

D) TWO INTRIGUING COINCIDENCES

- The seesaw scales needed for Leptogenesis versus the seesaw scales needed for ν masses.

↳ for a hierarchical spectrum of N_i : $M_{N_1} \ll M_{N_{2,3}}$

$$\begin{aligned} E_{N_1} &= -\frac{3}{16\pi} \sum_{B=2,3} \frac{M_{N_1}}{M_{Nb}} \frac{\text{Im}[(Y_\nu Y_\nu^*)_{1B}]^2}{|Y_\nu Y_\nu^*|_{111}} \quad \left(M_\nu = \frac{n^2}{2} Y_\nu^T (M_N)^{-1} Y_\nu \right) \\ &= -\frac{3}{8\pi} \frac{M_{N_1}}{n^2} \frac{\text{Im}[(Y_\nu M_\nu^* Y_\nu)_1]}{|Y_\nu Y_\nu^*|_{111}} \quad \left(Y_\nu = \frac{\sqrt{2}}{n} (M_N)^{1/2} R (M_\nu^{\text{diag}})^{1/2} \right) \\ &= -\frac{3}{16\pi} \frac{M_{N_1}}{n^2} \frac{\sum m_{\nu_j}^2 \text{Im}(R_{1j}^2)}{\sum m_{\nu_i} |R_{1i}|^2} \quad \left(R_{1j} = \frac{V_{PMNS}^+}{V_{PMNS}} \right) \end{aligned}$$

$$\Rightarrow E_{N_1} \leq \frac{3}{16\pi} \frac{M_{N_1}}{n^2} (m_{\nu_3} - m_{\nu_1}) \leq \frac{3}{16\pi} \frac{M_{N_1}}{n^2} \underbrace{\sqrt{\Delta m_{13}^2}}_{0.05 \text{ eV}}$$

$$\Rightarrow \text{requiring } \frac{m_B}{B} = \frac{28}{79} \cdot E_{N_1} \left(\frac{M_{N_1}}{n} \right) \underset{T \gg m_{N_1}}{\underset{\leq 1}{\text{}}} \cdot \mathcal{N} \approx 9 \cdot 10^{-21}$$

$$\Rightarrow \underline{M_{N_1} > 4 \cdot 10^8 \text{ GeV}} \quad \text{Davidson-Ibarra 02.}$$

↳ this value (obtained for $\text{Im}(R_{1j}^2) \sim 1/R_{1j}^2 \sim 1$) or higher values of M_{N_1} (for smaller R_{1j}) fit very well with seesaw scales needed for ν masses (in particular scales where seesaw mass suppression of M_ν is fully operative)

↳ N.B.: for $M_{N_{2,3}} \sim M_{N_1}$, for example one can have successful leptogenesis for much smaller values of M_{N_1} (for large entries of R)

T.H. et al 03

- The v mass scale value versus n and m_{Planck} scales coincidence :

$$K_N \equiv \frac{\Gamma_{N_1}^{\text{TOT}}}{H(T=M_{N_1})} \geq \frac{m_\nu^{\text{MIN}}}{10^{-3} \text{eV}} \quad \epsilon \sim 10^{-3} \text{eV} = \frac{17.8 \pi n^2}{m_{\text{Planck}}} \\ \uparrow \quad \uparrow \\ \propto \frac{1}{M_{N_1}^2} \frac{Y_{N_1}^T M_{N_1} Y_\nu}{m_{\text{Planck}}} \quad \propto Y_\nu^T M_{N_1}^{-1} Y_\nu.$$

\Rightarrow given the $m_\nu \leq 1.7 \text{eV}$ bound from KATRIN this means the washout is naturally limited:

- $\Gamma_{N_1}/H(T=M_{N_1}) \leq 1$ cannot be much violated
 - \hookrightarrow real coincidence because 10^{-3}eV scale is determined by $n = 246 \text{ GeV}$ and $m_{\text{Planck}} = 10^{18} \text{GeV}$
 - $\hookrightarrow m_\nu \sim \text{keV}$ for example would kill Leptogenesis

- In fact for $M_{N_1} \ll M_{N_2,3}$ successful Leptogenesis requires $m_\nu \leq 0.12 \text{eV}$ & in this case $m_{\nu_1} \sim m_{\nu_2} \sim m_{\nu_3} \sim m_\nu$
 - \hookrightarrow due to 2 effects :

