# Towards an experimental test for a quantum extension of Einstein's Equivalence Principle





# Motivation

# Approaches to a quantum equivalence principle

# **Quantum Aspects of the Equivalence Principle**

On Gravity's Role in Quantum State Reduction

On the Equivalence Principle in Quantum Theory \*

Y. Aharonov

Roger Penrose<sup>1,2</sup>

Claus Lämmerzahl\*

Laboratoire de Gravitation et Cosmologie Relativiste, Université Pierre et Marie Curie,

CNRS/URA 679, F - 75252 Paris Cedex 05, France

Department of Physics, University of South Carolina, Columbia, South Carolina<sup>1</sup>

and G. Carmi<sup>2</sup>

Department of Physics, St. John's University, Jamaica, New York

#### Implementation of the Quantum Equivalence Principle

Lucien Hardy

Quantum mechanics and the equivalence principle

P.C.W. Davies

### Quantum formulation of the Einstein equivalence principle

Magdalena Zych <sup>™</sup> & Časlav Brukner

Can quantum probes satisfy the weak equivalence principle?

Luigi Seveso<sup>a</sup>, Matteo G. A. Paris<sup>a,b</sup>

Sagnac interferometer and the quantum nature of gravity

Chiara Marletto<sup>1,2,3</sup> and Vlatko Vedral<sup>1,2,3</sup>

Einstein's Equivalence principle for superpositions of gravitational fields and quantum reference frames

Flaminia Giacomini<sup>1,2,\*</sup> and Časlav Brukner<sup>3,4</sup>

A Simple Test of the Equivalence Principle(s) for Quantum Superpositions

Patrick J. Orlando, Robert B. Mann, Xavan Modi, and Felix A. Pollock 1, \*

Einstein's equivalence principle in quantum mechanics revisited

Michael Nauenberg

# Classical equivalence principles

# (At least) three different principles

Weak
Equivalence
Principle
(WEP)

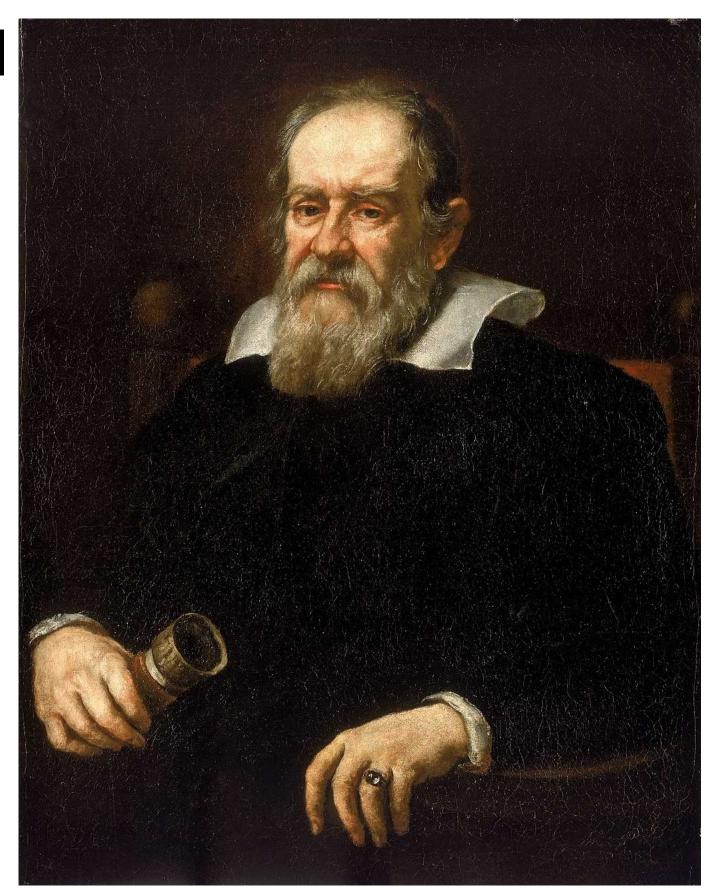
Einstein's
Equivalence
Principle
(EEP)

Strong
Equivalence
Principle
(SEP)

# Weak Equivalence Principle

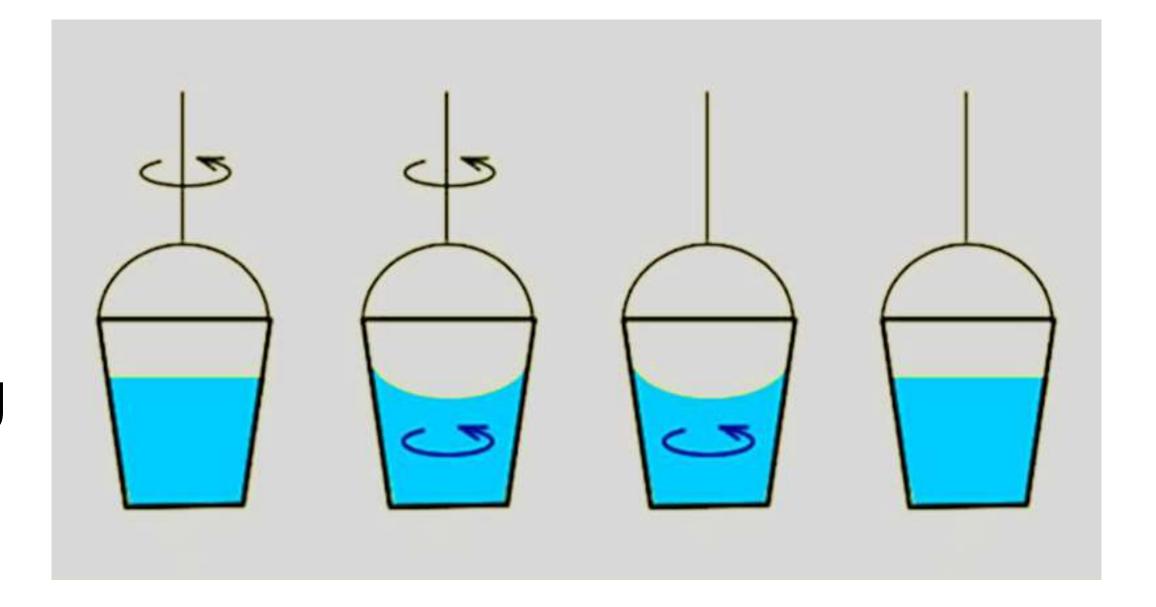
- The weak equivalence principle asserts the equivalence between inertial mass and gravitational mass: it's the universality of free fall.
- Alternatively, it can be formulated saying that a constant acceleration is undistinguishable from a constant gravitational field.

$$m_i a = m_g g$$
 $m_i = m_g$ 
 $a = g$ 



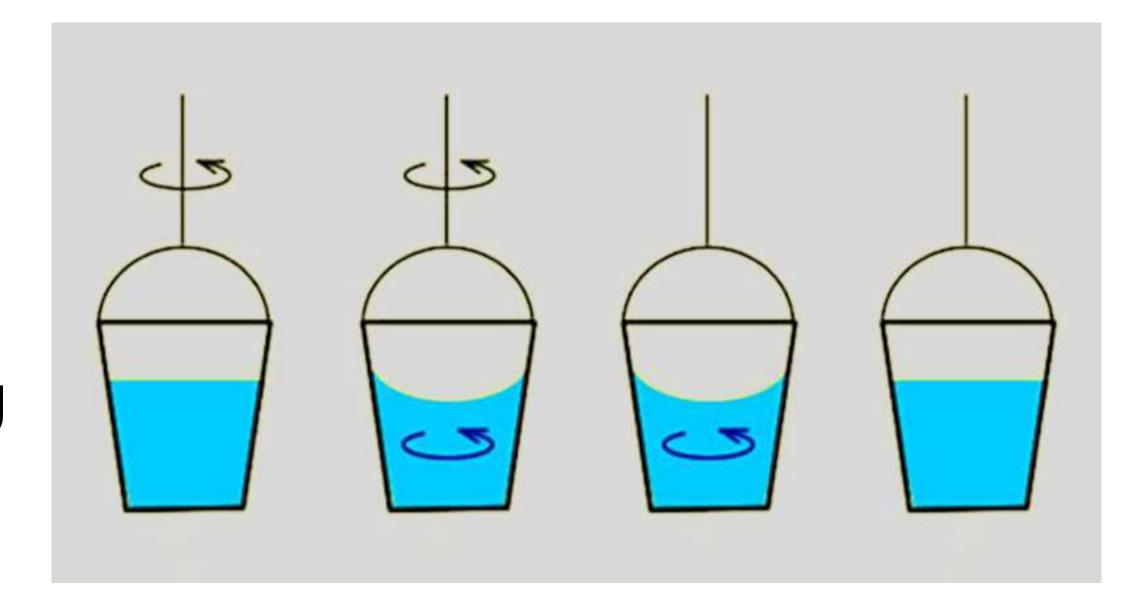
### Newton's bucket or the absoluteness of motion

- According to Newton, this experiment shows that motion (in particular, acceleration) is absolute.
- From this idea, he defines his notion of inertial frame: inertial frames are moving with constant velocity with respect to an absolute space



### Newton's bucket or the absoluteness of motion

- According to Newton, this experiment shows that motion (in particular, acceleration) is absolute.
- From this idea, he defines his notion of inertial frame: inertial frames are moving with constant velocity with respect to an absolute space



### Ernst Mach (1983):

When, accordingly, we say that a body preserves unchanged its direction and velocity in space, our assertion is nothing more or less than an abbreviated reference to the entire universe.

- Einstein understands that inertia and the gravitational field are the same object: it is the gravitational field that defines what frame is inertial, hence the gravitational field is not a force, but provides the structure of spacetime.
- The inertial frames are the freely falling frames. This is the core of Einstein's equivalence principle, what he called his *happiest thought*.

- Einstein understands that inertia and the gravitational field are the same object: it is the gravitational field that defines what frame is inertial, hence the gravitational field is not a force, but provides the structure of spacetime.
- The inertial frames are the freely falling frames. This is the core of Einstein's equivalence principle, what he called his *happiest thought*.

In any and every locally inertial frame, anywhere and anytime in the universe, all the (nongravitational) laws of physics must take on their familiar non-relativistic form.

 The Einstein's Equivalence Principle prescribes a metric theory for gravity, and gives a recipe to write equations of motion:

$$\frac{d^2X^{\mu}}{d^2T^2} = 0$$

$$\frac{d^2x^{\lambda}}{dt^2} + \tilde{\Gamma}^{\lambda}_{\nu\alpha} \frac{dx^{\nu}}{dT} \frac{dx^{\alpha}}{dT} = 0$$

$$\tilde{g}_{\mu\nu} = \frac{\partial X^{\rho}}{\partial x^{\mu}} \frac{\partial X^{\sigma}}{\partial x^{\nu}} \eta_{\rho\sigma}$$





Weinberg, Steven. Gravitation and Cosmology: Principles and Applications of the General Theory of Relativity (Wiley 1972).

