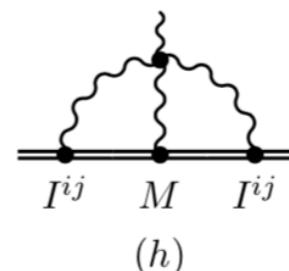
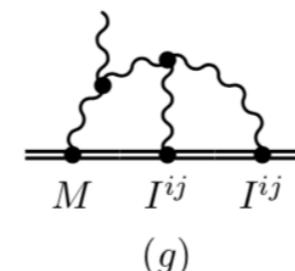
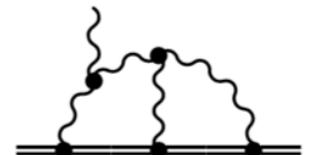
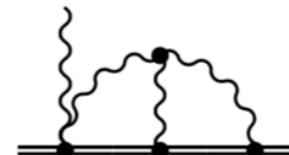
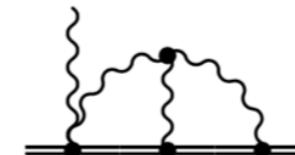


Leading logs from renormalisation group analysis at arbitrary high PN order



+

Goldberger, Ross, Rotstein



$$E^{bare} = Z(\mu) E^{ren}(\mu)$$

$$\frac{d \log E(\mu)}{d \log \mu} = -\frac{2G^2}{5} \left[2Q_{ij}^{(1)} Q_{ij}^{(5)} - 2Q_{ij}^{(2)} Q_{ij}^{(4)} + Q_{ij}^{(3)} Q_{ij}^{(3)} \right] M$$

$$E_{\text{leading-}(\log)^n} = -\frac{M\nu x}{2} \left[\frac{64\nu}{15} \sum_{n=1}^{+\infty} \frac{6n+1}{n!} (4\beta_Q)^{n-1} x^{3n+1} (\log x)^n \right]$$

$$\beta_Q = -\frac{214}{105}$$

Perfect agreement

Kavanagh et al.

self-force $m_2 \ll m_1$

vs

PN EFT

+

RG flow

+

1st law

$$c_{\log^3}^{10PN} = \frac{445503488}{496125}$$

$$c_{\log^4}^{13PN} = -\frac{5017776128}{2083725}$$

$$c_{\log^5}^{16PN} = \frac{133151707332608}{27348890625}$$

$$c_{\log^6}^{19PN} = -\frac{68019046365134848}{8614900546875}$$

$$c_{\log^7}^{22PN} = \frac{67666082665077932032}{6331951901953125}$$

$$\beta_Q = -\frac{214}{105}$$

$$\frac{64}{15} \frac{6n+1}{n!} (4\beta_Q)^{n-1}$$

$$n = 3 \dots 7$$