## Critical dynamics of natural swarms

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UNLP







### Natural swarms: experiments Statics Dynamics

### Theory

Field theories with alignment Field theory with alignment and activity The missing ingredient: inertia

## Swarms of midges

Diptera: Chironomidae and Diptera: Ceratopogonidae



[Attanasi et al. PLOS Comp. Biology 2014]

## Space correlations

Swarms are not a gas of insects



[Attanasi et al. PLOS Comp. Biology 2014]

▶ 
$$r_0 \sim 20$$
-30 cm,  $r_1 \sim 5$  cm,  $r_c \sim 2$ -5 cm, swarm size  $\sim 1$  m  
▶  $r_0$  is not quite  $\xi$   
▶ ls  $r_0$  large?

## Space correlations are scale-free

When looking at swarms of different size,  $r_0$  grows as L



Midges

# Space-time correlations



$$\begin{split} C(r,t) &= \langle \delta \boldsymbol{v}(r_0,t_0) \cdot \delta \boldsymbol{v}(r_0+r,t_0+t) \rangle \\ &= \left\langle \sum_{i,j}^N \frac{\delta \boldsymbol{v}_i(t_0) \cdot \delta \boldsymbol{v}_j(t_0+t) \delta \left[r - r_{ij}(t_0,t)\right]}{\sum_{k,l} \delta \left[r - r_{kl}(t_0,t)\right]} \right\rangle_{t_0}, \\ r_{ij}(t_0,t) &= |\boldsymbol{x}_i(t_0) - \boldsymbol{x}_j(t_0+t)| \end{split}$$

## Dynamic scaling Critical slowing down



Swarms obey scaling hypothesis τ<sub>k</sub> = k<sup>-z</sup>g(kξ) = ξ<sup>z</sup>(ξk)<sup>-z</sup>g(kξ)
 z ≈ 1.37 ± 0.11 ⇒ unknown universality class

## Underdamped relaxation



$$f(x) = \left. -\frac{1}{x} \log C(x) \right|_{x=t/\tau}$$



### Natural swarms: experiments

Statics Dynamics

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# A theory of natural swarms

Must account for:

- Scale-free correlations (static scaling)
- $\blacktriangleright$  Critical slowing down (dynamic scaling) with  $z pprox 1.37 \pm 0.11$
- Underdamped dynamics (damped oscillations)

Required ingredients:

- 1. Ferromagnetism (alignment)
- 2. Activity
- 3. Inertia

## Ferromagnetism

classical ferromagnet

$$\frac{d\boldsymbol{\sigma}_i}{dt} = J \sum_{ij} n_{ij} \boldsymbol{\sigma}_j + \xi_i, \qquad |\boldsymbol{\sigma}| = 1, \quad \xi: \text{ Gaussian noise}$$
$$\downarrow \text{ coarse-graining}$$

model A:

$$\begin{split} &\frac{\partial \psi}{\partial t} = -\Gamma \frac{\delta \mathcal{H}}{\delta \psi(x,t)} + \xi(x,t), \\ &\mathcal{H} = \int \! d^d x \, \left\{ (\nabla \psi)^2 + r \psi^2 + u \psi^4 \right\}, \qquad \text{(Landau-Guinzburg)} \end{split}$$



$$z\approx 2 ~ \mathrm{vs} ~ z_{\mathrm{exp}} = 1.37 \pm 0.11$$

[Hohenberg, Halperin, Ma 1972]

## Activity Vicsek: active ferromagnet

$$\begin{aligned} \frac{d\boldsymbol{\sigma}_i}{dt} &= J \sum_{ij} n_{ij} (\{\boldsymbol{x}\}) \boldsymbol{\sigma}_j + \xi_i, \\ \frac{d\boldsymbol{x}_i}{dt} &= v_0 \boldsymbol{\sigma}_i. \\ & \downarrow \text{ coarse-graining} \end{aligned}$$

Toner-Tu (1998) field theory (active model A)

$$D_t \boldsymbol{v} = -\Gamma \frac{\delta \mathcal{H}}{\delta \psi(x,t)} - \nabla P + \xi(x,t),$$
$$\partial_t \rho + \nabla(\rho \boldsymbol{v}) = 0$$

Material derivative  $D_t = \partial_t + \lambda(\boldsymbol{v}\cdot 
abla).$ 

# Incompressible Toner-Tu continuous transition

The field-theory version of the (incompressible) Vicsek model

$$\frac{\partial \boldsymbol{v}}{\partial t} + \lambda(\boldsymbol{v} \cdot \nabla)\boldsymbol{v} = \Gamma \nabla^2 \boldsymbol{v} + \Gamma r \boldsymbol{v} + \Gamma u v^2 \boldsymbol{v} + \boldsymbol{\nabla} P + \boldsymbol{\xi},$$

noise: 
$$\langle \xi_{\alpha}(\boldsymbol{x},t)\xi_{\beta}(\boldsymbol{x}',t')\rangle = 2\tilde{\Gamma}\delta(\boldsymbol{x}-\boldsymbol{x}')\delta(t-t')\delta_{\alpha\beta}$$

Incompressibility condition  $\nabla \cdot v = 0$  fixes pressure P.