 The Einstein's Equivalence Principle prescribes a metric theory for gravity, and gives a recipe to write equations of motion:

$$\frac{d^2X^{\mu}}{d^2T^2} = 0$$

$$\frac{d^2x^{\lambda}}{dt^2} + \tilde{\Gamma}^{\lambda}_{\nu\alpha} \frac{dx^{\nu}}{dT} \frac{dx^{\alpha}}{dT} = 0$$

$$\tilde{g}_{\mu\nu} = \frac{\partial X^{\rho}}{\partial x^{\mu}} \frac{\partial X^{\sigma}}{\partial x^{\nu}} \eta_{\rho\sigma}$$

• Einstein's Equivalence Principle:

$$\tilde{\Gamma}^{\lambda}_{\nu\alpha} \to \Gamma^{\lambda}_{\nu\alpha}$$

$$\tilde{g}_{\mu\nu} \rightarrow g_{\mu\nu}$$





Weinberg, Steven. Gravitation and Cosmology: Principles and Applications of the General Theory of Relativity (Wiley 1972).

# Different aspects of Einstein's Equivalence Principle

 Weak Equivalence Principle (WEP): The local effects of motion in a curved spacetime (gravitation) are indistinguishable from those of an accelerated observer in flat spacetime.

# Different aspects of Einstein's Equivalence Principle

- Weak Equivalence Principle (WEP): The local effects of motion in a curved spacetime (gravitation) are indistinguishable from those of an accelerated observer in flat spacetime.
- Local Lorentz Invariance (LLI): The outcome of any local nongravitational experiment is independent of the velocity of the freely-falling reference frame in which it is performed.

# Different aspects of Einstein's Equivalence Principle

- Weak Equivalence Principle (WEP): The local effects of motion in a curved spacetime (gravitation) are indistinguishable from those of an accelerated observer in flat spacetime.
- Local Lorentz Invariance (LLI): The outcome of any local nongravitational experiment is independent of the velocity of the freely-falling reference frame in which it is performed.
- Local Position Invariance (LPI): The outcome of any local nongravitational experiment is independent of where and when in the universe it is performed.

# The Strong Equivalence Principle

• EEP: In any and every locally inertial frame, anywhere and anytime in the universe, all the (nongravitational) laws of physics must take on their familiar non-relativistic form.

In any and every locally inertial frame, anywhere and anytime in the universe, all the (gravitational or not) laws of physics must take on their familiar non-relativistic form.

 The Strong Equivalence Principle singles out metric-only theories, such as General Relativity.

# Quantum equivalence principles

# Quantum Interpretations of WEP

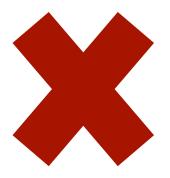
The trajectories of freely falling bodies are independent of their composition

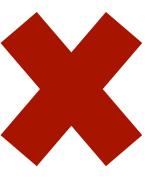


# Quantum Interpretations of WEP

The trajectories of freely falling bodies are independent of their composition

The equations of motion of freely falling bodies do not depend on their mass



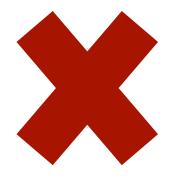


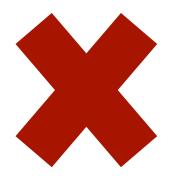
# Quantum Interpretations of WEP

The trajectories of freely falling bodies are independent of their composition

The equations of motion of freely falling bodies do not depend on their mass

The equations of motion under constant gravitational field are indistiguishable from the ones with constant acceleration





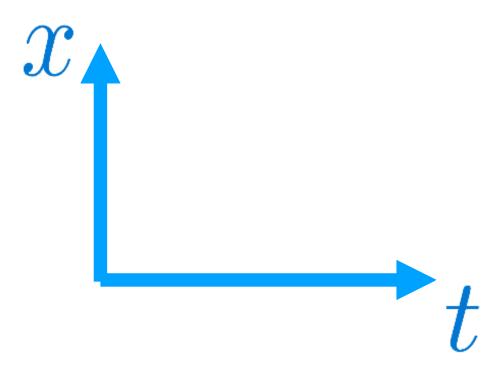


# Quantum Reference Frames

# Motivation

## An operational approach to reference frames

Reference frames are typically treated as an abstraction of rods and clocks.

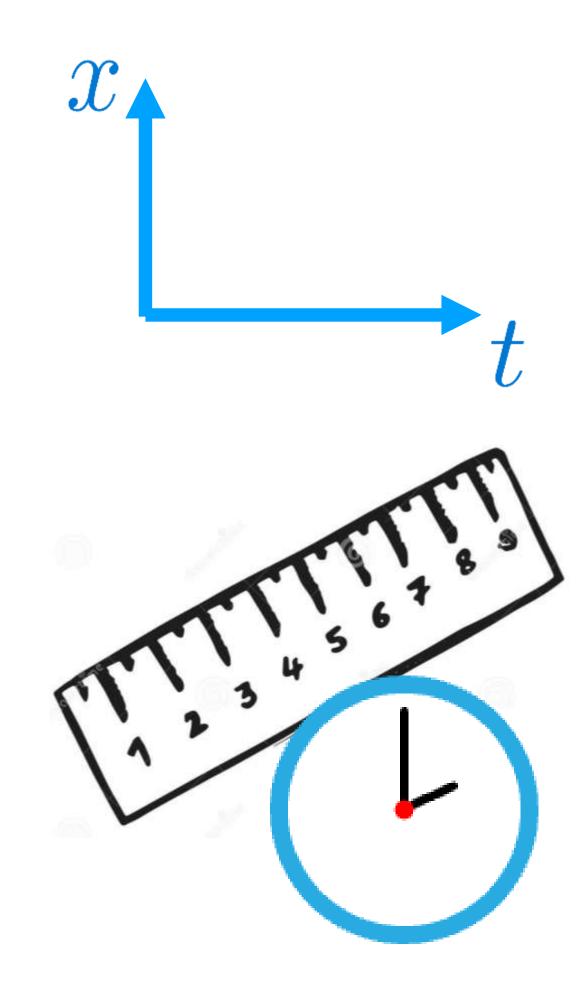


# An operational approach to reference frames

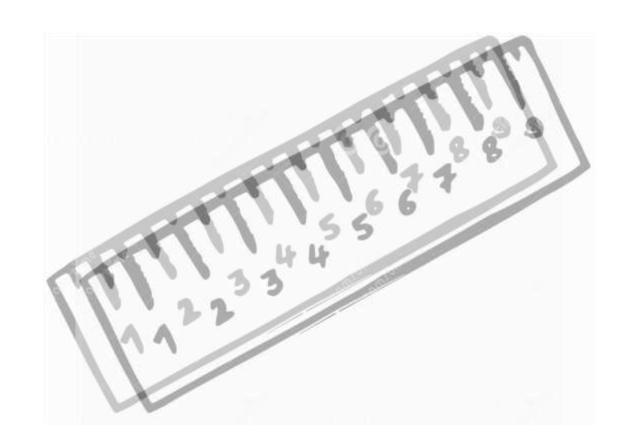
Reference frames are typically treated as an abstraction of rods and clocks.

Nevertheless, they are always attached to actual physical systems, and as such they must obey dynamical laws.

And in particular, they must obey quantum mechanics.

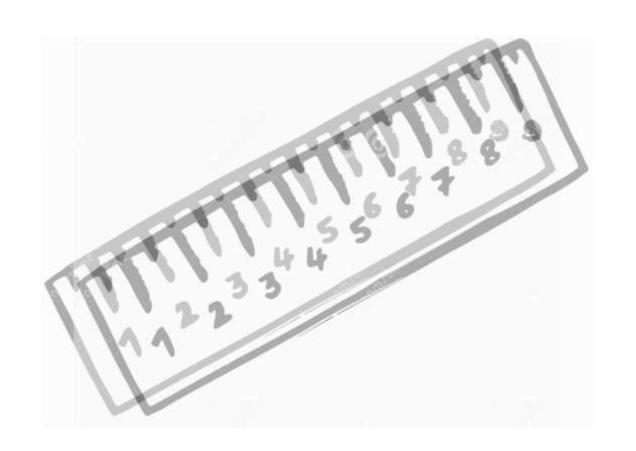


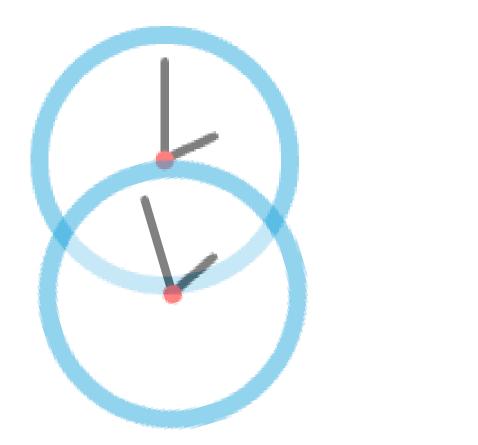
# A quantum ruler and a quantum clock



Given a ruler that is in quantum superposition of different locations, can we still use it to describe position?

# A quantum ruler and a quantum clock





Given a ruler that is in quantum superposition of different locations, can we still use it to describe position?

Given a clock that is in quantum superposition of different heights in a gravitational field, what time will tell?

# Experiments



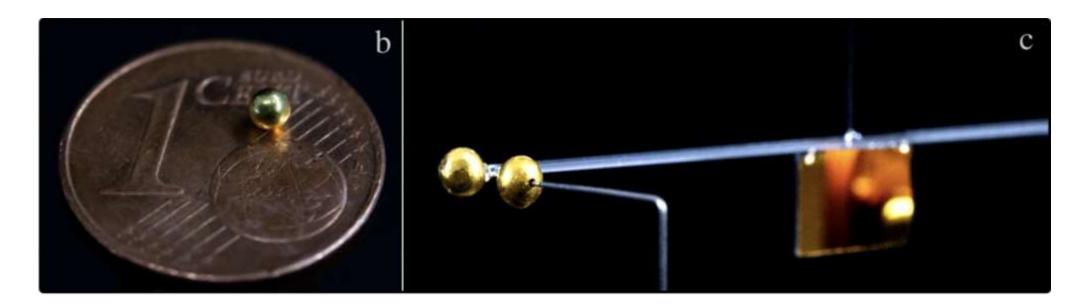
#### **General Relativity and Quantum Cosmology**

[Submitted on 21 Sep 2020 (v1), last revised 2 Mar 2021 (this version, v2)]

#### Measurement of Gravitational Coupling between Millimeter-Sized Masses

Tobias Westphal, Hans Hepach, Jeremias Pfaff, Markus Aspelmeyer

90 mg



# Experiments



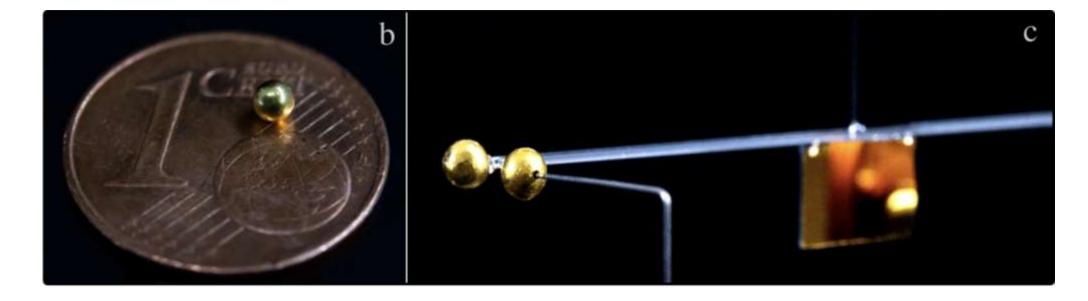
#### **General Relativity and Quantum Cosmology**

