

Illinois Center for Advanced Studies of the Universe



Dynamic magneto-chiral instability in a semiconductor

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arXiv:2502.05170

9th Conference on Chirality, Vorticity, and Magnetic Fields in Quantum Matter

São Paulo, Brazil

July 2025

Motivation: the chiral magnetic effect



Chirally imbalanced plasma of massless fermions in a magnetic field

Motivation: the chiral magnetic effect



Motivation: the chiral magnetic effect





electrically conductive crystalline materials









Chiral magnetic effect in Weyl semimetals



Chiral magnetic effect in Weyl semimetals







$\delta \vec{B} \bullet \uparrow \bullet$

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Self-amplifying magnetic fluctuations Y. Akamatsu & N. Yamamoto, PRL 111, 052002 (2013)



Self-amplifying magnetic fluctuations Y. Akamatsu & N. Yamamoto, PRL 111, 052002 (2013)





Essential ingredients:

- $\vec{J} \propto \vec{B}$
- Maxwell's equations

Self-amplifying magnetic fluctuations Y. Akamatsu & N. Yamamoto, PRL 111, 052002 (2013)

The chiral anomaly is not necessary!

Stellar dynamo: "alpha effect"

Helical turbulence in rotating bodies

 \longrightarrow macroscopic $ec{J} \propto ec{B}$

Can generate astrophysical magnetic fields

Neutron star merger (cross section) E. R. Most, PRD 108, 123012 (2023)



'turbulence"

magneti

Self

Y. Akan

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Crystal symmetry requirements

<u>Magneto-chiral current</u>: $J_i = \sigma_{B,ij} B_j$



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Crystal symmetry requirements

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Crystal symmetry requirements Time-reversal symmetric materials Magneto with chiral crystal structure **Spatial** ane Stru rall hiral semiconductor tellurium 2502.05170 $\Rightarrow \sigma_{B,ij} = \begin{pmatrix} \sigma_B^{\perp} & 0 & 0 \\ 0 & \sigma_B^{\perp} & 0 \\ 0 & 0 & \sigma^{\parallel} \end{pmatrix} | \text{effec}$ effect, Nick Abboud | U lpaign

THz emission from photoexcited Te

Y. Huang, NA, et al., 2502.05170



THz emission from photoexcited Te

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- Sum of four exponentially decaying oscillations...
 - → Impurity levels K. Natori et al., JPSJ 34, 1263 (1973)

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THz emission from photoexcited Te

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• Sum of four exponentially decaying oscillations...

ightarrow Impurity levels

K. Natori et al., JPSJ 34, 1263 (1973)

...AND one growing oscillation!



A first attempt: condensed matter



A first attempt: condensed matter



plane wave solutions $\vec{B} \sim e^{i(\vec{k}\cdot\vec{r}-\omega t)}$ $\vec{k} \in \mathbb{R}^3$ $\mathrm{Im}[\omega] > 0 \implies$ instability

A first attempt: condensed matter







$$\begin{array}{l} \underline{ \text{Lorentzian oscillator model for charged impurities}}\\ \vec{J} \supset -\frac{n_{\mathrm{imp}}e^2}{m^*} \frac{i\omega}{(\omega_0^2 - \omega^2) - i\gamma\omega} \vec{E} \end{array}$$

Recall, the observed frequency of the instability is characteristic of an impurity level.



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Results



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A branch of the impurity polariton is unstable!

- Using reasonable parameter values, some measured on the sample
- Agrees with experiment to order of magnitude
- Damping of the impurity oscillators can render the blue mode stable again

Steady-state vs. transient chiral transport



- Consequence of chiral anomaly
- Property of emergent massless carriers

e.g. X. Huang et al. (2015); Q. Li et al. (2016)



Conclusions & Outlook

- We propose a simple electrodynamic model in which an impurity polariton is rendered dynamically unstable by magneto-chiral transport effects in transiently excited chiral materials.
- Experimental results highlight the potential to manipulate THz-range radiation in far-from-equilibrium chiral materials, including wave amplification.
- Motivates a systematic study of nonlinear dynamics including background fields in far-from-equilibrium microscopic descriptions of chiral semiconductors.