[Chen, Toner, Lee, New J Phys 2015]

# Incompressible Toner-Tu continuous transition

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Hohenberg-Halperin model A plus activity

# Incompressible Toner-Tu

The field-theory version of the (incompressible) Vicsek model

$$\frac{\partial \boldsymbol{v}}{\partial t} + \lambda(\boldsymbol{v} \cdot \nabla)\boldsymbol{v} = \Gamma \nabla^2 \boldsymbol{v} + \Gamma \boldsymbol{r} \boldsymbol{v} + \Gamma \boldsymbol{u} \boldsymbol{v}^2 \boldsymbol{v} + \boldsymbol{\nabla} P + \boldsymbol{\xi},$$
  
noise:  $\langle \xi_{\alpha}(\boldsymbol{x}, t) \xi_{\beta}(\boldsymbol{x}', t') \rangle = 2\tilde{\Gamma} \delta(\boldsymbol{x} - \boldsymbol{x}') \delta(t - t') \delta_{\alpha\beta}$ 

Incompressibility condition  $\boldsymbol{\nabla} \cdot \boldsymbol{v} = 0$  fixes pressure P.

stirred fluid [Forster, Nelson and Stephen PRA 1977] plus alignment

# Incompressible Toner-Tu continuous transition

The field-theory version of the (incompressible) Vicsek model

$$\frac{\partial \boldsymbol{v}}{\partial t} + \lambda(\boldsymbol{v} \cdot \nabla)\boldsymbol{v} = \Gamma \nabla^2 \boldsymbol{v} + \Gamma r \boldsymbol{v} + \Gamma u v^2 \boldsymbol{v} + \boldsymbol{\nabla} P + \boldsymbol{\xi},$$

noise: 
$$\langle \xi_{\alpha}(\boldsymbol{x},t)\xi_{\beta}(\boldsymbol{x}',t')\rangle = 2\tilde{\Gamma}\delta(\boldsymbol{x}-\boldsymbol{x}')\delta(t-t')\delta_{\alpha\beta}$$

Incompressibility condition  $\nabla \cdot v = 0$  fixes pressure P.

Note  $\lambda \neq 1$ ,  $\tilde{\Gamma} \neq \Gamma$ 





BUT 
$$z_{exp} = 1.37 \pm 0.11$$



activity: 
$$\alpha = \lambda^2 \left( \frac{\tilde{\Gamma}}{\Gamma^3} \right) \Lambda^{-\epsilon}$$
, alignment:  $\hat{u} = u \left( \frac{\tilde{\Gamma}}{\Gamma} \right) \Lambda^{-\epsilon}$ 

## Dispersion relations Clue to missing inertia



 $\omega \sim ck + iDk^2, \qquad {\rm vs} \qquad \omega \sim iDk^2$ 

### Theory of dynamic critical phenomena

#### P. C. Hohenberg

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#### B. I. Halperin\*

theory with purely

irreversible

 $\cdot \omega = iDk^2$ 

terms

 $\cdot z \approx 2$ 

Department of Physics, Harvard University, Cambridge, Mass. 02138

An introductory review of the central ideas in the modern theory of dynamic critical phenomena is followed by a more detailed account of creent developmens in the field. The concepts of the conventional theory, mode-coupling, scaling, universality, and the renormalization group are introduced and are illustrated in the context of a simple example—the phase separation of a symmetric binary fluid. The renormalization group is then developed in some detail, and applied to a variety of systems. The main renormalization group theories uscessfully acplain available experimental data at the critical phenot pure fluids, and binary mixtures, and at many magnetic phase transitions, but that a number of discrepancies exist with data at the superfluid transition of 'He.

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#### Reviews of Modern Physics, July 1977

#### •••

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theories with reversible mode-coupling terms  $\cdot z = 1.5$  $\cdot \omega = iDk^2 \pm ck$ 

## Inertia Model A vs model G

$$\frac{\partial \boldsymbol{v}}{\partial t} = -\Gamma \frac{\delta \mathcal{H}}{\delta \boldsymbol{v}} + \boldsymbol{\xi}(\boldsymbol{x}, t), \qquad \qquad \frac{\partial \boldsymbol{v}}{\partial t} = +g\boldsymbol{v} \times \frac{\delta \mathcal{H}}{\delta \boldsymbol{s}} - \Gamma \frac{\delta \mathcal{H}}{\delta \boldsymbol{v}} + \boldsymbol{\xi}_{\boldsymbol{v}}(\boldsymbol{x}, t), \\ \frac{\partial \boldsymbol{s}}{\partial t} = -g\boldsymbol{v} \times \frac{\delta \mathcal{H}}{\delta \boldsymbol{v}} - \Lambda \frac{\delta \mathcal{H}}{\delta \boldsymbol{s}} + \boldsymbol{\xi}_{\boldsymbol{s}}(\boldsymbol{x}, t),$$

Reversible dynamics generated by Poisson brackets

$$\{s_{\mu}, v_{\nu}\} = g\epsilon_{\mu\nu\gamma}v_{\gamma}, \qquad \{s_{\mu}, s_{\nu}\} = g\epsilon_{\mu\nu\gamma}w_{\gamma},$$

The spin s(x,t) is the generator of the rotations of the primary field v(x,t). Spin is conserved  $\{s, \mathcal{H}\} = 0$ ,  $\Lambda = \lambda \nabla^2$ .