[Submitted on 21 Sep 2020 (v1), last revised 2 Mar 2021 (this version, v2)]

#### Measurement of Gravitational Coupling between Millimeter-Sized Masses

Tobias Westphal, Hans Hepach, Jeremias Pfaff, Markus Aspelmeyer

90 mg



Published: 23 December 2015

### Quantum superposition at the half-metre scale

T. Kovachy, P. Asenbaum, C. Overstreet, C. A. Donnelly, S. M. Dickerson, A. Sugarbaker, J. M. Hogan & M. A. Kasevich

Nature 528, 530–533 (2015) Cite this article

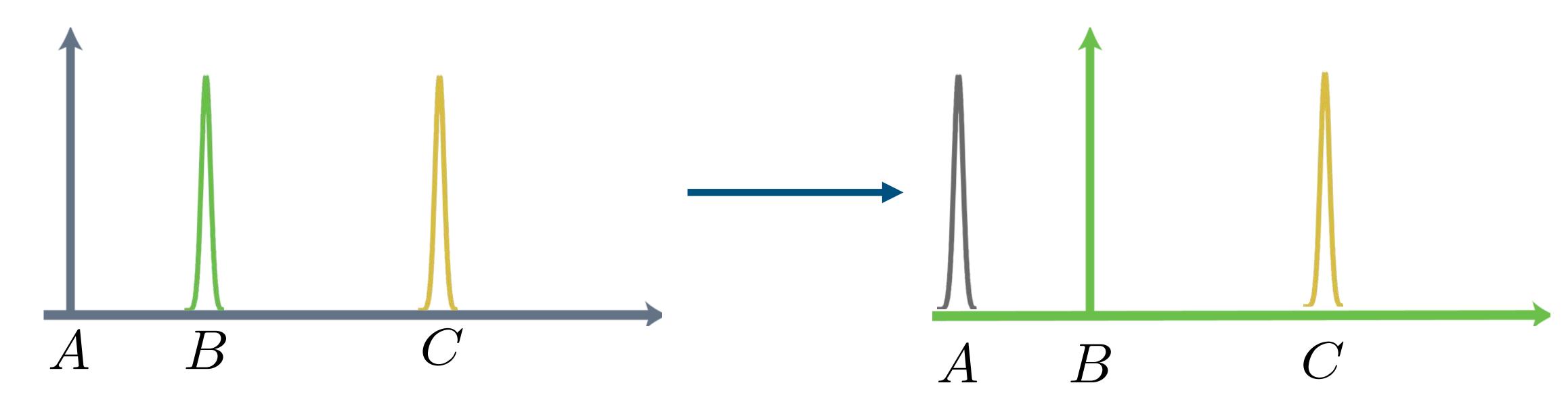
# (Some) literature on QRFs

- 1712.07207
   2006.07298
   2107.07545
   2201.03120
   2308.10967
- 1809.00556 2008.03317 **2109.01405** 2203.06755 **2308.12912**
- 1809.05093 2011.01951 2109.06184 **2207.00021 2402.10267**
- 1811.08228
   2012.13754
   2110.13199
   2209.02214
   2405.00114
- 1908.10165 2012.15769 2110.13824 2211.01657 2406.15838
- 1911.04903 **2101.11628** 2112.00046 2212.14081 2406.19448
- 2004.14292 2103.01232 **2112.03303 2302.11629**
- 2006.06364 2103.05022 **2112.11473** 2308.09131

# An introduction to QRFs

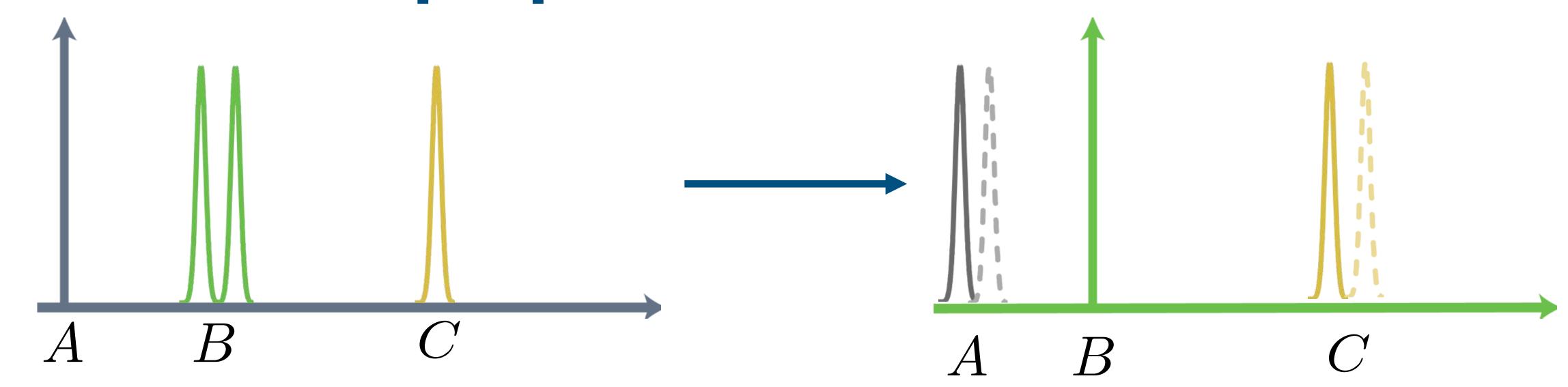
F. Giacomini, E. Castro-Ruiz, C. Brukner, *Quantum mechanics and the covariance of physical laws in quantum reference frames*. Nat Commun 10, 494 (2019).

### Classical RF transformations

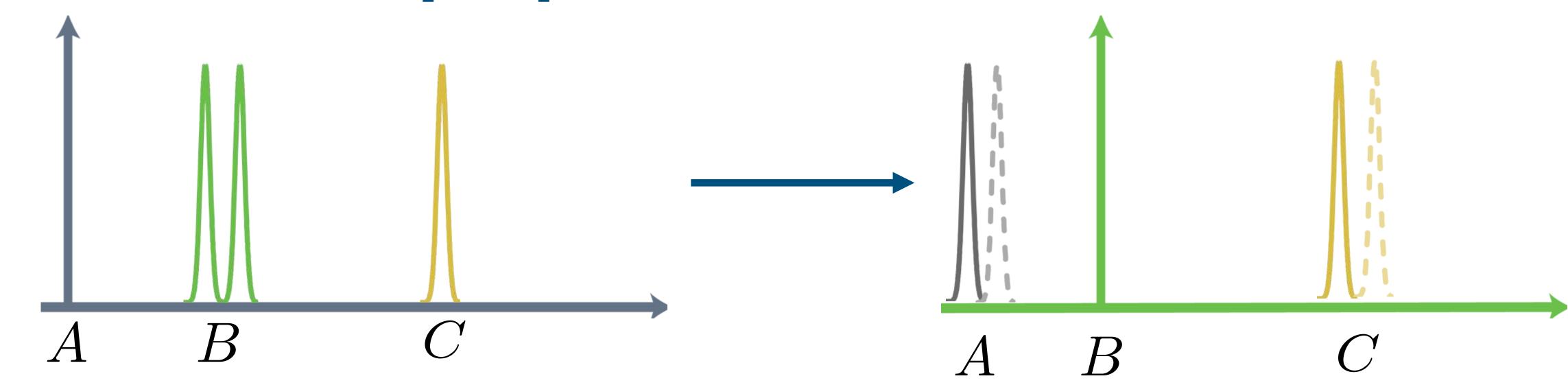


$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A |x_1\rangle_B |x_2\rangle_C$$

$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A |x_1\rangle_B |x_2\rangle_C \qquad |\psi\rangle_{ABC}^{(B)} = |-x_1\rangle_A |0\rangle_B |x_2 - x_1\rangle_C$$

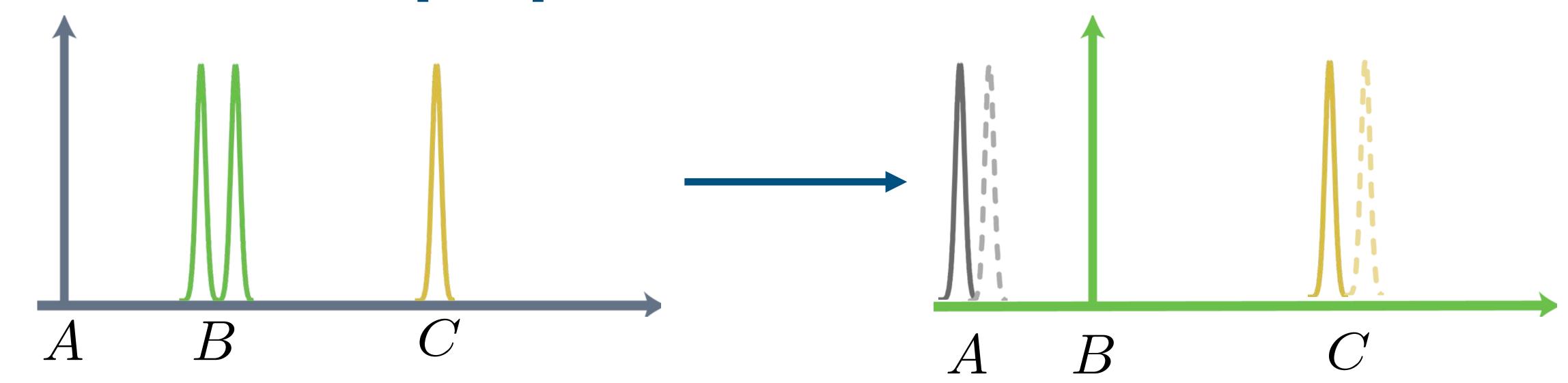


$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A \frac{|x_1\rangle_B + |x_2\rangle_B}{\sqrt{2}} |x_3\rangle_C$$



$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A \frac{|x_1\rangle_B + |x_2\rangle_B}{\sqrt{2}} |x_3\rangle_C$$

