Causal and stable magnetohydrodynamics for an ultrarelativistic two-component gas

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9th Conference on Chirality, Vorticity and Magnetic Fields in Quantum Matter - São Paulo, Brazil

WHY RELATIVISTIC MAGNETOHYDRODYNAMICS?

Magnetars



Heavy-ion collisions



European Space Agency

 $B\sim 10^{15}\,{
m G}$

ALICE Collaboration

 $\textit{B} \sim 10^{19}\,{\rm G}$

Locally neutral, non-resistive plasma

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• Conservation laws:

$$\partial_{\mu}T^{\mu\nu} = 0 \implies T^{\mu\nu} = \left(\varepsilon + \frac{B^2}{2}\right)u^{\mu}u^{\nu} - \Delta^{\mu\nu}\left(P + \frac{B^2}{2}\right) + \pi^{\mu\nu} - B^{\mu}B^{\nu}.$$

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- Maxwell's equations:

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- Equations for the dissipative currents:
 - Additional relations are required for closure.
 - Phenomenology, microscopic calculations, ...

Traditional Israel-Stewart theory with corrected energy-momentum tensor

$$\dot{\pi}^{\langle\mu\nu\rangle} + \pi^{\mu\nu} = \mathbf{2}\eta\sigma^{\mu\nu} + \cdots.$$

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Longitudinal approximation for the shear-stress tensor

$$\pi^{\mu
u}pprox \pi_{bb}\left(b^{\mu}b^{
u}+rac{1}{2}\Xi^{\mu
u}
ight).$$

M. Chandra et al., Astrophys. J. 810, 162 (2015).

Bounds on nonlinear causality have been investigated.

I. Cordeiro et al., Phys. Rev. Lett. 133, 091401 (2024).

Assuming a two-component gas of classical particles described by a Boltzmann equation

$$\begin{aligned} k^{\mu}\partial_{\mu}f_{\mathbf{k}}^{+} + q^{+}k_{\nu}F^{\mu\nu}\frac{\partial}{\partial k^{\mu}}f_{\mathbf{k}}^{+} &= C[f_{\mathbf{k}}^{+}, f_{\mathbf{k}}^{-}],\\ k^{\mu}\partial_{\mu}f_{\mathbf{k}}^{-} + q^{-}k_{\nu}F^{\mu\nu}\frac{\partial}{\partial k^{\mu}}f_{\mathbf{k}}^{-} &= C[f_{\mathbf{k}}^{-}, f_{\mathbf{k}}^{+}].\end{aligned}$$

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Method of moments

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14-moment approximation

Express the distribution function solely in terms of the usual hydrodynamic variables.

W. Israel and J. M. Stewart, Ann. Phys. (N.Y.) 118, 341 (1979). G. S. Denicol et al., Eur. Phys. J. A 48, 170 (2012).

Factorize the single-particle distribution as

$$f_{\mathbf{k}}^{\pm} = f_{0\mathbf{k}}^{\pm} + \delta f_{\mathbf{k}}^{\pm} \implies \delta f_{\mathbf{k}}^{\pm} = \frac{f_{0\mathbf{k}} \left(\pi_{\mu\nu}^{\pm} k^{\mu} k^{\nu} \right)}{2 \left(\varepsilon^{\pm} + P^{\pm} \right) T^{2}}.$$

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Obtain coupled equations for the total and relative shear-stress tensor