- $\bullet \quad \epsilon_{N_1} < \frac{3}{4\pi} \frac{M_{N_1}}{n^2} (m_{\nu_3} - m_{\nu_1}) \downarrow \text{when } m_\nu \uparrow$
 - $\hookrightarrow m_{\nu_3} - m_{\nu_1} = \frac{2.5 \cdot 10^{-3} \text{eV}^2}{2 m_\nu}$

- $\bullet \quad K_{N_1} \geq \frac{m_\nu^{\text{MIN}}}{10^{-3} \text{eV}} \uparrow \text{when } m_\nu \uparrow = m_\nu^{\text{MIN}} \sim m_\nu \text{ for } m_\nu > 0.05 \text{eV}$

- \hookrightarrow however : - to get a quasi-degenerate m_ν spectrum from hierarchical M_{N_1} spectrum implies tuning (cancellation in M_ν & $Y_\nu^T M_{N_1}^{-1} Y_\nu$)
 - \hookrightarrow more natural is to get it from quasi-

degenerate spectrum of $M_{N_1} \Rightarrow$ in this case no more relevant upper bound on m_N
 ↳ resonance in E_N : see below

T.H. et al 03

- flavor effect also relax this bound (see below)

\Rightarrow to sum up so far: the scales ($m_N, v, m_{\text{planet}}$) and the $\frac{m_B}{v}$ observed value match very well!!

↳ it is nevertheless difficult to make any precise prediction:

$$E_{N_1} = -\frac{3}{16\pi} \frac{M_{N_1}}{v^2} \frac{\sum m_{N_j}^2 \operatorname{Im}(R_{1j}^2)}{\sum m_{N_i} |R_{1i}|^2}$$

↳ depends on low energy m_N value but the phases leading to leptogenesis are the 3 "high-energy" ones

+ dependence on 3 real parameters in R which are also "high-energy"
 + dependence in unknown M_{N_1}

E) FLAVOR LEPTOGENESIS

↳ so far all results: counting the # of lepton created and destroyed independently of their e, μ or τ flavor

↳ justified for $T \gtrsim 10^{12} \text{ GeV}$; e^-, μ^-, τ^- indistinguishable in the thermal bath: same gauge inter. and SM charged inter. out of equilibr.

↳ \Rightarrow the N_1 couples to a single $\tilde{\ell} \propto Y_{N_1 e} e + Y_{N_1 \mu} \mu + Y_{N_1 \tau} \tau$ and this combination remains coherent afterwards: one counts # of $\tilde{\ell}$ created/destroyed \Rightarrow a single Boltzmann eqn.

↳ However: for $T \lesssim 10^{12} \text{ GeV}$: $\Gamma_\tau^{\text{SM}} > H \ll \text{SM } \tau \text{ Yukawa}$

↳ thermal bath distinguishes τ flavor from other flavor

\Rightarrow 2 Boltzm. eqns: \rightarrow for # of τ
 \downarrow
 \rightarrow for # of e, μ

↳ each one with its flavor asym.:

$$E_{N_d} = \frac{\Gamma(N \rightarrow \tilde{\ell}_d H) - \Gamma(N \rightarrow \overline{\tilde{\ell}_d} \bar{H})}{\Gamma_{N_d}^{\text{TOT}}} \quad d = \tau, e, \mu$$

for $T \lesssim 10^9 \text{ GeV}$: $\Gamma_\mu^{\text{SM}} > H$

↳ 3 Boltzm eqns: $\begin{cases} \uparrow \text{for # of } \tau \\ \rightarrow \text{for # of } \mu \\ \rightarrow \text{for # of } e \end{cases}$

N.B.: in fact more precisely the Y_τ Yukawa int. breaks the $\tilde{\ell}$ coherence if a τ undergoes a Y_τ interact. on average before it goes in an inverse decay $\tau + h \rightarrow N_1$

Blanchet, Dibani, Raffelt 06

Flavor typical effects:

- Reduction of washout effect: in strong washout case: $\Gamma_{N_1} \gg 1$

↳ in one flavor approx.: strong washout

↳ in several flavor case: possibility of

less washout: e.g. if $\Gamma(N_1 \rightarrow L_\mu H) \gg \Gamma(N \rightarrow L_\tau H)$: τ asym. less washed-out

↳ change result by factor of a few (typically)

↳ does not change on $M_{N_1} \gtrsim 4 \cdot 10^8$ GeV bound obtained anyway in no washout case.