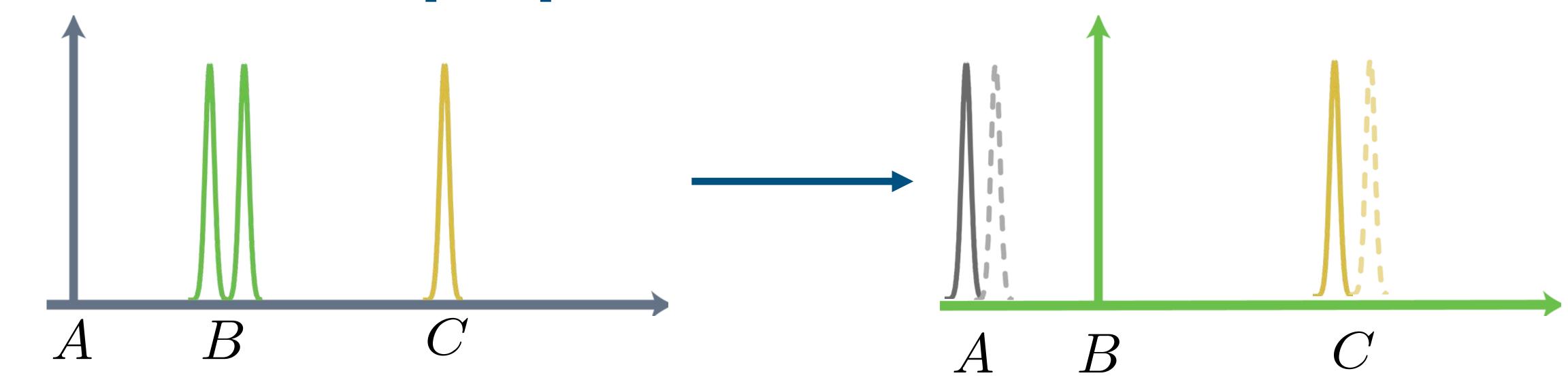
$$|\psi\rangle_{ABC}^{(B)} = |0\rangle_B \frac{|-x_1\rangle_A |x_3 - x_1\rangle_C + |-x_2\rangle_A |x_3 - x_2\rangle_C}{\sqrt{2}}$$



$$|\psi\rangle_{ABC}^{(A)} = |0\rangle_A \frac{|x_1\rangle_B + |x_2\rangle_B}{\sqrt{2}} |x_3\rangle_C$$

$$|\psi\rangle_{ABC}^{(B)} = |0\rangle_B \frac{|-x_1\rangle_A |x_3 - x_1\rangle_C + |-x_2\rangle_A |x_3 - x_2\rangle_C}{\sqrt{2}}$$

In its frame, B is localized



$$|\psi\rangle_{BC}^{(A)} = \frac{|x_1\rangle_B + |x_2\rangle_B}{\sqrt{2}} |x_3\rangle_C$$

$$|\psi\rangle_{AC}^{(B)} = \frac{|-x_1\rangle_A |x_3 - x_1\rangle_C + |-x_2\rangle_A |x_3 - x_2\rangle_C}{\sqrt{2}}$$

### Quantum Reference Frames

(Ideal) QRF tranformations are **coherent superposition** of CRF transformations

### Given:

- 1. A locally compact group G
- 2. A Hilbert space  $\mathcal{H} = L^2(G)$  with orthonormal basis  $\{|g\rangle\}_{g \in G}$ , i.e.  $\langle g|g'\rangle = \delta_{g^{-1} \circ g'}$
- 3. A unitary regular representation of G on  $\mathcal{H}$ , U(g), that is  $U(g)|g'\rangle = |g \circ g'\rangle$

#### Quantum Reference Frames

(Ideal) QRF tranformations are **coherent superposition** of CRF transformations

$$S^{(C)\to(A)}: \mathcal{H}_A \otimes \mathcal{H}_B \to \mathcal{H}_B \otimes \mathcal{H}_C$$
$$|\psi\rangle_{AB}^{(C)} \mapsto |\psi\rangle_{BC}^{(A)}$$

$$S^{(C)\to(A)} = \int_G dg |g^{-1}\rangle_{CA} \langle g| \otimes U_B^{\dagger}(g)$$

Quantum-controlled classical transformation

#### Frame-dependency of resources

$$|\psi\rangle_{BC}^{(A)} = \frac{|x_1\rangle_B + |x_2\rangle_B}{\sqrt{2}} |x_3\rangle_C \longrightarrow |\psi\rangle_{AC}^{(B)} = \frac{|-x_1\rangle_A |x_3 - x_1\rangle_C + |-x_2\rangle_A |x_3 - x_2\rangle_C}{\sqrt{2}}$$

Coherence and entanglement are frame-dependent features

$$C_e^{(A)} + \mathcal{E}_e^{(A)} = C_e^{(B)} + \mathcal{E}_e^{(B)}$$

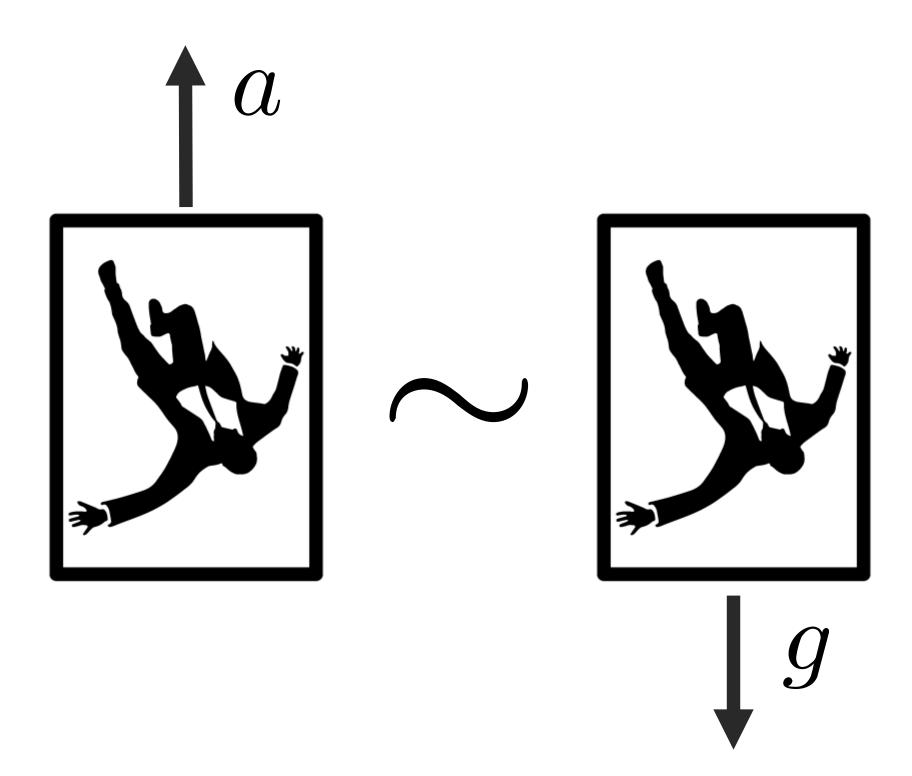
C. Cepollaro et al, The sum of entanglement and subsystem coherence is invariant under quantum reference frame transformations, arXiv:2406.19448

F. Giacomini, E. Castro-Ruiz, C. Brukner, *Quantum mechanics and the covariance of physical laws in quantum reference frames*. Nat Commun 10, 494 (2019).

#### The classical weak equivalence principle

#### Weak equivalence principle:

The physical effects seen in a uniform gravitational field are indistinguishable from those seen from a system in constant acceleration.

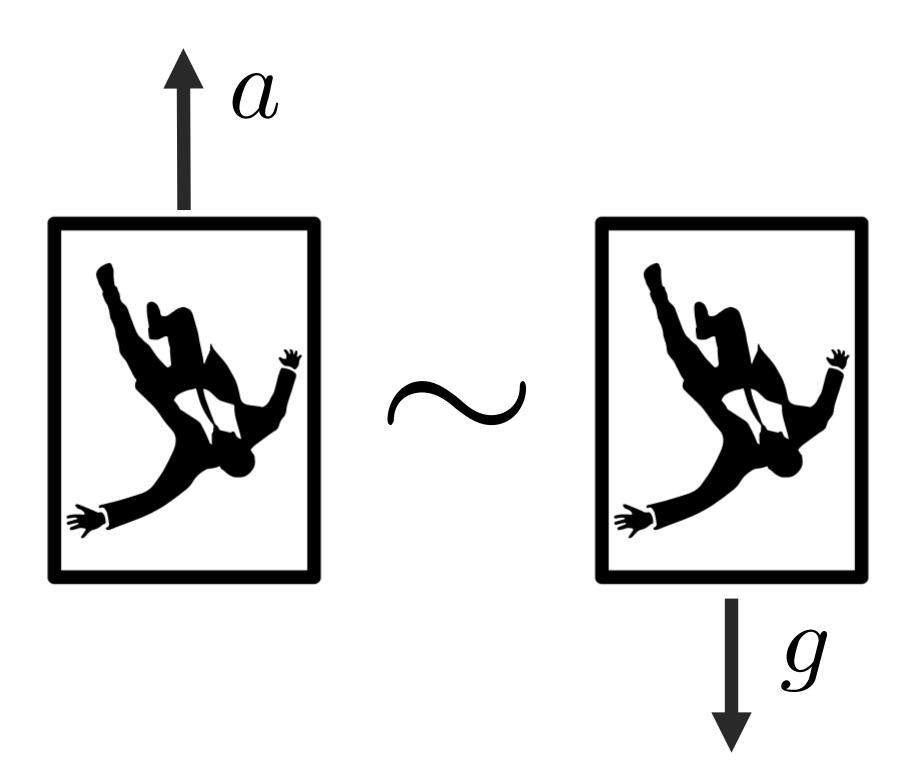


#### The classical weak equivalence principle

#### Weak equivalence principle:

The physical effects seen in a uniform gravitational field are indistinguishable from those seen from a system in constant acceleration.

Given two particles A and B, if A freely falls with constant g and B is free, in the RF of A, B is moving with a constant acceleration a =-g