Model G has z = 3/2 and  $\omega = iDk^2 \pm ck$ .

## Active model ${\sf G}$

or coarse-grained inertial spin model (ISM)

Add activity to model G:

$$\begin{split} &\frac{\partial \boldsymbol{v}}{\partial t} = +g\boldsymbol{v}\times\frac{\delta\mathcal{H}}{\delta\boldsymbol{s}} - \Gamma\frac{\delta\mathcal{H}}{\delta\boldsymbol{v}} + \boldsymbol{\xi}_{v}(\boldsymbol{x},t),\\ &\frac{\partial\boldsymbol{s}}{\partial t} = -g\boldsymbol{v}\times\frac{\delta\mathcal{H}}{\delta\boldsymbol{v}} - \Lambda\frac{\delta\mathcal{H}}{\delta\boldsymbol{s}} + \boldsymbol{\xi}_{s}(\boldsymbol{x},t), \end{split}$$

$$\partial_t \boldsymbol{v} o D_t \boldsymbol{v} = \partial_t \boldsymbol{v} + \gamma_v (\boldsymbol{v} \cdot \nabla) \boldsymbol{v},$$
  
 $\partial_t \boldsymbol{s} o D_t \boldsymbol{s} = \partial_t \boldsymbol{s} + \gamma_v (\boldsymbol{v} \cdot \nabla) \boldsymbol{s},$ 

Our theory:

$$\begin{split} D_t \boldsymbol{v}(\boldsymbol{x}, t) &= +g \boldsymbol{v} \times \frac{\delta \mathcal{H}}{\delta \boldsymbol{s}} - \Gamma \frac{\delta \mathcal{H}}{\delta \boldsymbol{v}} - \nabla P + \boldsymbol{\xi}_v(\boldsymbol{x}, t), \\ D_t \boldsymbol{s}(\boldsymbol{s}, t) &= -g \boldsymbol{v} \times \frac{\delta \mathcal{H}}{\delta \boldsymbol{v}} - \Lambda \frac{\delta \mathcal{H}}{\delta \boldsymbol{s}} + \boldsymbol{\xi}_s(\boldsymbol{x}, t), \\ \nabla \cdot \boldsymbol{v} &= 0 \\ \text{(Small) dissipation } \Lambda &= \eta + \lambda \nabla^2 \end{split}$$

$$\mathcal{H} = \int d^d x \, \left[ \frac{1}{2} \, (\nabla \boldsymbol{v})^2 + \frac{1}{2} r \boldsymbol{v}^2 + \frac{u}{4!} \boldsymbol{v}^4 + \frac{1}{2} s^2 \right]$$

RG flow At 1 loop



## New dynamic critical exponent Alignment + activity + inertia



# Numerical simulations

Inertial spin model

$$\begin{aligned} \frac{d\boldsymbol{v}_i}{dt} &= \frac{1}{\chi} \boldsymbol{s}_i \times \boldsymbol{v}_i, \\ \frac{d\boldsymbol{s}_i}{dt} &= \boldsymbol{v}_i \times \sum_j n_{ij}(t) \boldsymbol{v}_j - \frac{\eta}{\chi} \boldsymbol{s}_i + \boldsymbol{v}_i \times \boldsymbol{\xi}_i, \\ \frac{d\boldsymbol{x}_i}{dt} &= \boldsymbol{v}_i. \end{aligned}$$







## Conclusions

- Midge swarms live near a critical point:
  - velocity correlation length scales with size,
  - velocities are disordered.
- Swarms obey dynamic scaling: the correlation length rules dynamic behavior.
- $\blacktriangleright$  The dynamic critical exponent z is different from known dynamic models.
- ► Values of z:

 $z_{sim} = 1.35 \pm 0.04$ 

## Collaborators

Istituto Sistemi Complessi and/or Sapienza University (Roma):

- Andrea Cavagna (ISC Roma)
- Irene Giardina (Sapienza Roma)
- Stefania Melillo (ISC Roma)
- Leonardo Parisi (ISC Roma)
- ► Luca Di Carlo (Sapienza Roma → Princeton)
- ▶ Giulia Pisegna (Sapienza Roma → Göttingen)
- Mattia Scandolo (Sapienza Roma)

Natural swarms in 3.99 dimensions, Nature Physics 2023