- $N_{2,3}$ leptogenesis:

↳ in one flavor approx.: Υ created by N_2, N_3 decays washed out by N_1 interactions

↳ in several flavor case: not so true anymore if N_i 's couple to \neq flavors

- Initial condition dependence:

↳ in one flavor approx.: any preexisting τ asym would be easily washed out by N_i 's interactions

↳ in several flavor case: not so true anymore

- CP phases contributing to leptogenesis

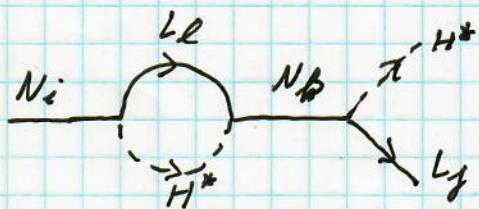
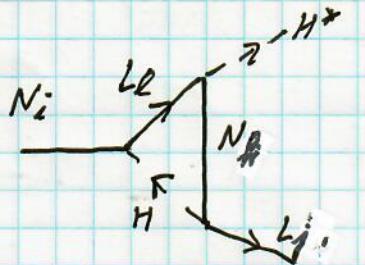
↳ in one flavor approx.: only the 3 high energy phases

↳ in several flavor case: the 3 low energy phases

$S_{CP} = \text{complicated}$ in M_Y , can also contribute \Rightarrow in general less low energy comb.
combination of 6 high predictive. Possibility to have leptogenesis with only energy phases

Dirac phase in M_Y, S_{CP} , but no reason to have only it at all

• Purely flavor neutral leptogenesis:



Anistizabal-Sierra,
Losada, Nandi 08
Gonzalez-Garcia,
Racker, Rivas 09
...

$$\Rightarrow \text{for each } L_j \text{ flavor: } \varepsilon_j^K = \frac{\Gamma(N \rightarrow L_j H^+) - \Gamma(N \rightarrow \bar{L}_j H)}{\Gamma_N^{\text{TOT}}}$$

\hookrightarrow in one flavor case: $\sum_j \varepsilon_j^K = 0 \Rightarrow \text{no leptogenesis}$

\hookrightarrow in several flavor case: $\varepsilon_2^K = -\varepsilon_{2+3}^K \neq 0$
 can be not washedout can be washedout $\Rightarrow \text{leptogenesis}$

\hookrightarrow generically subleading because suppressed by $\frac{m_{N_2}^2}{m_{N_{2,3}}^2}$

E) QUASI-DEGENERATE M_{N_i} SPECTRUM CASE: LOW SCALE LEPTOGENESIS

If at least 2 N_i have similar masses \Rightarrow resonance effect!

↳ for $M_{N_1} \sim M_{N_2}$:

Covi, Roulet,
Nissani 93
Flanz et al 93

$$\Rightarrow \mathcal{E}_{N_1} = \frac{1}{8\pi} \frac{\text{Im} [(Y_\nu Y_\nu^+)^2_{12}]}{(Y_\nu Y_\nu^+)^2_{11}} \cdot \frac{M_{N_1}}{M_{N_2}} \cdot S_{12}$$

$$S_{12} = \frac{M_{N_2}^2}{M_{N_1}^2 - M_{N_2}^2}$$

\Rightarrow need to include decay width:

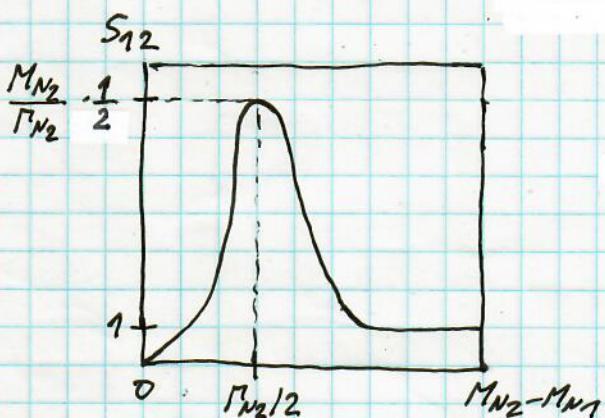
$$\left. \begin{array}{l} M^{(1)} M^{(0)*} \propto \frac{1}{M_{N_1}^2 - M_{N_2}^2 + i\Gamma_{N_2}^2 M_{N_1}} \\ M^{(0)} M^{(1)*} \propto \frac{1}{M_{N_2}^2 - M_{N_1}^2 + i\Gamma_{N_1}^2 M_{N_2}} \end{array} \right\} \Rightarrow S_{12} = \frac{M_{N_2}^2 (M_{N_1}^2 - M_{N_2}^2)}{(M_{N_2}^2 - M_{N_1}^2)^2 + \Gamma_{N_2}^2 M_{N_1}^2}$$