$$H = \frac{p_A^2}{2m_A} + \frac{p_B^2}{2m_B} + m_A g x_A$$

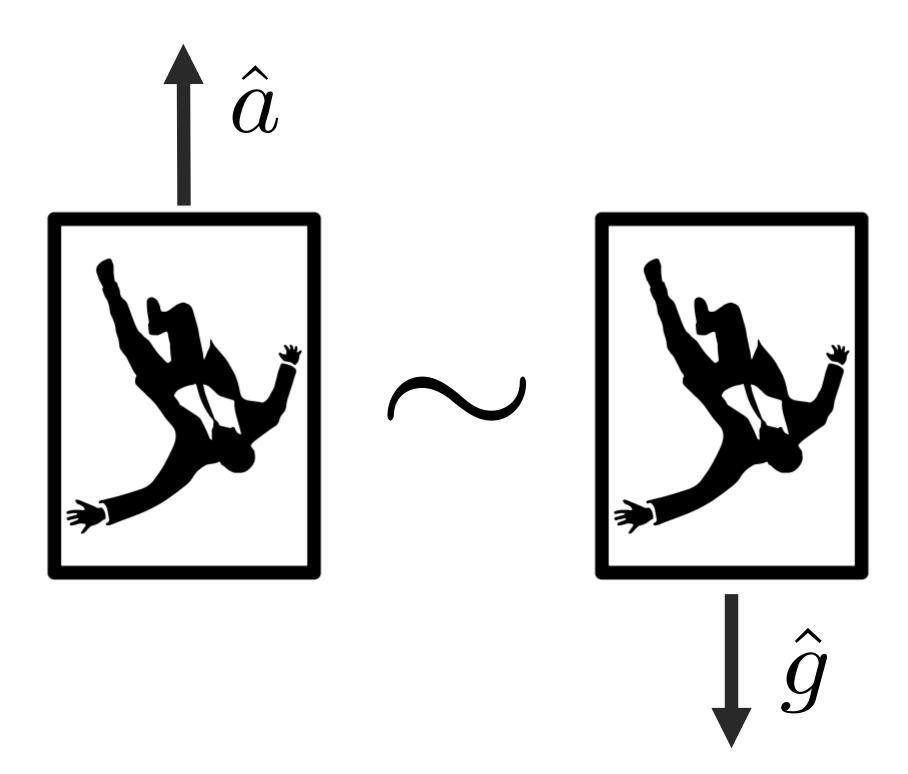
$$x'_{A,B}(t) = x_{A,B} - x_A(t)$$

$$H = \frac{p_A'^2}{2m_A} + \frac{p_B'^2}{2m_B} + m_B a x_B'$$

$$a = -g$$

#### Quantum weak equivalence principle:

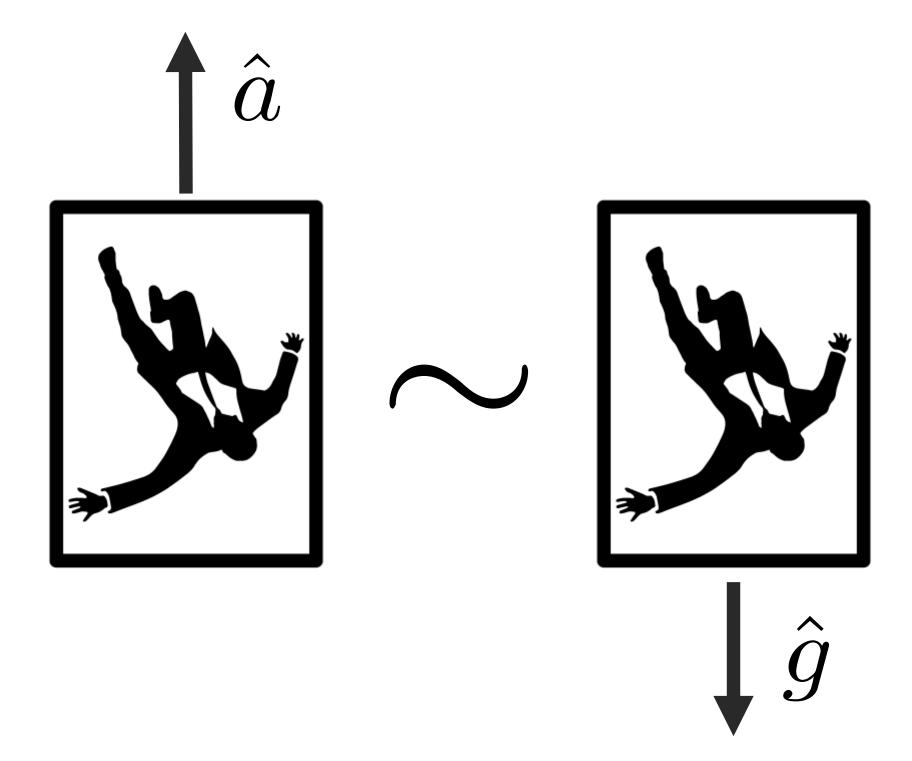
The physical effects seen in a superposition of uniform gravitational fields are indistinguishable from those as seen from a system in superposition of accelerations.



#### Quantum weak equivalence principle:

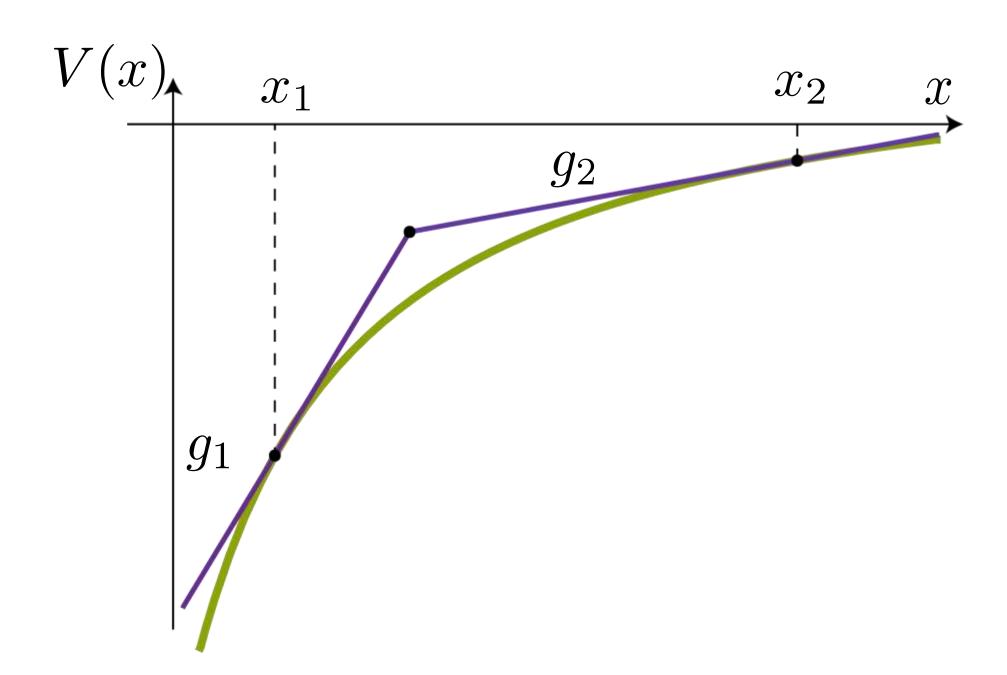
The physical effects seen in a superposition of uniform gravitational fields are indistinguishable from those as seen from a system in superposition of accelerations.

Given two particles A and B, if A freely falls in superposition of different g1, g2 and B is free, can we make a claim similar to the classical case?

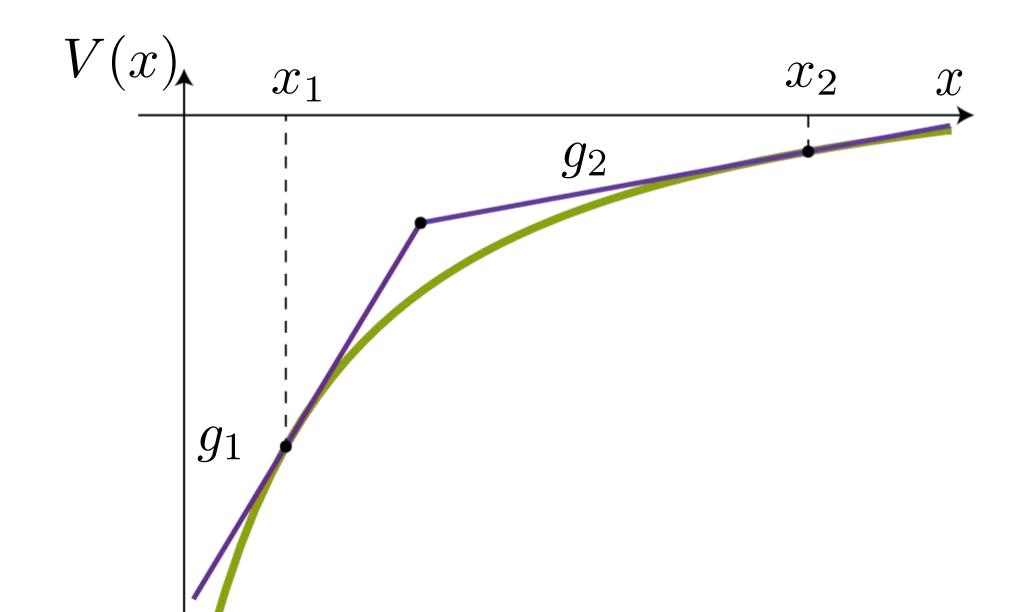


QRFs!

$$\hat{H}_{AB}^{(C)} = \frac{p_A^2}{2m_A} + \frac{p_B^2}{2m_B} + V(\hat{x}_A)$$



$$\hat{H}_{AB}^{(C)} = \frac{p_A^2}{2m_A} + \frac{p_B^2}{2m_B} + V(\hat{x}_A)$$



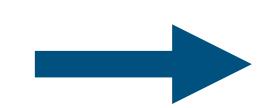
$$|\psi\rangle_{AB}^{(C)} = \frac{|x_1\rangle_A + |x_2\rangle_A}{\sqrt{2}} |0\rangle_B$$

$$\hat{g} = \frac{1}{m_A} \frac{dV(\hat{x}_A)}{d\hat{x}_A}$$

$$\hat{g}\frac{|x_1\rangle_A + |x_2\rangle_A}{\sqrt{2}} = \frac{g_1|x_1\rangle_A + g_2|x_2\rangle_A}{\sqrt{2}}$$

A evolves in a **superposition** of two gravitational accelerations

$$\hat{H}_{AB}^{(C)} = \frac{p_A^2}{2m_A} + \frac{p_B^2}{2m_B} + V(\hat{x}_A)$$

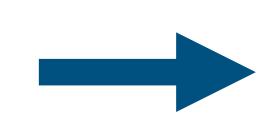


$$|\psi\rangle_{BC}^{(A)} = \frac{|-x_1\rangle_B |-x_1\rangle_C + |-x_2\rangle_B |-x_2\rangle_C}{\sqrt{2}}$$

$$|\psi\rangle_{AB}^{(C)} = \frac{|x_1\rangle_A + |x_2\rangle_A}{\sqrt{2}} |0\rangle_B$$

$$\hat{H}_{BC}^{(A)} = \frac{\pi_B^2}{2m_B} + \frac{\pi_C^2}{2m_C} + \frac{m_C}{m_A}V(-\hat{q}_C) - \frac{m_B}{m_A}\frac{dV}{d\hat{x}_A}\Big|_{-\hat{q}_C}\hat{q}_B$$

$$\hat{H}_{AB}^{(C)} = \frac{p_A^2}{2m_A} + \frac{p_B^2}{2m_B} + V(\hat{x}_A)$$



$$|\psi\rangle_{BC}^{(A)} = \frac{|-x_1\rangle_B |-x_1\rangle_C + |-x_2\rangle_B |-x_2\rangle_C}{\sqrt{2}}$$

$$|\psi\rangle_{AB}^{(C)} = \frac{|x_1\rangle_A + |x_2\rangle_A}{\sqrt{2}} |0\rangle_B$$

$$\hat{H}_{BC}^{(A)} = \frac{\pi_B^2}{2m_B} + \frac{\pi_C^2}{2m_C} + \frac{m_C}{m_A}V(-\hat{q}_C) - \frac{m_B}{m_A}\frac{dV}{d\hat{x}_A}\Big|_{-\hat{q}_C}\hat{q}_B$$

$$\hat{a}_B = -\hat{g}$$

When A falls in a superposition of gravitational accelerations, in the QRF of A, B evolves in a superposition of accelerations

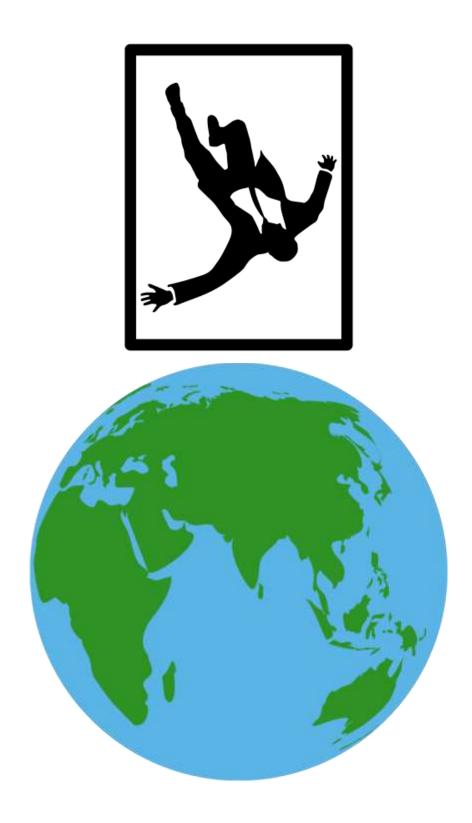
This is the Quantum Weak Equivalence Principle

# The Einstein Equivalence Principle for Quantum Reference Frames

- F. Giacomini, Č. Brukner, Einstein's Equivalence principle for superpositions of gravitational fields and quantum reference frames, arXiv:2012.137544
- F. Giacomini, Č. Brukner, Quantum superposition of spacetimes obeys Einstein's equivalence principle. AVS Quantum Sci. 1 March 2022; 4 (1): 015601.
- C. Cepollaro, F.Giacomini, Quantum generalisation of Einstein's equivalence principle can be verified with entangled clocks as quantum reference frames, 2024 Class. Quantum Grav. 41 185009

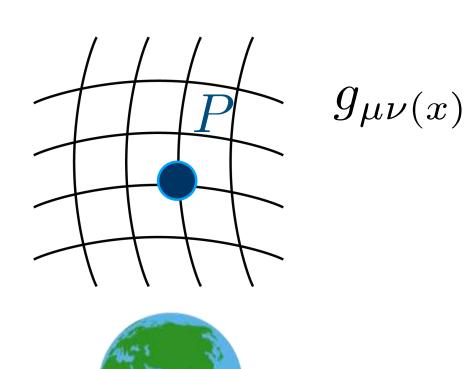
# The Einstein's Equivalence Principle (EEP)

In any and every **locally inertial frame**, anywhere and anytime in the universe, all the (nongravitational) laws of physics must take on their familiar non-relativistic form.