$$\begin{split} \dot{\pi}^{\langle\mu\nu\rangle} + \Sigma \pi^{\mu\nu} + \frac{2|q|B}{5T} b^{\lambda\langle\mu} \delta \pi_{\lambda}^{\nu\rangle} &= \frac{8}{15} \varepsilon \sigma^{\mu\nu} - \frac{4}{3} \pi^{\mu\nu} \theta - \frac{10}{7} \sigma^{\lambda\langle\mu} \pi_{\lambda}^{\nu\rangle} - 2\omega^{\lambda\langle\nu} \pi_{\lambda}^{\mu\rangle}, \\ \delta \dot{\pi}^{\langle\mu\nu\rangle} + \Sigma' \delta \pi^{\mu\nu} + \frac{2|q|B}{5T} b^{\lambda\langle\mu} \pi_{\lambda}^{\nu\rangle} &= -\frac{4}{3} \delta \pi^{\mu\nu} \theta - \frac{10}{7} \sigma^{\lambda\langle\mu} \delta \pi_{\lambda}^{\nu\rangle} - 2\omega^{\lambda\langle\nu} \delta \pi_{\lambda}^{\mu\rangle}. \end{split}$$

where $\pi^{\mu\nu} = \pi^{\mu\nu}_+ + \pi^{\mu\nu}_-$ and $\delta\pi^{\mu\nu} = \delta\pi^{\mu\nu}_+ + \delta\pi^{\mu\nu}_-$. K. Kushwah and G. S. Denicol, Phys. Rev. D 109, 096021 (2024).

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Is this theory suitable to describe physical systems?

What are the minimal requirements a theory must satisfy?

- Causality: perturbations travel with subluminal speed.
- Stability: perturbations decay exponentially with time.

These properties are intrinsically connected.

L. Gavassino, Phys. Rev. X 12, 041001 (2022).

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Compute the modes of the theory, $\omega(k)$, considering two cases:

- Transverse perturbations: orthogonal to the magnetic field.
- Longitudinal perturbations: parallel to the magnetic field.

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Compute the modes of the theory, $\omega(k)$, considering two cases:

- Transverse perturbations: orthogonal to the magnetic field.
- Longitudinal perturbations: parallel to the magnetic field.
- Constrain the transport coefficients such that the following conditions are satisfied:

Im
$$(\omega) > 0$$
, $\lim_{k \to \infty} \left| \frac{\partial \operatorname{Re} (\omega)}{\partial k} \right| \leq 1$.

Stability

Causality

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For small k,

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Hydrodynamic and nonhydrodynamic modes

- Hydrodynamic modes: long-lived, associated with conserved quantities.
- Nonhydrodynamic modes: short-lived, necessary for linear causality and stability.

Longitudinal perturbations

$$egin{aligned} & \omega_{\mathrm{nh}}(k) = i \Sigma - rac{4i}{15 \Sigma} k^2 + \mathcal{O}(k^3), \ & \omega_{\mathrm{nh}}(k) = rac{i}{2} \left[\Sigma + \Sigma' \pm \sqrt{(\Sigma - \Sigma')^2 - rac{4q^2 B_0^2}{25 T^2}} + \mathcal{O}(k^2), \ & \omega_{\mathrm{h}}(k) = \pm rac{1}{\sqrt{3}} k + rac{2i}{15 \Sigma} k^2 + \mathcal{O}(k^3), \ & \omega_{\mathrm{h}}(k) = \pm v_{\mathrm{A}} k + rac{2i}{5} rac{\Sigma'}{\left(1 + rac{B_0^2}{arepsilon_0 + P_0}
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Longitudinal perturbations

In the small wave number limit, the modes are

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- Diffusive part of Alfvén modes suppressed at large magnetic field.
- Terms that are absent in the longitudinal approximation.

LONGITUDINAL PERTURBATIONS

Comparison of the hydrodynamic modes



Transverse perturbations

Transverse perturbations

Complete theory

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Transverse perturbations

Complete theory

Longitudinal approximation

 $\omega = \pm \sqrt{\frac{1+3\mathcal{B}}{3(1+\mathcal{B})}}k.$

 $\omega = i\Sigma$

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- Modes are significantly distinct from Israel-Stewart-like theories.
- Causality and stability remain being fulfilled in the longitudinal approximation.
- Longitudinal approximation is justified in the limit of large *B*.
- Longitudinal approximation fails to capture the dynamics when *B* is not sufficiently large.