\Rightarrow if $M_{N_2} \gg M_{N_1}$: $S_{12} = 1$

if $M_{N_2} - M_{N_1} = \frac{\Gamma_{N_2}}{2}$ $\Rightarrow S_{12} = \frac{M_{N_2}}{\Gamma_{N_2}} \cdot \frac{1}{2} \gg 1$

if $M_{N_2} = M_{N_1} \Rightarrow S_{12} = 0$

huge resonance enhancement.



$\Rightarrow \bullet E_{N_1}$ not bounded anymore by $M_{N_1} \frac{(m_{\nu_3} - m_{\nu_1})}{v^2}$

$$\hookrightarrow E_{N_1} \leq \frac{1}{2} \quad E_{N_1} + E_{N_2} \leq 1$$

↑
at the peak

Pilaftsis 97, ...

\Rightarrow no more bound on M_{N_1} from E_{N_1}

\Rightarrow "low scale leptogenesis"

$\hookrightarrow M_{N_1} \sim M_{N_2} \sim v_{EW}$ perfectly possible

$\hookrightarrow N$ decay to LH only requires

$$M_{N_1} > m_L + m_H \sim 100 \text{ GeV}$$

\bullet No more relevant upper bound on m_ν : $m_\nu \sim 1 \text{ eV}$ perfectly possible

\hookrightarrow if $m_{\nu_1} \sim m_{\nu_2} \sim m_{\nu_3} \sim 1 \text{ eV}$: $\frac{m_{\nu_3} - m_{\nu_1}}{m_{\nu_3}} \sim 10^{-3}$ perfectly possible

\hookrightarrow easier to explain if $\frac{M_{N_1} - M_{N_2}}{M_{N_1}} \lesssim 10^{-3}$ too \Rightarrow resonance!

\bullet A proper treatment of resonant case: quantum Boltzmann equations bring corrections

Buchmiller et al
De Simone, 00
Riotto 07
Garny et al 09

\bullet Tests of low scale leptogenesis:

\hookrightarrow if $M_{N_1} \sim \text{TeV}$ one could naively expect a production of N 's at colliders

\hookrightarrow however $m_\nu \sim Y_\nu^2 v^2 / M_N \sim 0.1 \text{ eV}$ requires typically $Y_\nu \sim 10^{-7} \Rightarrow$ too small to produce N 's at colliders (N is $SU(2)_L \times U(1)_Y$: no gauge interactions).

\hookrightarrow 2 ways out one could think of:

- production of N 's from other interactions

↳ for example: L-R models : $w_R \begin{matrix} \nearrow f_N \\ \searrow g_R \end{matrix} \leftrightarrow w_L \begin{matrix} \nearrow f_L \\ \searrow g_L \end{matrix}$

$\Rightarrow N$'s can be produced at LHC via $w_R \begin{matrix} \nearrow \bar{q} \\ \searrow q \end{matrix} \leftrightarrow w_L \begin{matrix} \nearrow f_N \\ \searrow \bar{f}_N \end{matrix}$

↳ however: any large interact^o of N such as g_R ones will deeply thermalize TeV scale N 's:

$$\dots \frac{dY_N}{dz} = \dots \gamma_D (1 - \frac{Y_N}{Y_N^{\text{eq}}}) + \dots \gamma_{w_R} (1 - \frac{Y_N}{Y_N^{\text{eq}}}) + \dots$$

↑
e.g. $g\bar{q} \leftrightarrow w_R \leftrightarrow N \bar{N}$

\Rightarrow since $\gamma_{w_R} \gg \gamma_D \Rightarrow Y_N - Y_N^{\text{eq}}$ much more suppressed than with γ_D term only

$$\Rightarrow \dots \frac{dY_L}{dz} = \dots \underbrace{\gamma_D (1 - \frac{Y_N}{Y_N^{\text{eq}}})}_{\text{very suppressed}} E_N + \dots$$

very suppressed

↳ \Rightarrow not much product^o of Y_L

↳ this is not a washout effect but a N thermalizat^o effect

↳ \Rightarrow successful leptogenesis requires

Friese, T.M., Verlangen
08'

$m_{w_R} > 18 \text{ TeV} \Rightarrow N$ production at LHC through w_R hopeless in this case