C. W. Misner, K. S. Thorne, J. A. Wheeler, Gravitation (W. H. Freeman, 1973).

#### Locally inertial frames

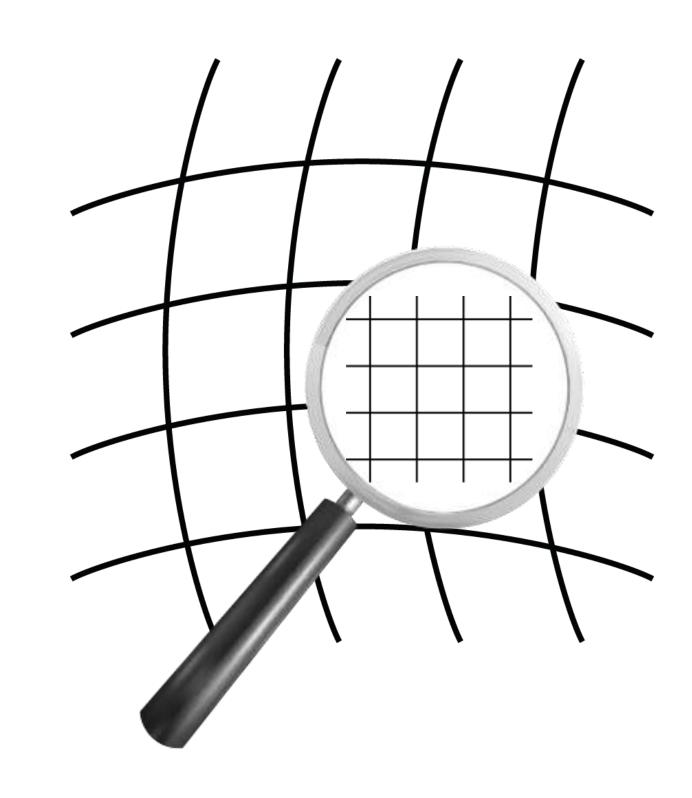


Change of coordinates:

- 1. Translation
- 2. Stretching

$$x^{\mu} = x^{\mu}(\xi^{\mu}) = x_{P}^{\mu} + \frac{\partial x^{\mu}}{\partial \xi^{\alpha}} \xi^{\alpha} + O(\xi^{2})$$
$$\tilde{g}^{\mu\nu}(\xi) = \frac{\partial x^{\mu}}{\partial \xi^{\alpha}} \frac{\partial x^{\nu}}{\partial \xi^{\beta}} g_{\alpha\beta}(x(\xi))$$

$$g_{\mu\nu}(x) \to \tilde{g}_{\mu\nu}(\xi) = \eta_{\mu\nu} + O(\xi^2)$$

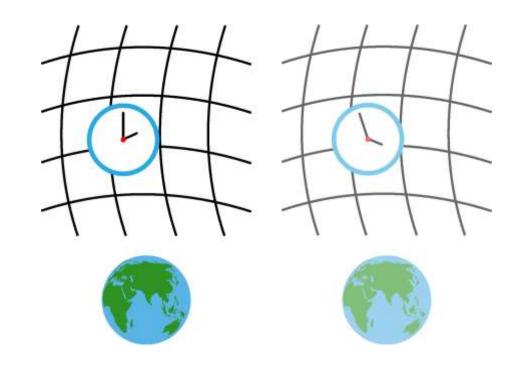


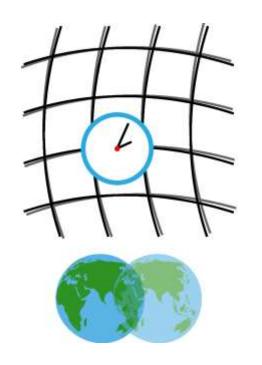
C. W. Misner, K. S. Thorne, J. A. Wheeler, Gravitation (W. H. Freeman, 1973).

#### Quantum locally inertial frame

A mass in spatial superposition generates a superposition of gravitational fields.

A clock that evolves in this scenario will get entangled with the gravitational field.





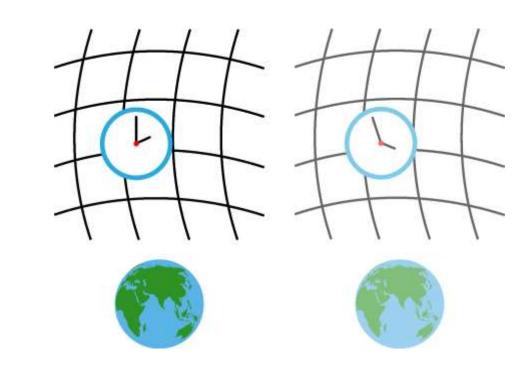
F. Giacomini, Č. Brukner, Einstein's Equivalence principle for superpositions of gravitational fields and quantum reference frames, arXiv:2012.137544

F. Giacomini, Č. Brukner, Quantum superposition of spacetimes obeys Einstein's equivalence principle. AVS Quantum Sci. 1 March 2022; 4 (1): 015601.

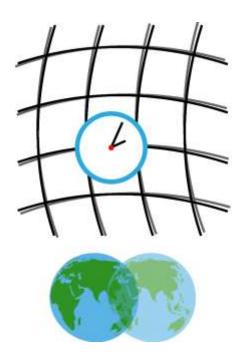
#### Quantum locally inertial frame

A mass in spatial superposition generates a superposition of gravitational fields.

A clock that evolves in this scenario will get entangled with the gravitational field.



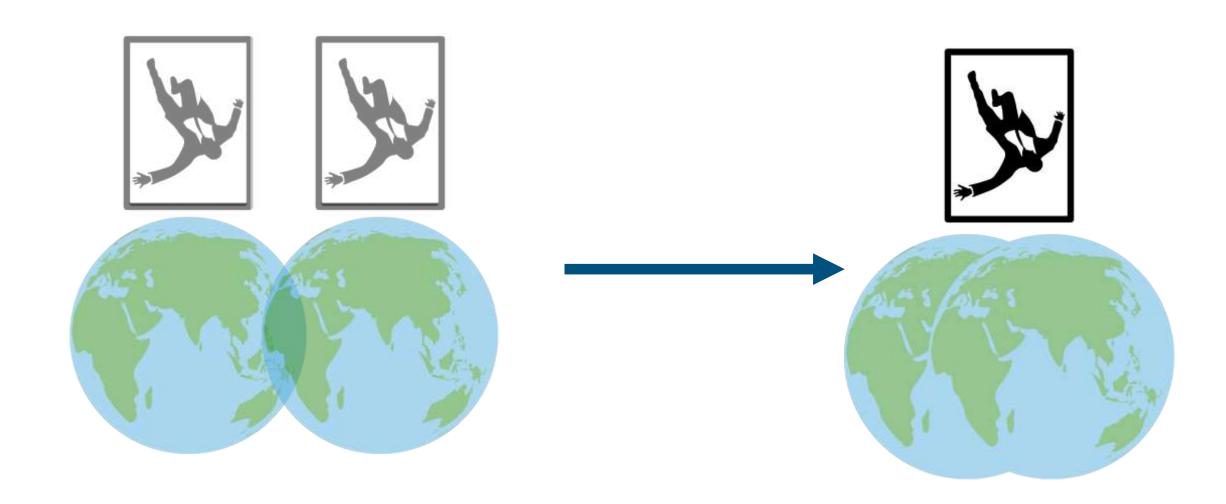
The **Quantum** Locally Inertial Frame (QLIF) of the clock is a frame associated to a quantum particle where the metric is locally flat and well defined.



- F. Giacomini, Č. Brukner, Einstein's Equivalence principle for superpositions of gravitational fields and quantum reference frames, arXiv:2012.137544
- F. Giacomini, Č. Brukner, Quantum superposition of spacetimes obeys Einstein's equivalence principle. AVS Quantum Sci. 1 March 2022; 4 (1): 015601.

#### The Einstein Equivalence Principle for QRFs

In any and every **Quantum** Locally Inertial Frame (**QLIF**), anywhere and anytime in the universe, all the (nongravitational) laws of physics must take on their familiar non-relativistic form.



F. Giacomini, Č. Brukner, Einstein's Equivalence principle for superpositions of gravitational fields and quantum reference frames, arXiv:2012.137544

F. Giacomini, Č. Brukner, Quantum superposition of spacetimes obeys Einstein's equivalence principle. AVS Quantum Sci. 1 March 2022; 4 (1): 015601.

#### The three aspects of the EEP for QRFs



The local effects of (quantum) motion in a superposition of uniform gravitational fields are indistinguishable from those of an observer in flat spacetime that undergoes a quantum superposition of accelerations.



The outcome of any local nongravitational experiment is independent of the velocity of the freely falling **quantum** reference frame in which it is performed.



The outcome of any local nongravitational experiment is independent of the position of the **quantum** reference frame in which it is performed.