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Bonus slides

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Directions for further work

- Our linear response model implicitly assumes fast a long-lived out-of-equilibrium state.
 - Plausible—carrier recombination is slow. Jnawali et al., Nat. Comm. 11, 3991 (2020)
 - Calls for a more microscopic approach, e.g. a kinetic theory.

Understand observed dependence on applied magnetic field

 Not yet taken into account

• Better experimentally resolve the polarization of the outgoing radiation.



10

time (ps)

15



Some experimental data



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Other enantiomorphic point groups

Point Groups matrix representation of Ξ Cartesian axes					
C_1	$\begin{bmatrix} T_{xx} & T_{xy} & T_{xz} \\ T_{yx} & T_{yy} & T_{yz} \\ T_{zx} & T_{zy} & T_{zz} \end{bmatrix}$	arbitrary			
C_2	$\begin{bmatrix} T_{xx} & T_{xy} & 0 \\ T_{yx} & T_{yy} & 0 \\ 0 & 0 & T_{zz} \end{bmatrix}$	two-fold rotation along \boldsymbol{z}			
D ₂	$\begin{bmatrix} T_{xx} & 0 & 0 \\ 0 & T_{yy} & 0 \\ 0 & 0 & T_{zz} \end{bmatrix}$	rotation axes along (x, y, z)	D_4	$\begin{bmatrix} T_{ } & 0 & 0 \\ 0 & T_{ } & 0 \\ 0 & 0 & T \end{bmatrix}$	four-fold rotation along z
C ₃	$\begin{bmatrix} T_{ } & -T & 0 \\ T^{-} & T_{ } & 0 \\ 0 & 0 & T_{zz} \end{bmatrix}$	three-fold rotation along z	C_6	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	six-fold rotation along z
D_3	$\left[egin{array}{cccc} T_{ } & 0 & 0 \ 0 & T_{ } & 0 \ 0 & 0 & T_{zz} \end{array} ight]$	three-fold rotation along z		$\begin{bmatrix} 0 & 0 & T_{zz} \\ T_{ } & 0 & 0 \end{bmatrix}$	
C_4	$\begin{bmatrix} T_{ } & -T^{-} & 0 \\ T^{-} & T_{ } & 0 \\ 0 & 0 & T_{-} \end{bmatrix}$	four-fold rotation along z	D ₆	$\begin{bmatrix} 0 & T_{ } & 0 \\ 0 & 0 & T_{zz} \end{bmatrix}$	six-fold rotation along z
ļļ		<u> </u>	Т	$\left[\begin{array}{cccc} T_0 & 0 & 0 \\ 0 & T_0 & 0 \\ 0 & 0 & T_0 \end{array}\right]$	(x, y, z) along crystal axes
			0	$\left[\begin{array}{rrrr} T_0 & 0 & 0 \\ 0 & T_0 & 0 \\ 0 & 0 & T_0 \end{array}\right]$	(x, y, z) along crystal axes

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Trying adding Hall effects?

- Classical Hall effect (in large B-field) swamps valley Hall
- No anomalous Hall effect in TR-invariant tellurium
- Classical Hall effect causes instability to propagate (Nishida 2023), but the magnitude of the effect is off by several orders without unrealistic fine tuning
 - Would lead to B-dependent growth rate, contrary to observations



More than one unstable mode

Tweak the parameters?

 \rightarrow sometimes multiple oscillatory instabilities



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Speculative mechanism for imbalance



<u>Time-reversal-invariant chiral crystal:</u>

→ Opposite-chirality Weyl nodes are not related by symmetries, so generically have different energies.

 \rightarrow Hence, photoexcitation by a linearly polarized source populates opposite-chirality nodes differently.

 \rightarrow Plus, after photoexcitation, the average rate of decay of the right-handed fermions generically differs from that of the left-handed ones.