- approximate L conservation frameworks: if one assumes (just so for the sake of testability) that there exists an extra $U(1)_L$ global symmetry that is broken only by a small amount, one can get small ν masses even though some of the Yukawa's are large.

example: 2N's framework with $L(N_1) = 1, L(N_2) = -1$

$$\Rightarrow M = \begin{pmatrix} V_L & N_1^c & N_2^c \\ 0 & Y_{NN} & 0 \\ Y_{NN} & 0 & M_N \\ 0 & M_N & 0 \end{pmatrix} \Rightarrow M_\nu = 0$$

↑
L conserved

\Rightarrow adding & perturbations:

$$M = \begin{pmatrix} 0 & Y_{NN} & Y'_N n \\ Y_{NN} & \mu' & M_N \\ Y'_N n & M_N & \mu \end{pmatrix}$$

$$\Rightarrow M_\nu \sim Y'_N \frac{1}{M_N} Y_N n^2 - Y'_N \frac{1}{M_N} \mu \frac{1}{M_N} Y_N n^2$$

\Rightarrow testable at LHC for $M_N \sim 1\text{TeV}, Y_N \sim 10^{-n}, n \sim 0-2$

testable by experiments searching for rare charged lepton flavor violating processes such as $\mu \rightarrow e\gamma, \mu \rightarrow ee, \mu \rightarrow e$ conversion in atomic nuclei \leftarrow rates are $\propto \frac{1}{M_N^4} \leftarrow$ dim-6 operat.

for leptogenesis such frameworks e.g. imply a suppression

Drewes et al
P. Hernandez et al
Garnbrot et al
 $\vdots \dots$

leptogenesis testable to some extend \leftarrow of E_N CP-asymmetries which requires L violation $\Rightarrow \alpha$ to pushing all parameters in right direction. 42. small $Y'_N, \mu \dots$ which can be compensated by large quasi-degeneracy of $M_{N_1, 2}$: $M_{N_1} \sim M_{N_2}$ predicted here: $M_{N_1} \sim M_{N_2} \sim \mu, \mu'$

GI LEPTOGENESIS WITH VERY SMALL VALUES OF M_N

- ↳ if $M_N \ll m_h, m_W, m_Z \Rightarrow$ no more 2 body N decays
- ↳ but $N_i \leftrightarrow N_j$ SP transitions can still be operative together with NLH Yukawa interactions inducing various processes, $H \rightarrow NL +$ scatterings...
- ↳ Leptogenesis from N oscillations: Akmedhov, Rubakov, Smirnov
- ↳ purely flavoneutral framework occurring at T typically 10^5 GeV $\Rightarrow T \gg m_{N_i}$: finite temperature effect important.
- ↳ generation of $L_i - \bar{L}_i$ asym and $N_i - \bar{N}_i$ asym with $\sum_i (L_i - \bar{L}_i) = 0 = \sum_i (N_i - \bar{N}_i)$ until a flavor asym is more washed out than another one $\Rightarrow \sum_i (L_i - \bar{L}_i) = -\sum_i (N_i - \bar{N}_i) \neq 0$: total lepton asym \neq hot total SM lepton asym \neq 0 since this latter one is transferred into B via sphalerons (at $T > 150$ GeV) unlike the N asym \Rightarrow baryogenesis.
- ↳ doesn't require huge quasi-degeneracy of N 's but some degeneracy is still needed.

↳ there is also a $\propto \frac{1}{T}$ contribution which is suppressed by $\frac{M_N}{T^2}$ but involves 2 powers less of $T_N \Rightarrow$ can dominate easily for larger quasi-degeneracy of N and occurs just before sphaleron decoupling: $T \approx 200$ GeV

T.H. Teresi¹⁷
Klanic¹⁸
Shaposhnikov¹⁹ ↳ both pure flavor ABS and $\propto \frac{1}{T}$ contributions can be considered in a density matrix formalism as well as decay/ scattering Boltzmann formalism

III. LEPTOGENESIS IN ALTERNATIVE SCENARIOS

↳ here will talk mainly about frameworks involving type-II and type-III seesaw

Type-II leptogenesis:

↳ type-II requires only one $\Delta_L \Rightarrow$ no CP-asymmetry:

$$\begin{array}{c} \Delta_{L_1} \xrightarrow{H} \Delta_L \xrightarrow{H} \chi_L \\ \Delta_{L_2} \xrightarrow{H} \end{array} \Rightarrow \text{Im}[\dots] = 0$$