C. Cepollaro, F. Giacomini, 2024 Class. Quantum Grav. 41 185009

# Towards a test for the principle

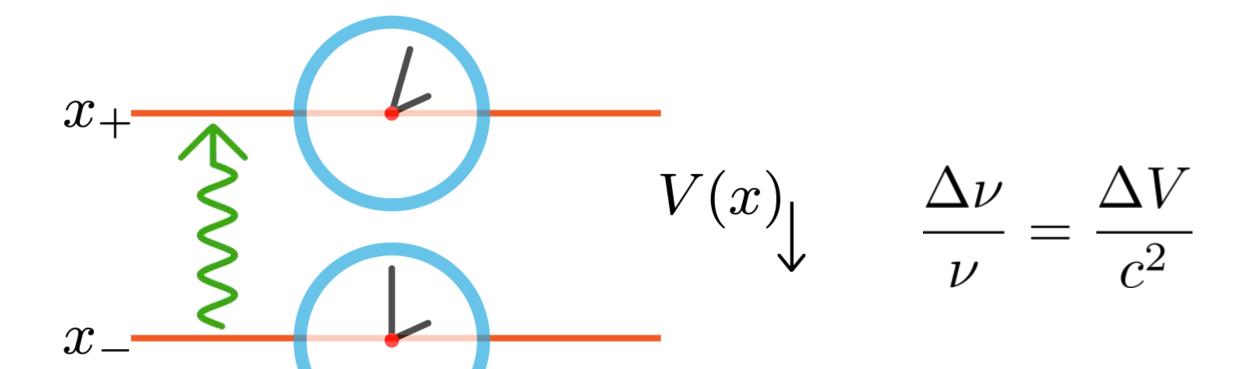
C. Cepollaro, F. Giacomini, 2024 Class. Quantum Grav. 41 185009

#### Classical tests of EEP

- WEP tests: Comparison between accelerations of two bodies of different composition in an external gravitational field.
- LLI tests: Tests of special relativity (e.g., Michelson-Morley experiment).
- LPI tests: Gravitational redshift experiment.

#### Classical tests of EEP

- WEP tests: Comparison between accelerations of two bodies of different composition in an external gravitational field.
- LLI tests: Tests of special relativity (e.g., Michelson-Morley experiment).
- LPI tests: Gravitational redshift experiment.



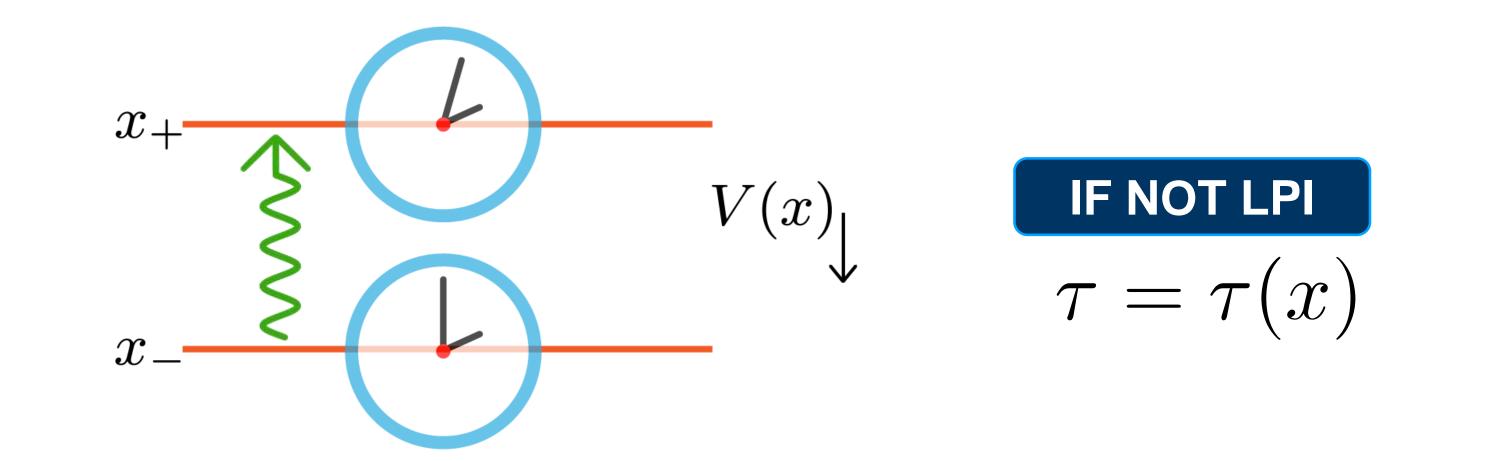
The locally inertial frames momentarily at rest with each clock are related through Lorentz boosts

The doppler effect gives the time dilation factor

#### A classical test for LPI



The outcome of any local nongravitational experiment is independent of the position of the reference frame in which it is performed.



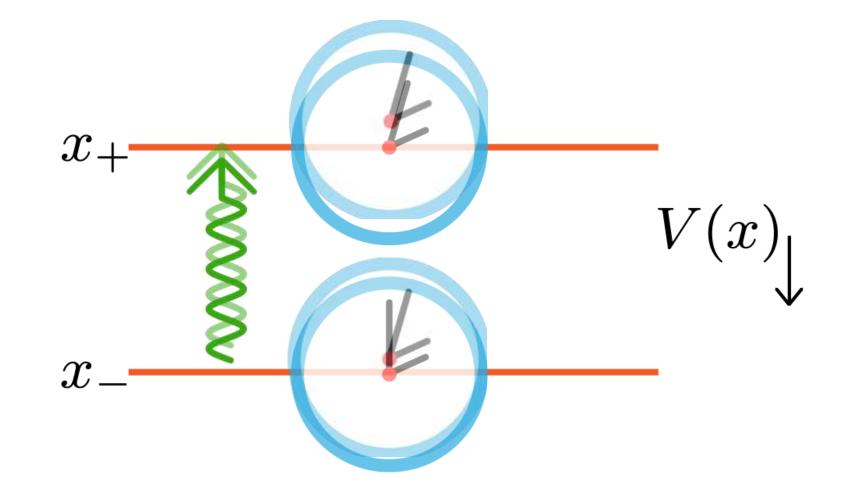
$$\frac{\Delta \nu}{\tau = \tau(x)} = \frac{\Delta \nu}{\nu} = (1 + \alpha(x)) \, \frac{\Delta V}{c^2}$$

LPI can be verified by measuring  $\alpha(x)$ 

#### A quantum test for Q-LPI



The outcome of any local nongravitational experiment is independent of the position of the quantum reference frame in which it is performed.



$$\tau = \tau(\hat{x})$$

$$\frac{\Delta \nu}{\tau = \tau(\hat{x})} = \frac{\Delta \nu}{\nu} = (1 + \alpha(\hat{x})) \, \frac{\Delta V}{c^2}$$

Q-LPI can be verified by measuring  $\alpha(\hat{x})$ 

#### Atomic clocks

External + Internal DOF

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) \qquad m \to \hat{M} = m + \frac{H_I}{c^2}$$

Newtonian QM + mass-energy equivalence

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) + \hat{H}_I \left( 1 + \frac{V(\hat{\mathbf{x}})}{c^2} - \frac{\hat{\mathbf{p}}^2}{2m^2c^2} \right)$$

- 1. C. Lämmerzahl, On the equivalence principle in quantum theory. Gen Relat Gravit 28, 1043–1070 (1996).
- 2. M. Zych, F. Costa, I. Pikovski, Č. Brukner, Quantum interferometric visibility as a witness of general relativistic proper time. Nat Commun 2, 505 (2011).
- 3. M. Zych, Č. Brukner, Quantum formulation of the Einstein Equivalence Principle, Nature Phys 14, 1027–1031 (2018).

#### Atomic clocks

#### External + Internal DOF

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) \qquad m \to \hat{M} = m + \frac{H_I}{c^2}$$

$$m \rightarrow \hat{M} = m + \frac{H_I}{c^2}$$

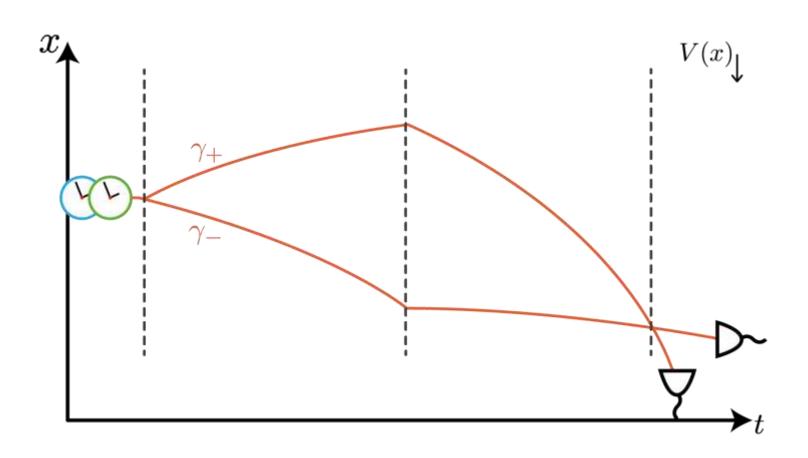
Newtonian QM + mass-energy equivalence

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) + \hat{H}_I \left( 1 + \frac{V(\hat{\mathbf{x}})}{c^2} - \frac{\hat{\mathbf{p}}^2}{2m^2c^2} \right)$$

$$\omega' = \omega \left( 1 + \frac{V(x)}{c^2} - \frac{p^2}{2m^2c^2} \right) = \omega \frac{d\tau}{dt} \qquad \qquad \omega = \frac{\Delta E}{\hbar} \qquad \Delta E$$

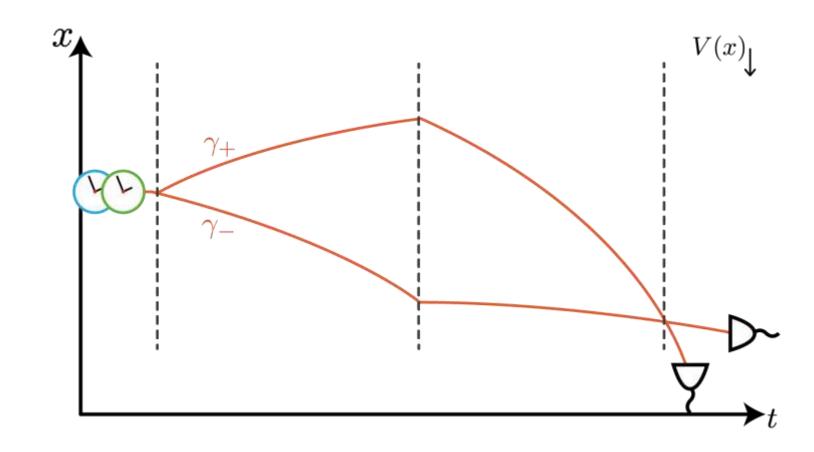
- 1. C. Lämmerzahl, On the equivalence principle in quantum theory. Gen Relat Gravit 28, 1043–1070 (1996).
- 2. M. Zych, F. Costa, I. Pikovski, Č. Brukner, Quantum interferometric visibility as a witness of general relativistic proper time. Nat Commun 2, 505 (2011).
- 3. M. Zych, Č. Brukner, Quantum formulation of the Einstein Equivalence Principle, Nature Phys 14, 1027–1031 (2018).

$$|\psi_0\rangle_{AB} = \frac{|x_+\rangle_{\mathbf{A}}|x_+\rangle_{\mathbf{B}} + |x_-\rangle_{\mathbf{A}}|x_-\rangle_{\mathbf{B}}}{\sqrt{2}} |\tau_{in}\rangle_{C_A} |\tau_{in}\rangle_{C_B}$$



$$|\psi_0\rangle_{AB} = \frac{|x_+\rangle_{\mathbf{A}}|x_+\rangle_{\mathbf{B}} + |x_-\rangle_{\mathbf{A}}|x_-\rangle_{\mathbf{B}}}{\sqrt{2}} |\tau_{in}\rangle_{C_A} |\tau_{in}\rangle_{C_B}$$

$$|\psi(t)\rangle_{AB} = \frac{1}{\sqrt{2}}(|x_{+}\rangle_{\mathbf{A}}|x_{+}\rangle_{\mathbf{B}}|\tau_{+}\rangle_{C_{A}}|\tau_{+}\rangle_{C_{B}} + e^{i\sum_{j=A,B}\Delta\phi_{j}}|x_{-}\rangle_{\mathbf{A}}|x_{-}\rangle_{\mathbf{B}}|\tau_{-}\rangle_{C_{A}}|\tau_{-}\rangle_{C_{B}})$$

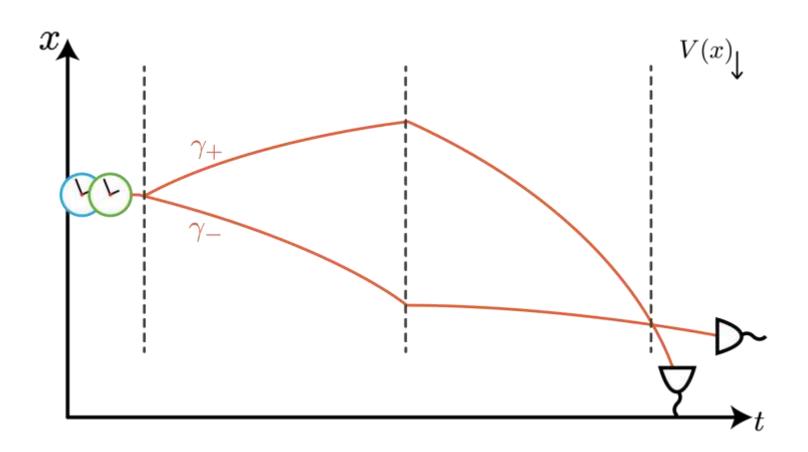


$$|\psi_0\rangle_{AB} = \frac{|x_+\rangle_{\mathbf{A}}|x_+\rangle_{\mathbf{B}} + |x_-\rangle_{\mathbf{A}}|x_-\rangle_{\mathbf{B}}}{\sqrt{2}} |\tau_{in}\rangle_{C_A} |\tau_{in}\rangle_{C_B}$$

$$|\psi(t)\rangle_{AB} = \frac{1}{\sqrt{2}}(|x_{+}\rangle_{\mathbf{A}}|x_{+}\rangle_{\mathbf{B}}|\tau_{+}\rangle_{C_{A}}|\tau_{+}\rangle_{C_{B}} + e^{i\sum_{j=A,B}\Delta\phi_{j}}|x_{-}\rangle_{\mathbf{A}}|x_{-}\rangle_{\mathbf{B}}|\tau_{-}\rangle_{C_{A}}|\tau_{-}\rangle_{C_{B}})$$

$$\Delta \phi_j = \frac{1}{\hbar} \int_{\Delta \gamma} dt \left( m_j c^2 + \frac{\mathbf{p}_j^2}{2m_j} + m_j V(\mathbf{x}_j) \right)$$

$$|\tau_{\pm}\rangle_{C_j} = e^{-\frac{i}{\hbar} \int_{\gamma_{\pm}} dt \left(1 + \frac{V(\mathbf{x}_j)}{c^2} - \frac{\mathbf{p}_j^2}{2m_j^2 c^2}\right) \hat{H}_{int}^j} |\tau_{in}\rangle_{C_j}$$

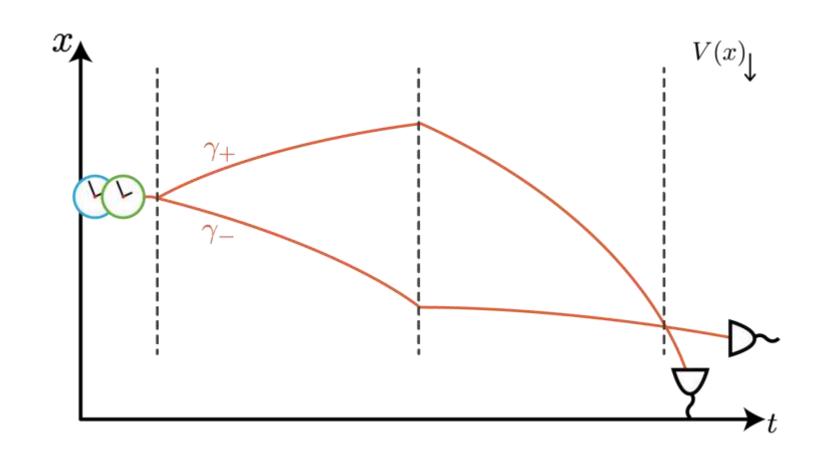


$$|\psi_0\rangle_{AB} = \frac{|x_+\rangle_{\mathbf{A}}|x_+\rangle_{\mathbf{B}} + |x_-\rangle_{\mathbf{A}}|x_-\rangle_{\mathbf{B}}}{\sqrt{2}} |\tau_{in}\rangle_{C_A} |\tau_{in}\rangle_{C_B}$$

$$|\psi(t)\rangle_{AB} = \frac{1}{\sqrt{2}}(|x_{+}\rangle_{\mathbf{A}}|x_{+}\rangle_{\mathbf{B}}|\tau_{+}\rangle_{C_{A}}|\tau_{+}\rangle_{C_{B}} + e^{i\sum_{j=A,B}\Delta\phi_{j}}|x_{-}\rangle_{\mathbf{A}}|x_{-}\rangle_{\mathbf{B}}|\tau_{-}\rangle_{C_{A}}|\tau_{-}\rangle_{C_{B}})$$

$$\Delta \phi_j = \frac{1}{\hbar} \int_{\Delta \gamma} dt \left( m_j c^2 + \frac{\mathbf{p}_j^2}{2m_j} + m_j V(\mathbf{x}_j) \right)$$

$$|\tau_{\pm}\rangle_{C_j} = e^{-\frac{i}{\hbar}\int_{\gamma_{\pm}} dt \left(1 + \frac{V(\mathbf{x}_j)}{c^2} - \frac{\mathbf{p}_j^2}{2m_j^2 c^2}\right) \hat{H}_{int}^j} |\tau_{in}\rangle_{C_j}$$



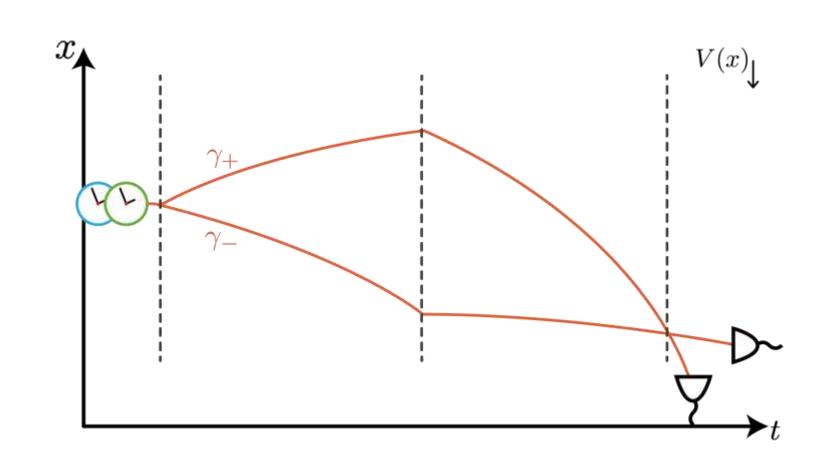
$$|D_{\pm}\rangle_{\mathbf{AB}} = \frac{|x_{+}\rangle_{\mathbf{A}}|x_{+}\rangle_{\mathbf{B}} \pm |x_{-}\rangle_{\mathbf{A}}|x_{-}\rangle_{\mathbf{B}}}{\sqrt{2}}$$

$$P_{\pm} = \frac{1}{2} \left( 1 \pm |\langle \tau_{+} | \tau_{-} \rangle|^{2} \cos \left( 2(\Delta \phi + \varphi') \right) \right)$$

#### The test for EEP for QRFs

#### IF NOT Q-LPI

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) + \hat{H}_I \left( 1 + (1 + \alpha(\hat{x})) \frac{V(\hat{\mathbf{x}})}{c^2} - \frac{\hat{\mathbf{p}}^2}{2m^2c^2} \right)$$

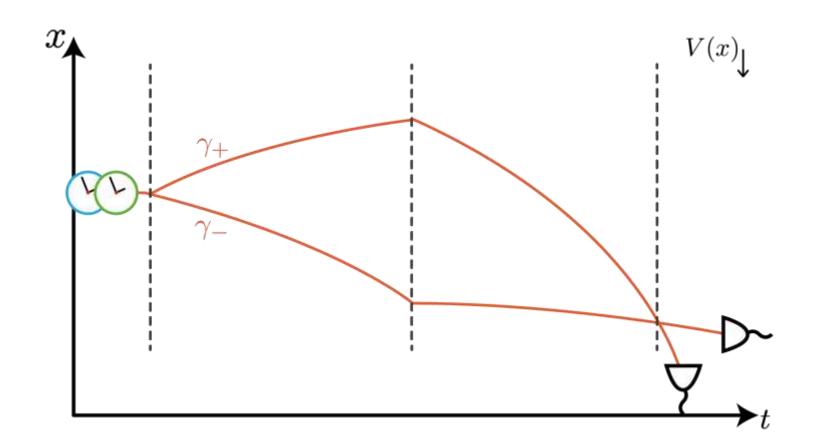


#### The test for EEP for QRFs

#### IF NOT Q-LPI

$$\hat{H} = mc^2 + \frac{\hat{\mathbf{p}}^2}{2m} + mV(\hat{\mathbf{x}}) + \hat{H}_I \left( 1 + (1 + \alpha(\hat{x})) \frac{V(\hat{\mathbf{x}})}{c^2} - \frac{\hat{\mathbf{p}}^2}{2m^2c^2} \right)$$

$$\langle \tau_{+}|\tau_{-}\rangle' = \langle \tau_{+}|\tau_{-}\rangle - \frac{i}{\hbar} \int_{\Delta\gamma} dx \, \frac{V(\mathbf{x}_{\gamma})}{c^{2}} \alpha(\mathbf{x}_{\gamma}) \langle \tau_{+}|\hat{H}|\tau_{-}\rangle$$



$$P'_{\pm} = \frac{1}{2} \left( 1 \pm |\langle \tau_{+} | \tau_{-} \rangle'|^{2} \cos \left( 2(\Delta \phi + \varphi') \right) \right) \neq P_{\pm}$$

Atomic clock interferometers can be used to measure vioaltions of the EEP for QRFs

# Summary

 There are different definitions of equivalence principle, depending on their interpretation one ends up with different compatibility with QM

- There are different definitions of equivalence principle, depending on their interpretation one ends up with different compatibility with QM
- Quantum Reference Frames (QRFs) generalize classical RFs for particles that can be in quantum superposition

- There are different definitions of equivalence principle, depending on their interpretation one ends up with different compatibility with QM
- Quantum Reference Frames (QRFs) generalize classical RFs for particles that can be in quantum superposition
- QRFs can be used to define a quantum extension of the Einstein's Equivalence Principle

- There are different definitions of equivalence principle, depending on their interpretation one ends up with different compatibility with QM
- Quantum Reference Frames (QRFs) generalize classical RFs for particles that can be in quantum superposition
- QRFs can be used to define a quantum extension of the Einstein's Equivalence Principle
- This EEP for QRFs can be tested with atomic clock interferometers

# Thank you!