Type-III leptogenesis:

T.H., Lin, Notarni, Papucci,
Strumia 03

$$\begin{array}{c} \Sigma_i \xrightarrow{H} \Sigma_B \\ H \quad L \end{array} + \begin{array}{c} \Sigma_i \xrightarrow{H} \Sigma_B \\ H \quad L \end{array} \Rightarrow M_{\Sigma_1} > 1.5 \cdot 10^{10} \text{ GeV} \quad (M_{\Sigma_1} \ll M_{\Sigma_2, 3})$$

$$M_{\Sigma_1} > 1.6 \text{ TeV} \quad (M_{\Sigma_2} \sim M_{\Sigma_3})$$

Strumia 09

Type-I + Type II leptogenesis: T.H., G. Senjanovic, 03'

$$\begin{array}{c} N_1 \xrightarrow{\Delta_L} \chi_L \\ N_2 \xrightarrow{\Delta_L} - \\ N_3 \xrightarrow{\Delta_L} - \end{array} + \begin{array}{c} \Delta \xrightarrow{H} N_1 \\ H \quad L \end{array} \Rightarrow M_\Delta > 3 \cdot 10^9 \text{ GeV}$$

T.H., Raidal, Strumia 07

$$\Rightarrow M_{N_1} > 4 \cdot 10^8 \text{ GeV} \quad (M_{N_1} \ll M_{N_2, N_3})$$

↳ T.H., Senjanovic 03
Antusch, King 04

↳ this framework is very well motivated in SO(10)

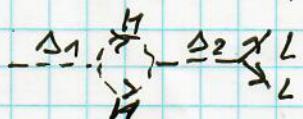
GUT: all fermion SM generators in a single SO(10) representation: 16 representations

↳ contains also a $N \Rightarrow$ type-I seesaw!

↳ M_N Majorana mass if generated from renormalizable coupling requires a 126 scalar representation:
this 126 scalar repres. contains a $\Delta_L \Rightarrow$ type-II seesaw!

Type-II + Type-II

TH, Raridash, Stremnia 07



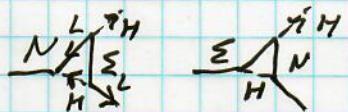
Mo, Sarker 98

$$\Rightarrow M_{\Delta_1} > 3 \cdot 10^{10} \text{ GeV} \quad (M_{\Delta_1} \ll M_{\Delta_2})$$

$$M_{\Delta_1} > 7.6 \text{ TeV} \quad (M_{\Delta_1} \approx M_{\Delta_2})$$

Stremnia 09

Type-II + Type III



: pure vertex

$$M_{\Delta_1} > 4 \cdot 10^{10} \text{ GeV} \quad (M_{\Delta_1} \ll M_{\Sigma})$$

$$\Rightarrow M_{\Sigma} > 7.5 \cdot 10^{10} \text{ GeV} \quad (M_{\Sigma} \ll M_W)$$

↳ natural in SU(5) GUT

Bajc et al 07'

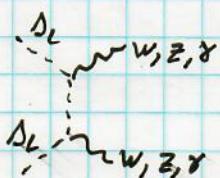
Fleuret-Perez et al 07'

with a single 24 fermion

repres. which contains αN and $\alpha \Sigma$

- From decay of a triplet (Δ_1 or Σ): lower bounds on decaying triplet mass above are somewhat higher than on M_{N_1} in type-I seesaw

↳ triplets unlike N 's have $SU(2)_L$ gauge interactions:



: fast interactions: $g \approx 1$

\Rightarrow suppresses $Y_\Delta - Y_\Delta^{\text{EQ}}$ in $\frac{dY_\Delta}{dz}$ Boltz. equ.

\Rightarrow suppresses Y_L asymmetry product

$$\text{in } \frac{dY_L}{dz} = \frac{Y_D(1 - Y_D)}{Y_D^{\text{EQ}}} E_\Delta + \dots$$

suppressed

↳ such a suppression occurs if gauge interact. are in thermal equil. $\frac{\gamma_W}{m_{\Delta_1}^{\text{EQ}} H} > 1 \iff T \lesssim 10^{14} \text{ GeV}$ and if $\gamma_W > Y_D \iff$ if a Δ_1 undergo on average a gauge inter. before undergoing a decay. \Rightarrow even on the resonance peak one still needs $M_{\Delta_1}, M_\Sigma > 1.6 \text{ TeV}$